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COMMUNICATION FROM THE OBSERVATORY AT LEIDEN.

Discussion of 122, mostly new, variable stars in or near the constellation Corona Australis,

by *H. van Gent*.

1. This paper is a continuation to *B. A. N.* 227. The 66 variable stars, mostly cluster variables, discussed there were considered to be only a fraction of the variables which could be found on the available plate material. This conclusion was arrived at from the few rediscoveries amongst the objects found by comparing 9 pairs of plates in the blink microscope of the Union Observatory, Johannesburg, South Africa (*B. A. N.* 227, Vol. VI, p. 164). Therefore it was thought worth while to continue the search for variables on the same plate material. For this purpose 22 more pairs of plates have been compared by the writer in the blink microscope of the Kapteyn Astronomical Laboratory at Groningen. The blink microscope, constructed by Zeiss, is in its essential parts fully identical to the one at the Union Observatory, so that the discoveries of variable stars with both instruments can be added together into one homogeneous collection. In total 200 objects were added to the 109 already discovered with the Johannesburg blink microscope. Three of these proved to be identical with two new asteroids which during the search were noted as variable stars too faint to be visible on one plate of the plate pair in examination. A number of objects proved to be only false objects and particles of dust. Some objects were found to be certainly variable but were so faint that they could be estimated only on the best plates of the plate collection, and only when they were in maximum light. They have been excluded from the discussion. When an object after a few estimates turned out to be a long period variable, it was put aside, as the interval of time covered by the series of plates, viz. 2694 days, is too short, and the plates are unfavourably distributed for a good determination of the mean period and the mean light curve for this type of variable. Only when there was a hope of the star being a δ Cep variable with a long period the estimates have been carried out on the whole

plate material. In this way 7 long period variables have been estimated on all the plates. For these an exception was made and they have been included in the main list (Table 3). In total 122 variables have been listed; they are distributed over the different types of variability as follows:

- 97 cluster variables
- 12 eclipsing variables
- 7 long period variables
- 2 δ Cep variables
- 2 irregular variables
- 1 SS Cyg type variable
- 1 ultra short period variable.

2. The results of the comparison of the 22 plate pairs in the Groningen blink microscope have been collected in Table 1. The first column contains the current number of the variable, the second column the type of variability, the third column the plate pair on which the variable was discovered with the indication whether it was found bright on the earlier plate (*e*) or on the later plate (*l*). The fourth column contains the interval between the J. D. Hel. Grw. Astr. M. T. of the two plates and the fifth column gives the observed difference in steps of the brightness of the variable on the two plates. If the variable has been found more than once the plate pairs on which it was rediscovered have been added in the sixth column.

This table is a continuation to Table 1 in *B. A. N.* 227. Three variables of the new list, viz. No. 35, 42 and 221, although rediscovered by me at Groningen, had already been found by me in the Johannesburg blink microscope. A number of the collection of variables in *B. A. N.* 227 have been rediscovered by me at Groningen; an account of these rediscoveries is given in Table 1a. This table completes the sixth

TABLE I.

var. No.	type	found on plate pair	Δd	Δs	also found on plate pair	var. No.	type	found on plate pair	Δd	Δs	also found on plate pair	var. No.	type	found on plate pair	Δd	Δs	also found on plate pair
35	cluster	D, l	26.99	10.4	Q	170	cluster	O, e	1.06	5.1		217	cluster	T, l	3.98	9.0	
42	cluster	D, l	26.99	7.8	K, W, Y, Z	172	cluster	O, l	1.06	4.1		218	cluster	T, l	3.98	6.5	U
112	cluster	J, l	.96	9.3		173	cluster	O, l	1.06	6.8		219	cluster	T, l	3.98	7.6	
115	cluster	J, l	.96	8.8		174	cluster	O, e	1.06	8.5	V, W	220	cluster	U, l	3.01	7.8	
118	W UMa	J, l	.96	5.8		176	cluster	O, l	1.06	10.2		221	cluster	B, e	4.89	11.0	U
119	cluster	J, l	.96	9.7	U	177	Algol	O, l	1.06	5.1		222	cluster	U, l	3.01	12.0	
120	cluster	J, l	.96	7.8		179	cluster	P, e	.97	7.0	T	223	cluster	U, e	3.01	5.9	
121	cluster	J, e	.96	5.2		180	cluster	P, e	.97	3.5		224	cluster	U, e	3.01	12.5	
122	cluster	J, l	.96	7.8		181	cluster	P, e	.97	6.7		226	irregular	U, l	3.01	5.3	
123	cluster	J, e	.96	5.3	W, Z	182	cluster	P, e	.97	3.2	α, δ	231	cluster	W, e	2.99	3.0	
124	irregular	J, e	.96	9.9		183	cluster	P, e	.97	6.5	U, δ	232	Algol	W, e	2.99	19.2	
125	cluster	J, l	.96	11.5		184	cluster	P, e	.97	5.3		235	cluster	W, e	2.99	8.7	
126	cluster	J, l	.96	5.2	P	186	cluster	P, e	.97	8.5	U	236	cluster	W, e	2.99	7.5	
128	cluster	K, l	.70	11.5		187	cluster	P, e	.97	3.5		237	cluster	W, l	2.99	8.6	
130	cluster	K, e	.70	5.1	S	188	cluster	P, e	.97	3.1	U	238	cluster	W, l	2.99	8.3	
131	δ Cep	K, e	.70	10.3		189	cluster	Q, l	3.74	6.9		239	cluster	W, l	2.99	10.4	α
133	Algol	K, e	.70	12.5	M	190	cluster	Q, l	3.74	6.8	ϵ	240	cluster	W, e	2.99	8.8	Y
134	cluster	K, e	.70	10.3	S	191	cluster	Q, l	3.74	7.7	R	242	W UMa	J, l	.96	3.6	W
135	cluster	K, e	.70	2.5	Q, W	192	cluster	Q, e	3.74	9.8		245	cluster	X, l	5.30	3.3	δ
136	δ Cep	L, l	1.83	12.3	R, U	193	cluster	Q, e	3.74	9.1		246	Algol	X, l	5.30	12.0	
138	cluster	L, l	1.83	8.5		194	cluster	Q, l	3.74	8.7		248	cluster	X, l	5.30	5.9	
139	cluster	L, e	1.83	7.9	α	195	cluster	R, l	.79	7.3		250	cluster	X, l	5.30	7.9	Y
141	cluster	L, l	1.83	8.8		196	cluster	R, e	.79	8.6	W	253	SS Cyg	V, e	8.01	14.7	Y, δ
143	cluster	L, l	1.83	10.8	N	198	cluster	R, e	.79	5.4		255	cluster	Y, l	3.08	8.3	
144	cluster	L, l	1.83	7.7		199	cluster	S, e	1.21	7.5	V	268	cluster	α, l	6.17	2.0	
145	cluster	L, l	1.83	9.5		200	cluster	S, l	1.21	11.7	W	272	?	α, l	6.17	5.9	
146	cluster	L, l	1.83	10.8	N, Q, Z	201	cluster	S, l	1.21	9.5	W	273	cluster	α, l	6.17	5.7	
147	cluster	L, e	1.83	10.9	N, T, U, V	202	cluster	S, l	1.21	7.7		274	longper.	β, e	14.86	10.0	
149	W UMa	L, l	1.83	4.6		203	cluster	S, l	1.21	5.4	W	277	longper.	β, l	14.86	7.0	
150	cluster	M, e	4.72	6.2	V	204	cluster	S, e	1.21	6.5	γ, δ	280	longper.	β, l	14.86	7.3	
154	cluster	M, e	4.72	8.1	T	205	cluster	T, e	3.98	9.1		283	cluster	γ, l	15.85	8.7	
155	cluster	M, e	4.72	9.2	O	206	Algol	T, l	3.98	5.3		288	longper.	γ, l	15.85	5.0	
157	Algol	M, l	4.72	7.8		207	cluster	T, l	3.98	9.6		300	longper.	δ, e	19.00	3.6	ϵ
160	Algol	N, e	.28	10.7	Q	208	cluster	T, l	3.98	6.6							
161	cluster	N, l	.28	>8.5	Q, V, W	209	cluster	T, l	3.98	10.5		302	Algol	δ, l	19.00	3.24	
162	cluster	N, e	.28	10.8	Q, S, Z, α, β, δ	210	cluster	T, e	3.98	8.6							
164	cluster	N, e	.28	9.0		211	W UMa	T, l	3.98	4.0	U	303	cluster	δ, l	19.00	4.3	
165	cluster	O, e	1.06	5.1		212	cluster	T, l	3.98	4.9		309	cluster	ϵ, l	17.85	5.7	
167	cluster	O, e	1.06	6.5		213	cluster	T, l	3.98	4.7	δ	310	cluster	ϵ, l	17.85	10.1	
168	cluster	O, l	1.06	11.0		214	cluster	T, l	3.98	10.2		312	longper.	ϵ, l	17.85	6.6	
169	cluster	O, l	1.06	9.0		215	cluster	T, l	3.98	5.1		316	longper.	ϵ, l	17.85	7.3	
						216	cluster	T, l	3.98	7.2	U, W, Z, ϵ						

TABLE 1a.

var. No.	also found on plate pair	var. No.	also found on plate pair
2	R	48	J
3	β	53	Z
4	Y, δ	57	M, O, T, U
5	M	60	N, R, U
8	Z	61	J, X
9	J	66	P
11	J, L, O, V, ϵ	67	Z
12	J, α, ϵ	72	N, R
15	T, W, X	79	T
18	Z	81	M, Z
20	N, S	84	Q
24	S	86	R
30	K	90	K, ϵ
33	J, N	93	K
43	P	104	X

column of Table 1, *B. A. N.* 227, up to date; together with the sixth column of Table 1 in this paper a complete account of the rediscoveries is obtained for the $66 + 122 = 188$ variables of *B. A. N.* 227 and this paper together.

Table 2 contains the J. D. Hel. Grw. Astr. Mean Times for each plate pair and the number of variables found on it. The last column of Table 2 contains the number of variables which had already been found on an earlier plate pair of the list. Table 2 in this paper is a continuation to Table 2 in *B. A. N.* 227; in order to bring this latter one up to date the number of variables found should be increased from 13 to 14 and from 6 to 8 for the plate pairs B and D respectively.

TABLE 2.

plate pair	J. D. Hel. Astr. Times of plate pair compared in blink microscope	Mean Greenwich	number of variables found	number of rediscoveries
J	2426561.4594	with 2426562.4211	18	6
K	6563.5266	6564.2308	10	4
L	6561.4809	6563.3119	11	1
M	6562.5070	6567.2228	8	4
N	6562.2492	6562.5285	11	7
O	6565.4490	6566.5134	12	4
P	6566.2765	6567.2443	12	3
Q	6562.4856	6566.2287	12	6
R	6563.4192	6564.2094	9	6
S	6562.2922	6563.5051	11	5
T	6562.2707	6566.2523	21	6
U	6563.2904	6566.2987	17	12
V	5499.3399	5507.3505	7	6
W	5794.5192	5797.5104	20	12
X	5437.2307	5442.5327	7	3
Y	5791.4180	5794.4970	6	5
Z	5441.3790	5442.5770	10	10
α	5385.4398	5391.6051	8	5
β	5507.3723	5522.2362	5	2
γ	5508.3851	5524.2388	3	1
δ	5480.2793	5499.2748	10	7
ε	5420.4951	5438.3471	10	6

3. In Table 3 the right ascension and declination for each variable have been given. They have been derived in the same way as described in *B. A. N.* 227. On pages 49 to 53 diagrams are given of the surroundings for all the variables. At the right hand bottom corner the size of each square in minutes of arc is indicated. The variable has been marked by an open circle; on each diagram the sequence of comparison stars used has been indicated. Variables No. 300 and 316 are in one diagram; they are so close together that the same set of comparison stars could be used for both of them.

Amongst the 188 variables of *B. A. N.* 227 and this paper the following identities with known or suspected variables have been found:

<i>B. A. N.</i> 227 ZINNER ¹⁾ INNES ²⁾			
var. 5	=	I472	= 48
4		I475	49
9			ZZ Sgr
3		I527	70
18			UV Cr A
11			N. G. C. 6723, BAILEY var. 4
1		I601	165
98		I635	99

this paper	ZINNER	INNES
var. 274	=	I453 = 40
283	?	I486 52 ^a

¹⁾ *Astr. Abh.*, Ergänzungsheft zu den *Astr. Nachr.*, Bd. 8 No. 1.
²⁾ *Union Observatory Circular* 37.

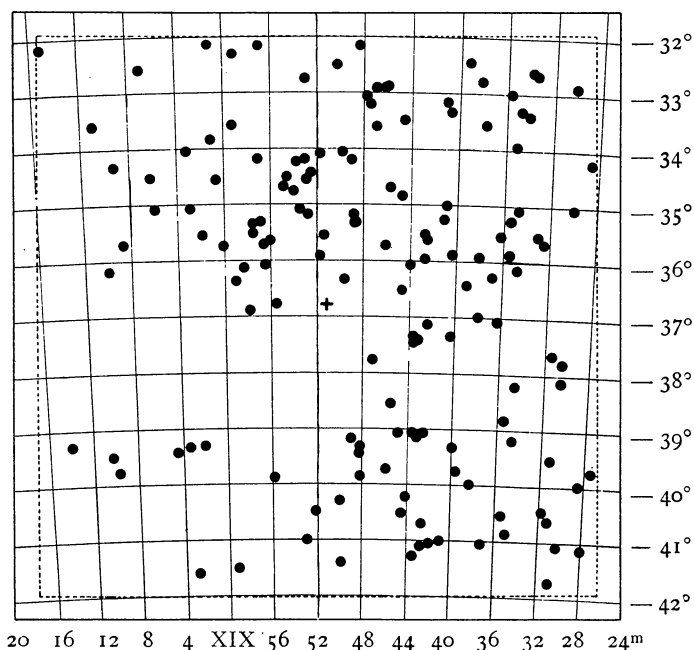
(continued)

this paper	ZINNER	INNES
200	I523	69
310		UW Cr A
186	I548	160
121 ?		AB Cr A
162		N. G. C. 6723, BAILEY var. 15
213		" " " I
214		" " " II
238		CF Sgr
300		VW Cr A
316		VX Cr A
239 ?	I593	163
240		V 413 Sgr
194	I623	170
302		BQ Sgr
174		V 343 Sgr
135		V 436 Sgr

Nearly all the variables are faint and it is not always possible to be sure of the identification. A sketch of the surroundings, as given on pp. 49 to 53, will preclude misidentifications. The periods given by INNES in *U. O. C.* 37 give no clue for the certainty of the identification, as they are often considered by INNES himself as doubtful, the material used by him being not very suitable for the determination of short periods.

The diagram on p. 165 in *B. A. N.* 227 showing the distribution of the cluster variables over the plate field is completed for the cluster variables of the list in this paper, as shown by fig. 1. As several plates

FIGURE 1.



have a slightly different centre, the diagram covers more than the field of one plate. The border of the field of the majority of the plates is indicated by a dotted line.

The cluster variables of *N. G. C.* 6723 have been omitted. The diagram again shows a marked concentration of the cluster variables towards the Milky Way. If a straight line is drawn through the centre of the plate field parallel to the Milky Way, the number of cluster variables on the half of the plate that is nearest to the Milky Way is 114 against 44 on the opposite half. This indicates a strong increase in the space density for these variables when coming nearer to the Milky Way.¹⁾ It should be remembered however that in the neighbourhood of the irregular variable R CrA there is a region of heavy obscuring clouds extending over several square degrees, which may hide a number of cluster variables lying behind it.

4. About the treatment of the material the remarks made in *B. A. N.* 77 by E. HERTZSPRUNG, in *B. A. N.* 194 by W. E. KRUYTBOSCH and in *B. A. N.* 227 by myself may be referred to.

The total number of estimates used in this paper is 34566 or on the average 283 for each variable. The variables 237, 193 and 126 were estimated also on the series of 308 plates centered on τ Sgr, discussed by FERWERDA in *B. A. N.* 231, these three variables occurring on the overlapping region of the two fields. The total number of photographs examined is $122 \times 397 + 924 = 49358$.

The main data for each variable have been collected in Table 3. The type of variability has been indicated in column 3. The type of the ultra short period variable No. 272 has been considered as uncertain. For several variables with nearly sinusoidal, symmetrical lightcurves it was difficult to be sure whether it is a W UMa or a cluster variable of BAILEY's c -type. The number of estimates is indicated in column 4; the period in column 5. When this period was derived by least squares, its mean error has been added. The reciprocal period used in the computing machine to derive the phases of the estimates, has been given in column 6. Column 7 gives the number of epochs used (maxima in the case of the cluster variables, δ Cep variables and the ultra short period variable No. 272, and minima in the case of the Algol and W UMa variables). In column 8 the mean error of a single epoch is given.

The data for the least squares solutions have been collected in Table 5. The table contains the epochs used, the counting of the periods and the values O—C. In the cases of the 4 W UMa variables the

¹⁾ See also OORT, *B. A. N.* 238 p. 279.

odd and the even minima have both been used in the least squares solution; the period derived is the apparent period and should be doubled in order to obtain the period of revolution.

The phases of the observations were computed by the formula: phase = (J. D. — 2420000) \times reciprocal period. The observations were then arranged according to phase and means were taken in order to construct the lightcurves, diagrams of which have been given on pp. 54 to 60. An account of this is given in Table 6. The columns contain respectively: the number of observations used in forming the mean; the mean phase; the mean brightness in steps. In the case of the bright variable No. 302 the brightness has been given in magnitudes.

In the case of the W UMa variable No. 242 there was a distinct difference in depth of the odd and the even minima. Accordingly for this variable the phases with odd and even whole number have been separated, and the lightcurve represents the revolution period which is twice the period given in Table 3.

In column 10 of Table 3 the phase of maximum has been given; for the Algol and W UMa variables this number gives the phase of (principal) minimum. If four decimals have been given, this phase has been derived by least squares, assuming the lightcurve to be symmetrical (see *B. A. N.* 147 p. 179).

Sharper epochs can be derived from the rising branches than from the maxima. A point on the rising branch was defined in the way described in *B. A. N.* 227, p. 166. This point is characterized by the difference ΔP between its phase and the phase of the corresponding point of the same brightness on the descending branch. This quantity ΔP has been given in column 9, as well as the phase of the point on the rising branch defined by it.

The period of a few variables proved to be variable. In those cases the light curve is given in Table 6 for the years 1924—1925 (I); 1928—30 (II) and 1931 (III) separately. From these the following quadratic terms have been derived:

var.	quadratic term
143	$-002 \times 10^{-6} E^2$
155	$+005 \times 10^{-6} E^2$
176	$-007 \times 10^{-6} E^2$
184	$+03 \times 10^{-6} E^2$
200	$-005 \times 10^{-6} E^2$
222	$+003 \times 10^{-6} E^2$

In column 12 the maximum and minimum brightness of the lightcurve has been given in steps (zero-point arbitrary) as well as the range in steps. For var. 302 this has been given in magnitudes.

The last column of Table 3 shows the mean error of a single estimate, derived by the formula:

$$\text{m. e. of estimate} = \pm \sqrt{\frac{\sum (\Delta s)^2}{2 \times n}},$$

in which Δs is the difference in steps of two estimates following each other in phase and n the number of estimates.

