



Universiteit
Leiden
The Netherlands

Gravitational Wilson lines in 3D de Sitter

Castro, A.; Garnier, P.A.; Zukowski, C.

Citation

Castro, A., Garnier, P. A., & Zukowski, C. (2020). Gravitational Wilson lines in 3D de Sitter. *Journal Of High Energy Physics*, (7), 202. doi:10.1007/JHEP07(2020)202

Version: Publisher's Version

License: [Creative Commons CC BY 4.0 license](#)

Downloaded from: <https://hdl.handle.net/1887/139084>

Note: To cite this publication please use the final published version (if applicable).

RECEIVED: March 25, 2020

REVISED: July 5, 2020

ACCEPTED: July 6, 2020

PUBLISHED: July 28, 2020

Gravitational Wilson lines in 3D de Sitter

Alejandra Castro,^a Philippe Sabella-Garnier^b and Claire Zukowski^a

^a*Institute for Theoretical Physics, University of Amsterdam,
Science Park 904, Postbus 94485, 1090 GL Amsterdam, The Netherlands*

^b*Lorentz Institute, Leiden University,
Niels Bohrweg 2, 2333 CA Leiden, The Netherlands*

E-mail: a.castro@uva.nl, garnier@lorentz.leidenuniv.nl,
c.e.zukowski@uva.nl

ABSTRACT: We construct local probes in the static patch of Euclidean dS_3 gravity. These probes are Wilson line operators, designed by exploiting the Chern-Simons formulation of 3D gravity. Our prescription uses non-unitary representations of $so(4) \simeq su(2)_L \times su(2)_R$, and we evaluate the Wilson line for states satisfying a singlet condition. We discuss how to reproduce the Green's functions of massive scalar fields in dS_3 , the construction of bulk fields, and the quasinormal mode spectrum. We also discuss the interpretation of our construction in Lorentzian signature in the inflationary patch, via $SL(2, \mathbb{C})$ Chern-Simons theory.

KEYWORDS: Chern-Simons Theories, Classical Theories of Gravity, Wilson, 't Hooft and Polyakov loops

ARXIV EPRINT: [2001.09998](https://arxiv.org/abs/2001.09998)

Contents

1	Introduction	1
1.1	Overview	3
2	Chern-Simons formulation of Euclidean dS_3 gravity	3
3	Wilson lines in $SO(4)$ Chern-Simons	5
3.1	Non-unitary representations of $so(4)$	6
3.1.1	Singlet states	9
3.2	Wilson line and the Green's function	11
4	Local pseudofields from Wilson lines	13
4.1	Wilson line as an overlap of states	13
4.2	Construction of local basis	14
4.2.1	Wavefunction for the singlet states	17
4.2.2	Wick rotation and quasi-normal modes	18
5	Wilson lines in $SL(2, \mathbb{C})$ Chern-Simons	19
5.1	Chern-Simons formulation of Lorentzian dS_3 gravity	20
5.2	Construction of probes in $sl(2, \mathbb{C})$	20
5.3	Inflationary patch	22
6	Discussion	26
A	Conventions	29
B	Metric formulation of dS_3 gravity	29
B.1	Coordinates and patches	29
B.2	Geodesics and Green's functions in dS_3	32
C	Analytic continuation in the Chern-Simons formulation	34

1 Introduction

The Chern-Simons formulation of three-dimensional gravity seems more amenable to quantization than the more traditional metric formulation [1, 2]. One advantage is that the gauge theory formulation makes evident the topological nature of Einstein's theory in three dimensions. Also, Chern-Simons theory has inherently holographic properties: upon specifying a gauge group and boundary conditions on a 3-manifold with a boundary, the Chern-Simons theory can be viewed as dual to a conformal theory living on the boundary [3–5]. These features have propelled the use of the Chern-Simons formulation as a computational tool in perturbative gravity.

However, this alternative formulation of 3D gravity comes with a cost: local observables that are intuitive in a metric formulation — such as distances, surfaces, volumes, and local fields — are seemingly lost in Chern-Simons theory. To reintroduce this intuition, Wilson lines present themselves as reasonable objects in Chern-Simons that could restore portions of our geometric and local intuition [6]. In the early stages, it was clear that a Wilson line anchored at the boundary would correspond to a conformal block in the boundary theory [5, 7]; more recently this proposal has been made more precise and explicit for $SL(2)$ Chern-Simons theory [8–15]. In the context of AdS_3 gravity, where the relevant gauge group is $SO(2, 2)$, Wilson lines have been applied in a plethora of different contexts [16–20], with recent applications ranging from the computation of holographic entanglement entropy [21–27] to the probing of analytic properties of an eternal black hole [28, 29]. Applications of Wilson lines in Chern-Simons to flat space holography includes [30], and to ultra-relativistic cases [31, 32].

In the present work we will study $SO(4)$ Chern-Simons theory on a Euclidean compact manifold. This theory can be interpreted as a gravitational theory with positive cosmological constant, i.e. Euclidean dS_3 gravity. This instance is interesting from a cosmological perspective, where Chern-Simons theory could provide insights into appropriate observables in quantum cosmology. It is also powerful, since there is an extensive list of exact results in Chern-Simons theory for compact gauge group. Previous efforts that exploited this direction of Chern-Simons theory as a toy model for quantum cosmology include [33–38].

Our main emphasis is to interpret Wilson lines in $SO(4)$ Chern-Simons theory as local probes for dS_3 gravity, which follows closely the proposal in [6] for $SO(2, 2)$ Chern-Simons. The basic idea is as follows. We will consider a connection $\mathcal{A}$ valued on $so(4)$, and a Wilson line stretching from a point x_i to x_f :

$$W_{\mathcal{R}}(x_i, x_f) = \langle U_f | \mathcal{P} \exp \left(- \int_{x_i}^{x_f} \mathcal{A} \right) | U_i \rangle . \quad (1.1)$$

There are two important ingredients in defining this object. First we need to select a representation $\mathcal{R}$ of $so(4)$. This choice will encode the physical properties of the local probe, such as mass and spin. The second ingredient is to select the endpoint states $|U_{i,f}\rangle$: the freedom in this choice encodes the gauge dependence of $W_{\mathcal{R}}(x_i, x_f)$. More importantly, their choice will allow us to relate $W_{\mathcal{R}}(x_i, x_f)$ to the Euclidean Green’s function of a massive field propagating on S^3 . And while our choices are inspired by the analogous computations in AdS_3 gravity, they have a standing on their own. We will motivate and introduce the ingredients needed to have an interesting interpretation of (1.1) using solely $SO(4)$ Chern-Simons theory.

The interpretation of our results in the Euclidean theory will have its limitations if they are not analytically continued to Lorentzian signature. For example, recognising if the information contained in $W_{\mathcal{R}}(x_i, x_f)$ is compatible with causality necessitates a Lorentzian understanding of the theory. This is tied with the issue of bulk locality and reconstruction in de Sitter, which remains intriguing in cosmological settings. In the Chern-Simons formulation, the Lorentzian theory corresponds to a theory with gauge group $SL(2, \mathbb{C})$. We will present the basics of how to discuss our results in $SL(2, \mathbb{C})$ Chern-Simons theory,

and their relation to the Euclidean theory. One interesting finding is that our choice of representation in $SO(4)$ Chern-Simons naturally leads to quasinormal modes in the static patch of dS_3 when analytically continued.

1.1 Overview

In section 2, we review the Chern-Simons formulation of Euclidean dS_3 (EdS_3) gravity, establishing our conventions along the way.

In section 3, we describe Wilson lines in $SO(4) \simeq SU(2) \times SU(2)$ Chern-Simons. We show how the Green's function on EdS_3 of a scalar field of given mass can be described by a Wilson line evaluated in a non-unitary representation of the algebra, which we construct in detail. These unusual representations of $su(2)$ resemble the usual spin- l representation, with the important distinction that $-1 < l < 0$. And while it might be odd to treat l as a continuous (negative) parameter, these features will be key to recover local properties we attribute to dS_3 in Chern-Simons theory.

In section 4, we take this further and show how this description of the Wilson line can be used to define local states in the geometry. We present a map between states in the algebraic formulation and the value of a corresponding scalar pseudofield in the metric formulation, and we build an explicit position-space representation of the basis states. We also match the action of the generators of the algebra to the Killing vectors of the geometry. The local pseudofields constructed from the Wilson line continue to quasinormal modes in the static patch, and they are acted on by an $sl(2, \mathbb{R}) \times sl(2, \mathbb{R})$ inherited from our representations. This can be contrasted to a similar $sl(2, \mathbb{R})$ structure of the quasi-normal mode spectrum that was discovered and dubbed a “hidden symmetry” of the static patch in [39].

In section 5, we discuss how to analytically continue our results to Lorentzian dS_3 gravity, which is described by an $SL(2, \mathbb{C})$ Chern-Simons theory. We find that our exotic $so(4)$ representations analytically continue to a highest-weight representation of an $sl(2, \mathbb{R}) \times sl(2, \mathbb{R})$ slice of $sl(2, \mathbb{C})$.

In section 6, we highlight our main findings and discuss future directions to further explore quantum aspects of dS_3 gravity. Finally, appendix A collects some of our conventions for easy reference, and appendix B reviews some basic facts about the metric formulation of dS_3 . In appendix C, we give more details about how to construct an analytic continuation between the $SO(4)$ and $SL(2, \mathbb{C})$ Chern-Simons theories.

2 Chern-Simons formulation of Euclidean dS_3 gravity

For the purposes of setting up notation and conventions we begin with a short review of Chern-Simons gravity, focusing on its relation to Euclidean dS_3 gravity. This is based on the original formulation of 3D gravity as a Chern-Simons theory [1, 2]; and related work on Euclidean dS_3 in the Chern-Simons formulation are [35, 38], although we warn the reader that conventions there might be different than ours. In appendix B we provide a review of the metric formulation of dS_3 gravity.

Consider Chern-Simons theory on $\mathcal{M} = S^3$ with gauge group $SO(4)$. This group manifestly splits into $SO(4) \simeq SU(2)_L \times SU(2)_R$, and in terms of its Lie algebra we use

generators L_a for $su(2)_L$ and $\bar{L}_a$ for $su(2)_R$, $a = 1, 2, 3$. Our conventions are such that

$$[L_a, L_b] = i\epsilon_{abc}L^c, \quad (2.1)$$

and similarly for the $\bar{L}_a$; we also set $\epsilon_{123} \equiv 1$. There is an invariant bilinear form given by the trace: we take

$$\text{Tr}(L_a L_b) = \text{Tr}(\bar{L}_a \bar{L}_b) = \frac{1}{2}\delta_{ab}. \quad (2.2)$$

Indices in (2.1) are raised with δ^{ab} .

The $SO(4)$ Chern-Simons action relevant for Euclidean dS₃ gravity is

$$S_E = S_{\text{CS}}[A] - S_{\text{CS}}[\bar{A}], \quad (2.3)$$

where

$$A = A_\mu^a L_a dx^\mu, \quad \bar{A} = \bar{A}_\mu^a \bar{L}_a dx^\mu, \quad (2.4)$$

and the individual actions are

$$S_{\text{CS}}[A] = -\frac{k}{4\pi} \int_{\mathcal{M}} \text{Tr} \left(A \wedge dA + \frac{2}{3} A \wedge A \wedge A \right), \quad (2.5)$$

and similarly for $S_{\text{CS}}[\bar{A}]$.

The relation to the first-order formulation of the Einstein-Hilbert action is as follows. The algebra that describes the isometries of Euclidean dS₃ is

$$\begin{aligned} [J_{ab}, P_c] &= -\delta_{ac}P_b + \delta_{bc}P_a, \\ [J_{ab}, J_{cd}] &= -\delta_{ac}J_{bd} + \delta_{bc}J_{ad} + \delta_{ad}J_{bc} - \delta_{bd}J_{ac}, \\ [P_a, P_b] &= -\Lambda J_{ab}, \end{aligned} \quad (2.6)$$

where $\Lambda = \frac{1}{\ell^2}$, and ℓ is the radius of the 3-sphere. Here P_a and J_{ab} are the generators of translations and rotations of the ambient $\mathbb{R}^4$, respectively. We also raise indices with δ^{ab} . It is convenient to define the dual

$$J_a = \frac{1}{2}\epsilon_{abc}J^{bc}, \quad J_{ab} = \epsilon_{abc}J^c. \quad (2.7)$$

In relation to the $su(2)$ generators, we identify

$$L_a = -\frac{i}{2}(J_a + \ell P_a), \quad \bar{L}_a = -\frac{i}{2}(J_a - \ell P_a). \quad (2.8)$$

The variables in the gravitational theory are the vielbein and spin connection,

$$e^a = e_\mu^a dx^\mu, \quad \omega^a = \frac{1}{2}\epsilon^a_{bc}\omega_\mu^{bc}dx^\mu. \quad (2.9)$$

The vielbein is related to the metric as $g_{\mu\nu} = e_\mu^a e_\nu^b \delta_{ab}$. We define the gauge field in terms of these geometrical variables as

$$A = i \left(\omega^a + \frac{1}{\ell} e^a \right) L_a, \quad \bar{A} = i \left(\omega^a - \frac{1}{\ell} e^a \right) \bar{L}_a. \quad (2.10)$$

Using (2.10), the action (2.3) becomes

$$S_E = \frac{k}{2\pi\ell} \int_{\mathcal{M}} \left[e^a \wedge \left(d\omega_a - \frac{1}{2} \epsilon_{abc} \omega^b \wedge \omega^c \right) - \frac{1}{6\ell^2} \epsilon_{abc} e^a \wedge e^b \wedge e^c \right], \quad (2.11)$$

which reduces the Einstein-Hilbert action with positive cosmological constant given the identification

$$k = \frac{\ell}{4G_3}. \quad (2.12)$$

The equations of motion from (2.3) simply give the flatness condition,

$$F = dA + A \wedge A = 0, \quad \bar{F} = d\bar{A} + \bar{A} \wedge \bar{A} = 0, \quad (2.13)$$

which are related to the Cartan and Einstein equation derived from (2.11) after using (2.10).

The background we will mostly focus on is S^3 , which we will cast as

$$\frac{ds^2}{\ell^2} = dr^2 + \cos^2 r d\tau^2 + \sin^2 r d\phi^2, \quad (2.14)$$

with $(\tau, \phi) \sim (\tau, \phi) + 2\pi(m, n)$ and $m, n \in \mathbb{Z}$; see appendix B.1 for further properties of this background. In the Chern-Simons language, the associated connections that reproduce the vielbein and spin connection are

$$\begin{aligned} A &= iL_2 dr + i(L_3 \cos r + L_1 \sin r)(d\phi + d\tau), \\ \bar{A} &= -iL_2 dr + i(L_3 \cos r - L_1 \sin r)(d\phi - d\tau). \end{aligned} \quad (2.15)$$

Note that we are using the same basis of $su(2)$ generators for both A and $\bar{A}$. This is convenient since we then can read off the metric as

$$g_{\mu\nu} = -\frac{\ell^2}{2} \text{Tr} \left[(A_\mu - \bar{A}_\mu) (A_\nu - \bar{A}_\nu) \right]. \quad (2.16)$$

The corresponding group elements that we will associate to each flat connection read¹

$$\begin{aligned} A &= g_L dg_L^{-1}, & g_L &= e^{-irL_2} e^{-i(\phi+\tau)L_3}, \\ \bar{A} &= \tilde{g}_R^{-1} d\tilde{g}_R, & \tilde{g}_R &= e^{i(\phi-\tau)L_3} e^{-irL_2}. \end{aligned} \quad (2.17)$$

This can be checked explicitly by using the following corollary of the Baker-Campbell-Hausdorff formula,

$$e^{-i\alpha L_a} L_b e^{i\alpha L_a} = \cos(\alpha) L_b + \sin(\alpha) \epsilon_{abc} L^c. \quad (2.18)$$

3 Wilson lines in SO(4) Chern-Simons

A gauge-invariant observable in Chern-Simons theory is the Wilson loop operator, which in the Euclidean theory with gauge group $SO(4) \simeq SU(2)_L \times SU(2)_R$ reads

$$W_{\mathcal{R}}(C) = \text{Tr}_{\mathcal{R}} \left(\mathcal{P} \exp \left(- \oint_C A \right) \mathcal{P} \exp \left(- \oint_C \bar{A} \right) \right), \quad (3.1)$$

¹The notation $\tilde{g}_R$ here will be justified and explained in section 3.2.

where C is a closed loop in the 3-manifold $\mathcal{M}$. Here $\mathcal{R}$ is a particular representation of the Lie algebra associated to the Chern-Simons gauge group. One of the challenges of the Chern-Simons formulation of 3D gravity is to build local probes in a theory that insists on being topological. Here we will design those probes by considering a Wilson *line* operator, i.e., we will be interested in

$$W_{\mathcal{R}}(x_i, x_f) = \langle U_f | \mathcal{P} \exp \left(- \int_{\gamma} A \right) \mathcal{P} \exp \left(- \int_{\gamma} \bar{A} \right) | U_i \rangle . \quad (3.2)$$

The curve $\gamma(s)$ is no longer closed but has endpoints x_i, x_f . This operator is no longer gauge-invariant, which is reflected in the fact that we need to specify states at its endpoint, denoted as $|U_i\rangle, |U_f\rangle$. In the following we will discuss representations $\mathcal{R}$ of $so(4)$, and suitable endpoint states, giving $W_{\mathcal{R}}(x_i, x_f)$ local properties we can naturally relate to the metric formulation. The representations we consider will differ from the unitary representations that are typically considered in $SU(2)$ Chern-Simons theory.²

Our strategy to select the representation and the endpoint states is inspired by the proposal in [6, 21], which is a prescription to use Wilson lines as local probes in AdS_3 gravity. The basic observation is to view $W_{\mathcal{R}}(x_i, x_f)$ as the path integral of a charged (massive) point particle. In this context the representation $\mathcal{R}$ parametrizes the Hilbert space for the particle, with the Casimir of $\mathcal{R}$ carrying information about the mass and spin (i.e., quantum numbers) of the particle [16, 17, 40]. With this perspective, our first input is to consider representations of $so(4)$ that carry a continuous parameter that we can identify with the mass of particle. As we will show in the following, this requirement will force us to consider non-unitary representations of the group which we will carefully construct.

