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## From frictional to viscous behavior: Three-dimensional imaging and rheology of gravitational suspensions

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We probe the three-dimensional flow structure and rheology of gravitational (nondensity matched) suspensions for a range of driving rates in a split-bottom geometry. We establish that for sufficiently slow flows, the suspension flows as if it were a dry granular medium, and confirm recent theoretical modeling on the rheology of split-bottom flows. For faster driving, the flow behavior is shown to be consistent with the rheological behavior predicted by the recently developed “inertial number” approaches for suspension flows.

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### I. INTRODUCTION

Flows of granular materials submersed in a liquid of unequal density have started to attract considerable attention [1–5] and are relevant in many practical applications [6]. These materials, which we will refer to as “gravitational” suspensions, clearly differ from density matched suspensions, which have been studied in great detail [7–10]. Gravitational suspensions exhibit sedimentation, large packing fractions and jamming of the material, which suggests a description similar to dry granular matter [11,12].

In the last two decades, various flow regimes have been identified for dry granular matter. Sufficiently slow flows are *frictional*: the ratio of shear (driving) to normal (confining) stresses becomes independent of flow rate if the material is allowed to dilate [12,13]. Faster flows are referred to as *inertial*: here the effective friction coefficient  $\mu$  depends on the so-called “inertial” number  $I$ , which is a nondimensional measure of the local flow rate [12,14,15].

For gravitational suspensions, the presence of liquid instead of gas as interstitial medium strongly affects the microscopic picture—how should we think of the flow of such suspensions? Pouliquen and co-workers proposed that the ratio of the strain rate and settling time,  $I_S$ , would play a similar role as the inertial number in dry granular flows [5]. They furthermore conjectured a dependence of the effective friction coefficient  $\mu$  on  $I_S$  similar to the dry case, and applied this rheological law to capture the behavior of underwater avalanches [16].

Here we test this picture by combining three-dimensional (3D) imaging and rheological measurements of the flow of gravitational suspensions in a so-called split-bottom geometry (Fig. 1). This geometry has two main advantages. First, the flow rate, which is the key control parameter in the inertial number framework, can be varied over several orders of magnitude, allowing us to access slow flows as seen in plane shear [3,17], faster flows as seen in gravity driven flows [5,18], and the crossover regime in between—something not achieved in previous studies of gravitational suspensions [3,5,17,18]. Second, extensive experimental and numerical work [19–24] has shown that the split-bottom geometry produces highly nontrivial slow dry granular flows. A simple

frictional picture is not sufficient to capture these flows [25,26], so that testing whether these profiles also arise in slowly sheared gravitational suspensions is a stringent test for similarities between slow dry flows and slow gravitational suspension flows.

### II. SETUP

The split-bottom geometry is sketched in Fig. 1(a), and consists of a square box, 15 cm in width with transparent acrylic walls, at the (rough) bottom of which a (rough) disk of radius  $R_s=4.5$  cm can be rotated at rate  $\Omega$ .

We use monodisperse acrylic spheres with a diameter  $d$  of 4.6 mm (Engineering Laboratories); all our results are qualitatively similar for 3.2 mm particles. The particles are suspended in a mixture of some 78% Triton X-100, 13 % water, and 9% ZnCl<sub>2</sub> (by weight) [31] with a fluorescent dye added (Nile Blue 690). The refractive indices of particles and fluid are approximately 1.49 and match closely—we adapted the

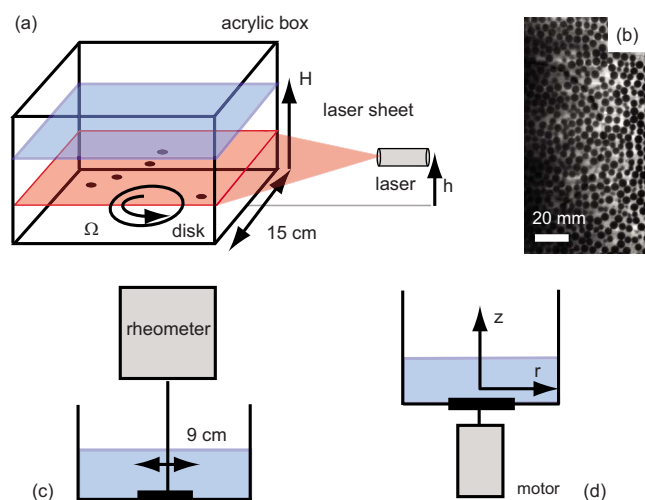


FIG. 1. (Color online) (a) The experimental setup used for flow visualization in a “split-bottom” geometry. (b) Example image of a single cross section, displaying half the box. (c) Geometry used for rheological measurements. (d) Geometry used for flow visualization.

