



Universiteit
Leiden
The Netherlands

Can the Blandford-Znajek mechanism power steady jets?

King, A.R.; Pringle, J.E.

Citation

King, A. R., & Pringle, J. E. (2021). Can the Blandford-Znajek mechanism power steady jets? *Astrophysical Journal Letters*, 918(1). doi:10.3847/2041-8213/ac19a1

Version: Not Applicable (or Unknown)

License: [Leiden University Non-exclusive license](#)

Downloaded from: <https://hdl.handle.net/1887/3263648>

Note: To cite this publication please use the final published version (if applicable).

CAN THE BLANDFORD-ZNAJEK MECHANISM POWER STEADY JETS?

A.R. KING^{1,2,3} AND J.E. PRINGLE⁴

¹*School of Physics and Astronomy, University of Leicester, Leicester, LE1 7RH, UK*

²*Astronomical Institute Anton Pannekoek, University of Amsterdam, Science Park 904, NL-1098 XH Amsterdam, Netherlands*

³*Leiden Observatory, Leiden University, Niels Bohrweg 2, NL-2333 CA Leiden, Netherlands*

⁴*Institute of Astronomy, University of Cambridge, Madingley Road, Cambridge CB3 0HA, UK*

ABSTRACT

We consider the Blandford-Znajek (BZ) mechanism for extracting black hole spin energy to drive astrophysical jets. In analyses of the BZ mechanism it is always assumed that the electric charge of the black hole remains zero. But, as noted by Wald and others, if the medium surrounding the black hole is an ionised plasma with mobile charges, then a spinning hole quickly acquires an electric charge. The effect of this charge is to nullify the electric field structures which drive the BZ mechanism – the electric and magnetic fields then obey $\mathbf{E} \cdot \mathbf{B} = 0$ everywhere. Since jets are now observed in a wide variety of classes of accreting objects, most of which do not contain a central black hole, it seems likely that the jet driving mechanism in all astrophysical objects uses energy directly from the accretion disc, rather than black hole spin.

Keywords: astrophysical jets – accretion discs – black hole physics

1. INTRODUCTION

Early maps of double-lobe radio galaxies (e.g. Mitton & Ryle, 1969) showed amorphous blobs of radio emission symmetrically placed each side of the central galaxy. Rees (1971) suggested that an unknown object in the galactic nucleus channels energy to the radio lobes through jets. There is now almost universal agreement that the central object in radio galaxies is a supermassive black hole (Rees 1984), and that the high energy activity in the nucleus is powered by accretion (Salpeter, 1964), most likely through an accretion disc (Lynden-Bell, 1969). Thus the relativistic jets seen to emanate from the nucleus are ultimately powered by accretion of matter on to the central black hole (see the reviews by Begelman, Blandford & Rees, 1984; Heckman & Best, 2014; Blandford, Meier & Readhead, 2019).

Blandford & Znajek (1977, hereafter BZ77) proposed a radical new mechanism in which the jets from galactic nuclei are powered by direct electromagnetic extraction of the spin energy from the central black hole. This is the Blandford–Znajek (BZ) mechanism.

In Section 2.1 we provide a brief overview of this mechanism, which invokes a spinning black hole situated in an aligned magnetic field. We note that the central black hole is assumed always to have zero net electric charge, that is, any current entering the hole is assumed to be balanced by a precisely opposing current entering elsewhere, so that there is an effective current through the hole. In Section 2.2 we draw attention to the analysis by Wald (1974; see also King et al., 1975; Petterson, 1975; Gibbons et al., 2013) in which he finds that if the black hole is surrounded by a plasma which contains mobile charges, it will acquire a specific electric charge. The charge is exactly such as to render the BZ mechanism inoperative. We provide a brief discussion in Section 3.