In the subsequent computations we will leave the connections A and $\bar{A}$ fixed, and quantize appropriately the point particle for our choice of $\mathcal{R}$. From this perspective, $W_{\mathcal{R}}(x_i, x_f)$ captures how the probe is affected by a given background characterized by A and $\bar{A}$. Here is where our choice of endpoint states will be crucial: our aim is to select states in $\mathcal{R}$ that are invariant under a subgroup of $so(4)$. Selecting this subgroup appropriately will lead to a novel way of casting local fields in the Chern-Simons formulation of dS_3 gravity.

3.1 Non-unitary representations of $so(4)$

Since $so(4) \simeq su(2)_L \times su(2)_R$, let us focus first on a single copy of $su(2)$. Recall that in our conventions, the $su(2)$ generators satisfy the algebra (2.1). The unique Casimir operator is the quadratic combination³

$$L^2 = L_1^2 + L_2^2 + L_3^2 . \quad (3.3)$$

We can build raising and lowering operators by defining

$$L_{\pm} \equiv L_1 \pm iL_2 , \quad L_0 \equiv L_3 . \quad (3.4)$$

²In the semi-classical regime, there is no principle in Chern-Simons theory that favors a choice of one representation over another. The choices we make, both for the representations and endpoint states, allow us to reproduce gravitational observables in de Sitter spacetime.

³Note that this definition of the Casimir discards the overall normalization of the bilinear form in (2.2).

For a compact group like $su(2)$ all unitary representations are finite dimensional and labelled by a fixed (half-)integer, the spin, as in the usual $SU(2)$ Chern-Simons theory. To introduce a continuous parameter, we need to build representations that are more analogous to the infinite-dimensional, highest-weight representations of $sl(2, \mathbb{R})$. This forces us to consider non-unitary representations, nevertheless a natural choice to make contact with local fields in dS_3 , as will show.

For unitary representations we would have that all of the L_a 's are Hermitian. Here we will relax this condition and choose generators that are not necessarily Hermitian. In particular, a consistent choice for a non-unitary representation that respects the Lie algebra is to take L_1, L_2 to be anti-Hermitian and L_3 to be Hermitian, which results in

$$L_{\pm}^{\dagger} = -L_{\mp}, \quad L_0^{\dagger} = L_0. \quad (3.5)$$

While it is not unique, this is the choice we will use to build a non-unitary representation. Notice that it is inconsistent to take all the generators to be anti-Hermitian, as this would violate the commutation relations (2.1). Informally, we only modify the construction of $su(2)$ representations as much as needed to obtain a continuous Casimir. As we see below, this modification is sufficient to obtain that property, with the rest of the construction mirroring the usual unitary case.

Our representation, despite its lack of unitarity, has to satisfy some minimal requirements which we will now discuss. We have a basis of vectors (states) that are joint eigenstates of L^2 and L_0 . These are denoted $|l, p\rangle$ with

$$L^2 |l, p\rangle = c_2(l) |l, p\rangle, \quad (3.6)$$

$$L_0 |l, p\rangle = (l - p) |l, p\rangle. \quad (3.7)$$

Here l labels the representation, i.e. controls the quadratic Casimir $c_2(l)$, and p labels the L_0 eigenvalue. Note that in a unitary representation we would use $m = l - p$, but we will find it more useful to use p as a label. We seek to build a representation such that the spectrum of L_0 is bounded (either from above or below), and that the norm squared of the states $|l, p\rangle$ is positive. To achieve these requirements, we build a representation by introducing a *highest weight state*. We define this state as

$$L_0 |l, 0\rangle = l |l, 0\rangle, \quad L_+ |l, 0\rangle = 0. \quad (3.8)$$

This in particular implies that we will create states by acting with L_- on $|l, 0\rangle$, and hence a basis for eigenstates is schematically given by $|l, p\rangle \sim (L_-)^p |l, 0\rangle$ with p a positive integer.⁴

Next we need to ensure that the norm of these states is *positive*; this will impose restrictions on the Casimir, and hence l . A useful identity in this regard is

$$\begin{aligned} |L_{\pm} |l, p\rangle|^2 &= -\langle l, p | L_{\mp} L_{\pm} |l, p\rangle \\ &= -c_2(l) + (l - p)(l - p \pm 1). \end{aligned} \quad (3.9)$$

⁴This follows from the commutation relation between $L_{\pm}$ and L_0 .

The minus sign in the first line comes from anti-Hermiticity in (3.5). In going from the first to the second line we used $L_{\mp}L_{\pm} = L^2 - L_0^2 \mp L_0$. The norm of $L_+|l, 0\rangle$ vanishing gives

$$c_2(l) = l(l+1), \quad (3.10)$$

relating the label l with the Casimir of the representation. Positivity of the norm of the first descendant requires

$$|L_-|l, 0\rangle|^2 = -2l > 0,$$

which clearly dictates that l is strictly negative. Any other state in the representation will be of the form

$$|l, p\rangle = \mathbf{c}_p(L_-)^p |l, 0\rangle, \quad (3.11)$$

where the normalization $\mathbf{c}_p$ is adjusted such that

$$\langle l, p' | l, p \rangle = \delta_{p', p}, \quad p = 0, 1, 2, \dots. \quad (3.12)$$

Demanding this relation leads to

$$L_+ |l, p\rangle = -\sqrt{p(p-2l-1)} |l, p-1\rangle, \quad (3.13)$$

$$L_- |l, p\rangle = \sqrt{(p+1)(p-2l)} |l, p+1\rangle. \quad (3.14)$$

The fact that the roles of the raising and lowering operators appear flipped, in other words L_+ lowers and L_- raises p , simply results from our convention in (3.6). If we had labelled states by their eigenvalue of L_0 they would raise and lower in the same way as the usual unitary $sl(2, \mathbb{R})$ representations. The minus sign in (3.13) is more fundamental. It was not present in highest weight representations of $sl(2, \mathbb{R})$; here it is necessary for the action of $L_{\pm}$ to be consistent with the $su(2)$ commutation relations.⁵

In the unitary case, representations are finite-dimensional since there is an upper bound for p . Additionally, the Casimir is strictly positive, and l is constrained to be either integer or half-integer. These constraints all come from demanding the positivity of squared norms. For our non-unitary representations, relaxing the requirement of Hermiticity means that p is not bounded and the Casimir is not necessarily positive. Our choices also lead to a spectrum of L_0 unbounded from below, whose eigenstates are (3.11)–(3.12). We also note that the Casimir is allowed to be negative since $l < 0$; in particular, for the range $-1 < l < 0$ we have

$$-\frac{1}{2} < c_2 < 0. \quad (3.15)$$

Our representation has a well-defined character too. Suppose we have a group element $M \in \text{SU}(2)$ which can be decomposed as

$$M = V^{-1} e^{i\alpha L_0} V. \quad (3.16)$$

Its character is simply given by

$$\text{Tr}(M) = \sum_{p=0}^{\infty} \langle l, p | e^{i\alpha L_0} | l, p \rangle = \frac{e^{i\alpha(l+1)}}{e^{i\alpha} - 1}. \quad (3.17)$$

⁵Normalization only determines $L_{\pm}|l, p\rangle$ up to a phase.

Finally, notice that for a fixed Casimir there are actually two distinct representations labelled by the two solutions for l in (3.10). These solutions are

$$l_{\pm} = -\frac{1 \pm \sqrt{1 + 4c_2}}{2}. \quad (3.18)$$

One representation has $-1 < l_+ < -\frac{1}{2}$ while the other has $-\frac{1}{2} < l_- < 0$, and each of these representations will be coined as $\mathcal{R}_{\pm}$. The role of $\mathcal{R}_{\pm}$ will become important later, when we compare the Wilson line to the Euclidean Green's function, and in the construction of local pseudofields. In particular, we will see that both representations are necessary to generate a complete basis of solutions for local fields in dS_3 .

3.1.1 Singlet states

Returning to $so(4) \simeq su(2)_L \times su(2)_R$, let's add a set of operators $\bar{L}_a$ with the same commutation relations as the unbarred ones and which commute with them:

$$[L_a, \bar{L}_b] = 0. \quad (3.19)$$

In the following we will be interested in building a state $|U\rangle$, assembled from the non-unitary representations of $su(2)$, that is invariant under a subset of the generators in $so(4)$. These states, denoted *singlet states*, will serve as endpoint states which we will use to evaluate the Wilson line (3.2). This construction is motivated by the derivations for $so(2, 2) \simeq sl(2, \mathbb{R})_L \times sl(2, \mathbb{R})_R$ presented in [6]. Here we will review the derivation as presented there, adapted appropriately to $so(4)$.

Singlet states of $so(4)$ can be constructed as follows. Consider a group element $U \in \text{SU}(2)$, and define the rotated linear combination

$$Q_a(U) = L_a + D_a^{a'}(U) \bar{L}_{a'}, \quad (3.20)$$

where $D_a^{a'}$ corresponds to the adjoint action of the group; see appendix A for our conventions. We define a state $|U\rangle$ through its annihilation by $Q_a(U)$,

$$Q_a(U) |U\rangle = 0. \quad (3.21)$$

In other words, $|U\rangle$ is a state that is invariant under a linear combination of $so(4)$ generators specified by $Q_a(U)$. This equation is crucial: the inclusion of both copies of $su(2)$ will ensure that the states $|U\rangle$ will prevent a factorization in our observables, and will allow us to interpret our choices in the metric formulation.

There are two interesting choices of $|U\rangle$ for which it is useful to build explicit solutions to (3.21). We refer to our first choice as an *Ishibashi* state: it is defined by selecting a group element $U = \Sigma_{\text{Ish}}$ such that

$$D_k^{k'}(\Sigma_{\text{Ish}}) L_{k'} = \Sigma_{\text{Ish}} L_k \Sigma_{\text{Ish}}^{-1} = -L_{-k}, \quad (3.22)$$

where we are using the basis (3.4), and therefore $k = -, 0, +$. The corresponding group element is

$$\Sigma_{\text{Ish}} = e^{\frac{\pi}{2}(L_+ - L_-)} = e^{i\pi L_2}. \quad (3.23)$$

The corresponding singlet state, i.e., Ishibashi state, is the solution to

$$(L_k - \bar{L}_{-k}) |\Sigma_{\text{Ish}}\rangle = 0. \quad (3.24)$$

This equation has a non-trivial solution for the non-unitary representations built in section 3.1. Consider the basis of states in (3.11)–(3.12) for each copy of $su(2)$ of the form

$$\sum_{p, \bar{p}} a_{p, \bar{p}} |l, p\rangle \otimes |\bar{l}, \bar{p}\rangle, \quad (3.25)$$

with coefficients $a_{p, \bar{p}}$, as an ansatz for $|\Sigma_{\text{Ish}}\rangle$. The $k = 0$ condition in (3.24) sets $l = \bar{l}$, and $k = \pm$ will give $a_{p, \bar{p}} = (-1)^p \delta_{p, \bar{p}}$, up to an overall normalization independent of p . The resulting state is

$$|\Sigma_{\text{Ish}}\rangle = \sum_{p=0}^{\infty} (-1)^p |l, p\rangle, \quad (3.26)$$

where $|l, p, \bar{p}\rangle \equiv |l, p\rangle \otimes |\bar{l}, \bar{p}\rangle$.

The second choice will be coined *crosscap* state. In this instance, we select $U = \Sigma_{\text{cross}}$ such that

$$D_k^{k'}(\Sigma_{\text{cross}}) L_{k'} = \Sigma_{\text{cross}} L_k \Sigma_{\text{cross}}^{-1} = -(-1)^k L_{-k}, \quad (3.27)$$

which leads to the group element

$$\Sigma_{\text{cross}} = e^{\frac{i\pi}{2}(L_+ + L_-)} = e^{i\pi L_1}. \quad (3.28)$$

Using (3.27) in (3.21), the crosscap state satisfies

$$(L_k - (-1)^k \bar{L}_{-k}) |\Sigma_{\text{cross}}\rangle = 0, \quad (3.29)$$

and in terms of the non-unitary $su(2)$ representations the solution to these conditions are

$$|\Sigma_{\text{cross}}\rangle = \sum_{p=0}^{\infty} |l, p\rangle. \quad (3.30)$$

In contrast to the Virasoro construction, it is important to emphasise that here we don't have an interpretation of (3.24) and (3.29) as a boundary condition of an operator in a CFT_2 as in [41, 42]. We are using (and abusing) the nomenclature used there because of the resemblance of (3.24) and (3.29) with the CFT_2 conditions, and its close relation to the $so(2, 2)$ states used in [6]. In this regard, it is useful to highlight some similarities and key differences in $so(4)$ relative to $so(2, 2)$. A similarity is that our choice to use p rather than the eigenvalue of L_0 to label the states in the non-unitary representation was precisely motivated to make the states match with those in $sl(2, \mathbb{R})$. However, one difference is that the group elements (3.23) and (3.28) differ by a factor of i in the exponent compared to their $sl(2, \mathbb{R})$ counterparts in [6]. Also we note that, unlike in the $sl(2, \mathbb{R})$ case, the relative phase in the state now appears in the Ishibashi state rather than the crosscap state. This is due to the extra minus sign in (3.13).

Another important property of the singlet states is their transformation under the action of $SU(2)$ group elements. Consider $G(L) \in SU(2)_L$, and $\bar{G}(R^{-1}) \in SU(2)_R$ for each copy appearing in $SO(4)$. A simple manipulation shows that

$$G(L)\bar{G}(R^{-1})Q_a(U)|U\rangle = D_a^{a'}(L^{-1})Q'_a(LUR)G(L)\bar{G}(R^{-1})|U\rangle = 0. \quad (3.31)$$

Thus we have

$$G(L)\bar{G}(R^{-1})|U\rangle = |LUR\rangle. \quad (3.32)$$

This identity will be used heavily in the following derivations.

3.2 Wilson line and the Green's function

We now come back to evaluating the Wilson line (3.2). We select as endpoints states

$$|U_i\rangle = |U_f\rangle = |\Sigma\rangle, \quad (3.33)$$

with the choice of $|\Sigma\rangle$ being either

$$|\Sigma_{\text{Ish}}\rangle \text{ or } |\Sigma_{\text{cross}}\rangle, \quad (3.34)$$

i.e., the Ishibashi (3.26) or crosscap (3.30) state. From this perspective we can view (3.2) as

$$W_{\mathcal{R}}(x_i, x_f) = \langle \Sigma | G(L)\bar{G}(R^{-1}) | \Sigma \rangle, \quad (3.35)$$

where we identify

$$G(L) = \mathcal{P}\exp\left(-\int_{\gamma} A\right), \quad \bar{G}(R^{-1}) = \mathcal{P}\exp\left(-\int_{\gamma} \bar{A}\right). \quad (3.36)$$

Given the properties of our singlet states, we can easily evaluate (3.35) as follows,

$$\begin{aligned} W_{\mathcal{R}}(x_i, x_f) &= \langle \Sigma | G(L)\bar{G}(R^{-1}) | \Sigma \rangle \\ &= \langle \Sigma | G(L\tilde{R}) | \Sigma \rangle \\ &= \sum_{p=0}^{\infty} \langle l, p | G(L\tilde{R}) | l, p \rangle \\ &= \frac{e^{i\alpha(l+1)}}{e^{i\alpha} - 1}. \end{aligned} \quad (3.37)$$

In the second line we used (3.32) to move the right group element R to the left, where

$$\tilde{R} \equiv \Sigma R \Sigma^{-1}. \quad (3.38)$$

To obtain the third line in (3.37) we use the explicit form of the states given by (3.26) and (3.30), where both the Ishibashi and cross cap state report the same answer. Finally in the last equality we used the formula for the character in (3.17), where α in this case is defined via the equation

$$L\Sigma R\Sigma^{-1} = V^{-1}e^{i\alpha L_0}V, \quad (3.39)$$

i.e., assuming we can diagonalise the left hand side, α captures the eigenvalue of the group element in the inner product.