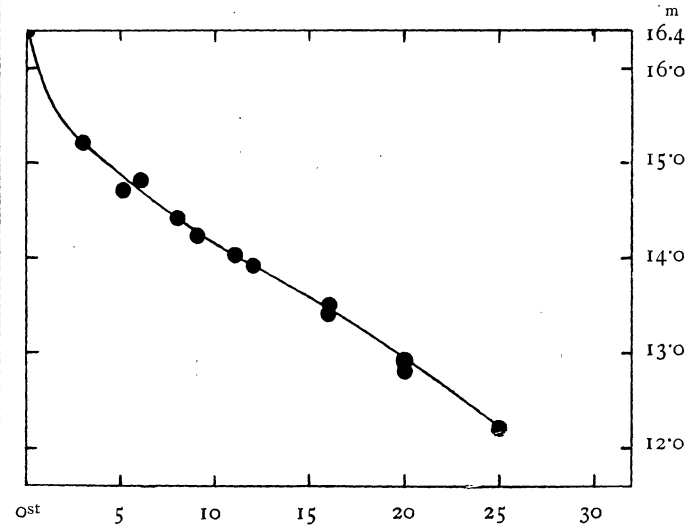
Column 11 contains the maximum and minimum brightness of the variables in magnitudes. In order to obtain these magnitudes the same procedure was followed as given in *B. A. N.* 227, p. 167. The brightness of all comparison stars was estimated in an absolute scale of steps on the same plate as used for this purpose in *B. A. N.* 227 taking the plate fog as zero point. For a number of these comparison stars the magnitude was derived by star counts on the enlargements used for measuring the positions of the variables (size of the enlargements: .27 square degree). The number of stars brighter than a comparison star were counted and this number was reduced to a square degree and compared with the tables for $\log N_{m, \beta, \lambda}$ in *Groningen Publ.* 43. The result is given in the following table:

variable No.	comparison star	brightness in absolute scale of steps	number of stars brighter than comparison star	$\log N_{m, \beta, \lambda}$	galactic longitude	galactic latitude	magnitude
160	a	20	32	2.08	330°	-16°	12.8 ^m
	b	16	61	2.36			13.4
	c	11	124	2.66			14.0
	d	8	198	2.87			14.4
	e	5	256	2.98			14.7
274	A	25	18	1.83	328°	-13°	12.2
	B	20	39	2.16			12.9
	a	16	75	2.45			13.5
	b	12	124	2.66			13.9
	c	9	162	2.78			14.2
	d	6	261	2.99			14.8
e	3	493	3.26	15.2			

The magnitudes were plotted against these estimates in steps in Fig. 2 and a curve has been drawn representing the relation between them. This curve is very similar to the one in *B. A. N.* 227 p. 167; the value of a step expressed in magnitudes has increased slightly however. The relation is linear till about 15^m; from there the value of a step increases rapidly.

The limiting magnitude of the plate is 16^m.4. In *B. A. N.* 235 W. CHR. MARTIN says that the limiting magnitude of the best plates of the series discussed there is about 15^m.5. These plates have been taken

FIGURE 2.



also with the Franklin-Adams telescope with an exposure time of 30^{min}, and are of the same brand, viz. Ilford Zenith, as the plates taken in 1928-31 of the series used here, so that the two plate series are fully comparable. Therefore one of the best plates of MARTIN's series was selected by myself and four squares of $\frac{1}{2}$ degree side were taken in the middles of the four squares into which the plate can be divided. All stars visible in them were counted; the total number obtained in this way was 3880 in an area of one square degree. Comparison with the tables of *Groningen Publ.* 43 gives 16^m.4 as the resulting limiting magnitude.

With the aid of the curve in Fig. 2 the estimates of the brightness of the comparison stars in steps from the plate fog have been reduced to magnitudes. These magnitudes are shown in the third column of Table 4.

For the derivation of these magnitudes a difference in limiting magnitude for the plate as mentioned above does no harm, as long as the same plate is used for the determination of the limiting magnitude and for the estimates of the brightness of the comparison stars in steps from the plate fog. It only shows that differences in sensitivity for different emulsions of the same plate brand may lower the limiting magnitude by an amount of nearly a magnitude. This explains why several faint variable stars, when discovered on a set of the most sensitive plates of the series, could not be estimated on the majority of the plates, as mentioned under 1.

The estimates of brightness of the variables were made in an arbitrary scale of steps, the variable being ordinarily enclosed between two of the comparison stars. Thus each plate yielded a value for the difference in steps between two successive comparison stars. The means of these differences have been taken in

order to obtain the brightness in steps for each comparison star of the sequence. These brightnesses have been given in column 2 of Table 4. The zero-point for each sequence is arbitrary.

5. Though the plate material is very suitable for cluster variables, several difficulties were encountered in determining the periods. In the following table the stars are given for which an erroneous period had at first been tried:

Var.	erroneous reciprocal period	correct reciprocal period	difference
	d^{-1}	d^{-1}	d^{-1}
1	3'298	2'298	1'000
5	3'095	2'092	1'003
7	3'223	2'220	1'003
8	2'930	1'927	1'003
13	2'499	1'496	1'003
18	3'116	2'114	1'002
47	2'644	1'642	1'002
48	3'153	2'151	1'002
189	3'0226	3'0198	'0028
208	2'691	1'689	1'002

In the plate material itself three periods are present, viz. the day, the synodical month and the year. If the plates cover only a small fraction of the day, resp. synodical month or year, there is always the possibility that for a variable a wrong period is derived of the type:

$$\frac{m}{P} \pm \frac{n}{P'} = \frac{1}{P''}$$

in which m and n are integers, and P , P' and P'' denote the true period, the period of the plate material (day, month or year) and the erroneously derived period respectively.

The month comes in because with the Franklin-Adams telescope ($\frac{O}{f} = \frac{1}{4.5}$) no observations could be taken when the moon was brighter than first or last quarter, the plate becoming too much fogged by scattered moonlight in half an hour exposure time and the variables being too faint to be obtained with a considerably shorter exposure time. The observations however generally cover more than half a month at a stretch, plates having been taken when the moon was still or already under the horizon. The year is the least dangerous periodicity, the fraction of the year occupied by observations being considerably more than half. Still in the case of var. 189 it was evidently the period of a year that gave rise to a wrong period. The other 9 cases arose from the daily period in the observations, and are all of the type $\frac{1}{P} + \frac{1}{P'} = \frac{1}{P''}$.

The differences between the two reciprocal periods are not exactly 1, as should be the case when the observations had been taken at intervals of exactly a mean solar day (observations taken at always the same mean time), but in the mean 1'0025. For observations taken at intervals of a sidereal day (observations taken at always the same hour angle) this difference should be 1'0027.

As the observations never extend over more than $10\frac{1}{2}$ hours at a stretch (the greatest eastern hour angle was $5^{\text{h}}02^{\text{m}}$, the greatest western one $5^{\text{h}}28^{\text{m}}$, both for the middle of the exposure time of a plate), it is the interval of the day that causes most of the troubles in determining the period. From observations all made at the same hour angle (or at the same mean time) it is impossible to derive the right period, as a wrong period of the type mentioned above will give the variable's correct lightcurve with the correct range and no more scattering of the points along the lightcurve than the correct period does.

A wrong period of the type $\frac{m}{P} - \frac{n}{P'} = \frac{1}{P''}$ will give a lightcurve which is symmetrical to the lightcurve derived from the right period with respect to the y -axis; for a symmetrical lightcurve (W UMa and Algol type for instance) this will not give a means to detect a wrong period of this type, but for a cluster variable it does.

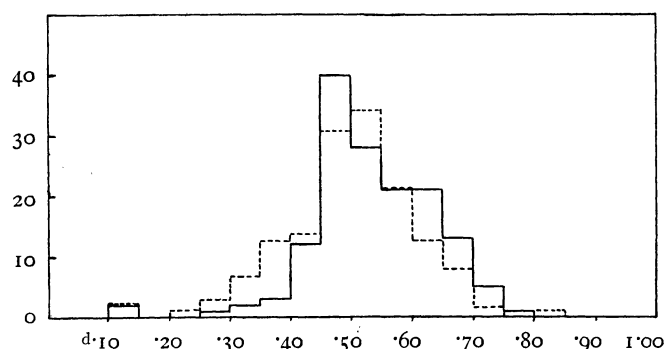
The troubles in determining the period of a variable when the observations have been made at regular intervals are essentially the same as in the case of spectroscopic binaries (see O. STRUVE, *Pop. Astr.* **36**, p. 411). Only observations taken at great hour angles east and west, or, much better, observations from observatories at greatly different longitudes will exclude wrong periods. As regarding the variables discussed in this paper the long series of plates taken during one night (up to $\cdot^{\text{d}}39$ difference between the first and the last plate of the same night, and up to 17 half hour exposures covering $\cdot^{\text{d}}35$ without other interruption than the time necessary for changing plates) proved indispensable in order to sift out the right period, and this measure of precaution cannot strongly enough be recommended.

6. As the number of cluster variables extracted from the plates is so much more complete now than in *B. A. N.* 227, a new comparison was made between the frequency of the periods of the cluster variables in PRAGER's catalogue 1932 and the collection of cluster variables of *B. A. N.* 227 and this paper together. The result is laid down in the following table. The variables in the globular cluster *N. G. C.* 6723 have been omitted.

Interval of period	number of variables in Prager's catalogue 1932	number of variables in Prager's catalogue 1932, reduced to 149	number of variables B.A.N. 227 and this paper
d d			
.10 — .15	4	2.3	2
.15 — .20	0	.0	0
.20 — .25	2	1.1	0
.25 — .30	5	2.9	1
.30 — .35	12	6.8	2
.35 — .40	22	12.6	3
.40 — .45	24	13.7	12
.45 — .50	54	30.8	40
.50 — .55	60	34.3	28
.55 — .60	37	21.1	21
.60 — .65	22	12.6	21
.65 — .70	14	8.0	13
.70 — .75	3	1.7	5
.75 — .80	0	.0	1
.80 — .85	2	1.1	0
	261	149.0	149

From the last two columns a graph (Fig. 3) was drawn showing the two distributions; full line: distribution from this paper; dotted line: distribution from PRAGER's catalogue.

FIGURE 3.



It is evident that the curve of this paper shows a preference for the longer periods. For periods under $d.40$ my curve has only 8 stars, the curve derived from Prager's catalogue 25.7 stars. For periods above $d.60$ these numbers are 40 and 23.4 respectively. It is possible that this difference (an excess of 17 stars for periods under $d.40$ against a shortage of 17 stars for periods above $d.60$ on a total of 149 stars) is due to wrong periods of the type described under 5 being present amongst the stars in PRAGER's catalogue. As shown by the table on p. 26, when two different periods come into consideration it is the shorter one that is likely to be wrong.

The frequency curve derived in *B. A. N.* 227, p. 168 shows a gap at $d.54$; this gap has disappeared in the curve derived here. The explanation of this gap is very probably as follows. Of the 57 cluster variables used for the curve in *B. A. N.* 227, 38 variables have been detected by comparing plate pairs

with intervals of $d.98$, $1d.03$, $1d.96$ and $3d.01$. These intervals are so close to 1, 2 and 3 days that stars with a period near $d.5$ had a very small chance of being detected. This holds especially for the plate pairs with difference $d.98$ and $1d.03$, by which 24 cluster variables were discovered. By also comparing plate pairs taken in very different hour angles (from $-d.36$ to $+d.30$ difference in hour angle) and comparing plates with a much longer interval, the material discussed here can be considered much more free from selective effect with respect to the period.

7. The richness of the plate field in cluster variables is conspicuous. Excluding the two ultra short period variables No. 78 and 272, and also the seven variables found in the globular cluster *N. G. C.* 6723, there have been found 147 cluster variables in the region, taking the discoveries of *B. A. N.* 227 and this paper together. The question will now be considered up to which degree the search has exhausted the number of cluster variables present on the plates. An idea can be got about this degree of exhaustion from the number of times that each cluster variable has been independently found. Denoting by α the probability that a variable is detected by comparing one plate pair, and by N the number of variables present in the plate field, all having the same discovery chance, and by n the number of plate pairs compared, the following numbers of variables will be found 0, 1, 2... n times:

$$a_0 = N(1-\alpha)^n \quad \dots \text{0 times found}$$

$$a_1 = N(1-\alpha)^{n-1} \alpha n \quad \dots \text{1 time found}$$

$$a_2 = N(1-\alpha)^{n-2} \alpha^2 \frac{n(n-1)}{1 \cdot 2} \quad \dots \text{2 times found}$$

etc.

$$a_n = N \alpha^n \quad \dots \text{n times found}$$

in which:

$$a_0 + a_1 + a_2 + \dots + a_n = N$$

Now the discovery chance α will certainly depend on the range of the lightcurve and the apparent magnitude of the variable. The lightcurves were assumed to be identical for the same range, which is a reasonable supposition to simplify the computation. As discussed under 5 the plate pairs have been selected in such a way that preference or avoidance for the discovery of a variable of a certain period may be considered as excluded. By these assumptions α will depend on the range and the apparent magnitude only, so that the collection of cluster variables was divided into groups according to range and magnitude, assuming each group to have the same discovery chance α .

Two numbers of the series a_1, a_2, \dots, a_n are sufficient to determine the two unknowns α and N , whereas $a_1 \dots a_n$ are all known observed quantities, so that the problem is "overdetermined". Therefore the discovery chance α and the total number of variables N were derived for each group from the total number of variables found and the average number of times that each of these was found.

The first quantity can be found from the observations by the formula:

$$A_1 = a_1 + a_2 + \dots + a_n$$

and the second quantity by the formula:

$$G_1 = \frac{a_1 + 2a_2 + 3a_3 + \dots + na_n}{a_1 + a_2 + a_3 + \dots + a_n} \quad (1)$$

The average number of times that each variable of the group, including the undiscovered ones, was found is:

$$G_0 = \frac{0a_0 + 1a_1 + 2a_2 + \dots + na_n}{a_0 + a_1 + a_2 + \dots + a_n} = \alpha n. \quad (2)$$

From (1) and (2) we see:

$$G_1 = \alpha n \frac{N}{N - a_0} = \frac{\alpha n}{1 - (1 - \alpha)^n}. \quad (3)$$

By this formula α can be computed from G_1 .

As $N - a_0 = A_1$, the following formula for N results from (3):

$$N = \frac{A_1 G_1}{\alpha n}. \quad (4)$$

The number of plate pairs compared was 31. Of these plate pair E with an interval of 0.02 was blinked only in order to find asteroids. Only one variable, No. 30, was found on it by accident. This plate pair has been omitted, so that the computation was made for $n = 30$. This gives the following table representing the relation (3):

α	G_1
·140	4·246
·130	3·961
·100	3·132
·090	2·869
·080	2·614
·070	2·368
·060	2·133
·050	1·910
·040	1·699
·030	1·502
·025	1·409
·020	1·321
·015	1·235
·010	1·154

Of the 147 cluster variables 5 had ranges of $m.6$ and $m.5$. These 5 stars have all been discovered only once. Applying the formulae (3) and (4) to find α and N , these become 0 and ∞ respectively. This group has therefore been excluded. For the other 142 cluster variables the data have been collected from Table 1 (*B. A. N.* 227 and this paper) and 1a. Table 7 gives the groups into which they have been divided according to median apparent magnitude and range, with the resulting values for A_1, G_1, α and N .

TABLE 7.

apparent magnitude $m \quad m$ $13.25 - 13.85$	$1 \times$ found:	range $m \quad m$ 2.25 - 1.55	range $m \quad m$ 1.55 - 1.05	range $m \quad m$ 1.05 - .75	
2	0	0 stars	1 stars	2 stars	
3	1	$A_1 G_1 = 8$	1	$A_1 G_1 = 5$	
4	0	$G_1 = 4.000$	0	$G_1 = 1.667$	
5	1	$\alpha = .1314$	0	$\alpha = .0384$	
6	0	$N = 2.0$	0	$N = 4.3$	
		$A_1 = 2$	$A_1 = 4$	$A_1 = 3$	
$m \quad m$ $13.85 - 14.45$	$1 \times$ found:	range $m \quad m$ 2.25 - 1.55	range $m \quad m$ 1.55 - 1.05	range $m \quad m$ 1.05 - .75	range $m \quad m$.75 - .65
2	5	1 stars	6 stars	5 stars	3 stars
3	1	$A_1 G_1 = 31$	4	$A_1 G_1 = 11$	1
4	3	$G_1 = 2.818$	1	$G_1 = 1.375$	0
5	1	$\alpha = .0880$	3	$\alpha = .0231$	0
6	0	$N = 11.7$	0	$N = 15.9$	0
		$A_1 = 11$	1	$N = 17.0$	0
			$A_1 = 15$	$A_1 = 8$	$A_1 = 4$

TABLE 7 (continued).

m $14.45-15.05$	m $1 \times$ found: 2 3 4 5 6	m range $2.25-1.35$ 13 stars 9 4 $A_1 G_1 = 43$ 0 $G_1 = 1.654$ 0 $\alpha = .0377$ 0 $N = 38.0$ <hr/> $A_1 = 26$	m range $1.35-1.15$ 8 stars 10 1 $A_1 G_1 = 35$ 1 $G_1 = 1.750$ 0 $\alpha = .0424$ 0 $N = 27.5$ <hr/> $A_1 = 20$	m range $1.15-.75$ 22 stars 7 0 $A_1 G_1 = 40$ 1 $G_1 = 1.333$ 0 $\alpha = .0207$ 0 $N = 64.4$ <hr/> $A_1 = 30$	m range $.75-.65$ 3 stars 1 0 $A_1 G_1 = 5$ 0 $G_1 = 1.200$ 0 $\alpha = .0128$ 0 $N = 13.0$ <hr/> $A_1 = 4$
m $15.05-15.65$	m $1 \times$ found: 2 3 4 5 6		m range $2.25-1.15$ 3 stars 4 0 $A_1 G_1 = 11$ 0 $G_1 = 1.571$ 0 $\alpha = .0335$ 0 $N = 10.9$ <hr/> $A_1 = 7$	m range $1.15-.75$ 6 stars 1 0 $A_1 G_1 = 13$ 0 $G_1 = 1.625$ 1 $\alpha = .0362$ 0 $N = 12.0$ <hr/> $A_1 = 8$	

8. Adding the numbers found for N horizontally from Table 7 the distribution of the cluster variables over the apparent magnitudes is obtained. Taking the absolute magnitude of cluster variables as $m.0$ ($\pi = ".1$ "), the magnitude limits of the groups have been converted into distances in parsecs. The plate field of $10^\circ \times 10^\circ$ records the contents of a square solid angle of $10^\circ \times 10^\circ$ in space, which is divided into 5 sections by the spheres representing the distance limits for the groups. The size of each of these sections in cubic kiloparsecs has been computed.