2. THE BLANDFORD–ZNAJEK MECHANISM

2.1. *The basic mechanism; the uncharged black hole*

Black holes are generally assumed to have zero net electric charge. The reason is that if for example a Schwarzschild (non-rotating) black hole is given a charge, the neighbourhood of the hole then acquires an electric field. If (and only if) charge separation is allowed, charged particles in the surrounding astrophysical plasma move parallel or antiparallel to the electric field. In this way the black hole selectively acquires charges so that it moves quickly towards zero net charge. Charge separation is well known to occur in electrical media in which the charge carriers are able to move independently. It occurs for example in electrolytes (Debye & Hückel, 1923) and in ionized astrophysical plasmas (Salpeter, 1954), and it affects the rates of nuclear burning in stars (e.g. Clayton, 1968). The timescale to reach zero net charge is typically quite short in realistic astrophysical environments (see, for example, the discussion by Cardoso et al., 2016). Thus, Blandford

(1987) asserts that “charged, Kerr–Newman black holes are irrelevant to astronomy”.

Soon after black holes were recognised as a realistic astrophysical possibility, Wald (1974) devised an elegant method to compute the effect on the electromagnetic field structure when a black hole, with zero net charge, is placed in a uniform magnetic field. King et al. (1975) extended this result to consider more general aligned magnetic field structures. Wald (1974) showed that for a spinning (Kerr) black hole, with zero net electric charge, and with spin aligned with an external uniform magnetic field $\mathbf{B}$ in a vacuum, the spin induces an electric field with $\mathbf{E} \cdot \mathbf{B} \neq 0$ (for a classical analogy, see Ruffini & Treves, 1973). Wald assumed a field uniform at infinity, but King et al. (1975) showed that the field structure near the hole is essentially the same for any realistic aligned field. Wald also noted that an electric field with $\mathbf{E} \cdot \mathbf{B} \neq 0$ would lead to movement of any external charged particles. King et al. (1975) showed that the induced electric field structure is such that $\mathbf{E} \cdot \mathbf{B}$ has opposite signs near the poles and near the equator of the spinning hole. This implies that charges of one sign would be attracted towards the poles, whereas charges of the opposite sign would be attracted to a band near the equator.

This result led BZ77 to consider a spinning black hole at the centre of an accretion disc with the disc providing the required currents to give rise to an aligned magnetic field. The Bardeen–Petterson effect (Bardeen & Petterson, 1975) ensures that the central disc and black hole spin are generally aligned, and so by symmetry the magnetic field should also be aligned with the spin. BZ77 then noted that if the black hole is surrounded by a conducting medium the structure of the induced electric field found by Wald (1974) and by King et al. (1975) must produce an effective electric current through the black hole, returning through the surrounding medium (see, for example, Thorne & Blandford, 1982; Blandford, 1987). Any dissipation of this current’s energy in the surrounding material then taps the spin energy of the black hole.

This is the BZ–mechanism. It provides, in principle, a mechanism for the continuous extraction of energy from a spinning black hole. BZ77 speculated that it could be used to power the astrophysical jets seen to emanate from active galactic nuclei (see also Blandford, Meier & Readhead, 2019). The luminosity L_{BZ} produced by this process is, on dimensional grounds,

$$L_{\text{BZ}} \sim a^2 \times \frac{B^2}{8\pi} \times \frac{4\pi R_s^3}{3} \times \frac{c}{R_s}. \quad (1)$$

for a black hole of mass M , Schwarzschild radius $R_s \sim 2GM/c^2$, dimensionless spin parameter a , where $0 \leq a^2 \leq 1$, placed in a magnetic field of strength B , i.e. approximately the magnetic energy contained by the formal ‘volume’ of the hole, emitted every light crossing

time of the hole, and moderated by the amount of spin energy available.