The interpretation of (3.37) in the metric formulation of dS₃ gravity is interesting. First, we observe that for a pair of $su(2)$ Chern-Simons connections,

$$A = g_L dg_L^{-1}, \quad \bar{A} = g_R^{-1} dg_R, \quad (3.40)$$

we have

$$G(L) = g_L(x_f) g_L(x_i)^{-1}, \quad \bar{G}(R^{-1}) = g_R(x_f)^{-1} g_R(x_i), \quad (3.41)$$

where we evaluated the path ordered integral for a path γ with endpoints (x_i, x_f) . For concreteness, we will make the choice

$$g_L = e^{-irL_2} e^{-i(\phi+\tau)L_3}, \quad \tilde{g}_R \equiv \Sigma g_R \Sigma^{-1} = e^{i(\phi-\tau)L_3} e^{-irL_2}, \quad (3.42)$$

which for $SU(2)_L$ is the group element associated to S^3 in (2.17). But it is important to stress that, with some insight, we are specifying $\tilde{g}_R$ rather than g_R , since this is all we need at this stage to evaluate $W_{\mathcal{R}}(x_i, x_f)$. Using (3.42) we find that the solution for α in (3.39) is

$$\cos\left(\frac{\alpha}{2}\right) = \cos(r_f) \cos(r_i) \cos(\tau_f - \tau_i) + \sin(r_f) \sin(r_i) \cos(\phi_f - \phi_i). \quad (3.43)$$

α , which labels the equivalence class of $L\Sigma R\Sigma^{-1}$, can then be related to the geodesic distance between points (x_i, x_f) on S^3 (see (B.31)):

$$\alpha = \pm 2\Theta + 4\pi n, \quad n \in \mathbb{Z}, \quad (3.44)$$

with n accounting for winding. As explained in appendix B.2, the propagator of a scalar field of mass m in dS₃ can be written as

$$G(\Theta) = \mathcal{G}_h(\Theta) + \mathcal{G}_{1-h}(\Theta),$$

$$\mathcal{G}_h(\Theta) = a_h \frac{e^{-2ih\Theta}}{e^{-2i\Theta} - 1}, \quad (3.45)$$

with

$$a_h = \frac{i}{2\pi\ell} \frac{1}{1 - e^{-4\pi i h}}, \quad h = \frac{1 + \sqrt{1 - (m\ell)^2}}{2}. \quad (3.46)$$

Equations (3.37) and (3.45) lead us to conclude that if we pick a representation $\mathcal{R} = \mathcal{R}_+$ in (3.18) with $l = -h$ then

$$W_{\mathcal{R}_+}(x_i, x_f) = \frac{1}{a_h} \mathcal{G}_h(\Theta). \quad (3.47)$$

Similarly, picking instead a representation $\mathcal{R} = \mathcal{R}_-$ in (3.18), where now $l = h - 1$, leads to

$$W_{\mathcal{R}_-}(x_i, x_f) = \frac{1}{a_{1-h}} \mathcal{G}_{1-h}(\Theta). \quad (3.48)$$

The full propagator can then be written as

$$G(\Theta) = a_h W_{\mathcal{R}_+}(x_i, x_f) + a_{1-h} W_{\mathcal{R}_-}(x_i, x_f). \quad (3.49)$$

$\mathcal{R}_\pm$ are the two possible representations with the same Casimir $c_2 = h(h-1) = -\frac{m^2 \ell^2}{4}$. We emphasize that, unlike in AdS_3 , we need to consider both of these representations to obtain the correct propagator. This is related to the fact that the de Sitter propagator is not simply given by the analytic continuation from AdS_3 due to differences in causal structures [43, 44].

Moving away from the specificity of group elements (3.42), for any pair of flat connections (3.40), we will have that $W_{\mathcal{R}_\pm}(x_i, x_f)$ gives the $\mathcal{G}_{h,1-h}(\Theta)$ contribution to the Green's function between the points (x_i, x_f) in the Euclidean space with metric

$$g_{\mu\nu} = -\frac{\ell^2}{2} \text{Tr} \left[(A_\mu - \Sigma \bar{A}_\mu \Sigma^{-1}) (A_\nu - \Sigma \bar{A}_\nu \Sigma^{-1}) \right]. \quad (3.50)$$

A proof of this statement, beyond the explicit computation done here for S^3 , follows step by step the derivations in [6] for $so(2, 2)$ adapted to $so(4)$. The geometric role of our singlet states is now more clear: Σ is the group element that controls how the right connection $\bar{A}$ acts as a left element relative to A , and vice-versa. These derivations also establish the gravitational Wilson line as a local probe of the Euclidean dS_3 geometry, and hence will allow us to investigate notions of locality in the Chern-Simons formulation of gravity.

4 Local pseudofields from Wilson lines

The aim of this section is to further extract local quantities from the gravitational Wilson line. We will focus on the background connections associated to the round 3-sphere for concreteness, and show how to build local pseudofields from the singlet states used in the previous section. We use the term “pseudofields” because while the objects we will build from a single irreducible representation $\mathcal{R}$ (either $\mathcal{R}_+$ or $\mathcal{R}_-$) are local, and behave in many ways like fields, both representations are needed to form a complete basis for local fields in dS_3 .

4.1 Wilson line as an overlap of states

Until now, we have described the Wilson line $W_{\mathcal{R}}(x_i, x_f)$ as the diagonal matrix element of an operator in a singlet state, as done in (3.35). For the purpose of building local probes, we want to rewrite this operator as a suitable overlap between states. From (3.41) we can write (3.35) as

$$W_{\mathcal{R}}(x_i, x_f) = \langle \Sigma | \bar{G}(g_R(x_f)^{-1}) G(g_L(x_f)) G(g_L(x_i)^{-1}) \bar{G}(g_R(x_i)) | \Sigma \rangle. \quad (4.1)$$

If our representation $\mathcal{R}$ used Hermitian generators, we would simply note that for unitary group elements, i.e.,

$$g_R^{-1} = g_R^\dagger, \quad g_L^{-1} = g_L^\dagger, \quad (4.2)$$

we would have $W_{\mathcal{R}}(x_i, x_f) = \langle U(x_f) | U(x_i) \rangle$ with $|U(x)\rangle = G(g_L(x)^{-1}) \bar{G}(g_R(x)) | \Sigma \rangle$. However, our representation is *non*-unitary, and hence these manipulations require some care.

Define the following state:

$$|U(x)\rangle = G(g_L(x)^{-1}) \bar{G}(g_R(x)) | \Sigma \rangle. \quad (4.3)$$

We will focus exclusively on the background introduced in (2.17). Because the representation we are using is non-unitary, we have

$$g_L(\tau, r, \phi)^\dagger = g_L(\tau, -r, \phi)^{-1} = g_L(\tau, r, \phi + \pi)^{-1} e^{i\pi L_3}, \quad (4.4)$$

and the same relation for g_R , which allow us to write the Wilson line as

$$W_{\mathcal{R}}(x_i, x_f) = \langle U(\tau_f, r_f, \phi_f + \pi) | U(\tau_i, r_i, \phi_i) \rangle. \quad (4.5)$$

In this equality we used

$$e^{i\pi L_3} e^{i\pi \bar{L}_3} |\Sigma\rangle \sim |\Sigma\rangle, \quad (4.6)$$

since both singlet states are annihilated by $Q_a(\Sigma)$.

4.2 Construction of local basis

Having written $W_{\mathcal{R}}(x_i, x_f)$ as an overlap of states, we now can start the process of defining a local pseudofield from $|U(x)\rangle$. The most natural way to split (4.5) is as done in (4.3). Still this has its inherent ambiguities: in defining $|U(x)\rangle$ we are splitting the cutting curve $\gamma(s)$ at some midpoint x_0 , the choice of which is a gauge freedom at our disposal. More concretely, a general definition of the state should be

$$|U(x)\rangle = G(g_L(x_0)g_L(x)^{-1})\bar{G}(g_R(x_0)^{-1}g_R(x))|\Sigma\rangle \quad (4.7)$$

where we restored the dependence on this midpoint split. At this stage it is not clear to us that one choice of $g_{L,R}(x_0)$ is better than any other, so for sake of simplicity we will select $g_{L,R}(x_0) = \mathbb{1}$, i.e. the identity element. Therefore we will be working with (4.3), and explore the local properties of $|U(x)\rangle$.

First, we expand $|U(x)\rangle$ in the eigenstate $|l, p, \bar{p}\rangle$ basis:

$$|U(x)\rangle = \sum_{p, \bar{p}=0}^{\infty} \Phi_{p, \bar{p}}^*(x) |l, p, \bar{p}\rangle, \quad (4.8)$$

which we can reverse as

$$\Phi_{p, \bar{p}}(x) = \langle U(x) | l, p, \bar{p} \rangle. \quad (4.9)$$

$\Phi_{p, \bar{p}}(x)$ will be our basis of local pseudofields that will support the local properties in $|U(x)\rangle$. To build this basis of eigenfunctions, we can translate the action of the generators L_a on the basis vectors into the action of differential operators ζ_a acting on $\Phi_{p, \bar{p}}$. Specifically, we will find $\zeta_a, \bar{\zeta}_a$ such that

$$\langle U(x) | L_a | l, p, \bar{p} \rangle = \zeta_a \langle U(x) | l, p, \bar{p} \rangle, \quad (4.10)$$

$$\langle U(x) | \bar{L}_a | l, p, \bar{p} \rangle = \bar{\zeta}_a \langle U(x) | l, p, \bar{p} \rangle. \quad (4.11)$$

Using (3.13) and (3.14), the differential operators must therefore satisfy

$$\zeta_+ \Phi_{p, \bar{p}} = -\sqrt{p(p-2l-1)} \Phi_{p-1, \bar{p}} \quad (4.12)$$

$$\zeta_- \Phi_{p, \bar{p}} = \sqrt{(p+1)(p-2l)} \Phi_{p+1, \bar{p}} \quad (4.13)$$

$$\zeta_0 \Phi_{p, \bar{p}} = (l-p) \Phi_{p, \bar{p}}, \quad (4.14)$$

and similarly for the barred sector. It follows that $\Phi_{p,\bar{p}}$ satisfies the Casimir equation,

$$(\nabla^2 + \bar{\nabla}^2) \Phi_{p,\bar{p}}(x) = 2l(l+1)\Phi_{p,\bar{p}}(x), \quad (4.15)$$

where $\nabla^2 = \delta^{ab}\zeta_a\zeta_b$, and $\bar{\nabla}^2 = \delta^{ab}\bar{\zeta}_a\bar{\zeta}_b$.⁶ Our strategy will be to build the differential operators for $(\zeta_a, \bar{\zeta}_a)$ based on (4.10)–(4.11), and then solve for $\Phi_{p,\bar{p}}(x)$ from the differential equations (4.12)–(4.15).

We will start by building the generators ζ_a for Euclidean dS_3 . It is convenient to cast the state in (4.3) as

$$\begin{aligned} |U(x)\rangle &= G(g_L(x)^{-1})\bar{G}(g_R(x))|\Sigma\rangle \\ &= G(g_L(x)^{-1}\tilde{g}_R(x)^{-1})|\Sigma\rangle \\ &= e^{i(\phi+\tau)L_3}e^{2irL_2}e^{-i(\phi-\tau)L_3}|\Sigma\rangle. \end{aligned} \quad (4.16)$$

In the second line we moved all the group elements to left, as in (3.37), and in the third line we used (3.42). Next, consider the action of partial derivatives on $\Phi_{p\bar{p}}(x) = \langle U(x)|l, p, \bar{p}\rangle$:

$$\begin{aligned} \partial_+ \langle U(x)|l, p, \bar{p}\rangle &= -i\langle U(x)|L_3|l, p, \bar{p}\rangle, \\ \partial_- \langle U(x)|l, p, \bar{p}\rangle &= i\cos(2r)\langle U(x)|L_3|l, p, \bar{p}\rangle + i\sin(2r)\cos(\theta_+)\langle U(x)|L_1|l, p, \bar{p}\rangle \\ &\quad - i\sin(2r)\sin(\theta_+)\langle U(x)|L_2|l, p, \bar{p}\rangle, \\ \partial_r \langle U(x)|l, p, \bar{p}\rangle &= 2i\cos(\theta_+)\langle U(x)|L_2|l, p, \bar{p}\rangle + 2i\sin(\theta_+)\langle U(x)|L_1|l, p, \bar{p}\rangle, \end{aligned} \quad (4.17)$$

where we introduced the coordinates

$$\theta_{\pm} = \phi \pm \tau, \quad \partial_{\pm} = \frac{\partial}{\partial\theta_{\pm}}. \quad (4.18)$$

Inverting the relationship between $\partial_a \langle U(x)|l, p, \bar{p}\rangle$ and $\langle U(x)|L_a|l, p, \bar{p}\rangle$ leads to

$$\begin{aligned} \zeta_1 &= -i\frac{\cos\theta_+}{\sin(2r)}(\partial_- + \cos(2r)\partial_+) - \frac{i}{2}\sin\theta_+\partial_r, \\ \zeta_2 &= i\frac{\sin\theta_+}{\sin(2r)}(\partial_- + \cos(2r)\partial_+) - \frac{i}{2}\cos\theta_+\partial_r, \\ \zeta_3 &= i\partial_+, \end{aligned} \quad (4.19)$$

or, in terms of $\zeta_{\pm} = \zeta_1 \pm i\zeta_2$,

$$\zeta_{\pm} = -ie^{\mp i\theta_+}(\csc(2r)\partial_- + \cot(2r)\partial_+) \pm \frac{1}{2}e^{\mp i\theta_+}\partial_r, \quad (4.20)$$

and $\zeta_0 = \zeta_3$. These are simply three of the Killing vectors for S^3 , which together satisfy one copy of the $su(2)$ algebra.

To do the equivalent calculation for the barred sector, we should instead write

$$\begin{aligned} |U(x)\rangle &= G(g_L^{-1}(x))\bar{G}(g_R(x))|\Sigma\rangle \\ &= \bar{G}[g_R(x)\Sigma^{-1}g_L(x)\Sigma]|\Sigma\rangle \\ &= \bar{G}[\Sigma^{-1}\tilde{g}_R(x)g_L(x)\Sigma]|\Sigma\rangle \\ &= \Sigma^{-1}e^{i\theta_-}\bar{L}_3e^{-2ir\bar{L}_2}e^{-i\theta_+}\bar{L}_3\Sigma|\Sigma\rangle. \end{aligned} \quad (4.21)$$

⁶We will find that $\nabla^2 + \bar{\nabla}^2 = -\frac{1}{2}\nabla_{S^3}^2$, where $\nabla_{S^3}^2$ is the ordinary Laplacian for EdS_3 .

This, after all, is the purpose of our definition of Σ : it lets us intertwine the two copies of $su(2)$. Therefore, the exact action of Σ on group elements will affect the result of this calculation. We have two choices of Σ , given in (3.22) and (3.27),

$$\Sigma_{\text{cross}} = e^{i\pi\bar{L}_1}, \quad \Sigma_{\text{Ish}} = e^{i\pi\bar{L}_2}. \quad (4.22)$$

Working out the effect of the Ishibashi state in (4.21) we find

$$\Sigma_{\text{Ish}}^{-1} e^{i\theta_- \bar{L}_3} e^{-2ir\bar{L}_2} e^{-i\theta_+ \bar{L}_3} \Sigma_{\text{Ish}} = e^{-i\theta_- \bar{L}_3} e^{-2ir\bar{L}_2} e^{i\theta_+ \bar{L}_3}, \quad (4.23)$$

in other words conjugation by Σ_{Ish} flips $\theta_{\pm} \rightarrow -\theta_{\pm}$ while leaving r fixed. For the crosscap state we instead find

$$\Sigma_{\text{cross}}^{-1} e^{i\theta_- \bar{L}_3} e^{-2ir\bar{L}_2} e^{-i\theta_+ \bar{L}_3} \Sigma_{\text{cross}} = e^{-i\theta_- \bar{L}_3} e^{2ir\bar{L}_2} e^{i\theta_+ \bar{L}_3}, \quad (4.24)$$

so that conjugation by Σ_{cross} flips $\theta_{\pm} \rightarrow -\theta_{\pm}$ and in addition $r \rightarrow -r$. From here on, the calculation to build $\bar{\zeta}_a$ is very similar to the unbarred case, but there will be differences depending on the choice of Σ . First, solving (5.29), for $\Sigma = \Sigma_{\text{Ish}}$ we find

$$\begin{aligned} \bar{\zeta}_1 &= -i \frac{\cos \theta_-}{\sin 2r} (\partial_+ + \cos 2r \partial_-) - \frac{i}{2} \sin \theta_- \partial_r, \\ \bar{\zeta}_2 &= -i \frac{\sin \theta_-}{\sin 2r} (\partial_+ + \cos 2r \partial_-) + \frac{i}{2} \cos \theta_- \partial_r, \\ \bar{\zeta}_3 &= -i \partial_-, \end{aligned} \quad (4.25)$$

or in terms of $\bar{\zeta}_{\pm} = \bar{\zeta}_1 \pm i\bar{\zeta}_2$,

$$\bar{\zeta}_{\pm} = -ie^{\pm i\theta_-} (\csc 2r \partial_+ + \cot 2r \partial_-) \mp \frac{1}{2} e^{\pm i\theta_-} \partial_r, \quad (4.26)$$

and $\bar{\zeta}_0 = \bar{\zeta}_3$. These are the three additional Killing vectors for S^3 , which are related to (4.19) by the replacement $\theta_{\pm} \rightarrow -\theta_{\mp}$ and $r \rightarrow -r$. Together the generators ζ_a satisfy the $su(2)_L$ algebra, while $\bar{\zeta}_a$ correspond to the generators of the second $su(2)_R$. Selecting $\Sigma = \Sigma_{\text{cross}}$ is not dramatically different: we will again obtain (4.25) with $r \rightarrow -r$, and that flips the overall sign in $\bar{\zeta}_{1,2}$. Hence we will again find the second copy of Killing vectors obeying $su(2)_R$; the difference at this stage between the two singlet states is an orientation of r that does not affect the interpretation of $(\zeta_a, \bar{\zeta}_a)$ as the six Killing vectors for S^3 .