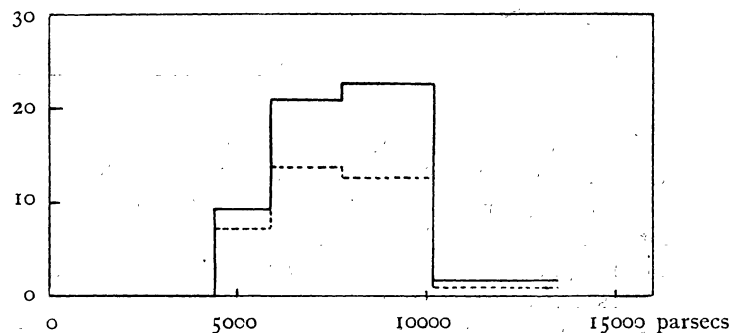
TABLE 8.

median apparent magnitude	distance (parsecs)	volume of section in cubic kiloparsecs	N	density in var. stars per cubic kiloparsec
m $m-13.25$	0-4467	.93	0 (0)	.0 (0)
$13.25-13.85$	4467-5888	1.21	1.2 (9)	9.3 (7.4)
$13.85-14.45$	5888-7762	2.76	57.6 (38)	20.9 (13.8)
$14.45-15.05$	7762-10233	6.32	142.9 (80)	22.6 (12.7)
$15.05-15.65$	10233-13490	14.48	22.9 (15)	1.6 (1.0)
$15.65-$	13490-			
			222.6 (142)	

By dividing the number of stars N by the volume of the corresponding section the density of cluster variables in stars per cubic kiloparsec is computed. Table 8 gives an account of this. The numbers when not corrected for incompleteness have been added in parentheses.

In Fig. 4 the density has been given as ordinate, the distance as abscissa. The dotted line represents

FIGURE 4.



the density from the uncorrected numbers of cluster variables.

From the figure it is clear that the cloud of cluster variables in the square solid angle represented by the plate field begins at about 4500 parsecs, has its maximum density at about 8000 parsecs, and ends at about 13000 parsecs.

The galactic longitude and latitude of the centre of the region discussed are 327° and $-18\frac{1}{2}^\circ$ respectively. This longitude corresponds with the longitude of the centre of the galactic rotation, so that the direction of the region cuts the axis of galactic rotation $18\frac{1}{2}^\circ$ south of the galactic plane. If the cloud of cluster variables is spherically situated round the galactic centre, this centre should be at a distance of 8400 parsecs.

It is interesting to compare this result with that obtained by H. SHAPLEY and Miss H. H. SWOPE in *Harv. Reprint 52* for Milky Way Field 185. This field is situated at $\lambda = 322^\circ$, $\beta = +8^\circ$, ten degrees closer to the Milky Way than the region discussed in this paper, but on the opposite side of it. SHAPLEY and SWOPE

1) SHAPLEY, *Star clusters*, p. 135.

derive a distance of 14400 parsecs for the centre of density of their region; their numbers of cluster variables however have not been corrected for incompleteness and have not been reduced to densities. The first correction will increase their distance, the latter however decrease it considerably. The difference between the brightest and the faintest apparent median magnitudes of the cluster variables mentioned in *Harv. Repr. 52* is $2^m.2$; of those discussed in this paper $2^m.4$, about the same. The field discussed in *Harv. Repr. 52* covers 70 square degrees; multiplying the numbers for M. W. F. 185 by $\frac{100}{70}$ to make them comparable with the results derived in this paper, the following table is obtained:

photographic median magnitude	number of cluster variables, this paper	reduced number of cluster variables, <i>Harv. Repr. 52</i>
m		
$m - 13.25$	0	0
$13.25 - 13.85$	9	0
$13.85 - 14.45$	38	11.4
$14.45 - 15.05$	80	8.6
$15.05 - 15.65$	15	41.4
$15.65 -$	0	50.0
	142	111.4

The average median magnitude of the cluster variables in M. W. F. 185 is $15^m.9$, in the field discussed here $14^m.6$, or $1^m.3$ brighter. This difference is too large to be caused by a difference in zeropoint of the magnitude scales only; very probably absorption comes in, and to a much greater extent in M. W. F. 185 than in the CrA field owing to its smaller galactic latitude.

OORT¹⁾ derives a distance of 10000 parsecs for the galactic centre; an earlier determination²⁾ gives 6300 parsecs; both determinations have been made from rotational effects in radial velocities.

Considerable doubt has been thrown on the zeropoint of the period-luminosity relation (see for instance KIPPER, *A. N.* 5775; GERASIMOVITCH, *A. J.* 951; CECCHINI, *Merale Contrib.* 14), SHAPLEY's absolute magnitudes, also adopted by me, being about $1^m.1$ too bright. This would mean that all distances derived by the period-luminosity relation have to be reduced by a factor .6. The point of greatest space density for cluster variables as derived here would consequently be reduced to 5000 parsecs distance.

If any sensible absorption has to be taken into account, this number has to be decreased further.

9. The numbers N obtained in Table 7, when

¹⁾ *B. A. N.* 159, p. 279.

²⁾ *B. A. N.* 132, p. 88.

added vertically, give the frequency of the cluster variables as a function of their range. As the limits in range do not correspond for the groups of Table 7, the distribution has been assumed to be uniform in each group. Adding the numbers N under this assumption the following table results:

TABLE 9.

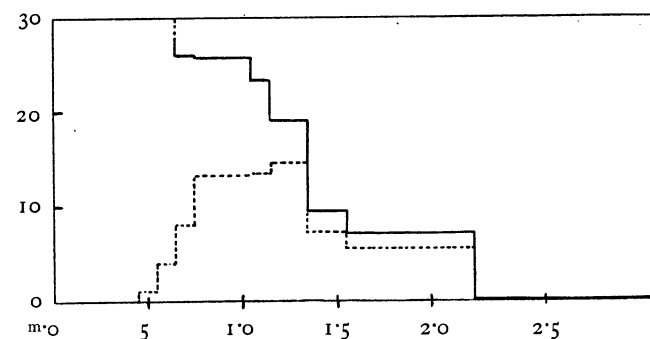
range	number of variables per $m.1$ range	number of variables per $m.1$ range, uncorrected
m		
$2.25 - 1.55$	7.2	5.4
$1.55 - 1.35$	9.2	7.3
$1.35 - 1.15$	19.2	14.4
$1.15 - 1.05$	23.5	13.3
$1.05 - .75$	25.8	13.2
$.75 - .65$	26.0	8
$.65 - .55$	—	4
$.55 - .45$	—	1

The data for the 5 stars with a range of $m.6$ and $m.5$, though not included in Table 7, have been added in Table 9.

In Fig. 5 these numbers are represented graphically; full line: column 2 of Table 9, dotted line: last column of the table.

The figure clearly shows the bad influence of a

FIGURE 5.



small range upon the discovery chance and consequently upon the completeness. The dotted line shows a maximum frequency at a range of $1^m.25$; the full line strongly suggests that the maximum frequency occurs for a range smaller than $m.7$, so that the maximum of the dotted curve at $1^m.25$ is only due to the small ranges having a very small discovery chance. The mean range for the 147 cluster variables discussed here is $1^m.16$, exactly the same value as obtained by STICKER (*Zs. f. Astroph.*, Bd. 2, p. 393) from PRAGER's catalogue 1931.

The full line is rising continually from range $2^m.25$ to $m.65$. The existing part of the frequency curve would even not preclude the value $m.0$ for the maximum frequency. This latter value might mean: all stars are

cluster variables, but only those with fairly large ranges have so far been discovered. The drawn part of the full curve represents 224.6 cluster variables; the total number of stars in the plate field till $15^m.65$ is about 200 000, so that the missing part of the curve between $m.0$ and $m.65$ would have to procure the remainder, about 199 800 variables. This is improbable; the curve can however have a maximum at $m.0$ and contain only a part of the 200 000 stars brighter than $15^m.65$ in the plate field — for instance only the stars of spectral types A and F. All known cluster variables belong to these two types, except one or two of type B 9.

GERASIMOVITCH (*Zs. f. Astroph. Bd. 2, p. 85*) discusses the frequency curve of the ranges of δ Cep variables. For those with a period between 1^d and 10^d he finds, after correction for incompleteness, a maximum for a range of $1^m.10$ (*loc. cit. p. 93*). It should be well remembered that GERASIMOVITCH's correction factors are derived in a geometrical way, and assuming a threshold value ϵ of $m.5$ for the difference in brightness on two plates necessary to discover the variable (see also *C. R. de l'Ac. des Sc. de l'U. R. S. S., 1931, p. 93*). The corrections in this paper are derived from the data in Table 1 and are, contrary to GERASIMOVITCH's correction factors, empirical ones. Comparison between the values derived for the discovery chance α in Table 7 and the ones derived by GERASIMOVITCH (*loc. cit., p. 96*) shows that the values derived in Table 7 are exceedingly small. Consequently after the 30 comparisons stated in Table 2 only 63 percent of the cluster variables with ranges over $m.65$ have been found, whereas GERASIMOVITCH states: "in an ideal case practically all Cepheids are fished out after the examination of 4–5 pairs of plates" (*loc. cit., p. 97*). The explanation of this is not a very high threshold value ϵ for the 30 plate pairs blinked here (compare the values Δs in Table 1), but the fact that many variables, although having a difference in brightness on a set of two plates greater than ϵ , remain undiscovered because they are simply overlooked. Thus the values for the probability of discovery α in Table 7 are the geometrical probabilities multiplied by a factor smaller than 1 depending on the attention and speed with which the plates were compared.

10. The statistical results obtained under 8 and 9 suggest the following desiderata for future investigation of the great cloud of cluster variables that is obviously present at galactic longitude 325° .

In the first place fainter variables should be included in order that the curve of Fig. 4 can be extended to greater distances and the decrease of the space

density for the cluster variables at great distances can be better shown. For this purpose the plates should penetrate one or two magnitudes deeper; unless much faster plates can be obtained a bigger lens than the Franklin-Adams one ($a = 254$ mm) will be required for this. It is dangerous to try to obtain the same result by making the exposure time much longer than 30^{min} , as the range of the variables as shown by the plates will decrease as the exposure time increases.

To extend the frequency curve of Fig. 5 to smaller ranges a procedure should be found by which also variables with small ranges should have a good probability of discovery. By plates showing stars one or two magnitudes fainter than the plate material discussed here the majority of the cluster variables would be brought into a steeper part of the gradation curve, which will increase their probability of discovery. Another helpful measure might be found in using plates with a very steep gradation and developing them with hard developers. In the blink microscope only the best plate pairs with best similarity of images should be compared.

11. Several of the lightcurves shown on pp. 54 to 60 and p. 181 to 184 in *B. A. N. 227* show a small decrease in brightness directly before the rise.

BAILEY's types *a* and *b* (*Harv. Ann. 38, p. 132*) cannot be separated according to period; amongst the periods above $d.4$ both types occur at random. The periods under $d.4$ however are all of BAILEY's type *c* and have all small ranges.

Comparison of the lightcurves of variables of small ranges with those of great ranges shows that the fraction of the period occupied by the rising branch is the smaller, the greater the range of the variable is.

The same remarks are made by GROSZE (*A. N. 5901*) in a discussion of the cluster variables in the globular cluster Messier 53; this again shows that there is no difference in behaviour for cluster variables whether they occur in globular clusters or are separate individuals far from any cluster.

12. The individual variables give rise to the following further remarks:

Var. 123, 154, 162, 213, 214. These variables are situated in the globular cluster N. G. C. 6723 = Δ 573, together with var. 11 and 12 of *B. A. N. 227*. On the maps of their surroundings the dense part of the cluster is marked by a circle. The mean of the median magnitudes of these 7 variables is $14^m.8$, corresponding to a distance for the cluster of 9100 parsecs. In SHAPLEY's list (*Star Clusters, p. 227*) the distance is given as 12300 parsecs. On the plates (scale: 1 mm. = $183''$)

the cluster is very crowded and the estimates were difficult and influenced by stars close to the variable. The identifications with BAILEY's variables (p. 23) could easily be made with the aid of *Harvard Annals* 38, plate XI.

Var. 124. This variable is irregular. Considerable time has been spent to find a period but without success. The maxima are unequal in brightness. The variability is unquestionable, the range being over 2^m .

Var. 131, 136. These variables have the uncommon periods of $1^d.0$ and $1^d.1$ respectively; the latter one has an exceptionally great range.

Var. 135, 240. These variables, although occurring on the overlapping part of the regions mentioned under 4 have been estimated on the 397 CrA photographs only. This was done on purpose to be fully independent from FERWERDA's results for these two stars, which are identical with his variables x and l respectively (*B. A. N.* 231). Comparison shows that the lightcurves as found in the two investigations are very similar. The value of a step as found from Table 3 is slightly smaller than FERWERDA's value. The periods, though derived independently from two entirely different series of plates, correspond to the sixth decimal. The apparent magnitudes as derived by FERWERDA (by comparison with Selected Area 159) are however about 1^m brighter than those derived here.

Var. 141. Two separate least squares solutions were made with the epochs given in Table 5, representing the maxima brighter than comparison star b and equal to b respectively. The two resulting periods were $^d.5144936 \pm ^d.0000024$ and $^d.5145022 \pm ^d.0000047$ respectively. The weighted mean was taken in order to obtain the period given in Table 3.

Var. 146, 217. A faint star is situated close to the variable; only in minimum the two stars are separately visible. In maximum the disturbing star is absorbed in the image of the variable. Consequently the range should be slightly decreased.

In cases where the disturbing star is too close to be seen separately even in minimum, only the combined brightness can be estimated. The effect will be that the derived lightcurve is incorrect and should be increased in range. Especially the minimum suffers a certain amount of flattening. It is impossible to distinguish these cases from the lightcurve alone; an elongation of the image, especially in minimum, is the only indication by which the variable's lightcurve and range can be distrusted.

Var. 162. Close to the variable is a bright star; the images are in contact when the variable is bright, which may have affected the estimates.

Var. 204. The maxima are certainly unequal in brightness.

Var. 226. A period of about $3\frac{1}{2}^d$ was tried, but proved unsatisfactory. Nothing better could be found. The star is an irregular variable.

Var. 253. This star is an irregular variable of the SS Cygni type. No definite period could be found. There are long maxima, which last about 8^d , and short maxima of one day only.

Var. 272. This variable has a period of $^d.1076$ only. The 5 variables with shortest known periods,¹⁾ all due to the collaboration between Johannesburg and Leiden, are now:

	period	author	<i>B. A. N.</i>
VV Pup	$\cdot 0697$	VAN GENT	214
KU Cen	$\cdot 0800$	MARTIN	232
$19^h 2^m 58^s.5 - 35^\circ 5' 7''$	$\cdot 1076$	VAN GENT	243
AI Vel	$\cdot 1116$	HERTZSPRUNG	224
AQ CrA	$\cdot 1187$	VAN GENT	227

Var. 280. A period of 64^d succeeded to represent all the observations satisfactorily, except those of the year 1925. Nothing better could be found.

Var. 302. Of this Algol variable 6 minima have been found. The counting of the periods was in full agreement with the epoch and period derived in *Harv. Circ.* 238 by IDA E. WOODS and MARTHA B. SHAPLEY. The number of minima and the interval of time covered by the plate material used by these authors was much greater than that used here. Moreover no observations were found to be on the steep parts of the rising and falling branches, so that there was no hope of improving the Harvard period by the aid of the present material. Consequently the period of *Harv. Circ.* 238 was used to derive the phases. The magnitude scale for the comparison stars was derived by the aid of estimates on a plate with two exposures of equal length taken with and without grating respectively. The difference in magnitude between free lens exposure and first order grating image was assumed to be $2^m.5$; as compared with the magnitudes used in *Harv. Circ.* 238 the relation between the Harvard scale and the one here proved to be linear, $1^m.00$ Harvard scale being equal to $1^m.06$ scale of this paper. The range of the variable was found to be $3^m.4$; the range given in *Harv. Circ.* 238 is $2^m.84$. Owing to the brightness of the variable the mean error of a single estimate is very small, only $\pm ^m.035$.

The periods given for the Algol variables are all the periods of the light variation. In the cases of var. 177 and 206 the possibility exists that these have to be doubled in order to obtain the revolution period.

¹⁾ The period of about $^d.06$ for ϵ^2 Lyrae, derived by E. A. FATH (*Pop. Astr.* 40, p. 88), is still open to doubt.

I want to express my thanks to prof. E. HERTZ-SPRUNG, who derived many of the periods, to prof. P. J. VAN RHIJN for his permission to use the Groningen blink microscope and to prof. R. PRAGER for comparing the list of variable stars in this paper

with his card catalogue. Mr. KOOREMAN has measured the positions of the variables, Mr. DE HAAS has made the drawings for the diagrams; both had a large share in the computing work for the least squares solutions.

TABLE 3.