recipe from [28]. The fluids viscosity  $\eta_f$  is  $0.3(\pm 0.05)$  Pa s, and the difference in density between the fluid and the particles is about  $100 \text{ g/m}^3$ .

The particle motion is visualized by illuminating the suspension with a thin ( $<200 \text{ }\mu\text{m}$ ) laser sheet [3,29,30]. The laser (Stocker Yale, 635 nm) is aligned parallel to the bottom of the box [Fig. 1(a)] and mounted on a z-stage which allows the illumination of slices of the suspension at different heights  $h$  [30]. Image acquisition is done with a triggered 12 bit cooled charge coupled device camera, and contrast is sufficient to image half of the box [Fig. 1(b)]. We use a particle-image-velocimetry-like (PIV) method to obtain the normalized azimuthal velocity  $\omega(r, z) = v_\theta / (r\Omega)$  in slices of constant  $z$ . Combining these slices, we reconstruct the full angular velocity field as function of radius and depth for a range of driving rates. An overview of the imaging technique will be published elsewhere [31].

Rheological experiments were carried out by driving the disk from above with a rheometer (Anton Paar MCR 501)—see Fig. 1(c). Velocimetry measurements were done by driving the disk from below with a dc motor—see Fig. 1(d). There is always at least half a centimeter of fluid above the suspension to ensure that the surface tension of the fluid will not affect the dilation [13] of the packing.

### III. CONSTITUTIVE EQUATION

We derive the constitutive equation for our suspension in the modified “inertial number” approach [2]. The typical rearrangement time scale for the particles in the suspension, given the viscosity and relative density of the particles becomes  $d/v_{\text{inf}} = \eta_f / P\alpha$ , where  $v_{\text{inf}}$ ,  $P$ , and  $\alpha$  are settling velocity, pressure and porosity, so that the inertial number becomes:  $I_S = \frac{\dot{\gamma}\eta_f}{P\alpha}$  [2]. The shear stress  $\tau$  is then written as  $\tau = \mu(I)P$ , with  $\mu(I)$  an empirical friction function. For small  $I$ ,  $\mu(I)$  can be expanded:  $\mu(I) = \mu_0 + \mu_1 I$  [2,5,32], with  $\mu_0$  and  $\mu_1$  empirical values. Combining this with the expression for  $I_S$ , we arrive at

$$\tau = \mu_0 P + \mu_1 \frac{\eta_f \dot{\gamma}}{\alpha}. \quad (1)$$

Thus, to lowest order, the local stress in a gravitational suspension is a linear combination of a frictional stress and a purely viscous stress [5]. This is reminiscent of the rheology of a Bingham fluid, in that slow flows are rate independent while faster flows become dominated by simple viscous drag. There is, however, a crucial difference: for slow driving, the shear stresses are predicted to be *proportional* to the pressure, while only for faster flows, the shear stresses become asymptotically *independent* of pressure.

### IV. FLOW PROFILES

In Figs. 2(a)–2(f) we compare, for a range of driving rates, the measured flow fields,  $\omega_S(r, z)$  (panel b–e) with predicted flow fields for dry granular media (panel a) and Newtonian flow (panel f). We fix the particle filling height at 23 mm ( $H/R_S \approx 0.5$ ). Clearly the flow structure progressively changes for faster flow rates, as the defining charac-