Now we would like to find explicit expressions for $\Phi_{p,\bar{p}}$. The procedure for either Σ_{Ish} or Σ_{cross} would produce the same special functions, with the difference being an overall normalization that depends on $(p, \bar{p})$. For concreteness we will just focus on Σ_{Ish} .

We can construct the pseudofields by first solving for a highest weight state $\Phi_{0,0}$, and then acting with $(\zeta_-)^p$ and $(\bar{\zeta}_-)^{\bar{p}}$ on this solution to generate $\Phi_{p,\bar{p}}$. This will give a position-space representation of our abstract states $|l, p, \bar{p}\rangle$. The highest weight state satisfies

$$\zeta_3 \Phi_{0,0} = \bar{\zeta}_3 \Phi_{0,0} = l \Phi_{0,0}, \quad (4.27)$$

$$\zeta_+ \Phi_{0,0} = \bar{\zeta}_+ \Phi_{0,0} = 0. \quad (4.28)$$

These equations are solved by

$$\Phi_{0,0}(r, \tau, \phi) = \langle U(x) | l, 0, 0 \rangle = e^{-2il\tau} \cos^{2l}(r). \quad (4.29)$$

The descendant states are then given by

$$\begin{aligned} \Phi_{p,\bar{p}}(r, \tau, \phi) &= c_{p\bar{p}} e^{-2il\tau} \cos^{2l}(r) e^{i(p\theta_+ - \bar{p}\theta_-)} \tan^{\bar{p}-p}(r) P_p^{\bar{p}-p, -(2l+1)}(1 + 2\tan^2(r)), \\ c_{p\bar{p}} &= (-1)^p \sqrt{\frac{p!(\bar{p} - (2l+1))!}{\bar{p}!(p - (2l+1))!}}, \end{aligned} \quad (4.30)$$

where here $P_n^{\alpha,\beta}(x)$ is a Jacobi polynomial. These satisfy (4.12)–(4.14) and their barred analogues.

4.2.1 Wavefunction for the singlet states

Where does our singlet state $|\Sigma\rangle$ sit on S^3 ? This question is ambiguous, since the answer depends on a choice of gauge. In the context of the discussion presented here, positions will depend on how one selects the midpoint in (4.7). Still it is instructive to answer it for the simple purpose of illustrating what our prior choices imply.

Consider first the Ishibashi state $|\Sigma_{\text{Ish}}\rangle$. To see the position of this state in S^3 , it is very clear that at $r = 0$, we have

$$\Phi_{p,\bar{p}}(\tau, r = 0, \phi) = (-1)^p e^{-2i\tau(l-p)} \delta_{p,\bar{p}}, \quad (4.31)$$

which follows from (4.30). This is to be expected since $p \neq \bar{p}$ introduces a ϕ dependence which we know is absent at $r = 0$. Therefore, we can write

$$|U(\tau, r = 0)\rangle = \sum_p (-1)^p e^{2i\tau(l-p)} |l, p, p\rangle, \quad (4.32)$$

which at $\tau = 0$ is simply the Ishibashi state (3.26). Thus we see that our Ishibashi state lives at $(r = 0, \tau = 0)$. If we had constructed a basis of $\Phi_{p,\bar{p}}$ from the $(\zeta_a, \bar{\zeta}_a)$ obtained from the crosscap states rather than the Ishibashi states, we would have seen that the crosscap state sits at $(r = 0, \tau = 0)$.

The wave function we would attribute to the Ishibashi state can also be explicitly calculated:

$$\begin{aligned} \langle \Sigma_{\text{Ish}} | U(x) \rangle &= \frac{(\cos(\Theta_{\text{NPole}}) - i \sin(\Theta_{\text{NPole}}))^{2l+1}}{2i \sin(\Theta_{\text{NPole}})} \\ &= \frac{e^{-2il\Theta_{\text{NPole}}}}{1 - e^{2i\Theta_{\text{NPole}}}}, \end{aligned} \quad (4.33)$$

where Θ_{NPole} is the geodesic distance (B.31) between x and $r' = 0, \tau' = 0$ —the North Pole of the three-sphere.⁷

Still we stress that the values of τ and r are somewhat artificial. For instance, in (4.32) the crosscap state can be seen to be related to the Ishibashi state by a simple shift in τ . This is a reflection of the fact that there is considerable gauge freedom in how we describe solutions.

⁷This corresponds to the North Pole of the S^2 time slices for Euclidean time $\tau' = 0$. It is a point on Penrose diagrams, not a line.

4.2.2 Wick rotation and quasi-normal modes

Before proceeding to discuss $SL(2, \mathbb{C})$ Chern-Simons theory, i.e. the Lorentzian formulation of dS_3 gravity, it is instructive to interpret our Euclidean results in Lorentzian signature. We will simply now use a Wick rotation of the metric formulation to provide a first interpretation of our results. As described in appendix B, the metric analytic continuation is implemented by taking

$$t \rightarrow -i\ell\tau. \quad (4.34)$$

The Wick-rotated $\Phi_{p,\bar{p}}$ in (4.30) are therefore

$$\begin{aligned} \Phi_{p,\bar{p}}(r, t, \phi) &= c_{p\bar{p}} e^{i\ell(\bar{z}-z)} \cos^{2l}(r) e^{i(pz-\bar{p}\bar{z})} \tan^{\bar{p}-p}(r) P_p^{\bar{p}-p, -(2l+1)}(1 + 2 \tan^2(r)), \\ c_{p\bar{p}} &= (-1)^p \sqrt{\frac{p!(\bar{p} - (2l+1))!}{\bar{p}!(p - (2l+1))!}}, \end{aligned} \quad (4.35)$$

with $z \equiv \phi + it$, and $\bar{z} \equiv \phi - it$. In terms of the more familiar hypergeometric functions and radial coordinate $u \equiv \sin(r)$, we have (using that $\Phi_{p,\bar{p}} = e^{2i(p-\bar{p})\phi} \Phi_{\bar{p},p}$):

$$\begin{aligned} \Phi_{\omega,k}(r, t, \phi) &= c_{p\bar{p}} \left(\frac{\frac{\omega+|k|}{2} + l}{\frac{\omega-|k|}{2} + l} \right) (1-u^2)^{-\omega/2} u^{|k|} e^{-ik\phi} e^{-\omega t} {}_2F_1 \left(\frac{|k|-\omega}{2} - l, \frac{|k|-\omega}{2} + l + 1; |k| + 1; u^2 \right), \\ \omega &= p + \bar{p} - 2l > 0, \quad k = \bar{p} - p. \end{aligned} \quad (4.36)$$

Note that instead of oscillating in time, these functions are now purely decaying. In fact, the $\Phi_{\omega,k}$ are exactly (up to normalization) the quasi-normal modes of dS_3 [39, 45]. As discussed in section 3.2, given a scalar field of mass m , there are two representations $\mathcal{R}_\pm$ that have the same Casimir: one with $l = -h$ and one with $l = h - 1$. These two representations have different characters (and thus Wilson lines), and both are needed to obtain the full Green's function: $G(\Theta) = a_h W_{\mathcal{R}_+} + a_{1-h} W_{\mathcal{R}_-}$. Each choice of l matches one of the two distinct sequences of quasi-normal modes in dS_3 . This reinforces the idea that both representations are needed to describe a bulk scalar field.

The Wick rotation can also be used to simply obtain Lorentzian Killing vectors from (4.19) and (4.25). These can then be re-organized in an $sl(2, \mathbb{C})$ representation in the following way:

$$\begin{aligned} -i\zeta_1 &\xrightarrow{\tau \rightarrow it/\ell} \mathcal{H}_1 = -\frac{\cos z}{\sin 2r} (\bar{\partial} + \cos 2r \partial) - \frac{1}{2} \sin z \partial_r, \\ -i\zeta_2 &\xrightarrow{\tau \rightarrow it/\ell} \mathcal{H}_2 = \frac{\sin z}{\sin 2r} (\bar{\partial} + \cos 2r \partial) - \frac{1}{2} \cos z \partial_r, \\ \zeta_3 &\xrightarrow{\tau \rightarrow it/\ell} \mathcal{H}_3 = i\partial, \\ i\bar{\zeta}_1 &\xrightarrow{\tau \rightarrow it/\ell} \bar{\mathcal{H}}_1 = \frac{\cos \bar{z}}{\sin 2r} (\partial + \cos 2r \bar{\partial}) + \frac{1}{2} \sin \bar{z} \partial_r, \\ -i\bar{\zeta}_2 &\xrightarrow{\tau \rightarrow it/\ell} \bar{\mathcal{H}}_2 = -\frac{\sin \bar{z}}{\sin 2r} (\partial + \cos 2r \bar{\partial}) + \frac{1}{2} \cos \bar{z} \partial_r, \\ -\bar{\zeta}_3 &\xrightarrow{\tau \rightarrow it/\ell} \bar{\mathcal{H}}_3 = i\bar{\partial}. \end{aligned} \quad (4.37)$$

The operators $(\mathcal{H}_a, \bar{\mathcal{H}}_a)$ have been normalised such that they form an $sl(2, \mathbb{R}) \times sl(2, \mathbb{R})$ algebra. More importantly, these operators have a simple action on the quasinormal modes. We can see this explicitly by reorganizing the operators into the combinations

$$\begin{aligned}\mathcal{H}_0 &= -\mathcal{H}_3, & \mathcal{H}_\pm &= \mathcal{H}_2 \mp i\mathcal{H}_1, \\ \bar{\mathcal{H}}_0 &= \bar{\mathcal{H}}_3, & \bar{\mathcal{H}}_\pm &= \bar{\mathcal{H}}_2 \pm i\bar{\mathcal{H}}_1.\end{aligned}\tag{4.38}$$

The quasinormal mode Φ_{00} is a highest weight state of our representation,

$$\mathcal{H}_+ \Phi_{00} = 0,\tag{4.39}$$

while the rest of the quasinormal modes obey

$$\begin{aligned}\mathcal{H}_0 \Phi_{p\bar{p}} &= (h + p) \Phi_{p,\bar{p}}, \\ \mathcal{H}_+ \Phi_{p\bar{p}} &= \sqrt{p(p + 2h + 1)} \Phi_{p-1,\bar{p}}, \\ \mathcal{H}_- \Phi_{p\bar{p}} &= \sqrt{(p + 1)(p + 2h)} \Phi_{p+1,\bar{p}},\end{aligned}\tag{4.40}$$

and similarly for the barred sector. In this expression we have $h = -l$,⁸ and hence the modes $\Phi_{p,\bar{p}}$ characterize a highest weight representations of $sl(2)$ with Casimir $h(h - 1)$. Furthermore, the (anti-)Hermitian properties of the $su(2)$ generators $L_{0,\pm}$ in (3.5) combined with the map in (4.37), dictate that the generators $\mathcal{H}_{0,\pm}$ have the usual Hermiticity properties. This makes the representations unitary when organized in terms of the $sl(2, \mathbb{R})$ basis.

The Wick rotation gives an interpretation for the algebraic structure of the quasinormal mode spectrum of the static patch. Our construction resonates with [39], where it was noticed that the quasinormal modes had a “hidden” $SL(2, \mathbb{R})$ symmetry, but the origin of this remained mysterious. A similar result was found in [46].

Finally, the quasinormal modes additionally satisfy the Casimir equation for our representations,

$$(\nabla^2 + \bar{\nabla}^2) \Phi_{p,\bar{p}}(x) = 2h(h - 1) \Phi_{p,\bar{p}}(x),\tag{4.41}$$

where $\nabla^2 = -\eta^{ab} \mathcal{H}_a \mathcal{H}_b$, and $\bar{\nabla}^2 = -\eta^{ab} \bar{\mathcal{H}}_a \bar{\mathcal{H}}_b$, so that $\nabla^2 + \bar{\nabla}^2 = -\frac{1}{2} \nabla_{dS_3}^2$ is the d’Alembertian on Lorentzian dS_3 . With the insight of the Wick rotation, the representation (4.40) will be our focus in the subsequent section as we study $SL(2, \mathbb{C})$ Chern-Simons theory.

5 Wilson lines in $SL(2, \mathbb{C})$ Chern-Simons

Everything we have discussed so far has been based on Euclidean dS_3 . In this section, we discuss how our construction can be translated to Lorentzian signature, guided by the properties of our representation under analytic continuation. Based on the Euclidean analysis, we will select a suitable representation of $sl(2, \mathbb{C})$, and implement this choice for the inflationary patch of dS_3 .

⁸We are focusing here on $\mathcal{R}_+$ for notational simplicity. Analogous results with $h \rightarrow 1 - h$ can be obtained for $\mathcal{R}_-$, which has $l = h - 1$.

5.1 Chern-Simons formulation of Lorentzian dS₃ gravity

We start from $SL(2, \mathbb{C})$ Chern-Simons theory with action

$$S_{CS}[\mathcal{A}] = \frac{is}{4\pi} \int \text{Tr} \left(\mathcal{A} \wedge d\mathcal{A} + \frac{2}{3} \mathcal{A} \wedge \mathcal{A} \wedge \mathcal{A} \right) - \frac{is}{4\pi} \int \text{Tr} \left(\bar{\mathcal{A}} \wedge d\bar{\mathcal{A}} + \frac{2}{3} \bar{\mathcal{A}} \wedge \bar{\mathcal{A}} \wedge \bar{\mathcal{A}} \right), \quad (5.1)$$

with $\mathcal{A}, \bar{\mathcal{A}} \in sl(2, \mathbb{C})$, and complex parameter s . The relation of (5.1) to Lorentzian dS₃ gravity was done in [47], and more recent discussions include [37, 48–50]. To build this gravitational interpretation, we expand the gauge fields over the generators $\mathcal{L}_a, \bar{\mathcal{L}}_a$ of $sl(2, \mathbb{C})$ as

$$\mathcal{A} = - \left(i\omega^a + \frac{1}{\ell} e^a \right) \mathcal{L}_a, \quad \bar{\mathcal{A}} = - \left(i\omega^a - \frac{1}{\ell} e^a \right) \bar{\mathcal{L}}_a, \quad (5.2)$$

where the $sl(2, \mathbb{C})$ generators can be related to the generators of $so(1, 3)$ isometries as

$$\mathcal{L}_a = \frac{i}{2} (\mathcal{J}_a + i\ell \mathcal{P}_a), \quad \bar{\mathcal{L}}_a = \frac{i}{2} (\mathcal{J}_a - i\ell \mathcal{P}_a). \quad (5.3)$$

They satisfy the algebra

$$\begin{aligned} [\mathcal{L}_a, \mathcal{L}_b] &= i\epsilon_{abc} \mathcal{L}^c, \\ [\bar{\mathcal{L}}_a, \bar{\mathcal{L}}_b] &= i\epsilon_{abc} \bar{\mathcal{L}}^c, \\ [\mathcal{L}_a, \bar{\mathcal{L}}_b] &= 0, \end{aligned} \quad (5.4)$$

with indices raised by η^{ab} , and we take the convention that $\eta^{11} = \eta^{22} = +1$ and $\eta^{33} = -1$. The trace is taken with the bilinear form

$$\text{Tr}(\mathcal{L}_a \mathcal{L}_b) = \text{Tr}(\bar{\mathcal{L}}_a \bar{\mathcal{L}}_b) = -\frac{1}{2} \eta_{ab}. \quad (5.5)$$

Using (5.2), the action (5.1) becomes

$$S_{EH} = \frac{s}{2\pi\ell} \int_{\mathcal{M}} \left[e^a \wedge \left(d\omega_a + \frac{1}{2} \epsilon_{abc} \omega^b \wedge \omega^c \right) - \frac{1}{6\ell^2} \epsilon_{abc} e^a \wedge e^b \wedge e^c \right]. \quad (5.6)$$

This reduces to the Einstein-Hilbert action with positive cosmological constant given the identification

$$s = \frac{\ell}{4G_3} \in \mathbb{R}. \quad (5.7)$$

It is important to note that $\mathcal{A}$ and $\bar{\mathcal{A}}$ are not independent variables. They are related by complex conjugation, and this relation depends on how we choose to relate $\mathcal{L}_a$ to $\bar{\mathcal{L}}_a$. For now it suffices to demand (5.5), which assures reality of the action (5.1), and we will constrain further the representation as we construct the appropriate probes.