Var. No.	α (1875)	δ (1875)	type	number of estimates	period	m. e.	reciprocal period	number of epochs	m. e. of single epoch	phase of rising br. ΔP	phase of max.	max.	min.	max. min. range	m. e. of single est.
	h m s	° ' "			d	d	d ⁻¹		d	P P P	m m	m m	st st st	st	
42	18 27 34.0	- 39 44.7	cluster	347	4683532	± 0.000016	2.135141	19	± 0.10	'03 '25 '11	12.7 14.4	12.6 14.7	.87 16.81 15.94	± 1.10	
226	18 27 49.7	- 39 8.6	irregular	361									3.0 14.6 17.6		
205	18 28 1.5	- 41 11.7	cluster	235	5985054	± 0.000024	1.670829	14	± 0.13	'28 '25 '35	14.1 15.3	14.0 15.1	5.72 15.11 9.39	± 1.06	
253	18 28 19.4	- 37 31.7	SS Cyg	286							13.9 15.6		3.3 19.9 16.6		
136	18 28 27.9	- 38 10.4	δ Cep	329	1.1276457	± 0.000071	.88680	17	± 0.26	'22 '30 '29	12.3 14.6		.88 21.55 20.67	± 1.57	
112	18 28 49.9	- 39 56.8	cluster	241	6566548	± 0.000044	1.52288	30	± 0.28	'63 '30 '72	14.0 15.1	2.44 14.07 11.63	± 1.57		
231	18 29 21.1	- 34 17.7	cluster	204	5631243	± 0.000057	1.775807	15	± 0.24	'82 '25 '88	14.6 16.2	4.97 11.30 6.33	± 1.15		
232	18 30 18.3	- 37 32.3	Algol	377	1.36545		.087986				13.1 14.8		6.71 26.53 19.82	$\pm .54$	
165	18 30 18.8	- 35 3.6	cluster	200	4624318	± 0.000017	2.162481	14	± 0.12	'59 '25 '66	14.5 16.4	1.45 7.30 5.85	$\pm .87$		
179	18 30 24.6	- 41 13.2	cluster	276	5424889	± 0.000045	1.843356	10	± 0.25	'90 '30 '99	13.9 15.1	3.96 12.35 8.39	± 1.47		
195	18 30 39.4	- 37 44.9	cluster	255	5470096	± 0.000035	1.828122	13	± 0.23	'04 '25 '10	14.3 15.4	2.72 10.05 7.33	± 1.03		
150	18 30 39.7	- 38 12.2	cluster	275	4904128	± 0.000014	2.039099	19	± 0.10	'46 '25 '53	14.0 15.1	2.82 12.67 9.85	± 1.19		
274	18 30 51.3	- 33 54.4	longper.	290							11.9 16.2		3.0 28.5 31.5		
128	18 31 37.2	- 40 37.6	cluster	210	5887684	± 0.000023	1.69846	5	± 0.09	'01 '30 '07	14.5 15.5	4.26 13.92 9.66	± 1.74		
138	18 33 41.3	- 35 39.2	cluster	236	5854805	± 0.000036	1.708	9	± 0.18	'05 '25 '08	14.5 15.2	4.41 11.30 6.89	± 1.46		
283	18 34 24.9	- 33 24.0	cluster	207	4572386	± 0.000034	2.187042	8	± 0.13	'70 '20 '75	14.9 16.4	3.67 11.40 7.63	± 1.22		
177	18 34 34.9	- 37 13.2	Algol	350	2.409662		.414996				14.0 14.6		5.30 9.63 4.33	$\pm .83$	
207	18 35 14.5	- 38 50.7	cluster	262	4826170	± 0.000016	2.072036	11	± 0.12	'45 '20 '49	14.1 15.3	4.83 14.68 9.85	± 1.51		
220	18 35 34.8	- 35 53.9	cluster	234	4800885	± 0.000022	2.082949	12	± 0.11	'19 '20 '27	14.2 15.5	2.72 11.40 8.38	± 1.02		
139	18 35 42.7	- 35 20.7	cluster	252	5416333	± 0.000020	1.846268	19	± 0.14	'23 '20 '29	14.3 15.5	5.33 18.45 13.12	± 1.38		
199	18 35 43.4	- 40 39.2	cluster	235	4432499	± 0.000018	2.256064	12	± 0.12	'88 '25 '94	14.4 15.3	7.02 14.70 7.68	± 1.27		
206	18 35 56.2	- 35 10.3	Algol	315	2.247540		.444931				14.2 14.7		5.96 11.87 5.91	± 1.00	
221	18 36 34.6	- 35 34.1	cluster	284	4579511	± 0.000018	2.183639	14	± 0.14	'28 '25 '38	14.1 15.5	1.62 14.13 12.51	± 1.75		
208	18 36 35.1	- 37 6.2	cluster	266	5922067	± 0.000034	1.688600	11	± 0.16	'69 '25 '75	14.4 15.5	1.78 11.09 9.31	± 1.28		
209	18 37 42.0	- 36 15.5	cluster	268	5242381	± 0.000033	1.907530	24	± 0.18	'27 '25 '32	14.2 15.6	3.63 14.22 10.59	± 1.22		
277	18 38 0.4	- 38 7.7	longper.	84							13.5 10.4		2.0 18.9 16.9		
222	18 38 38.1	- 35 58.2	cluster	284	5800326	± 0.000040	1.724041	24	± 0.23	'44 '30 '51	13.9 15.1	3.11 14.07 10.96	± 1.20		
115	18 38 39.6	- 32 49.4	cluster	277	5905392	± 0.000028	1.693368	19	± 0.18	'50 '25 '55	13.4 14.6	5.93 17.23 11.30	± 1.12		
130	18 38 46.6	- 40 0.8	cluster	300	4604108	± 0.000013	2.171973	23	± 0.13	'62 '25 '67	13.8 15.3	4.61 16.31 11.70	± 1.20		
245	18 39 41.1	- 32 29.8	cluster	273	6587352	± 0.000067	1.518061	15	± 0.29	'08 '30 '14	14.1 15.0	5.36 13.07 7.71	± 1.27		
189	18 39 45.2	- 39 45.2	cluster	324	3311464		3.019812				14.1 14.7		6.10 13.32 7.22	± 1.59	
168	18 40 08.4	- 39 33.1	cluster	267	4815783	± 0.000018	2.076505	31	± 0.17	'28 '25 '34	13.8 15.2	1.38 12.92 11.54	± 1.49		
141	18 40 34.5	- 35 51.6	cluster	203	5144954	± 0.000021	1.943652				14.6 15.7		5.90 15.93 10.03	± 1.20	
268	18 41 10.3	- 33 3.9	cluster	262	7117164	± 0.000064	1.405054	6	± 0.17	'40 '35 '50	14.3 15.2	4.49 10.89 6.40	± 1.12		
309	18 41 57.0	- 41 1.1	cluster	239	4810838	± 0.000022	2.078640	18	± 0.17	'30 '25 '37	14.7 15.3	6.00 13.40 7.40	$\pm .92$		
167	18 42 25.2	- 37 9.5	cluster	288	6124905	± 0.000082	1.63268	9	± 0.31	'11 '35 '24	14.1 15.1	5.96 13.45 7.49	± 1.45		
200	18 42 30.6	- 35 35.2	cluster	292	4778794	± 0.000023	2.092578	18	± 0.17	'65 '20 '70	13.9 15.1	3.68 14.90 11.22	± 1.26		
235	18 42 33.4	- 41 6.4	cluster	273	6556387	± 0.000036	1.525230	9	± 0.18	'84 '25 '90	14.4 15.4	2.58 11.67 8.69	± 1.03		
180	18 42 34.2	- 39 6.3	cluster	258	4437775	± 0.000020	2.253381	22	± 0.19	'27 '25 '34	14.3 15.3	3.08 11.39 8.31	± 1.51		
176	18 42 37.7	- 40 41.0	cluster	221	5057266	± 0.000038	1.977353	15	± 0.26	'80 '25 '88	14.5 15.5	1.38 8.30 6.92	± 1.14		
223	18 43 11.6	- 37 27.0	cluster	303	4833536	± 0.000022	2.068879	22	± 0.17	'44 '20 '48	14.0 15.3	.18 11.70 11.52	± 1.16		
181	18 43 48.8	- 39 4.8	cluster	281	6092638	± 0.000055	1.641325	18	± 0.36	'15 '45 '29	14.2 15.0	4.24 10.94 6.70	± 1.14		
310	18 43 53.1	- 37 29.4	cluster	305	5159405	± 0.000021	1.938208	25	± 0.18	'51 '25 '57	13.6 15.0	.79 13.17 12.38	± 1.16		
201	18 43 59.9	- 36 5.8	cluster	258	5253633	± 0.000027	1.903445	19	± 0.19	'15 '25 '21	14.3 15.6	2.21 11.51 9.30	± 1.24		
182	18 44 3.9	- 40 17.6	cluster	325	3337360		2.99638				14.2 14.9		5.39 10.09 4.70	± 1.38	
211	18 44 20.9	- 38 58.2	W UMa	277	2040833		4.89996				14.5 14.9		5.80 9.88 4.08	± 1.13	
160	18 44 39.3	- 33 45.5	Algol	375	1.641569	± 0.00010	.609173	5	± 0.14		13.4 14.4		5.84 16.52 10.68	$\pm .91$	
210	18 44 55.8	- 34 49.5	cluster	279	6014340	± 0.000027	1.662693	9	± 0.14	'89 '25 '95	13.6 15.1	1.18 14.31 13.13	± 1.23		
169	18 44 58.1	- 33 31.0	cluster	207	4862863	± 0.000064	2.056402	20	± 0.33	'30 '30 '40	14.6 15.5	3.58 10.46 6.88	± 1.34		
118	18 45 1.0	- 40 11.9	W UMa	302	4065867		2.4595				14.2 14.8		6.80 12.62 5.82	± 1.74	
170	18 45 9.5	- 36 31.8	cluster	238	6311588	± 0.000030	1.584387	21	± 0.21	'00 '25 '08	14.6 15.3	3.92 12.00 8.08	± 1.32		
280	18 45 12.2	- 39 58.1	longper.	340							13.7 15.2		1.6 12.2 10.6		
183	18 45 17.3	- 39 6.2	cluster	294	4590166	± 0.000017	2.178571	18	± 0.10	'98 '20 '04	14.2 15.6	4.11 15.49 11.38	± 1.33		
224	18 45 52.5	- 38 34.9	cluster	369	4709644	± 0.000017	2.123303	18	± 0.13	'80 '30 '87	13.2 14.2	.60 13.76 13.16	± 1.02		
35	18 46 27.1	- 35 43.5	cluster	291	5237549	± 0.000039	1.90929	16	± 0.21	'79 '25 '86	13.4 14.7	.81 14.52 13.71	± 2.03		
236	18 46 27.4	- 32 56.6	cluster	192	5124514	± 0.000083	1.951404	12	± 0.20	'37 '20 '41	14.6 15.5	5.07 11.10 6.03	± 1.10		
288	18 46 27.9	- 36 51.8	longper.	269							13.9 15.4		1.0 10.4 11.4		

TABLE 4 (continued).

s	m	b	s	m	s	m	b	s	m	d	s	m	d	s	m	b	s	m	b	s	m
223		c	8.7	13.3	239		c	9.2	14.0	e	8.6	15.0	e	22.3	14.7	c	5.5	14.7	c	5.5	14.7
a	0	d	14.4	13.7	a	0	d	13.4	14.2	e	11.8	15.4	e	25.5	15.2	c	8.4	14.8	c	8.4	14.8
b	4.4	e	19.7	14.0	b	4.5	e	17.4	14.7		268			277		d	12.2	15.2	d	12.2	15.2
c	8.0		27.2	14.8	c	10.3		248		b	0	13.8	a	0	13.3		310				
d	12.4	15.4	235		d	14.9	15.2	a	0	14.1	c	5.6	14.4	b	6.0	14.0					
	224		b	0	14.0	240		b	2.6	14.5	d	9.8	15.0	c	11.0	14.7					
a	0	13.2	c	5.4	14.7	b	0	13.7	c	6.0	14.7	d	15.9	15.4	d	15.9	15.4				
b	5.0	13.7	d	7.8	14.8	c	5.4	14.2	d	10.8	15.7	a	0	13.7							
c	13.8	14.2	e	11.8	15.4	d	7.9	14.7		250		b	5.0	14.1	a	0	13.7				
d	18.6	14.8		236		e	12.2	15.2	a	0	13.8	c	8.6	14.5	b	4.0	13.9				
	226		a	0	14.0	242		b	4.0	14.1	d	13.2	14.8	b	7.4	14.4					
A	0	13.0	b	4.8	14.5	b	0	14.2	c	8.7	14.8	d	13.1	15.4	c	11.2	15.0				
a	5.1	13.7	c	9.4	15.2	c	4.8	15.0	d	13.1	15.4	a	0	14.1	d	11.2	15.0				
b	10.4	14.1		237		d	8.0	15.2		253		b	4.6	14.5	a	0	14.4				
c	14.6	14.7	b	0	13.5	a	0	13.7	a	0	13.7	c	7.8	15.0	b	3.3	14.8				
	231		c	4.4	14.0	b	5.8	14.1	b	5.8	14.1	d	11.4	15.7	c	6.3	15.0				
a	0	13.9	d	9.6	14.7	c	0	13.7	c	9.7	14.4		274		d	9.4	15.7				
b	5.2	14.7	e	13.3	15.2	d	4.5	14.0	d	13.3	14.8	A	0	12.2		288					
c	8.4	15.2		238		e	8.6	14.5	e	16.9	15.2	B	5.2	12.9	a	0	14.0				
	232		a	0	14.2		11.8	14.8		255		a	9.9	13.4	b	3.8	14.4				
a	0	12.5	b	4.8	14.8	b	0	14.0	b	0	14.0	b	14.6	13.9	c	6.2	14.7				
			c	7.7	15.0	c	4.7	14.7	c	4.7	14.7	c	18.4	14.2	d	10.4	15.4				
			d	10.7	15.7	a	0	13.3													

TABLE 5.

35	112	d	t	d	2425498.255	2540	+	3	2426565.449	6178	-	7	2425532.310	3301	-	14					
2423883.575	o	+	005	2423990.273	o	-	005	278	2540	+	26	66.320	6180	-	4	61.234	3363	+	23		
4019.228	259	+	6	4017.169	41	+	24	5501.234	2545	+	29	08.280	2557	-	11	5707.505	3677	-	1		
20.256	261	-	14	28.321	58	-	36	08.280	2557	-	11	303	2557	+	12	19.608	3703	-	12		
5381.485	2860	-	24	30.289	61	+	1	303	2557	+	12	24.217	2584	-	19	630	3703	+	10		
5473.209	3035	+	43	5472.833	2260	-	20	24.217	2584	-	19	239	2584	+	3	64.352	3799	+	4		
74.231	3037	+	18	74.801	2263	-	33	239	2584	+	3	31.304	2596	-	18	6155.244	4638	-	3		
254	3037	+	41	78.737	2269	+	7	31.304	2596	-	18	326	2596	+	4	6552.443	5512	-	10		
99.340	3085	-	13	97.761	2298	-	20	6093.506	3548	-	10	4024.325	275	+	21	63.376	5514	-	8		
5500.378	3087	-	23	761	2298	+	4	6155.550	3653	+	28	25.347	277	-	9						
400	3087	-	1	761	2298	+	25	6562.378	4342	-	26	33.269	292	+	17	128					
08.258	3102	+	1	761	2298	+	47	6562.378	4342	-	26	4296.421	792	-	21	2423883.561	o	-	006		
5745.504	3555	-	14	761	2298	+	69	66.535	4349	-	2	97.446	794	-	49	5480.312	2712	+	5		
50.440	3658	-	25	5499.729	2301	-	7				5441.379	2967	+	60	5500.334	2746	+	9			
463	3658	-	2	729	2301	+	16				45.517	2975	-	13	6562.464	4550	+	1			
5808.361	3675	-	8	5520.721	2333	-	18				80.258	3041	-	13	6564.220	4553	-	9			
6566.252	5122	+	10	721	2333	+	23	119			279	3041	+	8							
				22.689	2336	-	7	2423883.546	o	+	015	301	3041	+	30	130					
				689	2336	+	15	87.285	239	+	2	323	3041	+	52	2423989.324	o	+	003		
				5714.897	2629	-	32	4017.225	308	-	11	5531.304	3138	-	26	4012.346	50	+	4		
				43.761	2673	-	12	30.263	338	+	4	5794.519	3638	-	1	18.321	63	-	6		
				761	2673	+	9	33.293	345	-	5	6093.506	4206	+	2	24.302	76	-	10		
				6091.441	3203	-	38	56.302	398	-	4	6155.638	4324	+	21	25.226	78	-	7		
				6153.105	3297	-	25	4261.636	871	-	4	6562.464	5097	-	45	250	78	+	17		
				6239.041	3428	-	23	5391.629	3474	+	4	486	5097	-	23	4296.421	667	+	6		
				6559.825	3917	+	5	5473.231	3662	-	7				5473.209	3223	+	16			
				825	3917	+	26	253	3662	+	16	125			231	3223	+	6			
				825	3917	+	48	75.420	3667	+	12	2423994.338	o	-	015	253	3223	+	28		
				6563.761	3923	-	36	79.303	3676	-	12	96.238	4	+	22	5500.378	3282	-	11		
				761	3923	-	15	325	3676	+	10	4016.247	47	-	4	400	3282	+	10		
				761	3923	+	7	99.275	3722	-	9	22.301	60	-	6	02.217	3286	-	14		
								296	3722	+	12	325	60	+	18	06.367	3295	-	8		
				115				5508.385	3743	-	16	27.410	71	-	22	08.212	3299	-	4		
				2423998.264	o	-	019	32.288	3798	+	11	29.297	75	+	1	25.248	3336	-	4		
				4018.345	34	-	16	5791.440	4395	o		30.222	77	-	6	6155.550	4705	-	4		
				21.337	39	+	23	97.510	4409	-	7	5391.605	2999	-	14	6562.550	5589	-	7		
				24.252	44	-	14	5808.361	4434	-	9	5499.251	3230	+	7	6563.462	5591	-	16		
				277	44	+	11	6090.527	5084	-	14	275	3230	+	30	484	5591	+	6		
								6563.290	6173	+	5										

TABLE 5 (continued)

d	t	d	d	t	d	I	145	150	d	t	d	I		
2426563	505	5591	+	27	2425501	234	2763	+	2423987	311	0	019		
66	229	5597	-	11	08	236	2775	-	89	273	4	9		
	252	5597	+	11	25	226	2804	-	90	273	6	10		
			134											
2423990	345	0	-	005	6156	400	3882	+	4012	334	51	2		
91	293	2	-	10	6241	268	4027	-	17	254	61	18		
40	33	244	90	0	6563	312	4577	+	5361	449	2802	-		
54	15	404	2990	+	139			2423998	239	0	-	008		
74	479	3114	-	30	2423881	520	0	+	264	0	+	18		
55	51	233	3275	+	3989	273	199	-	4017	225	29	-		
61	234	3296	+	11	4028	279	271	-	30	335	49	-		
57	07	527	3603	-	304	271	+	14	5442	511	2205	+		
16	592	3622	-	3	5415	404	2832	-	73	286	3030	+		
65	62	550	5397	-	427	2832	+	15	74	254	3032	+		
63	505	5399	-	8	5442	511	2882	+	98	266	3081	-		
65	428	5403	+	9	79	325	2950	0	99	263	3083	0		
66	363	5405	-	9	98	278	2985	-	55	25	3136	-		
			135											
2424012	346	0	-	021	99	363	2987	-	61	55	431	-		
42	96	421	619	-	5506	393	3000	-	62	42	458	+		
	445	619	+	9	31	304	3046	-	65	63	462	-		
54	42	317	3116	-	326	3046	+	4	66	342	3166	+		
65	342	3166	+	48	57	16	3388	+	74	457	3186	-		
74	457	3186	-	15	99	440	3541	+	479	3186	+			
75	420	3188	+	30	65	64	209	4953	75	420	3188	+		
82	265	3203	-	9	66	363	4957	-	99	251	3240	-		
99	251	3240	-	2	384	4957	0	0	275	3240	+			
	275	3240	+	21	406	4957	+	23	296	3240	+			
	296	3240	+	43	141									
55	09	329	3262	-	2424027	410	0	+	426	4	552	461		
22	236	3290	+	37	4264	552	461	-	536	1	449	2593		
27	223	3301	-	24	83	605	2636	+	57	16	571	3388		
24	248	3301	+	1	5479	280	2822	-	80	301	2824	-		
32	288	3312	-	7	80	301	2824	-	323	2824	+			
31	0	3312	+	15	5738	603	3326	+	60	90	506	4010		
60	93	506	4535	-	6090	506	4010	0	61	55	320	4136		
61	56	400	4672	-	6155	320	4136	-	34	1	436	+		
65	62	550	5557	-	6562	292	4927	-	5	314	4927	+		
63	462	5559	-	16	63	333	4929	+	7	66	406	4935		
48	4	5559	+	6	144									
66	229	5505	-	2	2424296	445	523	-	4	549	340	2861		
	252	5505	+	20	5499	340	2861	-	15	5500	378	2863		
			136											
2423880	546	0	-	029	32	288	2925	+	5	310	2925	+		
40	19	253	123	-	65	63	312	4929	-	34	355	4929	+	
20	395	124	-	8	65	63	312	4929	-	34	355	4929	+	
28	279	131	-	17	65	63	312	4929	-	34	355	4929	+	
	328	131	+	32	65	63	312	4929	-	34	355	4929	+	
42	94	433	367	+	65	63	312	4929	-	34	355	4929	+	
53	81	485	1331	+	65	63	312	4929	-	34	355	4929	+	
	508	1331	+	37	65	63	312	4929	-	34	355	4929	+	
91	605	1340	-	15	65	63	312	4929	-	34	355	4929	+	
62	9	1340	+	9	65	63	312	4929	-	34	355	4929	+	
54	42	317	1385	-	65	63	312	4929	-	34	355	4929	+	
57	95	328	1698	+	65	63	312	4929	-	34	355	4929	+	
65	63	211	2379	-	65	63	312	4929	-	34	355	4929	+	
	232	2379	-	12	65	63	312	4929	-	34	355	4929	+	
	250	2379	+	7	65	63	312	4929	-	34	355	4929	+	
	269	2379	+	25	65	63	312	4929	-	34	355	4929	+	
			138											
2423883	546	0	-	004	2423997	272	0	+	023	4019	228	39	-	
40	12	346	220	-	4019	228	39	-	21	253	39	+		
32	290	254	+	28	253	39	+	4	24	325	48	-		
			137											
2423877	528	0	-	019	4019	228	39	-	21	32	220	62	-	
	555	0	+	9	545	1	477	2578	+	13	80	235	2629	
39	91	268	218	+	99	462	7247	-	3	60	93	484	8444	
	92	289	218	+	57	19	630	6922	+	1	61	55	409	
40	30	360	293	-	40	508	7007	-	1	65	63	211	10356	
42	97	423	805	-	99	462	7247	-	3	41	9	10357	-	
	446	805	+	7	60	93	484	8444	-	22	66	384	10369	
55	01	285	3113	-	60	93	484	8444	-	22	66	384	10369	
24	239	3157	-	14	61	55	409	8096	-	1	65	63	211	
61	55	409	4367	+	65	63	211	10356	+	24	41	9	10357	
65	62	249	5147	+	66	384	10369	+	4	66	384	10369	+	
	63	269	5149	-	146									
66	406	5155	-	13	2423987	285	0	-	004	4018	271	54	-	
	428	5155	+	9	4018	271	54	-	6	22	301	61	+	
	449	5155	+	29	30	335	75	+	7	54	37	2527	-	
			146											
2423987	285	0	-	004	54	37	2527	-	1	54	37	2527	-	
	4018	271	54	-	37	2527	-	1	54	37	2527	-		
	22	301	61	+	37	2527	-	1	54	37	2527	-		
	30	335	75	+	37	2527	-	1	54	37	2527	-		
	54	37	2527	-	37	2527	-	1	54	37	2527	-		
	41	431	2534	-	37	2527	-	1	54	37	2527	-		
	75	306	2593	+	37	2527	-	1	54	37	2527	-		
	79	303	2600	-	37	2527	-	1	54	37	2527	-		
	325	2600	+	9	37	2527	-	1	54	37	2527	-		
	98	255	2633	+	37	2527	-	1	54	37	2527	-		
	55	25	2680	+	37	2527	-	1	54	37	2527	-		
	57	90	345	3142	-	1	57	90	345	3142	-			
	65	63	333	4489	+	3	57	90	345	3142	-			
			147											
2423879	512	0	+	013	2423879	512	0	+	013	80	546	2	-	
	80	546	2	-	80	546	2	-	13	42	96	421	787	
	42	96	421	787	42	96	421	787	+	30	97	446	789	
	5	97	446	789	5	97	446	789	-	5	54	10	369	
	27	54	10	369	27	54	10	369	+	27	37	423	2941	
	11	37	423	2941	11	37	423	2941	+	11	74	479	3011	
	14	74	479	3011	14	74	479	3011	-	14	80	301	3022	
	19	80	301	3022	19	80	301	3022	-	19	323	3022	+	
	4	323	3022	+	4	323	3022	+	4	323	3022	+		
	5	99	385	3058	5	99	385	3058	-	5	99	385	3058	
	6	55	07	329	3073	-	6	55	07	329	3073	-		
	10	57	06	521	3449	+	10	57	06	521	3449	+		
	1	65	61	481	5063	-	1	65	61	481	5063	-		
	14	65	62	528	5065	-	14	65	62	528	5065	-		
	9	55	0	5065	+	9	55	0	5065	+	9	55	0	
	2	65	66	252	5072	+	2	65	66	252	5072			