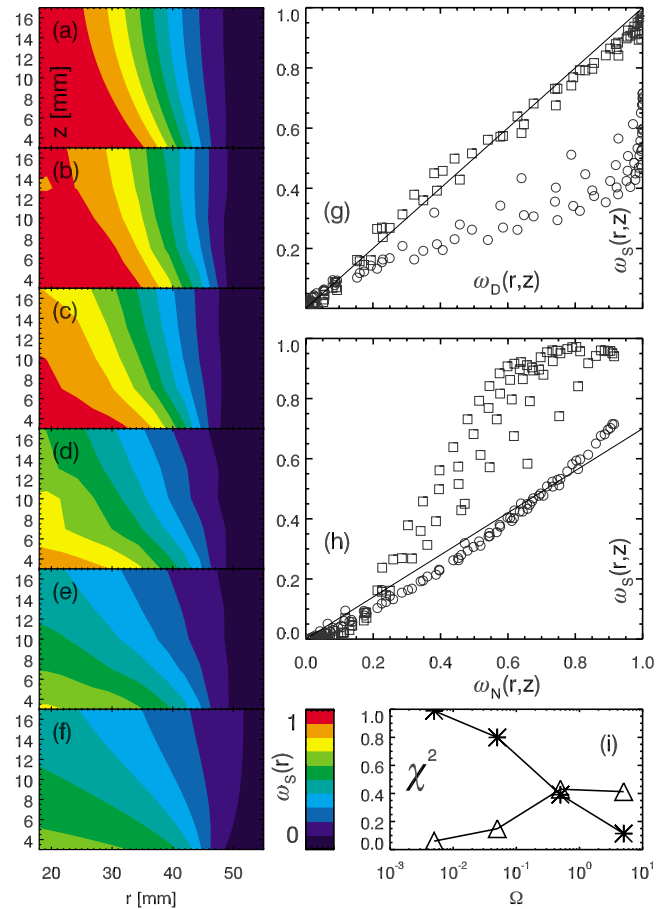


FIG. 2. (Color online) Frictional to viscous crossover in flow profiles. (a) The predicted flow field for dry granular flows  $\omega_D(r, z)$  from Refs. [19–25]. (b–e) Measured velocity fields  $\omega_S(r, z)$  at driving rates  $\Omega = 8.3 \times 10^{-5}$  rps (b),  $\Omega = 8.3 \times 10^{-4}$  rps (c),  $8.3 \times 10^{-3}$  rps (d) and  $8.3 \times 10^{-2}$  rps (e). (f) The Newtonian flow field  $\omega_N(r, z)$  calculated with the finite element method (see text). Note the similarity of (a) to (b) and (e) to (f). (g) Scatter plot comparison of  $\omega_D(r, z)$  (a) and  $\omega_S(r, z)$  for  $\Omega = 8.3 \times 10^{-5}$  rps ( $\square$ ) and  $\Omega = 8.3 \times 10^{-2}$  rps ( $\circ$ ). (h) Scatter plot comparison of  $\omega_S(r, z)$  and the flow field of a Newtonian flow  $\omega_N(r, z)$  for  $\Omega = 8.3 \times 10^{-5}$  rps ( $\square$ ) and  $\Omega = 8.3 \times 10^{-2}$  rps ( $\circ$ ). (i)  $\chi^2$  vs  $\Omega$  for comparison to granular ( $\triangle$ ) and Newtonian flow ( $\star$ ). The flow characteristics change from granular to Newtonian with increasing shear rate.

teristic of slow flows, the trumpetlike corotating inner core, disappears completely. This change is qualitative in nature, with a transition from concave to convex shapes of the isovelocity lines. In addition we note an increase of slip near the driving disk—while for slow flows, the normalized angular velocity  $\omega_S$  reaches 1 near the disk, for the fastest flow  $\omega_S$  has a maximum of 0.7.

The predicted flow field  $\omega_D(r, z)$  for slow dry flows with  $H/R_S \approx 0.5$  is shown in Fig. 2(a)—see Eqs. (1), (2), (6), and (7) in Ref. [25]. The similarity to the slowest flow profile, Fig. 2(b),  $\Omega = 8.3 \times 10^{-5}$  rps, is striking, and is confirmed in a scatter plot of  $\omega_S(r, z)$  vs  $\omega_D(r, z)$ , where all data for  $\Omega = 8.3 \times 10^{-5}$  rps (square) collapses on a straight line—see Fig. 2(g). We conclude that the flow profiles of slowly