5.2 Construction of probes in $sl(2, \mathbb{C})$

As in the Euclidean case, we would like to build probes in $SL(2, \mathbb{C})$ Chern-Simons theory via the Wilson line operator (3.2). The most natural choice is to simply implement the discrete highest weight representation we inferred in section 4.2.2 from the Euclidean theory. For a further motivation of this choice using an analytic continuation of the $SO(4)$ and $SL(2, \mathbb{C})$

Chern-Simons theories, see appendix C. In the language of the $SL(2, \mathbb{C})$ Chern-Simons, we will build this representation by using the $sl(2)$ generators⁹

$$\mathcal{L}_0 = -\mathcal{L}_3, \quad \mathcal{L}_\pm = \mathcal{L}_2 \mp i\mathcal{L}_1, \quad (5.8)$$

with algebra

$$[\mathcal{L}_0, \mathcal{L}_\pm] = \mp \mathcal{L}_\pm, \quad [\mathcal{L}_+, \mathcal{L}_-] = 2\mathcal{L}_0. \quad (5.9)$$

The highest weight representation in this basis satisfies

$$\begin{aligned} \mathcal{L}_0|h, p\rangle &= (h+p)|h, p\rangle, \\ \mathcal{L}_+|h, p\rangle &= \sqrt{p(p+2h+1)}|h, p-1\rangle, \\ \mathcal{L}_-|h, p\rangle &= \sqrt{(p+1)(p+2h)}|h, p+1\rangle, \end{aligned} \quad (5.10)$$

where p is a positive integer. For now, we take h to be a real parameter that controls the Casimir of the representation

$$\begin{aligned} -\eta^{ab}\mathcal{L}_a\mathcal{L}_b|h, p\rangle &= (\mathcal{L}_0^2 - \mathcal{L}_+\mathcal{L}_- - \mathcal{L}_-\mathcal{L}_+)|h, p\rangle \\ &= h(h-1)|h, p\rangle. \end{aligned} \quad (5.11)$$

Of course, we anticipate that this parameter will match $h = \frac{1+\sqrt{1-(m\ell)^2}}{2}$ (or the other solution which gives the same Casimir). In addition we demand the operators satisfy $\mathcal{L}_0^\dagger = \mathcal{L}_0$ and $\mathcal{L}_\pm^\dagger = \mathcal{L}_\mp$; this makes the representation unitary. For the barred sector we also select a highest-weight representation of $sl(2, \mathbb{R})$, which obeys

$$\begin{aligned} \bar{\mathcal{L}}_0|\bar{h}, \bar{p}\rangle &= (\bar{h}+\bar{p})|\bar{h}, \bar{p}\rangle, \\ \bar{\mathcal{L}}_+|\bar{h}, \bar{p}\rangle &= \sqrt{\bar{p}(\bar{p}+2\bar{h}+1)}|\bar{h}, \bar{p}-1\rangle, \\ \bar{\mathcal{L}}_-|\bar{h}, \bar{p}\rangle &= \sqrt{(\bar{p}+1)(\bar{p}+2\bar{h})}|\bar{h}, \bar{p}+1\rangle. \end{aligned} \quad (5.12)$$

The quadratic Casimir for this sector is

$$-\eta^{ab}\bar{\mathcal{L}}_a\bar{\mathcal{L}}_b|\bar{h}, \bar{p}\rangle = 2\bar{h}(\bar{h}-1)|\bar{h}, \bar{p}\rangle. \quad (5.13)$$

Singlet states in this case are defined in an analogous way as in section 3.1.1: we will consider two possible conditions

$$\begin{aligned} (\mathcal{L}_k - (-1)^k \bar{\mathcal{L}}_{-k})|\Sigma_{\text{cross}}\rangle &= 0, \\ (\mathcal{L}_k - \bar{\mathcal{L}}_{-k})|\Sigma_{\text{Ish}}\rangle &= 0, \end{aligned} \quad (5.14)$$

for $k = 0, \pm$, and the solutions are

$$\begin{aligned} |\Sigma_{\text{Ish}}\rangle &= \sum_{p=0}^{\infty} |h, p, p\rangle, \\ |\Sigma_{\text{cross}}\rangle &= \sum_{p=0}^{\infty} (-1)^p |h, p, p\rangle, \end{aligned} \quad (5.15)$$

⁹A similar discussion regarding representations of $sl(2, \mathbb{C})$ is discussed in [46]. One difference is that the authors take $sl(2, \mathbb{C}) \sim su(1, 1) \times su(1, 1)$.

where the singlet condition sets $h = \bar{h}$, and we are using $|h, p, \bar{p}\rangle \equiv |h, p\rangle \otimes |h, \bar{p}\rangle$. There is a difference in that the $(-1)^p$ factor appears for the crosscap state rather than for Ishibashi. This results from the fact that (5.10) and (5.12) do *not* contain a minus sign. In this sense they more closely resemble the AdS₃ rather than EdS₃ versions.

There is, however, a more important conceptual difference when we move to Lorentzian de Sitter. Recall that in EdS₃ the singlet states played a role in relating the two (barred and unbarred) copies of SU(2), which are initially independent; in the same way, here they allow us to relate two copies of SL(2, ℝ). Since in SL(2, ℂ) Chern-Simons theory the components $\mathcal{A}_a$ and $\bar{\mathcal{A}}_a$ are related by complex conjugation to ensure the reality of the Einstein-Hilbert action, the choice of a singlet state additionally picks out a reality condition on the fields propagating on the background created by $\mathcal{A}$ and $\bar{\mathcal{A}}$.

We can now evaluate the Wilson line. We are treating $sl(2, \mathbb{C})$ as two copies of $sl(2)$, as decomposed in (5.4), and hence we want to evaluate

$$W_{\mathcal{R}}(x_i, x_f) = \langle \Sigma | \mathcal{P} \exp \left(- \int_{\gamma} \mathcal{A} \right) \mathcal{P} \exp \left(- \int_{\gamma} \bar{\mathcal{A}} \right) | \Sigma \rangle, \quad (5.16)$$

where we selected the endpoint states to be one of the singlet states in (5.15): $|U_{i,f}\rangle = |\Sigma\rangle$. Writing this as group elements acting on each copy of $sl(2)$ we have

$$\begin{aligned} W_{\mathcal{R}}(x_i, x_f) &= \langle \Sigma | G(\mathcal{L}) \bar{G}(\mathcal{R}^{-1}) | \Sigma \rangle \\ &= \langle \Sigma | G(\mathcal{L} \tilde{\mathcal{R}}) | \Sigma \rangle \\ &= \sum_{p=0}^{\infty} \langle h, p | G(\mathcal{L} \tilde{\mathcal{R}}) | h, p \rangle \\ &= \frac{e^{ih\alpha}}{1 - e^{i\alpha}}, \end{aligned} \quad (5.17)$$

where

$$\begin{aligned} \mathcal{A} &= g dg^{-1}, & \mathcal{L} &\equiv g(x_f) g(x_i)^{-1}, \\ \bar{\mathcal{A}} &= \bar{g}^{-1} d\bar{g}, & \mathcal{R}^{-1} &\equiv \bar{g}(x_f)^{-1} \bar{g}(x_i), \end{aligned} \quad (5.18)$$

and $\tilde{\mathcal{R}} = \Sigma \mathcal{R} \Sigma^{-1}$. As before, we have defined α by assuming we can diagonalize the group element as

$$\mathcal{L} \Sigma \mathcal{R} \Sigma^{-1} = V^{-1} e^{i\alpha \mathcal{L}_0} V. \quad (5.19)$$

Other than the fact that we are using the states $|h, p, \bar{p}\rangle$ and generators $\mathcal{L}_a$ associated to our unitary Lorentzian representation rather than the states $|l, p, \bar{p}\rangle$ and generators L_a for the non-unitary Euclidean representation, everything proceeds as for the Euclidean case. In the end we can recognize that the Lorentzian Wilson line is just a character associated to our Lorentzian representations.

5.3 Inflationary patch

In this final portion we will consider the inflationary patch of dS₃ in order to illustrate our Lorentzian construction. The line element reads

$$\frac{ds^2}{\ell^2} = \frac{1}{\eta^2} (-d\eta^2 + d\omega d\bar{\omega}), \quad (5.20)$$

where $\eta > 0$, positive timelike infinity is located at $\eta \rightarrow 0$, and $w = x + iy$ is a complex variable. See appendix B.1 for a review of these coordinates.

For the inflationary patch, we use the group elements

$$g = e^{-\frac{iw}{\eta} \mathcal{L}_+} e^{\log \eta \mathcal{L}_0}, \quad \tilde{g} = e^{\log \eta \mathcal{L}_0} e^{\frac{i\bar{w}}{\eta} \mathcal{L}_-}. \quad (5.21)$$

These give connections

$$\mathcal{A} = g dg^{-1} = -\frac{d\eta}{\eta} \mathcal{L}_0 + \frac{idw}{\eta} \mathcal{L}_+, \quad \bar{\mathcal{A}} = \tilde{g}^{-1} d\tilde{g} = \frac{d\eta}{\eta} \mathcal{L}_0 + \frac{id\bar{w}}{\eta} \mathcal{L}_-. \quad (5.22)$$

In our conventions the Lorentzian metric is

$$g_{\mu\nu} = -\frac{\ell^2}{2} \text{Tr} [(\mathcal{A}_\mu - \bar{\mathcal{A}}_\mu)(\mathcal{A}_\nu - \bar{\mathcal{A}}_\nu)], \quad (5.23)$$

where here we are using the same generators for barred and unbarred connections. It is easy to check this reproduces (5.20).

As in the Euclidean case, we can define the local state from the group elements acting on the singlet state,

$$\begin{aligned} |U(x)\rangle &= G(g(x)^{-1}) \bar{G}(\bar{g}(x)) |\Sigma\rangle, \\ &= G(g(x)^{-1} \tilde{g}(x)^{-1}) |\Sigma\rangle, \end{aligned} \quad (5.24)$$

where $\tilde{g} = \Sigma \bar{g} \Sigma^{-1}$. Evaluating this using the group elements (5.21), we find

$$|U(x)\rangle = e^{-\log \eta \mathcal{L}_0} e^{\frac{iw}{\eta} \mathcal{L}_+} e^{-\frac{i\bar{w}}{\eta} \mathcal{L}_-} e^{-\log \eta \mathcal{L}_0} |\Sigma\rangle. \quad (5.25)$$

Now we will construct local pseudofields from the states $|U(x)\rangle$. We follow an exactly analogous procedure to the EdS₃ case in section 4.2, starting with expansion of the state over the states $|h, p, \bar{p}\rangle$ that form a basis for our unitary Lorentzian representations,

$$|U(x)\rangle = \sum_{p, \bar{p}=0}^{\infty} \Phi_{p, \bar{p}}^*(x) |h, p, \bar{p}\rangle. \quad (5.26)$$

Inverting this relation gives

$$\Phi_{p, \bar{p}}(x) = \langle U(x) | h, p, \bar{p} \rangle. \quad (5.27)$$

We can define a set of differential operators $\mathcal{H}_a$ and $\bar{\mathcal{H}}_a$ as

$$\langle U(x) | \mathcal{L}_a | h, p, \bar{p} \rangle = \mathcal{H}_a \langle U(x) | h, p, \bar{p} \rangle, \quad (5.28)$$

$$\langle U(x) | \bar{\mathcal{L}}_a | h, p, \bar{p} \rangle = \bar{\mathcal{H}}_a \langle U(x) | h, p, \bar{p} \rangle. \quad (5.29)$$

Taking derivatives of the pseudofield $\Phi_{p, \bar{p}}(x) = \langle U(x) | h, p, \bar{p} \rangle$, we find

$$\begin{aligned} \partial \langle U(x) | h, p, \bar{p} \rangle &= \frac{i}{\eta^2} \langle U(x) | \mathcal{L}_+ | h, p, \bar{p} \rangle - \frac{i\bar{w}^2}{\eta^2} \langle U(x) | \mathcal{L}_- | h, p, \bar{p} \rangle + \frac{2\bar{w}}{\eta^2} \langle U(x) | \mathcal{L}_0 | h, p, \bar{p} \rangle, \\ \bar{\partial} \langle U(x) | h, p, \bar{p} \rangle &= -i \langle U(x) | \mathcal{L}_- | h, p, \bar{p} \rangle, \\ \partial_\eta \langle U(x) | h, p, \bar{p} \rangle &= \frac{2i\bar{w}}{\eta} \langle U(x) | \mathcal{L}_- | h, p, \bar{p} \rangle - \frac{2}{\eta} \langle U(x) | \mathcal{L}_0 | h, p, \bar{p} \rangle, \end{aligned} \quad (5.30)$$

and from here we find

$$\begin{aligned}\mathcal{H}_+ &= -i(\eta^2\partial + \eta\bar{w}\partial_\eta + \bar{w}^2\bar{\partial}), \\ \mathcal{H}_- &= i\bar{\partial}, \\ \mathcal{H}_0 &= -\frac{\eta}{2}\partial_\eta - \bar{w}\bar{\partial}.\end{aligned}\tag{5.31}$$

These are three Killing vectors for the inflationary patch of dS_3 , whose boundary limits $\eta \rightarrow 0$ give one (barred) set of conformal generators.

The state $|U(x)\rangle$ can be equivalently be written in terms of the barred sector as

$$\begin{aligned}|U(x)\rangle &= \bar{G} [\Sigma^{-1}\tilde{g}(x)g(x)\Sigma] |\Sigma\rangle \\ &= \Sigma^{-1} e^{\log \eta \tilde{\mathcal{L}}_0} e^{\frac{i\bar{w}}{\eta} \tilde{\mathcal{L}}_-} e^{-\frac{iw}{\eta} \tilde{\mathcal{L}}_+} e^{\log \eta \tilde{\mathcal{L}}_0} \Sigma |\Sigma\rangle,\end{aligned}\tag{5.32}$$

where we have initially kept the state Σ arbitrary. Using the definitions (3.22) and (3.27) for the Ishibashi and crosscap states through their action on generators, for the Ishibashi state conjugation gives

$$\Sigma_{\text{Ish}}^{-1} e^{\log \eta \tilde{\mathcal{L}}_0} e^{\frac{i\bar{w}}{\eta} \tilde{\mathcal{L}}_-} e^{-\frac{iw}{\eta} \tilde{\mathcal{L}}_+} e^{\log \eta \tilde{\mathcal{L}}_0} \Sigma_{\text{Ish}} = e^{-\log \eta \tilde{\mathcal{L}}_0} e^{-\frac{i\bar{w}}{\eta} \tilde{\mathcal{L}}_+} e^{\frac{iw}{\eta} \tilde{\mathcal{L}}_-} e^{-\log \eta \tilde{\mathcal{L}}_0},\tag{5.33}$$

while for the crosscap state,

$$\Sigma_{\text{cross}}^{-1} e^{\log \eta \tilde{\mathcal{L}}_0} e^{\frac{i\bar{w}}{\eta} \tilde{\mathcal{L}}_-} e^{-\frac{iw}{\eta} \tilde{\mathcal{L}}_+} e^{\log \eta \tilde{\mathcal{L}}_0} \Sigma_{\text{cross}} = e^{-\log \eta \tilde{\mathcal{L}}_0} e^{\frac{i\bar{w}}{\eta} \tilde{\mathcal{L}}_+} e^{-\frac{iw}{\eta} \tilde{\mathcal{L}}_-} e^{-\log \eta \tilde{\mathcal{L}}_0}.\tag{5.34}$$

Restricting to the Ishibashi state for definiteness, we can follow a similar procedure and solve for the barred differential operators. We find

$$\begin{aligned}\bar{\mathcal{H}}_+ &= i(\eta^2\bar{\partial} + \eta w\partial_\eta + w^2\partial), \\ \bar{\mathcal{H}}_- &= -i\partial, \\ \bar{\mathcal{H}}_0 &= -\frac{\eta}{2}\partial_\eta - w\partial.\end{aligned}\tag{5.35}$$

Thus there is again a simple relation between the barred and unbarred differential operators. For the Ishibashi state the barred sector amounts to taking $w \leftrightarrow -\bar{w}$. The procedure can be repeated for the crosscap state, and in that case we must take $w \leftrightarrow \bar{w}$. We obtain from this a second set of Killing vectors whose $\eta \rightarrow 0$ limit matches onto the second (unbarred) set of conformal generators.

