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Search for cosmic neutrinos with ANTARES

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Appendices

A. Spectral functions

Several spectral functions have been introduced in this thesis. This appendix lists the main point source spectral models used to fit gamma-ray data. Typical units are $\text{GeV}^{-1} \text{ cm}^{-2} \text{ s}^{-1}$.

- **Power-law.** The power-law function can be expressed as:

$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^\gamma. \quad (\text{A.1})$$

where N_0 is the normalization factor, E_0 the energy scale and γ the spectral index. In neutrino astronomy the default spectral index is $\gamma = -2$.

- **Broken power-law.** The broker power-law function has the form:

$$\frac{dN}{dE} = N_0 \times \begin{cases} (E/E_b)^{\gamma_1} & \text{if } E < E_b \\ (E/E_b)^{\gamma_2} & \text{otherwise} \end{cases} \quad (\text{A.2})$$

where E_b is the breaking energy, i.e. the energy at which the spectral index changes from γ_1 to γ_2 .

- **Smoothly broken power-law.** The function is a low energy power-law with spectral index γ_1 which smoothly changes its spectral index to γ_2 at breaking energy E_b :

$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^{\gamma_1} \left(1 + \left(\frac{E}{E_b} \right)^{\frac{\gamma_1 - \gamma_2}{\beta}} \right)^{-\beta} \quad (\text{A.3})$$

where β is the smoothness of the break.

- **Exponential cut-off.** The function follows a power-law and rapidly decreases up to a maximum energy E_c :

$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^\gamma e^{-E/E_c} \quad (\text{A.4})$$

B. Run selection

In Chapter 6 we have briefly discussed the selection of the runs used in the analysis. In this appendix we describe this selection with more details.

Depending on the detector status and the optical background during data taking, physics runs are divided in the following four categories, called also Quality Basic (QB) flags [177]:

- **QB = 1.** Basic selection of runs for physics analyses.
- **QB = 2.** At least 80% of active OMs.
- **QB = 3.** Baseline rate ≤ 120 Hz and burst fraction ≤ 0.4 .
- **QB = 4.** Baseline rate ≤ 120 Hz and burst fraction ≤ 0.2 .

These run sets are cumulative, meaning that runs satisfying the condition $QB = i$ with $i = 2,3,4$ satisfy the condition $QB = i - 1$ as well. All runs with $QB \geq 1$ were chosen for the analysis. Among other criteria, the runs with $QB \geq 1$ require these conditions:

1. Apparent run duration close to the effective run duration: $0 \leq T_{app} - T_{eff} \leq 450$ s.
The apperent duration of a run is the time window between the start and the end of that run. The efficient duration corresponds to the livetime of the run, i.e. the product of the number of recorded frames and the frame duration (104.858 ms).
2. No runs which present synchronization problems.
3. No runs with double frames.
4. Muon rate between 0.01 Hz and 100 Hz.

Figure B.1 shows the QB flag for these runs.

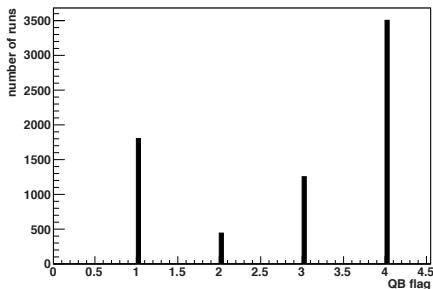


Figure B.1.: Number of selected runs for each of the QB flags.

B. Run selection

B.0.1. SCAN runs

For some runs the data taking conditions were not recorded in the database. This makes it very hard to reliably simulate these data. Hence, these runs, flagged as "SCAN", are not included in the final selection. The integrated livetime of the "SCAN" runs is roughly 60 days. It is likely that a large number of these runs were taken under stable conditions, but additional work is required to select such runs. Experts in the ANTARES collaboration are working towards a possible inclusion of "SCAN" runs in all the analyses.

C. Calibration and alignment

Each selected run is calibrated within the SeaTray framework (see for example Chapter five of [178] for a detailed description). The calibration is performed by reading all the necessary parameters from the ANTARES database. In this way for each run the most appropriate calibration set is chosen. Different calibrations were used for different data taking periods as summarised in Table C.1.

The alignment procedure is assigned using a standard calibration software (version 0.994). For a total of 82 runs no valid alignment was found. Hence, the reconstruction of these runs was not possible. However, this is not a big loss since the livetime covered by these runs is roughly few hours.

C. Calibration and alignment

First run	Last run	Label
25669	26796	2007:L5:v6.0
26770	27658	2007:L5:V6.0-bis
27659	28803	2007:L5:V6.1
28980	29761	2007:L5:V6.1-bis
29762	30427	2007:L5:V6.2
30508	31374	2007:L10:V7.0
31675	32491	2008:L10:V7.1
32529	33756	2008:L10:V7.2
34346	34976	2008:L12:V6.0
35000	36215	2008:L12:V6.1
36218	37475	2008::12:V6.2
37591	38759	2009:L12:V6.3-interlineoffset
38760	39589	2009:L12:V6.4
39590	40809	2009:L12:V6.6
40841	42425	2009:L12:V6.7
42477	42686	2009:L12:V7.1
42756	43282	2009:L12:V7.2
43285	44315	2009:L12:V7.3
45054	45565	2009:L12:V8.1
45459	47263	2010:L12:V2.0
47536	48064	2010:L12:V2.1
48484	49942	2010:L12:V2.2
50225	50955	2010:L12:V2.3
50958	52301	2010:L12:V2.4
52305	52853	2010:L12:V2.5
53074	53483	2010:L12:V2.7
53484	54045	2010:L12:V2.8
54049	54250	2010:L12:V2.9

Table C.1.: Versions of the calibration sets used and the range of runs which they were applied to.

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