TABLE 5 (continued).

d		t		d		d		t		d		d		t		d										
2425553	242	3122	+	20	2424012	242	93	+	29	2425532	288	3722	+	35	2426156	324	4326	-	3							
61	258	3138	+	16	20	395	109	-	32	5719	608	4145	-	16	6562	421	5099	-	12	2423883	575	0	+	025		
5719	608	3454	-	25	5360	498	2761	-	11	630	4145	+	6	63	462	5101	-	21	4029	225	246	-	8			
630	3454	-	3	5438	347	2915	+	20	91	418	4307	+	34	484	5101	+	1	5360	498	2494	-	15				
98	305	3611	-	22	41	379	2921	+	20	6093	506	4989	+	24	505	5101	+	22	83	605	2533	-	4			
6090	527	4194	-	20	51	455	2941	-	10	6563	419	6050	-	43	66	513	6057	-	50	5465	342	2671	+	8		
6562	206	5135	-	5	477	2941	+	12	198	2423880	546	0	+	015	23	362	2769	-	8	5507	372	2742	-	8		
66	229	5143	+	9	73	209	2984	+	15	3994	289	203	+	5	61	258	2833	-	13	23	362	2769	-	8		
67	223	5145	-	0	76	255	2990	+	29	4031	261	269	-	6	6155	244	3836	-	10	61	258	2833	-	13		
192	80	257	2998	-	11	79	280	2996	+	23	4297	423	744	-	14	6563	312	4525	+	27	66	252	4530	+	6	
2423988	344	0	-	005	82	265	3002	-	24	5438	347	2595	-	11	5475	306	2846	-	4	2423994	338	0	-	015		
89	324	2	+	4	239	3085	+	9	51	477	2619	-	15	98	278	2887	-	6	4025	299	59	+	16			
91	268	6	+	6	25	226	3087	-	15	74	457	2661	-	19	99	385	2889	-	20	5360	498	2606	-	19		
4020	395	66	+	3	248	3087	+	7	6562	357	4649	-	16	5508	385	2905	+	14	81	485	2646	-	2			
22	325	70	-	9	26	220	3089	-	31	63	462	4651	-	6	5719	630	3282	+	4	5410	344	2701	+	24		
25	250	76	+	3	27	248	3089	-	9	483	4651	+	15	97	510	3421	-	6	65	364	2806	-	1			
5410	369	2929	+	2	31	304	3099	0	2424016	247	0	-	018	6155	571	4060	-	13	73	253	2821	+	25			
41	431	2993	+	8	32	288	3101	-	27	271	0	+	6	6562	400	4786	-	4	74	254	2823	-	23			
74	457	3061	+	4	310	3101	-	5	28	255	27	+	22	421	4786	+	17	77	223	2924	-	2				
479	3061	+	26	5740	486	3513	-	17	5385	463	3089	-	1	63	527	4788	+	2	276	2823	+	2				
75	420	3063	-	4	508	3513	+	5	5442	632	3218	-	11	66	320	4793	-	7	298	2823	+	21				
79	303	3071	-	25	5904	223	3837	0	73	231	3287	+	3	342	4793	+	15	75	306	2825	-	19				
80	279	3073	0	6155	341	4334	-	21	5507	350	3364	-	8	204	2424298	423	0	+	008	329	2825	+	4			
5706	521	3539	+	1	6241	268	4504	+	3	51	233	3463	-	7	5385	416	2153	-	8	5507	307	2886	+	3		
19	630	3566	+	1	6566	207	5147	+	27	5740	508	3890	+	1	440	2153	+	16	329	2886	+	25				
40	5c8	3609	+	3	195	2423880	546	0	-	010	99	462	4023	+	2	5445	495	2272	-	9	27	223	2924	-	2	
6090	527	4330	-	21	3988	344	197	+	27	6566	276	5753	-	6	517	2272	+	13	5719	608	3291	-	12			
93	484	4336	+	23	4028	255	270	+	7	299	5753	+	17	74	276	2329	-	7	90	367	3426	-	25			
6561	481	5300	+	2	4292	440	753	-	14	200	2423883	575	0	-	008	79	303	2339	-	28	91	418	3428	-	23	
62	443	5302	-	7	98	446	764	-	25	4292	440	753	-	14	543	2789	+	15	80	323	2341	-	18			
464	5302	+	14	5385	416	2751	+	37	07	527	2791	-	11	5706	521	2789	-	7	5706	521	2789	-	7			
63	419	5304	-	2	5507	350	2974	-	12	99	440	2973	+	14	543	2789	+	15	5719	608	3291	-	12			
66	320	5310	-	14	372	2974	+	10	6155	409	3678	+	42	543	2789	+	15	6241	268	4286	+	31				
342	5310	+	8	5799	440	3508	-	25	6566	320	4492	-	20	07	527	2791	-	11	6566	252	4906	+	13			
193	2425437	423	0	-	026	6090	506	4040	+	31	205	2423878	530	0	-	002	5719	608	3291	-	12	91	418	3428	-	23
80	235	77	+	30	6155	550	4159	-	19	5706	521	2789	-	7	5719	608	3291	-	12	6155	244	4122	-	18		
5501	285	115	-	21	6561	459	4901	+	9	543	2789	+	15	5719	608	3291	-	12	56	324	4124	+	13			
51	244	205	-	37	62	528	4903	-	16	99	440	2973	+	14	5719	608	3291	-	12	6241	268	4286	+	31		
5719	529	508	0	583	443	+	17	196	2423883	575	0	+	017	6155	409	3678	+	42	6566	252	4906	+	13			
39	554	544	+	35	98	446	764	-	25	5797	488	4005	-	3	5706	521	2789	-	7	276	2823	+	21			
64	494	589	-	12	3990	345	241	+	34	5797	488	4005	-	3	5706	521	2789	-	7	75	306	2825	-	19		
94	508	643	+	17	4020	395	309	-	38	6155	409	4754	-	13	5706	521	2789	-	7	5507	350	3364	-	8		
99	486	652	-	2	21	299	311	-	19	56	400	4756	+	22	5706	521	2789	-	7	5706	521	2789	-	7		
5803	362	659	-	13	24	350	318	-	69	6563	505	5608	-	26	5706	521	2789	-	7	5706	521	2789	-	7		
08	350	668	-	23	29	249	329	-	43	66	384	5614	-	15	5706	521	2789	-	7	5706	521	2789	-	7		
32	258	711	+	8	273	329	-	19	201	2423883	575	0	-	030	5706	521	2789	-	7	5706	521	2789	-	7		
57	239	756	+	2	33	244	338	-	34	4012	321	245	+	2	5706	521	2789	-	7	5706	521	2789	-	7		
62	265	765	+	30	269	338	-	9	32	290	283	+	7	5706	521	2789	-	7	5706	521	2789	-	7			
63	349	767	+	4	56	302	390	-	10	4294	433	782	-	6	5706	521	2789	-	7	5706	521	2789	-	7		
93	341	821	+	11	4264	529	860	+	26	5361	449	2813	-	3	5706	521	2789	-	7	5706	521	2789	-	7		
6090	506	1176	+	54	92	440	923	+	30	5445	517	2973	+	7	5706	521	2789	-	7	5706	521	2789	-	7		
6155	420	1293	0	6	464	923	+	54	5502	217	3081	-	32	5706	521	2789	-	7	5706	521	2789	-	7			
61	516	1304	-	12	96	445	932	+	49	5716	571	3489	-	27	5706	521	2789	-	7	5706	521	2789	-	7		
6219	260	1408	-	16	5393	614	3409	+	8	592	3489	-	6	5706	521	2789	-	7	5706	521	2789	-	7			
39	278	1444	+	12	5438	347	3510	+	24	45	504	3544	+	11	5706	521	2789	-	7	5706	521	2789	-	7		
44	249	1453	-	15	369	3510	+	24	526	3544	+	33	5706	521	2789	-	7	5706	521	2789	-	7				
6562	443	2026	+	7	41	507	3517	+	61	64	427	3580	+	21	5706	521	2789	-	7	5706	521	2789	-	7		
63	538	2028	-	9	45	495	3526	+	62	97	510	3643	+	6	5706	521	2789	-	7	5706	521	2789	-	7		
66	310	2033	-	13	5500	311	3650	-	49	201	234	3652	-	11	5706	521	2789	-	7	5706	521	2789	-	7		
70	221	2040	+	11	334	3650	-	26	08	303	3668	-	30	5706	521	2789	-	7	5706	521	2789	-	7			
71	330	2042	+	10	334	3650	-	26</																		

TABLE 5 (continued).

213	d	t	d	2425719'630	3311	—	1	2426155'341	4447	+	5	224	d	t	d	2426243'290	1682	+	22
	d	t	d	6090'506	4011	—	2	6561'481	5293	—	10		d	t	d	6480'515	2186	—	10
2424029'249	0	+	'002	6565'276	4909	—	15	62'464	5295	+	13	2423992'289	0	—	'014	6570'425	2377	—	13
5420'495	2584	—	16					63'419	5297	+	7	4025'275	70	+	5	.446	2377	+	8
42'577	2625	—	9	217				66'299	5303	+	7	33'269	87	—	8	73'252	2383	—	11
'599	2625	+	13									.293	87	+	16				
73'275	2682	—	1	2423880'546	0	—	'018	221				5410'369	3011	—	8	238			
75'420	2686	—	10	83'546	4	—	15	2423996'264	0	+	'011	5500'311	3202	—	20	2423990'345	0	+	'013
80'258	2695	—	17	4016'247	181	+	51	4028'304	70	—	6	.334	3202	+	3	94'289	7	—	20
.279	2695	+	4	4293'436	551	—	18	.328	70	+	18	.356	3202	+	25	5393'614	2470	+	15
5501'285	2734	+	11	5381'485	2003	—	20	29'225	72	—	1	01'285	3204	+	12	5482'243	2626	+	17
08'258	2747	—	15	5507'350	2171	—	45	4293'460	649	—	4	02'217	3206	+	2	.265	2626	+	39
.280	2747	+	7	.372	2171	—	23	5415'427	3099	—	17	5740'508	3712	—	15	99'251	2656	—	19
22'277	2773	+	5	26'220	2196	+	91	65'342	3208	—	18	97'510	3833	+	1	5500'400	2658	—	6
23'362	2775	+	13	6156'324	3037	—	5	5508'385	3302	—	23	5808'339	3856	—	3	07'216	2670	—	8
5707'505	3117	+	18	6562'464	3579	—	11	09'329	3304	+	5	6155'431	4503	—	11	08'385	2672	+	25
6562'486	4705	—	3	63'250	3580	+	26	31'304	3352	—	1	6563'290	4595	+	16	24'239	2700	—	28
				65'449	3583	—	23	5798'305	3935	+	14	.312	5459	+	14	53'242	2751	—	0
214				66'229	3584	+	8	6563'527	5606	0		64'231	5461	—	8	5797'510	3181	—	25
2423883'575	0	—	'001	218				.548	5606	+	21					5808'339	3200	+	10
4012'371	241	+	30	2423880'517	0	+	'011	66'276	5612	+	1	231				6093'506	3702	—	22
4294'433	769	—	15	3992'264	244	+	5	222				2424030'222	0	—	'007	6562'228	4527	—	2
5415'427	2867	+	32	4024'325	314	+	5	2424029'249	0	+	'030	5420'562	2469	—	21	63'333	4529	—	33
38'347	2910	—	23	25'226	316	—	10	33'293	7	+	14	41'431	2506	+	12	.355	4529	—	11
.369	2910	—	1	30'263	327	—	11	5383'605	2335	+	10	42'533	2508	—	12	.398	4529	+	32
.417	2910	+	47	5441'379	3408	—	12	5451'455	2452	—	4	42'555	2508	+	10	66'229	4534	+	22
5500'311	3026	—	37	45'517	3417	+	4	.477	2452	+	18	80'279	2575	+	4				
.356	3026	+	8	51'455	3430	—	12	65'342	2476	—	38	98'278	2607	—	17	239			
22'236	3067	—	18	.477	3430	+	10	.364	2476	—	16	5507'350	2623	+	45	2423987'285	0	—	'010
61'234	3140	—	24	5507'329	3552	—	15	79'280	2500	—	20	51'233	2701	+	5	4031'333	63	+	40
5706'521	3412	—	64	.350	3552	+	6	.303	2500	+	3	5719'608	3000	+	6	4264'529	397	—	26
38'603	3472	—	40	5707'505	3989	+	12	5508'303	2550	+	1	94'497	3133	—	1	46'464	437	—	27
6093'506	4136	+	92	39'546	4059	—	7	22'236	2574	+	13	6155'431	3774	—	30	5385'463	2002	—	6
6155'431	4252	+	39	45'526	4072	+	19	5745'504	2959	—	31	6561'459	4495	—	14	5445'517	2088	—	14
6561'459	5012	+	4	94'497	4179	—	17	.526	2959	—	9	63'211	4498	+	48	5716'592	2476	+	87
62'507	5014	—	17	.519	4179	+	5	98'305	3050	—	13	66'513	4504	—	28	39'562	2509	+	10
66'252	5021	—	12	6155'431	4967	+	8	99'440	3052	—	38	235				.572	2509	+	20
215				6563'505	5858	—	1	.462	3052	—	16	2424020'305	0	—	'016	97'488	2592	—	31
2424019'253	0	+	'015	66'252	5864	—	2	6155'638	3666	+	20	22'277	3	—	10	.510	2592	—	9
30'222	18	+	10	219				6562'206	4367	—	15	4292'440	415	+	29	6562'228	3687	—	26
4298'446	458	—	2	2423990'345	0	+	'005	.228	4367	+	7	5551'233	2335	—	4	.249	3687	—	5
5437'231	2326	—	3	4017'225	49	+	13	63'376	4369	—	5	6156'400	3258	+	8	66'428	3693	—	17
79'280	2395	—	18	18'295	51	—	14	.398	4369	+	17	6562'206	3877	—	26	71'347	3700	+	14
.325	2395	+	27	29'297	71	+	20	66'276	4374	—	5	.228	3877	—	4	240			
99'385	2428	—	31	30'360	73	—	14	.299	4374	+	18	.249	3877	+	17	2424025'299	0	—	'009
5507'350	2441	+	9	4264'540	500	—	5	.342	4374	+	61	6563'548	3879	+	5	4293'460	569	—	9
26'220	2472	—	20	97'446	560	—	3	223								5391'629	2899	+	64
32'288	2482	—	48	5441'431	2646	+	1	2423990'345	0	+	'010	236				5441'507	3005	—	14
51'233	2513	—	2	42'533	2648	+	6	91'293	2	—	8	2425474'231	0	+	'005	46'217	3015	—	17
5791'440	2907	+	12	73'231	2704	—	6	92'264	4	—	4	5500'356	51	—	5	5561'234	3259	+	6
6155'431	3504	+	55	75'420	2708	—	11	4020'305	62	+	2	.400	51	+	39	5745'504	3650	+	3
6566'264	4178	—	2	79'280	2715	+	10	5420'562	2959	—	16	9716'592	473	—	24	64'352	3690	0	
216				5507'216	2766	—	23	73'253	3068	—	11	91'418	619	—	15	94'497	3754	—	17
2423965'374	0	—	'006	24'239	2797	0		.275	3068	+	11	.440	619	+	7	98'305	3762	+	20
4026'295	115	—	15	5805'339	3315	+	24	74'231	3070	+	1	94'497	625	—	11	6241'268	4702	—	25
.320	115	+	10	6093'495	3835	+	7	.253	3070	+	23	.519	625	+	11	42'244	4704	+	8
5337'632	2590	+	5	6562'378	4690	+	1	.276	3070	+	46	5808'339	652	—	5	6562'228	5383	—	11
61'449	2635	—	20	63'462	4692	—	12	99'340	3122	—	25	6156'274	1331	—	25	66'470	5392	—	11
5437'231	2778	—	3	66'218	4697	+	2	.363	3122	—	2	6563'211	2125	+	26	.492	5392	+	13
42'533	2788	+	1	220				5500'311	3124	—	20	64'209	2127	—	1	242			
.555	2788	+	23	2424020'395	0	+	'012	.334	3124	+	3	237				2423878'5296	0	—	'0006
73'253	2846	—	9	5480'323	3041	—	9	5501'285	3126	—	13	5451'455	0	—	'012	4298'4229	2866	+	97
.275	2846	+	13	82'243	3045	—	9	61'234	3250	0		5739'546	612	—	20	5420'4951	10525	+	12
82'265	2863	—	4	5738'603	3579	—	17	6090'506	4345	0		97'488	735	+	20	42'3173	10674	—	58
5508'236	2912	+	6	91'418	3689	—	12	6242'260	4659	—	19	5825'248	794	+	6	98'2783	11056	—	97
26'242	2946	—	3	.440	3689	+	11	6562'271	5321	+	12	6155'244	1495	+	6	5507'3723	11118	—	10
53'264	2997	—	2	6090'527	4312	+													