Now we can build solutions that explicitly realize our unitary representations. The highest weight state satisfies

$$\mathcal{H}_0\Phi_{0,0} = \bar{\mathcal{H}}_0\Phi_{0,0} = h\Phi_{0,0},\tag{5.36}$$

$$\mathcal{H}_+\Phi_{0,0} = \bar{\mathcal{H}}_+\Phi_{0,0} = 0,\tag{5.37}$$

and this equation is solved by

$$\Phi_{0,0}(\eta, w, \bar{w}) = \langle U(x)|h, 0, 0\rangle = \eta^{2h}(\eta^2 - w\bar{w})^{-2h}.\tag{5.38}$$

We can again build the descendents by lowering starting from this highest weight state. For the case $p > \bar{p}$ we find

$$\Phi_{p,\bar{p}}(\eta, w, \bar{w}) = b_{p,\bar{p}} \eta^{2h+2n} w^{p-\bar{p}} (\eta^2 - w\bar{w})^{-p-\bar{p}-2h} P_n^{(|p-\bar{p}|, -p-\bar{p}-2h)} \left(1 - \frac{2w\bar{w}}{\eta^2}\right), \quad (5.39)$$

where

$$b_{p,\bar{p}} = i^p (-i)^{\bar{p}} \sqrt{\frac{\bar{p}!(p+2h-1)!}{p!(\bar{p}+2h-1)!}}, \quad n = \frac{1}{2}(p + \bar{p} - |p - \bar{p}|). \quad (5.40)$$

For $\bar{p} > p$, the solution is $\Phi_{p,\bar{p}}(\eta, w, \bar{w}) = (-i)^p i^{\bar{p}} \Phi_{\bar{p},p}(\eta, \bar{w}, w)$. The solutions are again Jacobi polynomials $P_n^{\alpha,\beta}(x)$, however in this case n depends nontrivially on both quantum numbers $p, \bar{p}$. Just like the static patch quasinormal modes, these are eigenfunctions satisfying (4.40) and they solve the Klein-Gordon equation (4.41) in inflationary coordinates.

Restricting to $w = \bar{w} = 0$ at finite η , the solution for the pseudofield reduces to

$$\Phi_{p,\bar{p}}(\eta, 0, 0) = \eta^{-2(p+h)} \delta_{p,\bar{p}}. \quad (5.41)$$

This means we can write

$$|U(\eta, 0, 0)\rangle = \sum_p \eta^{-2(p+h)} |h, p, p\rangle, \quad (5.42)$$

which at $\eta = 1$ is simply the Ishibashi state, (5.15). Thus we see that our Lorentzian Ishibashi state lives at $w = \bar{w} = 0, \eta = 1$. By going to embedding coordinates (B.14), it is easy to see that, up to analytic continuation, this is the same bulk point as $r = 0, \tau = 0$ where the Ishibashi state was located in static coordinates. Of course, once again we note that there is nothing special about that point: it is simply the product of various gauge choices we made along the way.

Finally we turn to the Wilson line, which can be evaluated directly as

$$\begin{aligned} W_{\mathcal{R}}(x_i, x_f) &= \langle \Sigma | G(g(x_f)g(x_i)^{-1}) \bar{G}(\tilde{g}(x_f)^{-1}\tilde{g}(x_i)) | \Sigma \rangle \\ &= \langle \Sigma | G(g(x_f)g(x_i)^{-1}\tilde{g}(x_i)^{-1}\tilde{g}(x_f)) | \Sigma \rangle. \end{aligned} \quad (5.43)$$

Using (5.19) and the explicit inflationary group elements (5.21), we can solve for the parameter α describing the eigenvalue of the group element. We find

$$\cos\left(\frac{\alpha}{2}\right) = \frac{\eta_i^2 + \eta_f^2 - (w_f - w_i)(\bar{w}_f - \bar{w}_i)}{2\eta_i\eta_f}. \quad (5.44)$$

The right hand side is again just the invariant distance but now in inflationary coordinates (see appendix B.2). This is directly analogous to our analysis of the Euclidean case, where α was related to the invariant distance in Hopf coordinates. We again have

$$\alpha = \pm 2\Theta + 4\pi n, \quad n \in \mathbb{Z}. \quad (5.45)$$

We can now relate the Wilson line to a Green's function. Recall that the Lorentzian Wilson line was equal to a character of our representation,

$$W_{\mathcal{R}}(x_i, x_f) = \frac{e^{ih\alpha}}{1 - e^{i\alpha}}. \quad (5.46)$$

Using (5.45), we can convert this to a function of the invariant distance. After again defining

$$a_h = \frac{i}{2\pi\ell} \frac{1}{1 - e^{-4\pi i h}}, \quad (5.47)$$

we find that taking the irreducible representation $\mathcal{R}_+$ with $h = \frac{1 + \sqrt{1 - (m\ell)^2}}{2}$ leads to

$$W_{\mathcal{R}_+}(x_i, x_f) = \frac{1}{a_h} \mathcal{G}_h(\Theta). \quad (5.48)$$

As in the Euclidean case given by (3.49), to obtain the Green's function (B.30) it is necessary to use both representations $\mathcal{R}_{\pm}$ with highest weight h and $1 - h$,

$$G(\Theta) = a_h W_{\mathcal{R}_+}(x_i, x_f) + a_{1-h} W_{\mathcal{R}_-}(x_i, x_f). \quad (5.49)$$

6 Discussion

In this last section we highlight our main findings and discuss some interesting future directions.

Singlet states in 3D de Sitter. To summarize: the singlet states we constructed in section 3 take the form

$$|\Sigma\rangle = \sum_{p, \bar{p}} a_{p, \bar{p}} |l, p, \bar{p}\rangle, \quad (6.1)$$

where $|l, p, \bar{p}\rangle = |l, p\rangle \otimes |l, \bar{p}\rangle$ are basis vectors of a non-unitary representation of $su(2)$. One of the consequences tied to selecting this unconventional representation is that we have a continuous parameter that we can identify with the mass of particle: we take $-1 < l < 0$, and its relation to the mass is $4l(l+1) = -m^2\ell^2$. Although our discussion is limited to masses in the ranges $0 < m^2\ell^2 < 1$, our approach should be easily extendable to allow for arbitrary positive values of $m^2\ell^2$. Note that the representation we consider for Lorentzian de Sitter spacetime, which results from analytic continuation of the nonunitary $su(2)$ representation, is just the discrete series representation of $sl(2, \mathbb{R})$. We expect that extending the mass range would require building non-unitary representations of $su(2)$ that resemble the continuous series in $sl(2, \mathbb{R})$, which includes the principal and complementary series representations [46].

These singlet states are very reminiscent of the description of bulk local states in AdS. In [51–53], it was shown that a bulk field configuration at the centre of AdS corresponds to a crosscap state in the CFT. While there are certainly similarities between the two stories (emphasized by our choice of terminology for singlet states), there are also some notable differences. In the context of AdS/CFT, the crosscap states are states in the full Virasoro algebra, not just the global $sl(2, \mathbb{R}) \times sl(2, \mathbb{R})$ subalgebra. Furthermore, the CFT can be seen to set some bulk properties naturally through boundary conditions. These properties provide an external source for choices that otherwise seem arbitrary. For example, we found no obvious physical difference between the Ishibashi and crosscap states, because we had the freedom to relabel algebra generators. In AdS, these generators have an independent physical meaning in the boundary CFT that must be matched, hence the statement that the point at the origin must be a crosscap state rather than an Ishibashi state.

We also performed an analytic continuation and considered singlet states in the Lorentzian case, where for illustration, we focused on the inflationary patch of Lorentzian dS_3 . To describe gravity in Lorentzian de Sitter we were led to consider $SL(2, \mathbb{C})$ Chern-Simons theory. In this context, the choice of singlet state led to a natural reality condition for the $SL(2, \mathbb{C})$ Chern-Simons gauge fields. Lorentzian Wilson lines had a direct interpretation in terms of unitary $sl(2, \mathbb{R})$ representations that we motivated using an analytic continuation of our Euclidean $su(2)$ representations. Since the inflationary patch has a large amount of apparent symmetry, it would also be interesting to repeat our analysis for less symmetric bulks such as Kerr- dS_3 [54].

Bulk reconstruction in 3D de Sitter. The comparison to AdS/CFT naturally raises the question of bulk reconstruction. Consider our Lorentzian results for the inflationary patch. We now have an expression for pseudofields $|U(x)\rangle$ in terms of an abstract basis of states $|h, p, \bar{p}\rangle$ that mimics the discussion in AdS. And while a dS/CFT correspondence [55–57] is far from established, suppose for the sake of argument that we take seriously the idea that our states $|h, p, \bar{p}\rangle$ can be described as operators in a putative CFT, in other words that there is a state-operator correspondence that maps our states to operators inserted at the origin $w, \bar{w} = 0$: $|h, p, \bar{p}\rangle = \mathcal{O}(0, 0)|0\rangle$. Then the Ishibashi state

$$|\Sigma_{\text{Ish}}\rangle = \sum_{p=0}^{\infty} |h, p, p\rangle \quad (6.2)$$

can be expressed as

$$|\Sigma_{\text{Ish}}\rangle = \sum_{p=0}^{\infty} \frac{\Gamma(2h)}{\Gamma(p+1)\Gamma(p+2h)} \mathcal{H}_{-1}^p \bar{\mathcal{H}}_{-1}^p \mathcal{O}(0, 0)|0\rangle. \quad (6.3)$$

On the other hand, the Ishibashi state can be thought of as being localized at a particular bulk point, as seen in (5.42). This suggests that we can obtain pseudofields at arbitrary bulk points by acting on both sides of (6.3) with $sl(2, \mathbb{R})$ generators. On the bulk side, this could be interpreted as diffeomorphisms that move the point while on the boundary side there is a natural interpretation in terms of conformal transformations.

Thus, we are led to ask: is there then an analogue of the HKLL procedure [58, 59], where local fields in de Sitter can be thought of as a smearing of states in a region of a lower-dimensional surface? And is there an implementation of that procedure in Chern-Simons theory? To answer these questions, it is useful to compare to the existing literature on bulk reconstruction in de Sitter. A smearing function for the inflationary patch was constructed in [60], and further developments include [46, 61, 62]. Restricting to $d = 2$, the result is that a local scalar field Φ of mass m in the inflationary patch of dS_3 can be represented as

$$\begin{aligned} \Phi(\eta, w, \bar{w}) = \int_{|w'\bar{w}'| < \eta^2} dw' d\bar{w}' & \left[\frac{\Gamma(\Delta)}{\pi\Gamma(\Delta-1)} \left(\frac{\eta^2 - |w'\bar{w}'|}{\eta} \right)^{\Delta-2} \mathcal{O}_+(w+w', \bar{w}+\bar{w}') \right. \\ & \left. + \frac{\Gamma(2-\Delta)}{\pi\Gamma(1-\Delta)} \left(\frac{\eta^2 - |w'\bar{w}'|}{\eta} \right)^{-\Delta} \mathcal{O}_-(w+w', \bar{w}+\bar{w}') \right]. \end{aligned} \quad (6.4)$$

In de Sitter it was crucial to keep the contributions from not only a scalar operator $\mathcal{O}_+$ with scaling dimension $\Delta_+ = \Delta = 1 + \sqrt{1 - m^2 \ell^2}$ dual to Φ , but also the shadow operator $\mathcal{O}_-$ with scaling dimension $\Delta_- = 2 - \Delta = 1 - \sqrt{1 - m^2 \ell^2}$. Here it is necessary to have these two contributions for the two-point function of the field to reproduce the correct Green's function, (B.30), which differs substantially from AdS. The difference is related to the fact that the Euclidean Green's function we use for de Sitter is not simply the analytic continuation of the AdS Green's function, which would violate microcausality [43].

In our language the two terms come from considering the two representations with a fixed Casimir, with $l = -h$ and $l = h - 1$. Other than this subtlety, and assuming the existence of a state operator correspondence for the states in our representations, the computation of the contribution to a bulk local field for each set of operators in terms of smearing functions proceeds exactly analogously to the Poincaré case considered in [63]. All that is needed is to express the singlet state, translated to a point in the bulk, in terms of differential operators acting on CFT operators. This can then be converted into an integral representation in terms of smearing functions. There is however a need to have a more fundamental understanding of the role of $\mathcal{O}_+$ and $\mathcal{O}_-$ and its implications in dS quantum gravity.

Exact results in Chern-Simons theory. Chern-Simons theory on S^3 , with a compact gauge group, is exactly solvable using the techniques of non-abelian localization [64]. In particular, the Wilson loop expectation value can be computed exactly in this context [40, 65]. This suggests an extension of our semiclassical Euclidean results to a full quantum computation.

There are two crucial differences in our approach that prevent us from applying exact results directly. The first is that we consider Wilson line operators rather than loops, which means that our probes are not gauge invariant. Additionally, we compute the Wilson line for infinite dimensional (and subsequently non-unitary) rather than finite dimensional representations of $su(2)$. The choice of this peculiar representation is in fact intricately linked to the non-gauge invariance of the Wilson lines, as we required infinite dimensional representations to construct the singlet states describing the endpoints. In the semiclassical version these limitations did not end up presenting an obstruction to a generalization as in [6, 21], and so it would be interesting to implement techniques of localization to construct and quantify our Wilson line as a quantum operator.

It would be especially interesting to see if the quantization of the Wilson line sheds light on the necessity in de Sitter of using two representations $\mathcal{R}_\pm$, which from the CFT standpoint led us to consider an additional set of shadow operators. We saw that these were necessary in our framework to generate the complete set of quasinormal modes for de Sitter, and they are also crucial to reproduce the correct Green's function from a smearing function representation of a bulk local field. Moving beyond kinematics, one might hope that a quantization would help us define a Hilbert space that incorporates both representations and gives a definition for their overlap.

Acknowledgments

We are thank Costas Bachas, Monica Guica, Kurt Hinterbichler, Eva Llabrés, and Alex Maloney for useful discussions. This work is supported by the Delta ITP consortium, a program of the Netherlands Organisation for Scientific Research (NWO) that is funded by the Dutch Ministry of Education, Culture and Science (OCW). PSG thanks the IoP at the University of Amsterdam for its hospitality during this project and acknowledges the support of the Natural Sciences and Engineering Research Council of Canada (NSERC) [PDF-517202-2018]. AC thanks the Laboratoire de Physique Théorique de l'Ecole Normale Supérieure for their hospitality while this work was completed.

