



Erratum: “CONSTRAINING THE LOW-MASS SLOPE OF THE STAR FORMATION SEQUENCE AT $0.5 < z < 2.5$ ” (2014, *ApJ*, 795, 104)

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The level of precision of the polynomial coefficients in Table 1 is not adequate to reproduce the best-fit lines accurately in Figures 1 and 2 of the published article. This erratum provides an update to values in Table 1.

Table 1
Polynomial Coefficients

Redshift (5)	<i>a</i> (6)	<i>b</i> (7)	<i>c</i> (8)
$0.5 < z < 1.0$	-27.4020 ± 1.91	5.0215 ± 0.39	-0.2196 ± 0.02
$1.0 < z < 1.5$	-26.0395 ± 1.69	4.6179 ± 0.34	-0.1904 ± 0.02
$1.5 < z < 2.0$	-24.0415 ± 2.08	4.1693 ± 0.40	-0.1638 ± 0.02
$2.0 < z < 2.5$	-19.9898 ± 1.87	3.4422 ± 0.36	-0.1299 ± 0.02

Note. Corrected polynomial coefficients parameterizing the evolution of the $\log \Psi$ – $\log M_*$ relation from the median stacking analysis presented in Figure 1 (see Equation (2)) of Whitaker et al. (2014). When plotting the confidence intervals of the polynomial fits, note that the standard uncertainties in the polynomial coefficients are sign dependent.

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References

Whitaker, K. E., Franx, M., Leja, J., et al. 2014, *ApJ*, 795, 104

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