TABLE 5 (continued).

d	t	d	d	t	d	d	t	d	303	d	t	d													
2484	11240	+	35	6565	428	4635	-	4	19	253	287	+	7	4028	304	305	+	6							
5716	5924	12546	+	122	449	4035	+	17	32	266	408	-	2	4261	611	790	-	13							
6562	5070	18320	+	77	250			4261	611	2539	+	5	2423	877	528	0	-	026	4264	529	796	+	12		
245			2423	881	552	0	-	006	93	460	2835	-	1	3992	264	165	+	19	4264	529	796	+	19		
2424	028	255	0	-	037	5442	511	2776	+	8	5360	498	12750	-	13	94	338	168	+	8	5385	440	3126	+	5
4292	464	401	+	20	51	477	2792	-	23	85	486	12982	+	8	4017	225	201	-	44	5385	440	3126	+	28	
5442	599	2147	+	3	5508	280	2893	-	12	91	605	13039	-	8	26	320	214	+	15	5393	614	3143	+	0	
99	251	2233	+	4	5904	223	3597	+	71	5410	344	13213	+	5	33	244	224	-	12	5475	420	3313	+	22	
5501	234	2236	+	11	6561	459	4766	-	21	20	562	13308	0	5	269	224	+	13	5500	400	3365	-	14		
257	2236	+	34	481	4766	+	1	45	517	13540	-	13	5393	614	2181	+	50	5791	440	3970	-	30			
22	277	2268	-	26	6565	428	4773	+	12	51	477	13595	+	28	638	2181	+	64	6155	638	4727	-	13		
26	242	2274	-	13	66	513	4775	-	28	74	457	13809	-	23	5441	507	2250	-	19	6242	244	4907	-	2	
53	264	2315	+	1	255			80	258	13863	-	33	74	231	2297	+	35	6566	492	5581	-	4			
61	234	2327	+	66	2424	017	254	0	+	007	99	340	14040	0	76	255	2300	-	26	310					
5716	592	2563	-	38	29	249	22	-	17	5507	307	14114	+	3	5501	257	2336	-	48	2423	996	238	0	-	021
6562	421	3847	-	25	5383	605	2501	-	1	08	385	14124	+	5	285	2336	-	20	2423	996	238	0	+	5	
443	3847	-	3	5410	369	2550	-	7	27	223	14299	+	9	5508	211	2346	-	45	3997	296	2	+	5		
464	3847	+	18	42	599	2609	-	10	5739	562	16272	+	15	236	2346	-	20	4028	255	62	+	8			
66	384	3853	-	14	632	2609	+	23	6565	428	23946	+	7	258	2346	+	2	29	273	64	-	6			
246			5798	305	3260	+	38	273			280	2346	+	24	24	239	2369	-	4	30	335	66	+	24	
2424	026	2954	0	+	0032	6155	550	3914	-	14	5719	608	2650	+	42	95	328	2759	-	4	31	333	68	-	10
5360	4984	695	+	231	6566	363	4666	-	37	4297	423	456	+	3	99	462	2765	-	40	4264	529	520	-	19	
85	4398	708	+	85	384	4666	-	16	98	446	458	-	32	6093	484	3188	-	45	552	520	+	4			
5410	3686	721	-	186	406	4666	+	6	5337	608	2423	+	4	506	3188	-	23	93	436	576	-	5			
37	2528	735	-	100	71	347	4675	+	30	81	508	2506	+	13	6155	409	3277	+	16	460	576	+	19		
5508	2805	772	-	108	268			83	605	2510	-	6	6242	244	3402	-	36	94	454	578	-	18			
31	3258	784	-	17	2424	293	436	0	-	011	5794	497	3287	-	5	6563	398	3864	-	18	5438	347	2795	+	34
6562	2063	1321	+	63	5501	234	1697	+	5	283			65	449	3867	-	52	41	431	2801	+	23			
248			5799	440	2116	+	1	2424	292	440	0	-	011	66	207	3868	+	11	45	517	2809	-	19		
2424	025	226	0	+	005	462	2116	+	23	2424	292	440	0	-	011	229	3868	+	33	5501	234	2917	-	23	
5420	562	2546	+	6	6562	378	3188	-	21	5415	427	2456	-	2	252	3868	+	56	285	2917	+	28			
41	379	2584	-	3	400	3188	+	1	41	507	2513	+	15	309			5738	603	3377	+	13				
75	329	2646	-	32	272			5507	350	2657	+	16	2423	881	552	0	-	016	95	328	3487	-	15		
80	301	2655	+	8	2423	988	368	0	+	009	08	258	2659	+	10	3988	368	222	0	99	462	3495	-	9	
323	2655	+	30	98	264	92	+	4	24	239	2694	-	13	4016	247	280	-	24	5904	223	3698	+	16		
5502	217	2695	+	2	4012	346	223	-	12	6093	506	3939	-	8	271	280	-	0	6562	528	4974	-	19		
25	226	2737	-	7	272			6563	548	4967	-	7	309			550	4974	+	3						
5745	526	3139	-	23	272			6563	548	4967	-	7	309			550	4974	+	3						

TABLE 6.

n	phase	brightness	n	phase	brightness	n	phase	brightness	n	phase	brightness	n	phase	brightness	n	phase	brightness	n	phase	brightness	n	phase	brightness
P	35	st	P	st	P	st	P	st	P	112	st	P	st	P	st	P	st	P	st	P	st	P	st
10	017	6.09	10	017	6.09	10	037	7.75	10	521	15.64	10	726	2.79	10	244	17.08	10	732	11.01	10	732	11.01
10	049	7.31	10	049	7.31	10	069	2.28	10	553	16.12	10	758	2.44	9	262	16.44	10	778	12.67	10	778	12.67
10	085	8.90	10	085	8.90	10	094	1.32	10	590	16.28	10	787	4.83	10	277	16.57	10	840	13.82	10	840	13.82
10	118	9.74	10	118	9.74	10	117	.87	10	619	16.29	10	817	5.71	10	298	16.16	10	900	14.73	10	900	14.73
10	162	9.85	10	162	9.85	10	136	3.67	10	645	16.81	10	845	6.26	10	322	17.14	10	963	14.93			
10	205	10.88	10	205	10.88	10	172	4.35	10	666	16.34	10	873	6.70	10	356	17.06						
10	250	11.14	10	250	11.14	10	205	6.67	10	683	16.17	10	907	7.97	10	389	16.57						
10	283	12.86	10	283	12.86	10	231	7.55	10	722	16.57	10	940	7.90	10	419	17.23	11	018	11.99			
10	314	12.03	10	314	12.03	10	261	8.86	10	746	16.09	10	978	9.63	10	446	16.64	10	053	11.66			
10	341	12.72	10	341	12.72	10	292	9.46	10	777	16.57	10	395	12.90	10	471	15.91	10	075	11.53			
10	371	12.16	10	371	12.16	10	322	10.44	10	803	16.20	10	437	13.71	10	499	11.91	10	119	10.35			
10	400	12.06	10	400	12.06	10	356	12.29	10	846	15.96	10	501	13.02	10	529	7.29	10	169	10.29			
10	426	12.69	10	426	12.69	10	384	12.91	10	881	16.72	10	552	14.07	9	554	5.93	10	205	9.19			
10	454	12.87	10	454	12.87	10	414	13.26	10	920	16.75	10	599	12.13	10	584	6.85	10	236	8.14			
10	496	12.87	10	496	12.87	10	446	14.73	10	945	16.62	5	645	6.86	10	630	8.19	10	267	8.30			
10	536	12.75	10	536	12.75	10	469	13.67	9	966	16.21	5	664	4.36	10	670	9.57	10	305	8.25			
			10	012	13.94	10	500	15.02	9	988	15.18	10	694	2.80	10	703	10.71	10	340	8.54			

TABLE 6 (continued).

<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness									
	P	st		P	st		P	st		P	st		P	st		P	st									
10	'371	7:25	10	'435	7:12	10	'496	10:66	10	'644	12:47	10	'426	15:27	10	'697	4:92			10	'064	18:45				
10	'405	7:76	10	'462	6:93	10	'567	10:88	10	'665	12:97	10	'459	16:10	10	'727	4:89			10	'030	20:50	10	'095	18:09	
10	'434	7:01	10	'496	7:44	10	'611	11:24	10	'692	13:47	10	'490	15:94	10	'765	5:52			10	'066	20:93	10	'138	18:02	
10	'489	6:75	10	'520	7:76	10	'664	7:81	10	'724	14:36	10	'526	16:05	10	'794	5:51			10	'107	21:55	10	'196	14:18	
10	'538	6:96	10	'550	7:45	10	'720	6:33	10	'771	14:49	10	'559	15:86	10	'819	5:67			10	'144	21:36	10	'226	12:30	
10	'580	8:36	10	'576	8:32	10	'759	4:09	10	'816	14:32	10	'600	13:51	10	'846	6:83			10	'167	21:41	5	'242	9:68	
10	'608	6:80	10	'616	6:97	10	'785	3:88	10	'851	14:59	10	'631	9:99	10	'879	6:91			5	'195	18:54	5	'257	6:66	
10	'631	8:60	10	'641	7:69	10	'813	5:12	10	'887	14:10	10	'668	4:61	10	'916	5:71			5	'230	9:34	10	'287	5:33	
10	'661	8:31	10	'681	7:19	10	'843	4:75	10	'932	13:79	10	'697	4:68	9	'937	5:56			10	'264	1:71	10	'325	5:94	
10	'681	9:13	10	'704	8:09	10	'890	5:81	10	'979	13:46	10	'727	6:39	9	'964	5:40			10	'302	:88	10	'355	8:53	
10	'715	9:43	10	'726	8:15	10	'934	6:54			126	10	'763	7:02	9	'986	5:04			10	'348	3:92	10	'399	9:90	
10	'736	9:18	10	'760	7:82	11	'977	7:88				10	'795	8:73					10	'378	5:93	10	'469	12:56		
10	'768	10:64	10	'786	6:37				30	'028	8:04		10	'832	9:52			134		10	'409	7:23	10	'511	13:58	
10	'798	11:00	10	'815	6:71			123	30	'077	7:69		10	'860	9:75			10	'019	17:73	10	'434	8:84	10	'540	13:70
10	'837	10:75	10	'846	5:67	10	'037	8:68	30	'129	6:90		10	'889	11:71			10	'060	17:66	10	'462	9:86	10	'575	14:80
10	'868	11:87	10	'868	6:56	10	'094	7:20	30	'168	6:00		10	'923	11:90			10	'096	18:25	10	'479	10:44	10	'616	15:68
10	'899	12:62	10	'891	3:81	10	'132	7:13	30	'220	4:32		10	'945	12:68			10	'144	18:10	10	'498	10:82	10	'664	16:30
10	'925	12:01	10	'921	3:12	10	'180	4:44	30	'271	2:68		10	'982	13:66			10	'198	17:84	10	'522	11:91	10	'698	16:71
11	'961	12:11	11	'950	2:90	10	'213	3:92	30	'336	2:12							10	'245	18:05	10	'544	12:99	10	'733	16:81
10	'986	11:47	11	'975	2:52	10	'243	3:62	30	'398	2:25							10	'281	18:23	10	'566	14:52	10	'775	17:28
						10	'281	3:44	30	'460	2:20							10	'323	17:48	10	'603	14:67	10	'807	17:16
						10	'317	4:36	30	'531	2:63		14	'049	10:14			10	'360	18:25	10	'633	15:31	10	'848	16:83
						10	'356	4:21	30	'596	2:92		14	'117	10:09			10	'385	18:61	10	'660	16:51	10	'890	17:85
						10	'402	6:15	30	'657	4:03		14	'227	10:01			10	'411	17:16	10	'683	16:94	11	'940	17:38
						10	'444	5:90	30	'712	5:53		14	'274	10:22			10	'438	14:21	10	'707	16:62	11	'982	17:65
						10	'478	6:27	30	'773	6:06		14	'324	10:74			10	'467	11:85	10	'729	17:83			
						10	'520	6:89	30	'838	6:29		14	'382	10:76			10	'487	11:09	10	'750	18:24			
						10	'558	7:49	30	'910	7:04		14	'427	11:32			10	'511	7:39	10	'778	19:96	11	'025	13:30
						10	'600	6:94	28	'973	7:68		14	'465	9:67			10	'532	8:06	10	'810	19:69	11	'083	13:41
						10	'630	8:00					14	'493	8:46			10	'556	8:63	10	'833	19:83	11	'131	14:11
						10	'670	7:61					14	'511	7:74			10	'586	9:63	10	'860	19:93	10	'199	14:88
						10	'712	7:54	10	'022	7:72		14	'538	5:53			10	'613	9:48	10	'887	20:75	10	'271	14:17
						10	'760	8:73	10	'060	4:32		14	'575	4:08			10	'645	11:54	10	'914	20:87	10	'336	15:16
						10	'823	8:90	10	'092	4:26		14	'644	6:01			10	'684	12:07	10	'943	20:18	10	'403	15:22
						10	'872	8:31	10	'122	4:64		14	'790	7:87			10	'710	12:96	9	'973	21:23	10	'438	14:95
						10	'925	8:35	10	'157	6:55		14	'892	9:31			10	'735	14:28				10	'481	15:24
						10	'972	8:55	10	'186	7:62		14	'947	9:53			10	'761	14:19				10	'537	15:93
									10	'219	7:56							138		5	'005	10:54	10	'606	15:04	
									10	'275	8:64							10	'797	15:20	5	'033	9:20	10	'675	15:77
									10	'334	9:85							10	'832	15:21	10	'079	4:41	10	'718	14:91
						10	'015	13:71	10	'387	10:84		10	'013	4:98			10	'867	16:27	10	'108	4:53	10	'757	11:78
						10	'046	14:48	10	'452	11:83		10	'038	4:67			10	'899	16:59	10	'135	5:30	10	'792	6:43
						10	'068	14:17	10	'510	12:14		10	'061	4:37			10	'935	16:73	10	'169	5:92	10	'826	5:90
						10	'100	13:49	10	'544	11:94		10	'090	4:28			11	'979	17:23	10	'223	6:20	10	'853	7:24
						3	'113	13:50	10	'585	12:16		10	'117	4:91			10	'227	6:20	10	'291	7:75	10	'886	7:92
						3	'132	9:37	10	'641	13:12		10	'142	4:93			10	'223	6:20	10	'324	9:23	10	'922	9:43
						3	'142	6:13	10	'707	12:99		10	'168	4:70			10	'031	11:70	10	'353	8:21	10	'972	10:98
						10	'155	4:34	10	'774	12:90		10	'193	4:77			10	'082	5:55	10	'401	9:02			
						10	'187	:42	10	'837	13:32		10	'219	5:05			10	'127	4:21	10	'401	9:02			
						10	'218	—1:50	10	'837	13:32		10	'252	5:93			10	'174	5:59	10	'449	9:34			
						10	'258	1:10	10	'898	13:13		10	'277	6:85			10	'227	5:72	10	'491	10:30			
						10	'290	2:64	10	'948	13:92		10	'301	8:23			10	'283	9:47	10	'524	10:38	10	'031	12:05
						10	'314	2:83	10	'988	11:60		10	'327	13:68			10	'326	8:69	10	'569	11:30	9	'100	10:40
						10	'332	3:32					10	'352	16:11			10	'413	12:85	10	'621	11:22	3	'164	5:17
						10	'369	5:00					10	'383	14:99			10	'457							

TABLE 6 (continued).