A Conventions

In this appendix we collect some basic conventions related to the Lie group $SU(2)$ and its algebra. For the algebra we use generators L_a and $\bar{L}_a$, $a = 1, 2, 3$, and we have

$$[L_a, L_b] = i\epsilon_{abc}L^c, \quad (A.1)$$

with $\epsilon_{123} \equiv 1$. For the invariant bilinear form, we take

$$\text{Tr}(L_a L_b) = \frac{1}{2}\delta_{ab}. \quad (A.2)$$

Indices are raised with δ^{ab} . In the fundamental representation of $su(2)$, we have $L_a = \frac{1}{2}\sigma_a$ with the Pauli matrices given by

$$\sigma_1 = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}, \quad \sigma_2 = \begin{bmatrix} 0 & -i \\ i & 0 \end{bmatrix}, \quad \sigma_3 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}. \quad (A.3)$$

To make an explicit distinction between the group and the algebra, we denote $G(M)$ as group element, and L_a are the algebra generators as specified above. The general group action is given by

$$G(M^{-1})L_a G(M) = D_a^{a'}(M)L_{a'}, \quad \bar{G}(M^{-1})\bar{L}_a \bar{G}(M) = D_a^{a'}(M)\bar{L}_{a'}, \quad (A.4)$$

where $D's$ are the elements in the adjoint representation of $su(2)$. As expected for any group, we also have

$$G(M_1)G(M_2) = G(M_1 M_2), \quad D_a^b(M_1)D_b^c(M_2) = D_a^c(M_1 M_2). \quad (A.5)$$

B Metric formulation of dS₃ gravity

B.1 Coordinates and patches

Three-dimensional de Sitter is easily understood in terms of its embedding in four-dimensional Minkowski space:

$$-(X^0)^2 + (X^1)^2 + (X^2)^2 + (X^3)^2 = \ell^2. \quad (B.1)$$

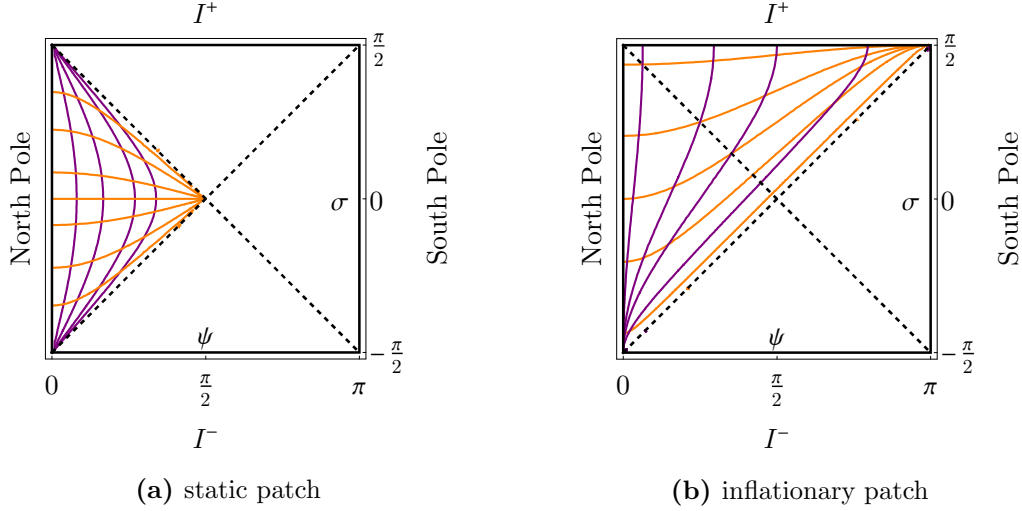


Figure 1. (Colour online) Penrose diagram of three-dimensional de Sitter space. Horizontal lines are slices of constant global time T (or σ), which correspond to 2-spheres. ψ is the polar angle on that sphere, so that each point on the diagram is a circle of radius $\sin \psi$. Vertical lines are slices of constant ψ . The top and bottom of the diagram are asymptotic timelike infinity, and the left and right edges are the North and South poles of the 2-spheres at each instant in global time. Constant t (orange) and r (or u , purple) slices on the static patch are shown on the static patch, with $r = 0$ at the North Pole and r increasing to $\frac{\pi}{2}$ at the horizon. Constant $\eta \geq 0$ (orange) and x (for $y = 0$, purple) slices are shown on the inflationary patch, with $\eta \rightarrow 0^+$ corresponding to positive timelike infinity and increasing to $+\infty$ at the horizon.

Global dS_3 corresponds to the following parametrization, which covers the whole space-time:

$$\begin{aligned}
 X^0 &= \ell \sinh(T/\ell), \\
 X^1 &= \ell \cosh(T/\ell) \cos \psi, \\
 X^2 &= \ell \cosh(T/\ell) \sin \psi \cos \phi, \\
 X^3 &= \ell \cosh(T/\ell) \sin \psi \sin \phi,
 \end{aligned} \tag{B.2}$$

with ψ and ϕ the polar and azimuthal coordinates of a two-sphere of unit radius. The metric is then

$$ds^2 = -dT^2 + \ell^2 \cosh^2(T/\ell) (d\psi^2 + \sin^2(\psi) d\phi^2). \tag{B.3}$$

The global time coordinate T , which has an infinite range, can be conformally rescaled:

$$\tan(\sigma) \equiv \sinh(T/\ell), \quad -\frac{\pi}{2} < \sigma < \frac{\pi}{2}. \tag{B.4}$$

After this rescaling, the metric is

$$ds^2 = \frac{\ell^2}{\cos^2 \sigma} (-d\sigma^2 + d\psi^2 + \sin^2 \psi d\phi^2). \tag{B.5}$$

With the metric in this form, it is easy to draw the Penrose diagram in figure 1.

Another useful parametrization of embedding coordinates is the following:

$$\begin{aligned} X^0 &= \sqrt{\ell^2 - u^2} \sinh(t/\ell), \\ X^1 &= \sqrt{\ell^2 - u^2} \cosh(t/\ell), \\ X^2 &= u \cos \phi, \\ X^3 &= u \sin \phi, \end{aligned} \tag{B.6}$$

for which the metric can be written as

$$ds^2 = - \left(1 - \frac{u^2}{\ell^2}\right) dt^2 + \frac{du^2}{1 - \frac{u^2}{\ell^2}} + u^2 d\phi^2. \tag{B.7}$$

This is the static patch of dS_3 . It has the advantage of making a timelike Killing vector manifest, at the cost of covering only a portion of the whole manifold. We can see which portion by relating the two parametrizations:

$$\begin{aligned} u &= \ell \cosh(T/\ell) |\sin(\psi)| = \ell \left| \frac{\sin(\psi)}{\cos(\sigma)} \right|, \\ \sinh^2(t/\ell) &= \frac{\sinh^2(T/\ell)}{1 - \cosh^2(T/\ell) \sin^2(\psi)} = \frac{\sin^2(\sigma)}{\cos^2(\sigma) - \sin^2(\psi)}. \end{aligned} \tag{B.8}$$

In particular, for the embedding coordinates to be real, we need $0 \leq u \leq \ell$. $u = 0$ corresponds to $\psi = 0$ and $u = \ell$ corresponds to $\sigma = \pm(\psi - \frac{\pi}{2})$, so that these coordinates cover the left wedge of the Penrose diagram (or the right wedge, but not both if the coordinates are to be single-valued). Trajectories of constant u or t are shown in figure 1a. A simple coordinate redefinition brings us to the coordinates used in the main text:

$$u = \ell \sin(r). \tag{B.9}$$

The embedding coordinates then take the form

$$\begin{aligned} X^0 &= \ell \cos(r) \sinh(t/\ell), \\ X^1 &= \ell \cos(r) \cosh(t/\ell), \\ X^2 &= \ell \sin(r) \cos(\phi), \\ X^3 &= \ell \sin(r) \sin(\phi), \end{aligned} \tag{B.10}$$

and the metric is

$$ds^2 = -\cos^2(r) dt^2 + \ell^2 dr^2 + \ell^2 \sin^2(r) d\phi^2. \tag{B.11}$$

It is instructive to go to Euclidean time in these coordinates: $t \rightarrow -i\tau\ell$,¹⁰ which leads to

$$\frac{ds^2}{\ell^2} = \cos^2(r) d\tau^2 + dr^2 + \sin^2(r) d\phi^2, \tag{B.12}$$

¹⁰Our Lorentzian metric has a mostly-+ signature. This fixes $t \rightarrow -i\tau\ell$ rather than $t \rightarrow +i\tau\ell$ in order to ensure that the equations of motion minimize the Hamiltonian rather than maximize it. The factor of ℓ is there to make the interpretation of τ as an angular coordinate manifest.

and

$$\begin{aligned} X^1 &= \ell \cos(r) \cos(\tau), \\ X^2 &= \ell \sin(r) \cos(\phi), \\ X^3 &= \ell \sin(r) \sin(\phi), \\ X^4 &= \ell \cos(r) \sin(\tau), \end{aligned} \tag{B.13}$$

where we've defined $X^4 = iX^0$. These coordinates are simply the Hopf coordinates for a three-sphere embedded in $\mathbb{R}^4$. Avoiding a conical singularity near $r = \frac{\pi}{2}$ requires that $\tau \sim \tau + 2\pi$, from which we can read off the inverse temperature of the horizon: $\beta = 2\pi\ell$.

Another parametrization of dS_3 gives coordinates on the inflationary patch:

$$\begin{aligned} X^0 &= -\ell \frac{\eta^2 - 1 - x^2 - y^2}{2\eta} \\ X^1 &= \ell \frac{\eta^2 + 1 - x^2 - y^2}{2\eta} \\ X^2 &= \ell \frac{x}{\eta} \\ X^3 &= \ell \frac{y}{\eta}. \end{aligned} \tag{B.14}$$

The metric in these coordinates is

$$\frac{ds^2}{\ell^2} = \frac{-d\eta^2 + dx^2 + dy^2}{\eta^2}. \tag{B.15}$$

With $0 < \eta < \infty$, these coordinates cover half of the space-time, with $\eta^{-1} = \frac{\cos \psi}{|\cos \sigma|} + \tan \sigma$. $\eta = 0^+$ corresponds to $\sigma = \frac{\pi}{2}$ (i.e. positive timelike infinity) and $\eta \rightarrow +\infty$ to $\sigma \rightarrow (\psi - \frac{\pi}{2})^+$. This is shown in figure 1b.

B.2 Geodesics and Green's functions in dS_3

We now write down the propagator for a scalar field in the static patch of three-dimensional de Sitter. We can exploit the symmetry of the system to write the wave equation in terms of the geodesic distance between two points. This is easier to do in Euclidean signature, where we consider S^3 described by embedding coordinates X^i given by equation (B.13). The only invariant quantity we can write out of two vectors X^i and Y^i is $X \cdot Y$. In fact, the geodesic distance between two points is simply

$$\ell\Theta = \ell \arccos \left(\frac{X \cdot Y}{\ell^2} \right). \tag{B.16}$$

The Euclidean propagator obeys:

$$\nabla^2 G(X, Y) - m^2 G(X, Y) = \delta^{(3)}(X - Y). \tag{B.17}$$

The propagator can only depend on coordinates through the quantity $\chi = \cos(\Theta)$. This implies

$$\begin{aligned}\nabla^2 G(X, Y) &= \nabla^2 G(\chi) \\ &= \frac{1}{\sin^2(\Theta)} \frac{d}{d\Theta} \left[(\sin(\Theta))^2 \frac{dG(\chi)}{d\Theta} \right] \\ &= (1 - \chi^2) \frac{d^2 G}{d\chi^2} - 3\chi \frac{dG}{d\chi}.\end{aligned}\tag{B.18}$$

Therefore, the homogeneous version of the wave equation is

$$(1 - \chi^2) \frac{d^2 G}{d\chi^2} - 3\chi \frac{dG}{d\chi} - (m\ell)^2 G(\chi) = 0.\tag{B.19}$$

This has solutions of the following form:

$$G(\chi) = c_1 \frac{P\left(\sqrt{1 - (m\ell)^2} - \frac{1}{2}, \frac{1}{2}; \chi\right)}{(\chi^2 - 1)^{1/4}} + c_2 \frac{Q\left(\sqrt{1 - (m\ell)^2} - \frac{1}{2}, \frac{1}{2}; \chi\right)}{(\chi^2 - 1)^{1/4}},\tag{B.20}$$

where P and Q are associated Legendre polynomials. These associated Legendre polynomials simplify precisely when the second argument is $1/2$:

$$\begin{aligned}P(x, 1/2, \cos(\Theta)) &= \sqrt{\frac{2}{\pi \sin(\Theta)}} \cos\left(\left(x + \frac{1}{2}\right)\Theta\right), \\ Q(x, 1/2, \cos(\Theta)) &= -\sqrt{\frac{\pi}{2 \sin(\Theta)}} \sin\left(\left(x + \frac{1}{2}\right)\Theta\right).\end{aligned}\tag{B.21}$$

Therefore, introducing two new undetermined constants, we have

$$G(\Theta) = \frac{A}{\sin(\Theta)} \left[\cos\left(\left(\sqrt{1 - (m\ell)^2}\right)\Theta\right) - C \sin\left(\left(\sqrt{1 - (m\ell)^2}\right)\Theta\right) \right].\tag{B.22}$$

Short distances correspond to χ approaching 1 from below, in other words taking $\Theta \sim \epsilon$. In that regime, we get

$$G(\Theta = \epsilon) = \frac{A}{\epsilon} + \dots.\tag{B.23}$$

In three dimensions, we know that a properly normalized Green's function has a short-distance divergence that goes as $-\frac{1}{4\pi d} = -\frac{1}{4\pi\ell\epsilon}$, so this fixes

$$A = -\frac{1}{4\pi\ell}.\tag{B.24}$$

Therefore, the Euclidean propagator is

$$G(\Theta) = \frac{i}{4\pi\ell} \left[(1 + iC) \frac{e^{2ih\Theta}}{1 - e^{2i\Theta}} - (1 - iC) \frac{e^{-2ih\Theta}}{1 - e^{-2i\Theta}} \right],\tag{B.25}$$

with

$$h = \frac{1 + \sqrt{1 - (m\ell)^2}}{2}.\tag{B.26}$$

Choosing the value of the integration constant C corresponds to picking a particular vacuum. The natural choice is $C = \cot(2h\pi)$. This removes the singularity at $\Theta = \pi$, in other words the singularity that appears on the lightcone of antipodal points. It is convenient to define

$$\mathcal{G}_h(\Theta) = \frac{i}{2\pi\ell} \frac{1}{1 - e^{-4\pi i h}} \frac{e^{-2ih\Theta}}{e^{-2i\Theta} - 1}. \quad (\text{B.27})$$

The Green's function satisfying the boundary conditions we've outlined above naturally splits into

$$G(\theta) = \mathcal{G}_h(\Theta) + \mathcal{G}_{1-h}(\Theta) \quad (\text{B.28})$$

$$= -\frac{\csc(2\pi h) \csc(\Theta) \sin(2h(\pi - \Theta) + \Theta)}{4\pi\ell} \quad (\text{B.29})$$

$$= -\frac{\Gamma(2h)\Gamma(2-2h)}{(4\pi)^{3/2}\Gamma(3/2)} {}_2F_1\left(2h, 2-2h; \frac{3}{2}, \frac{1+\cos(\Theta)}{2}\right), \quad (\text{B.30})$$

where the last line is manifestly the Green's function in the Euclidean vacuum of dS_3 [43].

We can write Θ explicitly in terms of the Hopf coordinates (B.13):

$$\cos(\Theta) = \cos(r_i) \cos(r_f) \cos(\Delta\tau) + \sin(r_i) \sin(r_f) \cos(\Delta\phi). \quad (\text{B.31})$$

Finally, note that if we want to analytically continue back to Lorentzian signature, we should take $\tau \rightarrow it/\ell$. In that case, we notice that $|\cos(\Theta)| > 1$ whenever points are timelike separated, in which case $i\ell\Theta$ corresponds to the proper time between the points. When the points are spacelike separated, $\ell\Theta$ remains the proper length between them.

In terms of inflationary coordinates (B.14), the invariant distance is

$$\cos(\Theta) = \frac{\eta_i^2 + \eta_f^2 - (\Delta x)^2 - (\Delta y)^2}{2\eta_i\eta_f} \quad (\text{B.32})$$

$$= 1 + \frac{(\Delta\eta)^2 - (\Delta x)^2 - (\Delta y)^2}{2\eta_i\eta_f}, \quad (\text{B.33})$$

where the distinction between spacelike-separated and timelike-separated points is manifest.

C Analytic continuation in the Chern-Simons formulation

Here, we provide more details on how to construct an analytic continuation between Euclidean and Lorentzian signature from the Chern-Simons perspective. The analytic continuation from Euclidean to Lorentzian signature is most easily understood in terms of these generators, which are simply related to rotations and boosts in embedding space.

The Euclidean Chern-Simons action, (2.3) and (2.5), can be written in unsplit form as

$$S_E = -\frac{k}{4\pi} \int \text{Tr} \left(\mathcal{A} \wedge d\mathcal{A} + \frac{2}{3} \mathcal{A} \wedge \mathcal{A} \wedge \mathcal{A} \right), \quad (\text{C.1})$$

where the gauge field is expanded in terms of the generators of Euclidean $so(4)$ isometries as

$$\mathcal{A} = e^a P_a + \omega^a J_a, \quad (\text{C.2})$$

with

$$[J_a, J_b] = -\epsilon_{abc} J^c, \quad (\text{C.3})$$

$$[J_a, P_b] = -\epsilon_{abc} P^c, \quad (\text{C.4})$$

$$[P_a, P_b] = -\Lambda \epsilon_{abc} J^c, \quad (\text{C.5})$$

and indices are raised with δ^{ab} . To construct an analytic continuation, we need to find a map to generators of the $so(1,3)$ algebra of isometries of Lorentzian dS_3 ,

$$[\mathcal{J}_a, \mathcal{J}_b] = \epsilon_{abc} \mathcal{J}^c, \quad (\text{C.6})$$

$$[\mathcal{J}_a, \mathcal{P}_b] = \epsilon_{abc} \mathcal{P}^c, \quad (\text{C.7})$$

$$[\mathcal{P}_a, \mathcal{P}_b] = -\Lambda \epsilon_{abc} \mathcal{J}^c, \quad (\text{C.8})$$

with indices raised by η^{ab} .