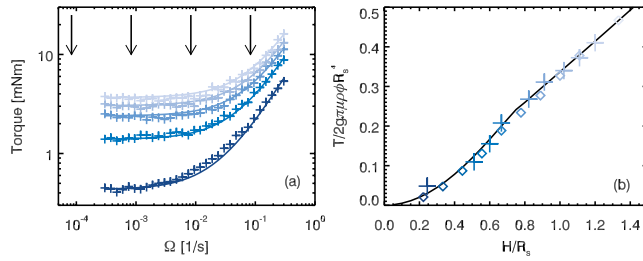


FIG. 3. (Color online) (a) Rheology of 4.6 mm acrylic particles in pure Triton X-100.  $T(\Omega)$  for a  $H/R_s = 0.24, 0.51, 0.60, 0.67, 0.82, 0.91, 1.0, 1.1, 1.2$ ; color (intensity) indicates  $H/R_s$ .  $\Omega = 3.0 \times 10^{-4}$  to 0.3. The curve is a fit of the form  $T = T_0 + C\Omega$  (see text). The four arrows indicate the driving rates where flow profiles were measured. (b)  $T_0(H)$ , the plateau values as a function  $H$  compared to the prediction from Eq. (2) with  $\mu = 0.57 \pm 0.03$  for the dry ( $\diamond$ ) and  $\mu = 0.59 \pm 0.03$  for the suspension ( $+$ ) case.

sheared gravitational suspension and dry granular media are indistinguishable.

The predicted flow field  $\omega_N(r, z)$  for Newtonian flows with  $H/R_s \approx 0.5$  is determined by a finite element software package (COMSOL) to solve the steady state Navier-Stokes equations for an incompressible Newtonian fluid, and is shown in Fig. 2(f). The similarity between the measured suspension flows for large  $\Omega$  and the Newtonian flow is, again, striking, and is confirmed in a scatter plot of  $\omega_S(r, z)$  vs  $\omega_N(r, z)$ , where all data for  $\Omega = 8.3 \times 10^{-2}$  rps (circles) tends to a straight line [Fig. 2(h)].

The crossover from frictional to viscous behavior can be quantified further by calculating as function of  $\Omega$  the total mean squared deviation ( $\chi^2$ ) obtained from a linear fit of the measured flow profiles  $\omega_S(r, z)$  to the predicted dry [ $\omega_D(r, z)$ ] and viscous [ $\omega_N(r, z)$ ] flows, as shown in Fig. 2(i). We conclude that the flow profiles of gravitational suspensions show a crossover from frictional, granular behavior to viscous flow upon increasing the driving rate.

## V. RHEOLOGY

Can we find the same crossover between these two regimes in the rheology? We measure the average driving torque as a function of filling height  $H$  and driving rate  $\Omega$ , in order to connect the rheology to the findings for the flow profiles discussed above [27]. Since index matching is not necessary, we use pure Triton X-100 as interstitial fluid; we use the same particles as before and keep the temperature fixed at 25 °C. In each experiment,  $\Omega$  is incremented from low to high values; each data point is obtained by averaging over three or more rotations (transients occur over much smaller strains).

In Fig. 3(a) we show  $T(H, \Omega)$  for several different suspension filling heights. We conclude that the trends in the rheology are similar for all filling heights. First, we observe a rate independent regime at small  $\Omega$ , which corresponds to the range where we observed flow profiles similar to the dry case. Moreover, the overall stress depends on filling height, which we will show below to be consistent with a pressure

dependence. Second, the stresses become rate dependent for  $\Omega \sim 0.01$  rps, and for larger rotation rates, the torque increases linearly with  $\Omega$ ; over the whole range of driving rates explored, the rheological data can be well fitted as  $T = T_0 + C\Omega$ , consistent with the behavior predicted by Eq. (1) [27]. We note here that a comparison of the measured torque for pure Triton and for the suspension yields that the effective viscosity of the suspension is only three to five times larger than  $\eta_f$ . This is far below than what would naively be expected from textbook formulas, e.g., Krieger-Dougherty. We have no explanation for this, but note that in the non-trivial split-bottom geometry, the suspension packing fraction varies throughout the material [25], which complicates the analysis.