<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness						
II			P			P			P			P			P			P					
P st			P st			P st			P st			P st			P st			P st			P st		
10	'029	12'32	10	'684	9'73	10	'771	10'16	10	'036	12'29	10	'179	7'35	10	'889	5'48	10	'604	8'29	10	'016	6'60
10	'078	11'42	10	'709	10'28	10	'817	10'25	10	'077	12'52	10	'225	8'17	10	'950	5'16	10	'634	8'60	10	'064	7'30
10	'121	12'57	10	'739	10'90	10	'860	11'51	10	'077	12'52	10	'273	7'69	11	'988	4'76	10	'670	8'56	10	'129	6'70
10	'171	8'85	10	'823	10'73	10	'892	11'92	10	'126	12'16	10	'345	8'81	160			10	'716	8'90	10	'210	6'90
5	'211	2'90	10	'852	10'36	10	'916	11'45	10	'178	12'30	10	'416	8'33	10	'014	5'87	10	'762	9'22	10	'267	7'10
5	'244	44	10	'882	10'87	11	'947	11'19	10	'228	12'67	10	'519	9'22	10	'038	6'64	10	'817	9'35	10	'311	7'30
5	'276	—'60	10	'907	10'09	11	'980	11'61	10	'274	12'33	10	'574	8'04	10	'058	6'86	10	'858	9'57	10	'359	7'20
5	'312	3'12	10	'933	9'88	147			10	'335	12'33	10	'632	8'57	10	'079	6'30	9	'928	9'20	10	'424	7'20
10	'357	2'80	10	'952	10'11	10	'056	11'26	10	'392	11'80	10	'727	—'11	10	'108	6'57	9	'954	9'26	10	'475	7'20
10	'418	4'47	11	'986	10'67	10	'115	12'16	5	'470	7'58	9	'792	—1'63	10	'135	6'35	9	'980	8'47	10	'530	7'30
10	'476	6'71	145			10	'158	12'21	10	'498	4'71	9	'870	1'22	10	'161	6'37	162			10	'587	5'25
10	'523	7'76	10	'014	7'83	10	'205	12'87	10	'516	4'31	9	'976	5'58	10	'187	6'55	10	'045	5'65	5	'617	2'80
10	'579	8'69	10	'050	8'82	10	'258	12'25	10	'538	2'82	III			10	'211	6'85	10	'117	7'19	5	'640	1'70
10	'626	10'60	10	'085	9'65	10	'314	12'06	10	'555	4'67	10	'205	7'02	10	'234	6'19	10	'163	7'42	10	'674	1'45
10	'673	10'71	10	'128	9'69	10	'364	12'57	10	'572	4'33	10	'402	8'28	10	'254	6'71	10	'212	7'15	10	'716	2'74
10	'730	11'21	10	'175	10'21	10	'398	11'96	10	'596	5'68	10	'549	8'54	10	'279	6'35	10	'324	8'06	10	'760	3'56
10	'834	11'27	10	'206	10'24	10	'438	12'18	10	'619	6'31	10	'650	7'89	10	'309	6'49	10	'267	6'88	10	'806	4'76
10	'908	12'41	10	'235	10'75	10	'496	11'97	10	'643	6'93	10	'775	2'66	10	'334	6'46	10	'324	8'06	10	'849	5'60
10	'944	11'25	10	'278	10'84	10	'542	12'48	10	'672	7'79	9	'792	—1'63	10	'374	6'48	10	'387	8'20	10	'899	5'31
11	'983	12'50	10	'322	11'97	10	'576	8'88	10	'700	7'48	9	'896	—'90	10	'404	6'55	10	'437	7'94	10	'944	6'40
III			10	'359	11'43	10	'610	4'25	10	'732	8'80	157			10	'431	6'72	10	'489	7'61	10	'975	5'85
7	'062	11'51	10	'409	11'56	10	'640	3'34	10	'766	9'45	10	'006	5'10	10	'459	6'40	10	'555	7'91	167		
2	'130	8'65	10	'457	11'43	10	'671	3'21	10	'799	10'26	10	'023	5'36	10	'484	6'58	10	'594	8'32	10	'028	13'17
1	'153	7'00	10	'499	11'58	10	'709	3'67	10	'830	10'39	10	'039	5'50	10	'508	6'39	10	'640	6'90	10	'062	11'72
1	'168	2'00	10	'563	11'40	10	'748	4'88	10	'865	11'21	10	'063	5'51	10	'534	6'60	5	'687	3'90	10	'098	9'60
1	'171	1'50	10	'621	11'64	10	'799	7'36	10	'899	11'88	10	'110	4'99	10	'572	6'91	5	'705	—'72	10	'142	9'14
1	'194	—2'00	10	'681	11'41	10	'856	7'69	10	'936	11'58	10	'142	5'10	10	'606	6'10	10	'735	—'82	10	'181	5'96
2	'210	—'50	10	'760	11'31	10	'916	9'64	10	'977	12'19	10	'110	4'99	10	'631	6'56	10	'815	—'39	10	'207	6'50
3	'247	0'07	5	'787	10'92	9	'988	10'17	154			10	'142	5'10	5	'661	8'26	10	'846	—'18	10	'245	6'10
4	'296	2'10	5	'804	8'96	149			10	'009	5'13	10	'177	5'37	5	'670	9'44	10	'881	1'73	10	'279	6'01
6	'357	5'80	5	'815	7'26	10	'015	3'95	10	'046	5'50	10	'203	5'20	5	'680	11'14	10	'909	1'89	10	'303	6'63
6	'423	5'72	5	'839	3'96	10	'042	4'22	10	'085	6'06	10	'221	4'79	5	'689	13'58	11	'951	3'93	10	'338	7'24
6	'504	8'83	9	'865	3'44	10	'042	4'22	10	'138	7'83	10	'234	5'06	5	'696	14'80	11	'988	5'45	10	'385	9'84
6	'618	10'32	9	'912	4'67	10	'075	3'85	10	'182	6'85	10	'256	4'77	5	'714	16'52	10	'427	9'81	10	'427	9'81
6	'716	11'57	9	'946	6'24	10	'107	3'45	10	'224	7'19	10	'287	5'08	5	'734	11'40	164			10	'461	9'31
6	'905	11'18	9	'981	7'89	10	'146	3'16	10	'260	9'00	10	'311	4'84	5	'752	8'58	10	'014	8'30	10	'500	10'85
144			146			10	'187	3'10	10	'289	8'76	10	'343	4'97	10	'771	6'63	10	'049	9'52	10	'540	11'02
10	'005	10'64	11	'017	11'33	10	'215	3'41	10	'327	8'81	10	'362	5'12	10	'805	6'18	10	'096	11'00	10	'574	11'04
10	'031	10'53	11	'063	12'20	10	'247	3'06	10	'363	8'73	10	'381	5'03	10	'836	6'21	10	'096	11'00	10	'607	10'73
10	'054	10'82	11	'098	11'69	10	'283	2'80	10	'412	10'04	10	'399	5'05	10	'857	6'46	10	'140	11'81	10	'644	12'36
10	'089	11'13	10	'128	10'78	10	'312	3'11	10	'452	9'40	10	'424	5'54	10	'885	6'24	10	'171	12'18	10	'676	12'04
10	'120	10'00	5	'165	8'20	10	'346	2'93	10	'489	9'30	10	'472	5'14	10	'912	5'84	10	'212	12'04	10	'707	12'08
5	'147	7'64	5	'180	3'58	10	'379	3'04	10	'520	10'33	10	'509	5'38	10	'937	6'83	10	'250	13'07	10	'743	11'81
5	'156	6'04	10	'211	—'17	10	'403	3'52	10	'560	9'97	10	'528	5'50	10	'964	6'51	10	'294	13'49	10	'769	12'96
5	'176	4'34	10	'236	—'81	10	'429	3'59	10	'610	10'50	10	'547	5'18	10	'992	6'36	10	'337	14'30	10	'797	12'42
5	'192	1'62	10	'259	—'03	10	'472	3'70	10	'647	10'10	5	'567	5'12	161			10	'381	13'79	10	'833	12'03
10	'209	—'20	10	'282	1'23	10	'502	4'25	10	'683	10'03	5	'591	6'92	10	'014	9'34	10	'439	13'93	10	'860	13'34
10	'234	—'40	10	'309	2'24	10	'535	4'25	10	'716	9'08	5	'603	7'98	10	'048	9'29	10	'496	13'21	10	'896	12'85
10	'263	—'06	10	'338	3'16	10	'582	5'47	10	'753	9'51	5	'614	10'72	10	'099	9'77	10	'532	14'09	10	'928	13'45
10	'292	—'88	10	'360	4'73	10	'609	6'44	10	'792	8'83	5	'627	12'16	10	'099	9'77	10	'587	13'91	9	'964	11'33
10	'333	1'92	10	'388	5'56	10	'642	6'51	10	'833	8'56	5	'640	12'74	10	'141	8'84	10	'643	13'76	9	'988	12'90
10	'361	3'81	10	'413	4'41	10	'667	6'86	10	'872	5'75	5	'651	13'34	10	'176	5'67	10	'683	13'60	168		
10	'394	3'91	10	'433	6'52	10	'696	6'70	10	'912	3'72	5	'659	12'62	10	'203	3'60	5	'708	13'34	10	'022	11'73
10	'429	4'44	10	'454	6'45	10	'720	6'64	10	'938	3'72	5	'665	11'28	10	'231	1'91	4	'728	9'			

TABLE 6 (continued).