For a ‘fiducial’ field strength (Begelman, Blandford & Rees, 1984; but see Ghosh & Abramowicz, 1997) of

$$B \sim 2 \times 10^4 M_8^{-1/2} \text{ G}, \quad (2)$$

this gives a luminosity

$$L_{\text{BZ}} \sim 2 \times 10^{45} a^2 M_8 \text{ erg s}^{-1}, \quad (3)$$

where M_8 is the mass of the black hole in units of $10^8 M_\odot$. This is comparable to the Eddington luminosity

$$L_E \sim 1.3 \times 10^{46} M_8 \text{ erg s}^{-1}. \quad (4)$$

BZ77 investigated this process using the force-free approximation in a charge-separated plasma in which particle inertia and interparticle collision terms can be ignored (see also Komissarov, 2004). Since then a number of authors have investigated this mechanism assuming that the surrounding medium could be modelled using the MHD approximation in which collision terms dominate, often by numerical means (see for example the reviews by Davis & Tchekhovskoy, 2020; Komissarov & Porth, 2021). However, in all these investigations, it is implicitly assumed that the movements of charges in any surrounding plasma do not permit a change in the net electrical charge of the black hole.

2.2. The acquisition of charge and its implication

The discussion of Section 2.1 above assumes, as is usually the case in astronomy, that the spinning black hole has zero net electric charge. However, in his seminal paper, Wald (1974) reasoned that a black hole would be surrounded by a standard astrophysical plasma in which the possibility of charge separation would exist. In this case, it is to be expected that the electric field drives the charges in such a way as to lead to a drop in electric potential along the magnetic field lines (cf. Komissarov, 2004). He argued further that for a spinning black hole in an aligned magnetic field $\mathbf{B}$ the movements of charge carriers (i.e. currents) induced by the electric field with $\mathbf{E} \cdot \mathbf{B} \neq 0$, together with the fact that the charge carriers are individually mobile, would lead to the hole selectively accreting net charge in such a way as to nullify the effects of the electric field. Specifically, he showed that once the net charge reaches the value

$$Q = 2BJ \quad (5)$$

in geometrized units (where $J = Ma$ is its total angular momentum), the electromagnetic 4-potential A^μ and therefore the electromagnetic field structure are identical to those of a non-spinning black hole embedded in the same uniform field. For such a field $\mathbf{E} \cdot \mathbf{B} = 0$ everywhere. This result was confirmed and generalised by

Petterson (1975). Since a non-spinning black hole has no spin energy to be tapped, it is evident that if a spinning black hole acquires the charge Q , the BZ mechanism can no longer operate. The charge Q is utterly negligible ($Q^2 \ll M^2$) in gravitational terms (Wald, 1974; Zajacek & Thursanov, 2018; Zajacek, 2019), so the spacetime metric is still to a high approximation uncharged Kerr, in agreement with the remark by Blandford (1987) quoted above. In particular the motion of uncharged particles is effectively identical to the case $Q = 0$. But charged particle motion is very different, and strongly influenced by the charge Q . This is another illustration of how extremely weak gravity is by comparison with electromagnetism. Thus, if the surrounding conducting medium is treated as a realistic space plasma which permits a net flow of charge into the black hole, then rather than providing a continuous process for removing spin energy from the hole, the induced electric currents are an initial transient effect which continues only until the black hole acquires the charge $Q = 2BJ$ (cf Zajacek & Thursanov 2018; Zajacek 2019)¹.

The energy released in this transient is, using the above numbers,

$$E_Q \sim L_{\text{BZ}} \times \frac{R_s}{c} \sim 2 \times 10^{48} a^2 M_8^2 \text{ erg}, \quad (6)$$

i.e. the transient emits the luminosity L_{BZ} for a time $\sim R_s/c \sim 10^3 M_8 \text{ s}$. Once the black hole acquires this charge, the torque on it vanishes and no more spin energy can be extracted.