We will now show that the height dependent torque for slow flows,  $T_0(H)$ , is well described by a prediction originally developed for slow dry flows [Fig. 3(b)]. From Eq. (1) it follows that the rheology should be determined by the local hydrostatic pressure and an effective friction coefficient  $\mu_0$ . Unger and co-workers [33] used these ingredients to predict  $r(z)$ , the center of the shear band of the dry split-bottom flow profiles, but their model also gives a prediction for  $T(H)$ ,

$$T(H) = 2g\pi\rho\phi\mu_0 \int_0^H (H-z)r^2 \sqrt{1 + (dr/dz)^2} dz. \quad (2)$$

Here  $\rho$  is the density of the particles, corrected for buoyancy in case of submersed particles,  $\phi$  is the average packing fraction ( $\sim 0.59$  [12]) and  $\mu_0$  is the effective friction coefficient. Minimization of Eq. (2) yields a prediction for  $T(H)$  which has not been tested previously.

As shown in Fig. 3(b) this prediction agrees very well with our measurements. The single fit parameter in the model allows to accurately extract a friction coefficient, which we estimate as  $\mu_0 \sim 0.59 \pm 0.03$ . We carried out the same measurement of  $T(H)$  on dry acrylic particles [Fig. 3(b)] and obtain a friction coefficient of  $\mu_0 = 0.57 \pm 0.03$ . The two friction coefficients are identical to within the experimental error, a fact also observed in Ref. [17]. This is strong evidence that in the slow driving rate limit the suspension behaves as a dry granular material and that lubrication and other hydrodynamic effects can be ignored. Furthermore, we can conclude that the simple frictional model by Unger correctly captures the overall stresses.

We have also tested the scaling in the viscous regime, by measuring the rheology of glass beads ( $\rho = 2.5 \times 10^3$  kg/m<sup>3</sup>) immersed to  $H/R_s = 0.4$  in glycerol for temperatures between 4 and 37 °C. The viscosity of the glycerol mixture varies more than a decade over this temperature range, and hence should change the rotation rate at which the viscous regime sets in. The results are shown in Fig. 4. Equation (1) requires that the data can be rescaled with the viscosity of the liquid  $\eta_f$ —this is indeed observed in the inset of Fig. 4. Note that the growth of torque with strain rate over the larger range probed here is somewhat slower than the simple linear prediction [27].



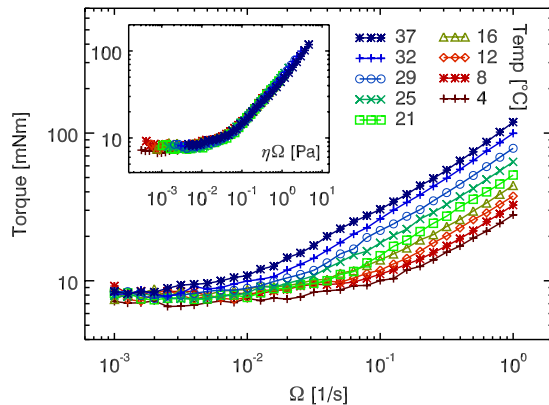


FIG. 4. (Color online) The rheology of glass beads in a glycerol mixture at different temperatures. The inset shows the same data, with the abscissa rescaled with the viscosity of the glycerol at the given temperatures.

## VI. CONCLUSIONS

Our main finding is that with increasing shear rate, a gravitational suspension crosses over from flowing like a dry granular material to flowing like a viscous liquid, consistent with recent modeling of suspensions based on the inertial number approach. We observe this both in the full three-

dimensional flow profile, which we revealed using an index matched scanning technique, and in rheological measurements. Most of our data can be understood based on simple scaling arguments (to obtain the “transition” shear rate) or elegant minimization principles (to obtain  $\mu$  from  $T(H)$ ). Our measurements indicate that the shape and width of the shear band in slow suspensions are the same as for slow dry granular flows. Whatever the physics beyond friction necessary to produce these flow profiles, our data shows that it is equally present in both dry granular and gravitational suspension flows. Still, a simple physical argument for the most prominent feature of split-bottom shear flows—the large width and error function shape of the shear zone [19–23]—remains elusive.

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