One possibility is given by the following:

$$\begin{aligned} \mathcal{J}_1 &= iP_1, \\ \mathcal{J}_2 &= iJ_2, \\ \mathcal{J}_3 &= -P_3, \\ \mathcal{P}_1 &= J_1, \\ \mathcal{P}_2 &= P_2, \\ \mathcal{P}_3 &= iJ_3. \end{aligned} \quad (\text{C.9})$$

Under this map, the $SO(4)$ bilinear form

$$\ell \text{Tr}(J_a P_b) = -\delta_{ab}, \quad \text{Tr}(J_a J_b) = \ell^2 \text{Tr}(P_a P_b) = 0 \quad (\text{C.10})$$

gets taken to

$$\ell \text{Tr}(\mathcal{J}_a \mathcal{P}_b) = -i \eta_{ab}, \quad \text{Tr}(\mathcal{J}_a \mathcal{J}_b) = \ell^2 \text{Tr}(\mathcal{P}_a \mathcal{P}_b) = 0. \quad (\text{C.11})$$

While the map we have constructed can be viewed as a map between real algebras, (C.11) is not an invariant bilinear form for real $SO(3,1)$. Indeed, the unique invariant bilinear form for $SO(3,1)$ is given by

$$\langle \mathcal{J}_a, \mathcal{J}_b \rangle = \eta_{ab}, \quad \langle \mathcal{P}_a, \mathcal{P}_b \rangle = -\Lambda \eta_{ab}, \quad \langle \mathcal{J}_a, \mathcal{P}_b \rangle = 0, \quad (\text{C.12})$$

rather than (C.11). In the Chern-Simons formulation for gravity one typically chooses a $\text{Tr}(J_a P_a)$ bilinear form for a reason, as the Chern-Simons theory defined using (C.12) does not reduce to Einstein gravity (see [2]). It is for this reason that we have considered a complexification to $SL(2, \mathbb{C})$. While the real $SO(3,1)$ algebra does not split as in (2.8), the complexification does split and therefore admits multiple bilinear forms, not only (C.12) but also (C.11).

With the map defined above, the bilinear form for the barred sector has the wrong sign:

$$\text{Tr}(\bar{\mathcal{L}}_a \bar{\mathcal{L}}_b) = +\frac{1}{2} \eta_{ab}. \quad (\text{C.13})$$

We can flip the sign while simultaneously multiplying the barred action by a minus sign. Combined with an analytic continuation of the Chern-Simons coupling,

$$s = ik, \quad (\text{C.14})$$

this takes us from (C.1) to the $\text{SL}(2, \mathbb{C})$ Chern-Simons action, (5.1).

Open Access. This article is distributed under the terms of the Creative Commons Attribution License ([CC-BY 4.0](https://creativecommons.org/licenses/by/4.0/)), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

References

- [1] A. Achucarro and P.K. Townsend, *A Chern-Simons Action for Three-Dimensional anti-de Sitter Supergravity Theories*, *Phys. Lett. B* **180** (1986) 89 [[INSPIRE](#)].
- [2] E. Witten, *(2 + 1)-Dimensional Gravity as an Exactly Soluble System*, *Nucl. Phys. B* **311** (1988) 46 [[INSPIRE](#)].
- [3] G.W. Moore and N. Seiberg, *Taming the Conformal Zoo*, *Phys. Lett. B* **220** (1989) 422 [[INSPIRE](#)].
- [4] S. Elitzur, G.W. Moore, A. Schwimmer and N. Seiberg, *Remarks on the Canonical Quantization of the Chern-Simons-Witten Theory*, *Nucl. Phys. B* **326** (1989) 108 [[INSPIRE](#)].
- [5] H.L. Verlinde, *Conformal Field Theory, 2-D Quantum Gravity and Quantization of Teichmüller Space*, *Nucl. Phys. B* **337** (1990) 652 [[INSPIRE](#)].
- [6] A. Castro, N. Iqbal and E. Llabrés, *Wilson lines and Ishibashi states in $\text{AdS}_3/\text{CFT}_2$* , *JHEP* **09** (2018) 066 [[arXiv:1805.05398](#)] [[INSPIRE](#)].
- [7] E. Witten, *Quantum Field Theory and the Jones Polynomial*, *Commun. Math. Phys.* **121** (1989) 351 [[INSPIRE](#)].
- [8] K.B. Alkalaev and V.A. Belavin, *Classical conformal blocks via AdS/CFT correspondence*, *JHEP* **08** (2015) 049 [[arXiv:1504.05943](#)] [[INSPIRE](#)].
- [9] E. Hijano, P. Kraus, E. Perlmutter and R. Snively, *Semiclassical Virasoro blocks from AdS_3 gravity*, *JHEP* **12** (2015) 077 [[arXiv:1508.04987](#)] [[INSPIRE](#)].
- [10] K.B. Alkalaev and V.A. Belavin, *Monodromic vs geodesic computation of Virasoro classical conformal blocks*, *Nucl. Phys. B* **904** (2016) 367 [[arXiv:1510.06685](#)] [[INSPIRE](#)].
- [11] M. Besken, A. Hegde, E. Hijano and P. Kraus, *Holographic conformal blocks from interacting Wilson lines*, *JHEP* **08** (2016) 099 [[arXiv:1603.07317](#)] [[INSPIRE](#)].
- [12] A.L. Fitzpatrick, J. Kaplan, D. Li and J. Wang, *Exact Virasoro Blocks from Wilson Lines and Background-Independent Operators*, *JHEP* **07** (2017) 092 [[arXiv:1612.06385](#)] [[INSPIRE](#)].
- [13] M. Besken, E. D'Hoker, A. Hegde and P. Kraus, *Renormalization of gravitational Wilson lines*, *JHEP* **06** (2019) 020 [[arXiv:1810.00766](#)] [[INSPIRE](#)].
- [14] Y. Hikida and T. Uetoko, *Conformal blocks from Wilson lines with loop corrections*, *Phys. Rev. D* **97** (2018) 086014 [[arXiv:1801.08549](#)] [[INSPIRE](#)].
- [15] E. D'Hoker and P. Kraus, *Gravitational Wilson lines in AdS_3* , [arXiv:1912.02750](#) [[INSPIRE](#)].

- [16] E. Witten, *Topology Changing Amplitudes in $(2+1)$ -Dimensional Gravity*, *Nucl. Phys. B* **323** (1989) 113 [[INSPIRE](#)].
- [17] S. Carlip, *Exact Quantum Scattering in $(2+1)$ -Dimensional Gravity*, *Nucl. Phys. B* **324** (1989) 106 [[INSPIRE](#)].
- [18] B.S. Skagerstam and A. Stern, *Topological Quantum Mechanics in $(2+1)$ -dimensions*, *Int. J. Mod. Phys. A* **5** (1990) 1575 [[INSPIRE](#)].
- [19] P. de Sousa Gerbert, *On spin and (quantum) gravity in $(2+1)$ -dimensions*, *Nucl. Phys. B* **346** (1990) 440 [[INSPIRE](#)].
- [20] C. Vaz and L. Witten, *Wilson loops and black holes in $(2+1)$ -dimensions*, *Phys. Lett. B* **327** (1994) 29 [[gr-qc/9401017](#)] [[INSPIRE](#)].
- [21] M. Ammon, A. Castro and N. Iqbal, *Wilson Lines and Entanglement Entropy in Higher Spin Gravity*, *JHEP* **10** (2013) 110 [[arXiv:1306.4338](#)] [[INSPIRE](#)].
- [22] J. de Boer and J.I. Jottar, *Entanglement Entropy and Higher Spin Holography in AdS_3* , *JHEP* **04** (2014) 089 [[arXiv:1306.4347](#)] [[INSPIRE](#)].
- [23] A. Castro and E. Llambres, *Unravelling Holographic Entanglement Entropy in Higher Spin Theories*, *JHEP* **03** (2015) 124 [[arXiv:1410.2870](#)] [[INSPIRE](#)].
- [24] J. de Boer, A. Castro, E. Hijano, J.I. Jottar and P. Kraus, *Higher spin entanglement and $\mathcal{W}_N$ conformal blocks*, *JHEP* **07** (2015) 168 [[arXiv:1412.7520](#)] [[INSPIRE](#)].
- [25] A. Hegde, P. Kraus and E. Perlmutter, *General Results for Higher Spin Wilson Lines and Entanglement in Vasiliev Theory*, *JHEP* **01** (2016) 176 [[arXiv:1511.05555](#)] [[INSPIRE](#)].
- [26] B. Chen and J.-q. Wu, *Higher spin entanglement entropy at finite temperature with chemical potential*, *JHEP* **07** (2016) 049 [[arXiv:1604.03644](#)] [[INSPIRE](#)].
- [27] A. Bhatta, P. Raman and N.V. Suryanarayana, *Holographic Conformal Partial Waves as Gravitational Open Wilson Networks*, *JHEP* **06** (2016) 119 [[arXiv:1602.02962](#)] [[INSPIRE](#)].
- [28] A. Castro, N. Iqbal and E. Llambres, *Eternal Higher Spin Black Holes: a Thermofield Interpretation*, *JHEP* **08** (2016) 022 [[arXiv:1602.09057](#)] [[INSPIRE](#)].
- [29] M. Henneaux, W. Merbis and A. Ranjbar, *Asymptotic dynamics of AdS_3 gravity with two asymptotic regions*, *JHEP* **03** (2020) 064 [[arXiv:1912.09465](#)] [[INSPIRE](#)].
- [30] R. Basu and M. Riegler, *Wilson Lines and Holographic Entanglement Entropy in Galilean Conformal Field Theories*, *Phys. Rev. D* **93** (2016) 045003 [[arXiv:1511.08662](#)] [[INSPIRE](#)].
- [31] A. Castro, D.M. Hofman and N. Iqbal, *Entanglement Entropy in Warped Conformal Field Theories*, *JHEP* **02** (2016) 033 [[arXiv:1511.00707](#)] [[INSPIRE](#)].
- [32] T. Azeyanagi, S. Detournay and M. Riegler, *Warped Black Holes in Lower-Spin Gravity*, *Phys. Rev. D* **99** (2019) 026013 [[arXiv:1801.07263](#)] [[INSPIRE](#)].
- [33] S. Carlip, *The Sum over topologies in three-dimensional Euclidean quantum gravity*, *Class. Quant. Grav.* **10** (1993) 207 [[hep-th/9206103](#)] [[INSPIRE](#)].
- [34] E. Guadagnini and P. Tomassini, *Sum over the geometries of three manifolds*, *Phys. Lett. B* **336** (1994) 330 [[INSPIRE](#)].
- [35] M. Bañados, T. Brotz and M.E. Ortiz, *Quantum three-dimensional de Sitter space*, *Phys. Rev. D* **59** (1999) 046002 [[hep-th/9807216](#)] [[INSPIRE](#)].

- [36] M.-I. Park, *Symmetry algebras in Chern-Simons theories with boundary: Canonical approach*, *Nucl. Phys. B* **544** (1999) 377 [[hep-th/9811033](#)] [[INSPIRE](#)].
- [37] T.R. Govindarajan, R.K. Kaul and V. Suneeta, *Quantum gravity on dS_3* , *Class. Quant. Grav.* **19** (2002) 4195 [[hep-th/0203219](#)] [[INSPIRE](#)].
- [38] A. Castro, N. Lashkari and A. Maloney, *A de Sitter Farey Tail*, *Phys. Rev. D* **83** (2011) 124027 [[arXiv:1103.4620](#)] [[INSPIRE](#)].
- [39] D. Anninos, S.A. Hartnoll and D.M. Hofman, *Static Patch Solipsism: Conformal Symmetry of the de Sitter Worldline*, *Class. Quant. Grav.* **29** (2012) 075002 [[arXiv:1109.4942](#)] [[INSPIRE](#)].
- [40] C. Beasley, *Localization for Wilson Loops in Chern-Simons Theory*, *Adv. Theor. Math. Phys.* **17** (2013) 1 [[arXiv:0911.2687](#)] [[INSPIRE](#)].
- [41] N. Ishibashi, *The Boundary and Crosscap States in Conformal Field Theories*, *Mod. Phys. Lett. A* **4** (1989) 251 [[INSPIRE](#)].
- [42] J.L. Cardy, *Boundary conformal field theory*, [hep-th/0411189](#) [[INSPIRE](#)].
- [43] R. Bousso, A. Maloney and A. Strominger, *Conformal vacua and entropy in de Sitter space*, *Phys. Rev. D* **65** (2002) 104039 [[hep-th/0112218](#)] [[INSPIRE](#)].
- [44] D. Harlow and D. Stanford, *Operator Dictionaries and Wave Functions in AdS/CFT and dS/CFT*, [arXiv:1104.2621](#) [[INSPIRE](#)].
- [45] A. Lopez-Ortega, *Quasinormal modes of D-dimensional de Sitter spacetime*, *Gen. Rel. Grav.* **38** (2006) 1565 [[gr-qc/0605027](#)] [[INSPIRE](#)].
- [46] A. Chatterjee and D.A. Lowe, *dS/CFT and the operator product expansion*, *Phys. Rev. D* **96** (2017) 066031 [[arXiv:1612.07785](#)] [[INSPIRE](#)].
- [47] E. Witten, *Quantization of Chern-Simons Gauge Theory With Complex Gauge Group*, *Commun. Math. Phys.* **137** (1991) 29 [[INSPIRE](#)].
- [48] J. Fjelstad, S. Hwang and T. Mansson, *CFT description of three-dimensional Kerr-de Sitter space-time*, *Nucl. Phys. B* **641** (2002) 376 [[hep-th/0206113](#)] [[INSPIRE](#)].
- [49] V. Balasubramanian, J. de Boer and D. Minic, *Notes on de Sitter space and holography*, *Class. Quant. Grav.* **19** (2002) 5655 [[hep-th/0207245](#)] [[INSPIRE](#)].
- [50] J. Cotler, K. Jensen and A. Maloney, *Low-dimensional de Sitter quantum gravity*, *JHEP* **06** (2020) 048 [[arXiv:1905.03780](#)] [[INSPIRE](#)].
- [51] H. Verlinde, *Poking Holes in AdS/CFT: Bulk Fields from Boundary States*, [arXiv:1505.05069](#) [[INSPIRE](#)].
- [52] M. Miyaji, T. Numasawa, N. Shiba, T. Takayanagi and K. Watanabe, *Continuous Multiscale Entanglement Renormalization Ansatz as Holographic Surface-State Correspondence*, *Phys. Rev. Lett.* **115** (2015) 171602 [[arXiv:1506.01353](#)] [[INSPIRE](#)].
- [53] Y. Nakayama and H. Ooguri, *Bulk Locality and Boundary Creating Operators*, *JHEP* **10** (2015) 114 [[arXiv:1507.04130](#)] [[INSPIRE](#)].
- [54] S. Deser and R. Jackiw, *Three-Dimensional Cosmological Gravity: Dynamics of Constant Curvature*, *Annals Phys.* **153** (1984) 405 [[INSPIRE](#)].
- [55] A. Strominger, *The dS/CFT correspondence*, *JHEP* **10** (2001) 034 [[hep-th/0106113](#)] [[INSPIRE](#)].

- [56] E. Witten, *Quantum gravity in de Sitter space*, in *Strings 2001: International Conference* (2001) [[hep-th/0106109](#)] [[INSPIRE](#)].
- [57] J.M. Maldacena, *Non-Gaussian features of primordial fluctuations in single field inflationary models*, *JHEP* **05** (2003) 013 [[astro-ph/0210603](#)] [[INSPIRE](#)].
- [58] A. Hamilton, D.N. Kabat, G. Lifschytz and D.A. Lowe, *Local bulk operators in AdS/CFT: A Boundary view of horizons and locality*, *Phys. Rev. D* **73** (2006) 086003 [[hep-th/0506118](#)] [[INSPIRE](#)].
- [59] A. Hamilton, D.N. Kabat, G. Lifschytz and D.A. Lowe, *Holographic representation of local bulk operators*, *Phys. Rev. D* **74** (2006) 066009 [[hep-th/0606141](#)] [[INSPIRE](#)].
- [60] X. Xiao, *Holographic representation of local operators in de Sitter space*, *Phys. Rev. D* **90** (2014) 024061 [[arXiv:1402.7080](#)] [[INSPIRE](#)].
- [61] D. Sarkar and X. Xiao, *Holographic Representation of Higher Spin Gauge Fields*, *Phys. Rev. D* **91** (2015) 086004 [[arXiv:1411.4657](#)] [[INSPIRE](#)].
- [62] D. Anninos, F. Denef, R. Monten and Z. Sun, *Higher Spin de Sitter Hilbert Space*, *JHEP* **10** (2019) 071 [[arXiv:1711.10037](#)] [[INSPIRE](#)].
- [63] K. Goto and T. Takayanagi, *CFT descriptions of bulk local states in the AdS black holes*, *JHEP* **10** (2017) 153 [[arXiv:1704.00053](#)] [[INSPIRE](#)].
- [64] C. Beasley and E. Witten, *Non-Abelian localization for Chern-Simons theory*, *J. Diff. Geom.* **70** (2005) 183 [[hep-th/0503126](#)] [[INSPIRE](#)].
- [65] C. Beasley, *Remarks on Wilson Loops and Seifert Loops in Chern-Simons Theory*, *AMS/IP Stud. Adv. Math.* **50** (2011) 1 [[arXiv:1012.5064](#)] [[INSPIRE](#)].