More recently numerical modelling of the BZ-mechanism has been undertaken using Particle-In-Cell (PIC) plasma methods, which permit independent mobility of individual charges. In principle these techniques should be able to test Wald’s fundamental hypothesis that a spinning black hole immersed in a magnetic field should acquire a net charge. The same technique has also been applied to ionised plasma surrounding rotating neutron stars (e.g. Kalapotharakos et al., 2018). However, in contrast to the neutron star case, where Kalapotharakos et al. (2018) treat the inner boundary with some care, noting that they ensure “current closure of charge carriers that reach the stellar surface”, those modelling the black hole case have so far chosen inner boundary conditions which either prevent or ignore the acquisition of charge by the black hole. For example, Parfrey et al. (2019, see also Crinquand et al, 2021) simply remove both the mass and the charge of any particle entering the black hole, and Hirotani et al. (2021) impose that $\mathbf{E} \cdot \mathbf{B} = 0$ and that both the radial component of the electric field and the meridional component of the magnetic field vanish at the inner boundary.

¹ For all conceivable boundary conditions far from the hole its net charge tends monotonically to Q .

3. DISCUSSION

The BZ mechanism is a standard cited mechanism for producing steady jets in objects that contain accreting black holes – galactic nuclei and some X-ray binaries. And indeed the process is often cited in papers which concern numerical MHD simulations of jets and outflows produced by magnetic accretion discs (see the review by Davis & Tchekhovskoy, 2020). We have argued above, in line with the original ideas of Wald (1974; see also Gibbons et al., 2013) that in a realistic space plasma which permits a net flow of charge into the black hole the BZ mechanism cannot tap the spin energy of a black hole continuously, and is therefore not a viable mechanism for powering continuous astrophysical jets.

This is primarily because the conducting medium surrounding the black hole should be treated as an astrophysical plasma with mobile charge carriers. When charge separation is allowed, along with the possibility of a net flux of charge into the black hole, any spinning black hole quickly acquires the net electrical charge Q , and the electric fields which drive the currents required for the BZ mechanism are nullified. The same process of charge separation which ensures that a non-rotating black hole has zero charge also ensures that a rotating hole acquires exactly the charge Q that makes the BZ mechanism inoperative. We have noted that these ideas need to be tested, for example using PIC plasma simulation techniques. For example, it might be that collective plasma effects serve to counteract the tendency of the black hole to acquire charge².

There are, of course, many other kinds of astrophysical objects which do not contain black holes and nevertheless produce jets (Burgarella et al., 1993; Smith, 2012). The jets emitted by young stellar objects are particularly spectacular (see the review by Ray & Ferreira, 2021). Thus the application of Occam’s razor³ has long suggested that the BZ mechanism, even if it were viable, is in fact not required for the production of astrophysical jets (Livio, 1996; see also Pringle, 1993; Price, Pringle & King, 2003). In addition, Russell, Gallo & Fender (2013) have shown that the jet production mechanism in binary X-ray sources is not consistent with the prediction of the BZ mechanism that the jet power should depend on the the square of the dimensionless jet spin parameter (see equation 3).

Thus, following Occam, if one is forced to choose a single mechanism capable of producing all astrophysical jets which emanate from accreting objects, then the most likely choice would be some form of MHD process resulting from poloidal magnetic field threading the accretion disc (Livio, 1996; Livio et al., 1999). A mechanism like this is already discussed by Blandford & Zna-

jek (1977), and early ideas on this process are given by Blandford & Payne (1982) and, in the protostellar case, by Pudritz & Norman (1983, 1986).

ACKNOWLEDGMENTS

We thank Bob Carswell, Gary Gibbons, Chris Nixon and Colin Norman for helpful comments, and the referee for a thoughtful report.

REFERENCES

- Bardeen, J.A. & Petterson, J.A., 1975, *ApJL* 195, L65
- Begelman, M.C., Blandford, R.D., & Rees, M.J., 1984, *Rev. Mod. Phys.*, 56, 255
- Blandford, R.D., 1987, in *Three Hundred Years of Gravitation*, Cambridge University Press, eds S. Hawking & W. Israel, pp. 277 – 329
- Blandford, R.D., Meier, D. & Readhead, A. 2019, *ARA&A*, 57, 467
- Blandford, R.D., & Payne, D.G., 1982, *MNRAS*, 199, 883
- Blandford, R.D., & Znajek, R. L. 1977, *MNRAS* 179, 433
- Burgarella, D., Livio, M., & O’Dea, C.P. (eds), 1993, *Astrophysical Jets*, Cambridge University Press
- Cardoso, V., Macedo C. F. B., Pani, P., & Ferrari, V., 2016, *JCAP*, 054
- Clayton, D.D., 1968, *Principles of Stellar Evolution and Nucleosynthesis*, Univ. of Chicago Press
- Crinquand, B., Cerutti, B., Dubus, G., Parfrey, K., Philippov, A., 2021, *A&A*, 650, A163
- Debye, P., & Hückel, E., 1923, *Physikalische Zeitschrift*, 24, 185
- Ghosh, P. & Abramowicz, M.A., 1997, *MNRAS*, 292, 887
- Gibbons, G. W., Mujtaba, A. H., & Pope, C. N. 2013, *Classical and Quantum Gravity*, 30, 125008
- Heckman, T.M., & Best, P. N., 2014, *ARA&A*, 52, 589
- Hirovani, K., Krasnopolsky, R., Shang, H., Nishikawa, K., Watson, M., 2021, *ApJ*, 908, 88
- Kalapotharakos, C., Brambilla, G. Timokhin, A. Harding, A. K. & Kazanas, D., 2018, *ApJ*, 857, 44
- King, A.R., Lasota, J.P., & Kundt, W., 1975, *Phys. Rev. D*, 12, 3037
- Komissarov, S.S., 2004, *MNRAS*, 350, 427
- Komissarov, S.S., & Porth, O., 2021, *NewAR*, 92, 101610
- Livio, M., 1997, in *Accretion Phenomena and Related Outflows* eds D.T. Wickramasinghe, L. Ferrario, & G.V. Bicknell, ASP Conference Series 121, 845
- Livio, M., Ogilvie, G.I., & Pringle, J.E., 1999, *ApJ*, 512, 100
- Lynden-Bell, D., 1969, *Nature*, 223, 690L
- Mitton, S., & Ryle, M., 1969, *MNRAS*, 146, 221

² We thank the referee for stressing this possibility.

³ ‘Do not multiply hypotheses’, or put simply, ‘don’t invent two theories for the same thing’.

- Parfrey, K., Philippov, A., Cerutti, B., 2019, Phys. Rev. Lett. 122c5101
- Petterson, J. A. 1975, Phys Rev D, 12, 2218
- Price, D.J., Pringle, J.E., & King, A.R., 2003, MNRAS 339, 1223
- Pringle, J.E., 1993, in *Astrophysical Jets*, Cambridge University Press, eds Burgarella, D., Livio, M., & O'Dea, C.P.
- Pudritz, R.E., & Norman, C.A., 1983, ApJ, 274, 677
- Pudritz, R.E., & Norman, C.A., 1986, ApJ, 301, 571
- Ray, T.P., & Ferreira, J. 2021, NewAR, 93, 101615
- Rees, M.J., 1971, Nature, 229, 312
- Rees, M.J., 1984, ARA&A, 22, 471
- Ruffini, R. & Treves, A. 1973, Astrophys. Lett., 13, 109
- Russell, D.M., Gallo, E., & Fender, R.P., 2013, MNRAS, 431, 405
- Salpeter, E.E., 1954, Aus J. Phys., 7, 373
- Salpeter, E.E., 1964, ApJ, 140, 796
- Smith, M.D., 2012, *Astrophysical Jets and Beams* Cambridge University Press
- Thorne, K.S., & Blandford, R.D., 1982, in *Proc IAU Symp 97* eds. Heeschen, D.S. & Wade, C.M., 255
- Wald, R.M., 1974, Phys Rev D, 10, 1680
- Zajacek, M, 2018, MNRAS, 480, 4408
- Zajacek, M., & Thursonov, A., 2019, The Observatory, 139, 231