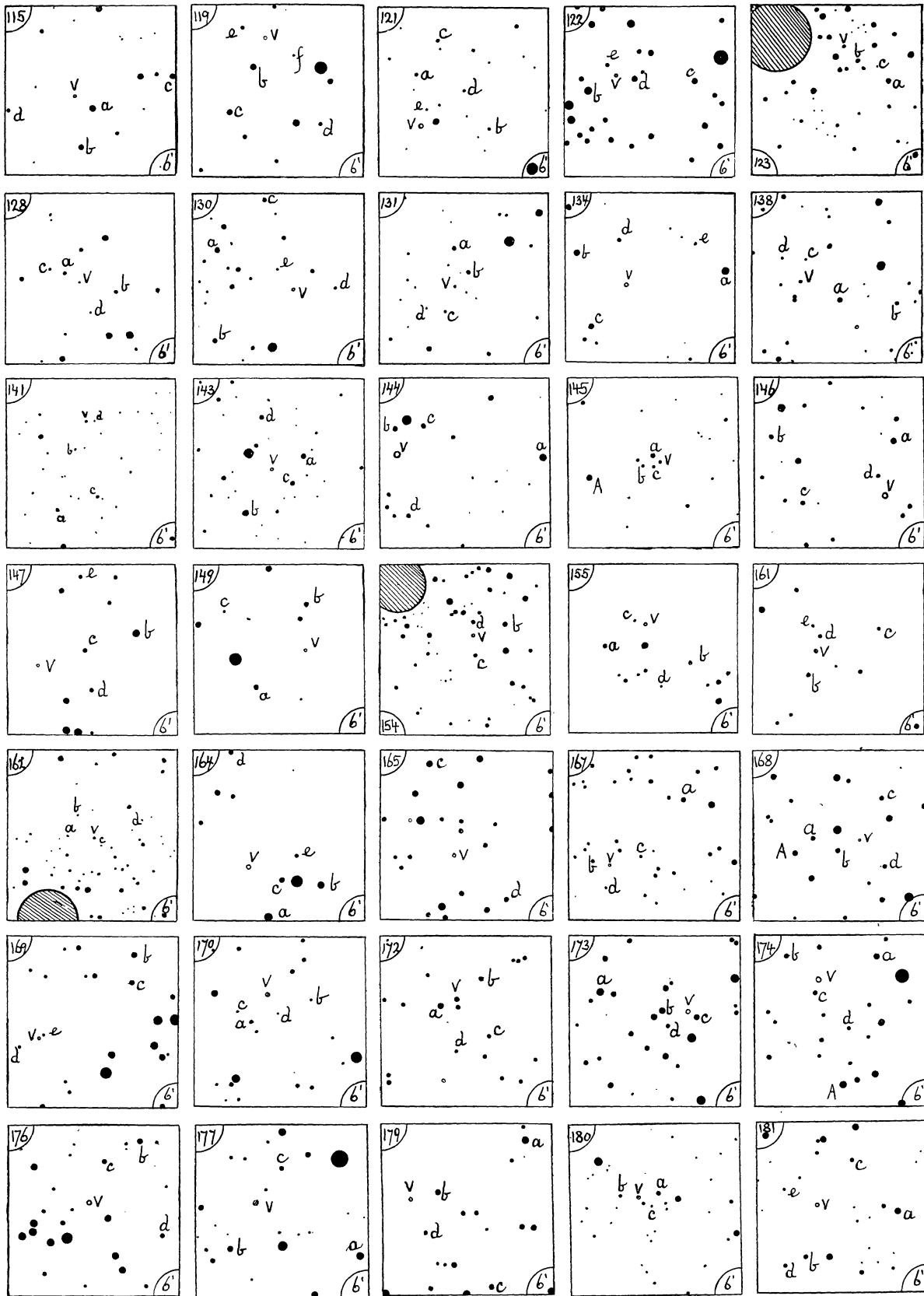
<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness							
	P	st		P	st		P	st		P	st		P	st		P	st		P	st							
10	'148	15'11	10	'075	14'00	10	'342	3'63	10	'463	6'32	10	'746	5'04	10	'866	7'85	10	'727	6'54	10	'710	9'93				
10	'206	14'85	10	'115	14'16	10	'368	5'27	10	'492	5'97	10	'785	6'38	10	'897	7'59	10	'750	7'79	10	'750	9'88				
10	'252	13'52	10	'156	14'42	10	'393	5'57	10	'532	6'93	11	'832	6'92	10	'919	7'39	10	'769	7'73	10	'794	9'89				
10	'288	9'87	10	'215	13'90	10	'424	7'41	10	'559	7'23	11	'884	7'43	11	'942	7'32	10	'791	7'93	10	'832	10'81				
10	'324	6'96	10	'259	14'36	10	'464	7'99	10	'606	6'57	11	'936	7'67	11	'958	8'29	10	'816	8'96	10	'875	9'98				
10	'360	5'72	10	'282	14'16	10	'503	9'15	10	'655	7'44	11	'983	7'70	11	'985	8'15	10	'851	9'56	9	'914	9'98				
10	'390	7'09	10	'321	14'18	10	'538	9'01	10	'694	8'07							10	'894	10'79	9	'945	10'40				
10	'420	7'72	10	'362	14'68	10	'571	10'64	10	'731	8'73		214			216		9	'923	9'43	9	'973	10'33				
10	'456	9'25	10	'397	14'12	10	'610	10'30	10	'768	8'76	10	'020	7'41	10	'033	12'36	9	'955	11'13							
10	'488	9'25	10	'429	11'71	10	'649	11'77	10	'807	9'37	10	'056	7'87	10	'086	12'87	9	'984	11'83							
10	'512	9'70	5	'454	9'90	10	'688	11'59	10	'843	9'38	10	'086	7'90	10	'124	12'97						220				
10	'540	10'29	5	'466	5'88	10	'729	12'31	10	'878	9'56	10	'107	7'49	12	'151	13'71			218		10	'020	10'00			
10	'579	11'80	10	'494	4'83	10	'761	12'71	9	'917	9'23	10	'137	7'82	10	'182	13'20					10	'017	9'22	10	'061	10'00
10	'610	11'87	10	'527	5'00	10	'794	12'97	9	'939	9'88	10	'165	8'00	10	'209	12'58	10	'054	9'34	10	'082	9'60	10	'106	11'10	
10	'638	11'89	10	'560	5'81	10	'828	12'90	9	'975	9'30	10	'190	8'13	10	'239	8'66	10	'082	9'60	10	'114	9'80	10	'156	9'41	
10	'670	12'60	10	'589	8'01	10	'857	12'67				10	'225	7'96	10	'268	2'85	10	'140	9'90	10	'140	9'90	10	'202	6'73	
10	'696	14'01	10	'615	7'97	10	'886	13'05		212		10	'273	7'69	10	'283	—'86	10	'178	9'52	10	'178	9'52	10	'229	4'07	
9	'717	13'21	10	'648	10'09	11	'923	13'03	10	'014	10'06	10	'308	7'82	10	'309	—1'50	10	'214	9'42	10	'214	9'42	10	'269	2'72	
9	'758	14'13	10	'683	10'34	11	'971	13'58	10	'046	10'35	10	'344	7'42	10	'331	—'94	10	'278	9'60	10	'278	9'60	10	'317	5'76	
9	'818	13'96	10	'726	11'57				10	'079	9'91	10	'372	8'20	10	'362	—'92	10	'324	9'80	10	'324	9'80	10	'366	6'38	
9	'876	14'94	10	'764	12'24		210		10	'110	10'36	10	'410	7'26	10	'382	'80	10	'364	10'00	10	'364	10'00	10	'405	7'92	
9	'942	14'07	10	'795	11'42	10	'015	3'85	10	'154	10'48	10	'447	7'42	10	'403	'96	10	'400	10'00	10	'400	10'00	10	'441	8'30	
	206		10	'832	12'61	10	'047	4'58	10	'188	10'58	10	'477	6'58	10	'418	1'99	10	'428	10'30	10	'428	10'30	10	'471	8'40	
10	'029	6'20	10	'863	12'74	10	'082	6'26	10	'225	11'05	10	'509	4'47	10	'445	3'50	10	'459	10'00	10	'459	10'00	10	'503	8'63	
10	'069	6'16	11	'912	13'61	10	'113	7'25	10	'274	10'48	10	'540	3'21	10	'471	4'76	10	'488	9'62	10	'488	9'62	10	'550	8'89	
10	'120	6'42	11	'966	13'45	10	'154	8'42	10	'304	9'55	10	'569	2'55	10	'496	5'78	10	'524	7'30	10	'524	7'30	10	'613	9'60	
10	'154	6'92		208		10	'204	9'32	10	'330	10'47	10	'603	3'30	10	'517	6'43	10	'554	4'98	10	'554	4'98	10	'747	9'90	
10	'182	6'45	10	'019	9'19	10	'292	11'09	10	'361	10'63	10	'631	1'86	10	'541	7'13	10	'566	2'08	10	'566	2'08	10	'787	10'20	
10	'207	6'49	10	'049	9'34	10	'322	11'41	10	'398	10'68	10	'658	4'76	10	'565	8'29	10	'581	'87	10	'581	'87	11	'835	9'67	
10	'229	5'98	10	'086	9'17	10	'342	12'07	10	'425	10'79	10	'689	3'60	10	'595	8'12	10	'610	1'17	10	'610	1'17	11	'878	9'86	
10	'250	6'13	10	'086	9'17	10	'342	12'07	10	'474	10'63	10	'716	4'97	10	'615	8'71	10	'635	2'16	10	'635	2'16	11	'923	9'95	
10	'279	6'78	10	'140	10'49	10	'366	11'64	10	'519	10'58	10	'744	5'78	10	'650	9'19	10	'669	3'87	10	'669	3'87	11	'973	9'95	
10	'299	6'31	10	'176	10'28	10	'395	11'79	10	'560	9'73	10	'772	6'62	10	'675	8'76	10	'698	5'06	10	'698	5'06				
10	'323	6'04	10	'204	10'60	10	'432	12'15	10	'602	5'92	10	'803	5'66	10	'704	11'22	10	'726	10'86	10	'726	10'86				
10	'350	6'27	10	'228	9'96	10	'465	12'46	10	'643	4'02	10	'839	7'32	10	'726	10'86	10	'739	5'48	10	'739	5'48				
10	'375	6'11	10	'260	10'49	19	'513	12'40	10	'685	5'03	10	'878	6'77	10	'765	12'20	10	'781	7'14	10	'781	7'14			221	
10	'407	5'97	10	'301	10'02	10	'577	13'09	10	'720	5'82	10	'933	7'49	10	'808	11'76	10	'810	7'35	10	'810	7'35	10	'020	13'42	
10	'474	6'05	10	'320	10'63	10	'615	12'97	10	'773	6'70	9	'975	7'67	10	'845	12'26	10	'848	7'84	10	'848	7'84	10	'053	13'00	
10	'508	6'19	10	'354	11'09	10	'647	12'54	10	'831	7'74				215		11	'874	11'97	10	'871	7'97	10	'089	13'39		
10	'533	5'96	10	'385	10'16	10	'685	12'74	10	'874	7'77	10	'015	9'19	11	'916	12'30	10	'900	8'17	10	'900	8'17	10	'123	13'84	
10	'566	6'30	10	'421	11'00	10	'732	12'74	10	'909	8'08	10	'045	8'99	11	'946	12'43	10	'927	8'66	10	'927	8'66	10	'158	13'55	
10	'610	5'97	10	'465	10'98	10	'767	13'88	11	'951	9'63	10	'077	9'95	11	'979	13'09	10	'966	8'92	10	'966	8'92	10	'198	14'13	
10	'648	6'35	10	'530	11'09	10	'795	14'31	11	'984	9'67	10	'110	11'19			217										
11	'683	6'35	10	'610	10'89	10	'831	13'52				10	'145	10'52	10	'012	11'90	10	'012	11'90	10	'012	11'90	10	'237	13'20	
3	'701	7'88	10	'650	9'54	9	'864	12'38		213		10	'201	11'96	10	'040	12'11	10	'040	12'11	10	'040	12'11	10	'265	11'52	
3	'719	11'30	10	'687	7'12	5	'895	5'86	10	'028	8'32	10	'246	12'59	10	'070	11'78	10	'070	11'78	10	'070	11'78	10	'300	5'50	
3	'729	10'63	10	'723	3'04	5	'913	4'46	10	'056	7'97	10	'270	12'79	10	'107	12'37	10	'045	11'14	10	'045	11'14	10	'328	3'66	
3	'738	11'43	10	'756	1'78	10	'936	2'72	10	'085	7'96	10	'299	12'13	10	'158	11'38	10	'075	10'92	10	'075	10'92	10	'361	2'70	
3	'742	11'63	10	'800	3'11	10	'955	1'18	10	'127	7'83	10	'299	12'13	10	'158	11'38	10	'109	9'44	10	'109	9'44	10	'382	1'62	
3	'751	11'87	10	'834	4'58	10	'987	2'45	10	'163	8'38	10	'328	13'49	10	'220	11'67	10	'142	5'44	10	'142	5'44	10	'415	3'21	
3	'761	10'37	10	'870	4'76				10	'198	7'60	10	'357	13'27	10	'259	12'75	10	'171	3'99	10	'171	3'99	10	'452	5'05	
3	'771	8'60	9	'902	5'59		211		10	'246	8'05	10	'385	13'30	10	'321	12'84	10	'197	1'58	10	'197	1'58	10	'481	6'63	
3	'778	10'17	9	'929	6'57	10	'023	7'93	10	'295	7'74	10	'414	13'91	10	'360	11'94	10	'225	1'88	10	'225	1'88	10	'515	7'30	
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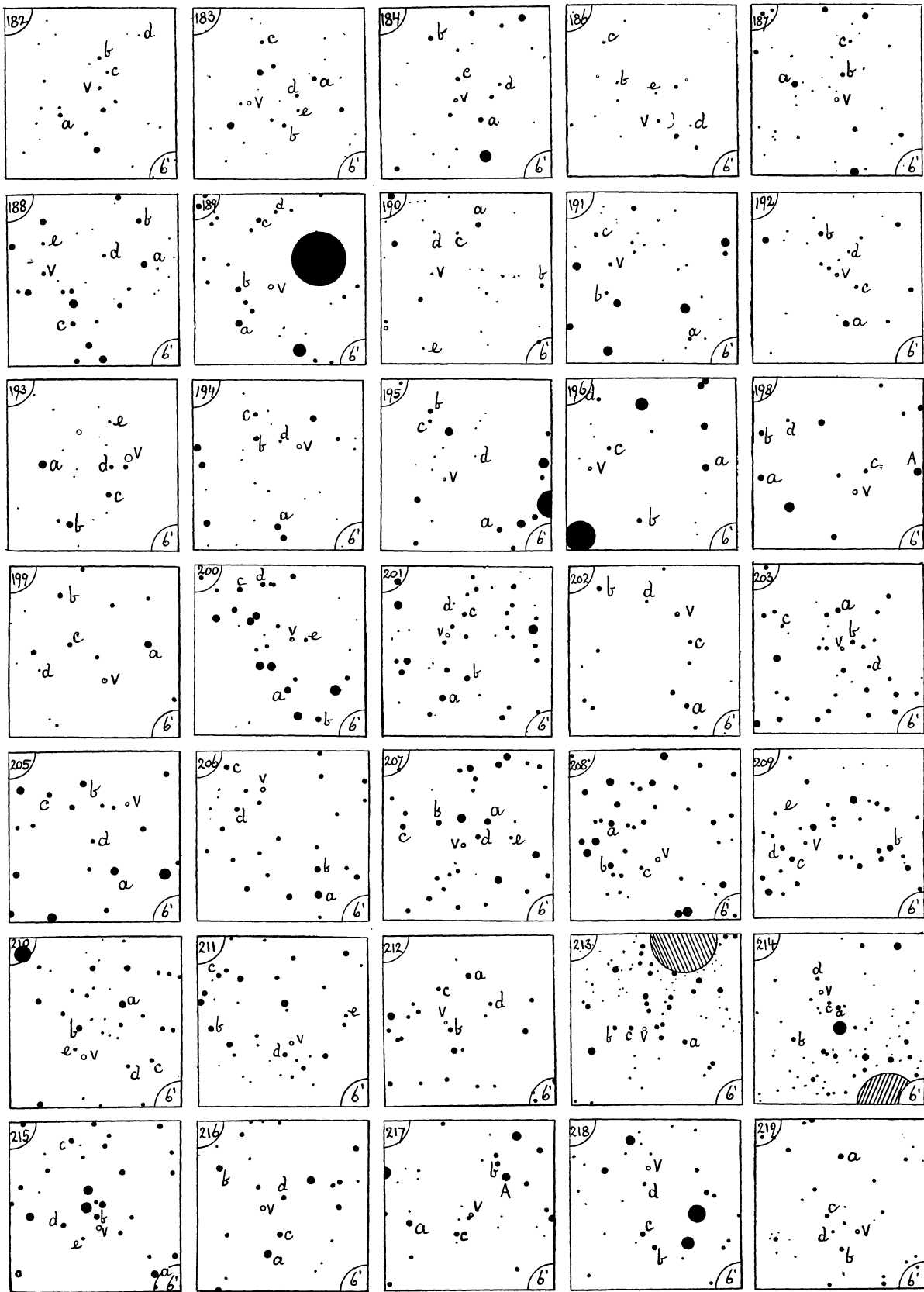
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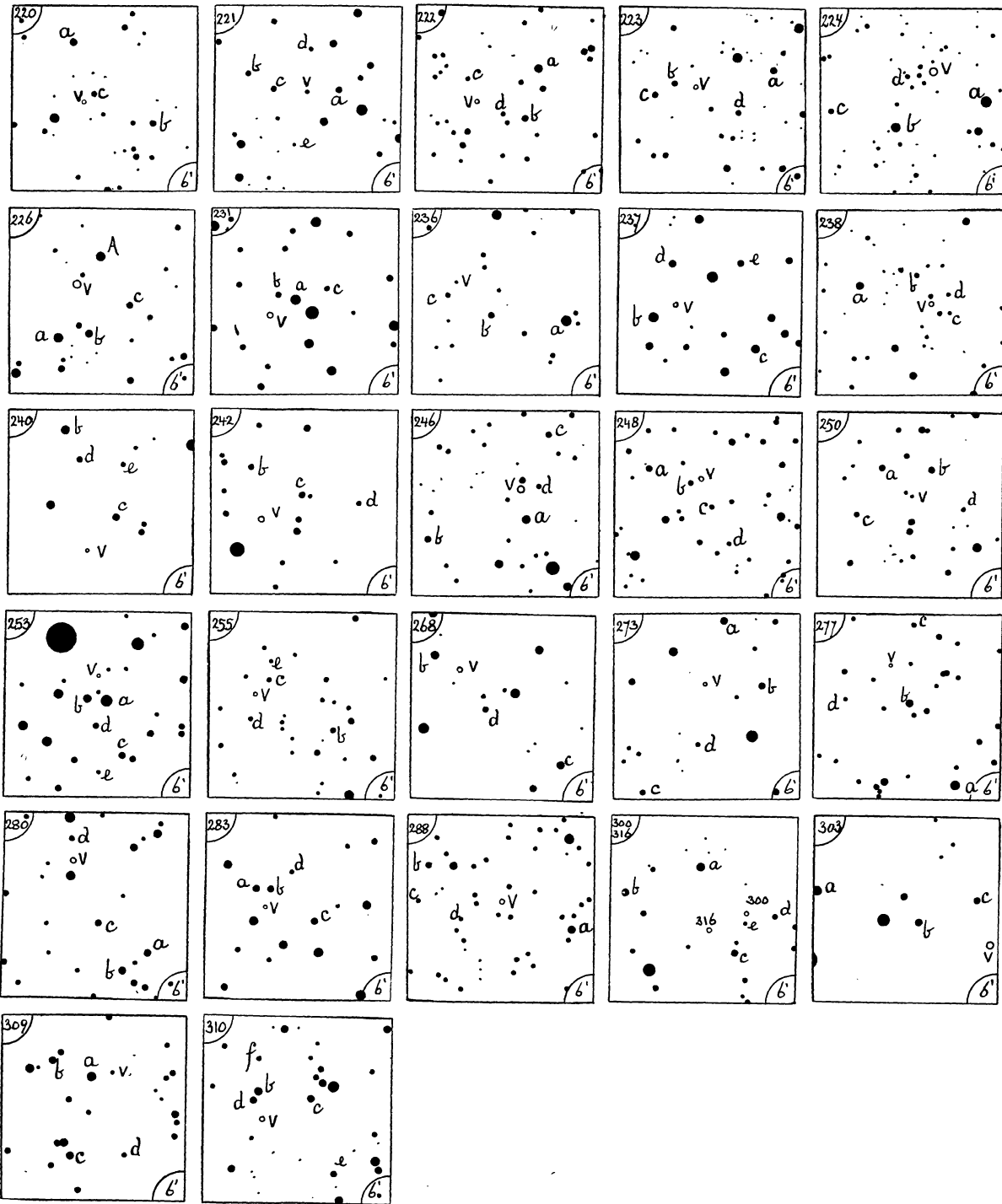
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I																									
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II + III			10	285 12:14	2	120 18:00			10	800 12:87	11	952 12:54	5	398 10:78											
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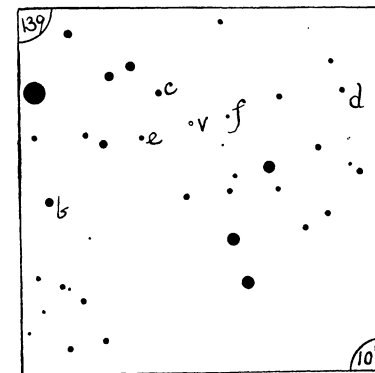
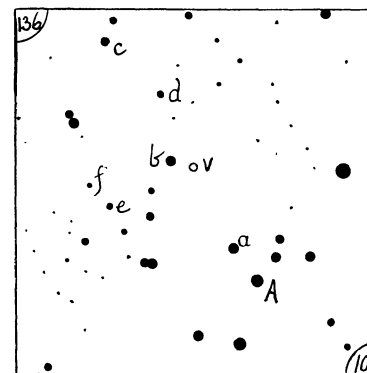
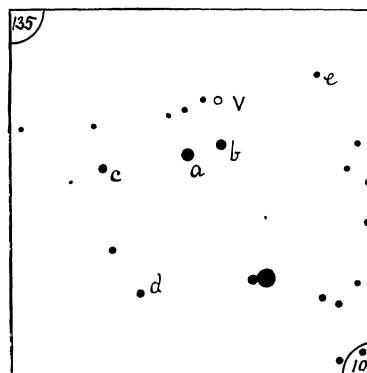
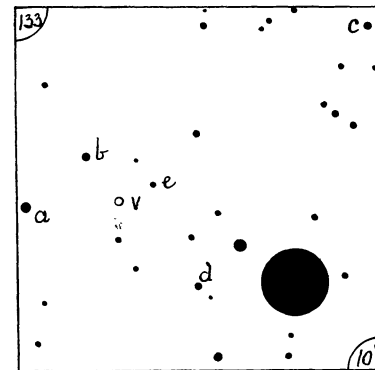
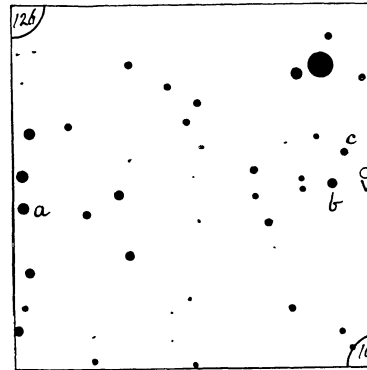
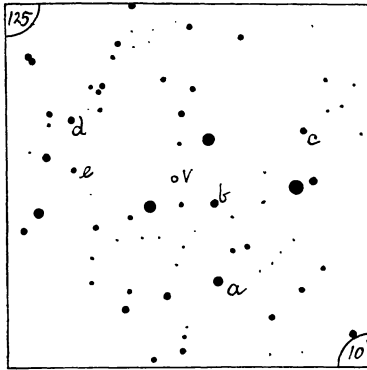
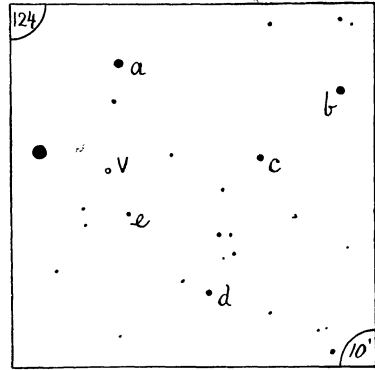
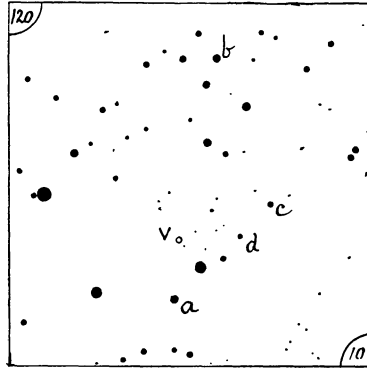
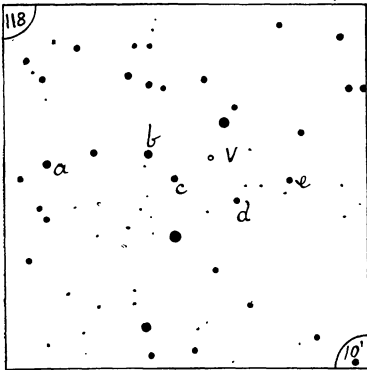
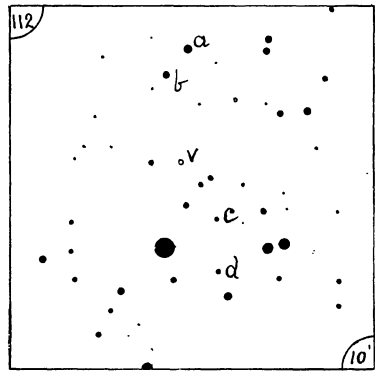
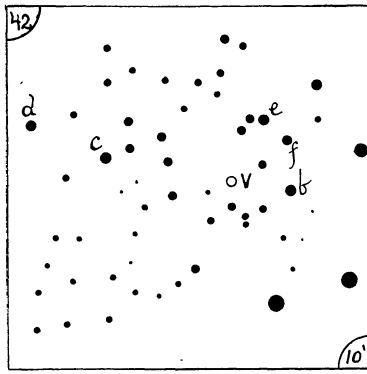
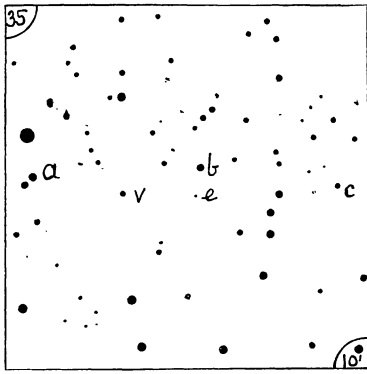
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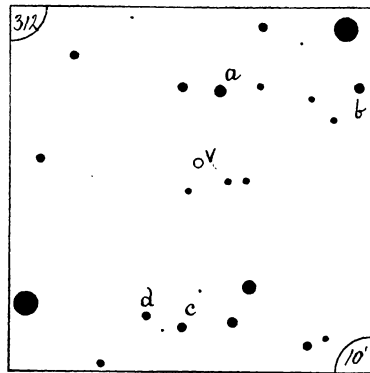
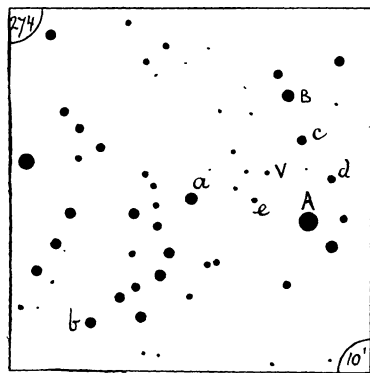
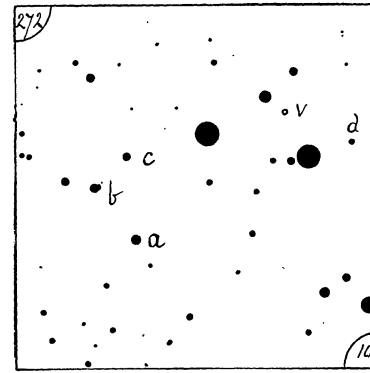
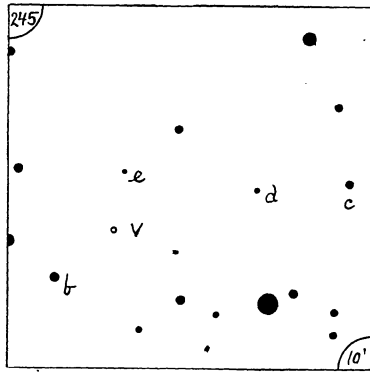
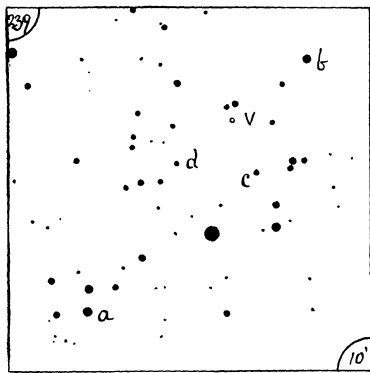
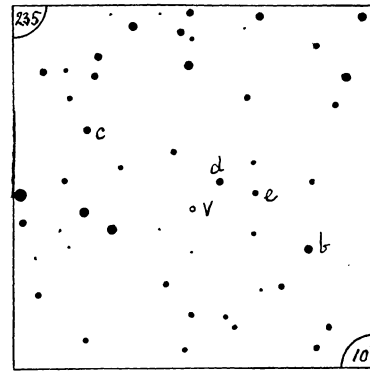
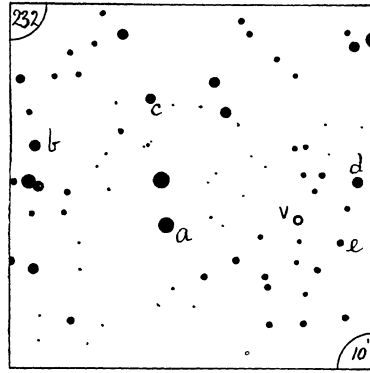
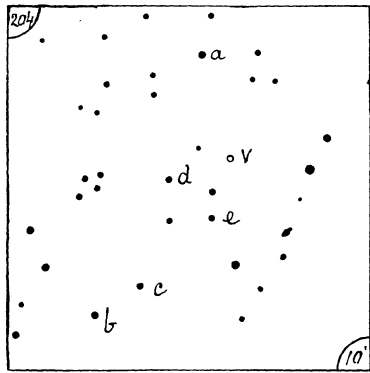
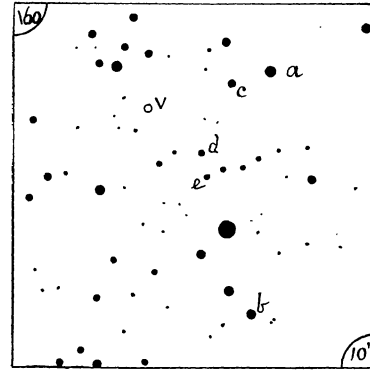
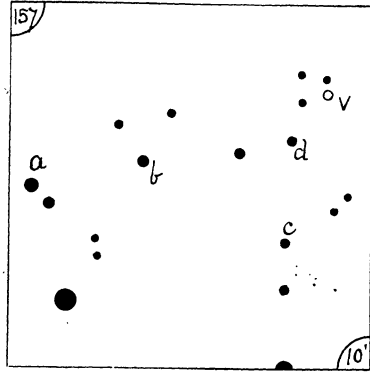
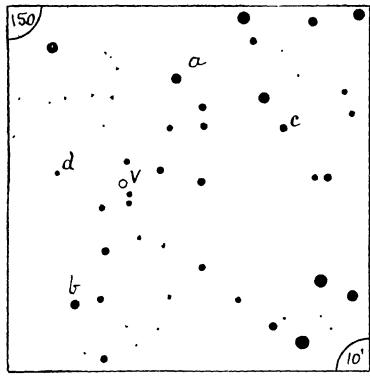
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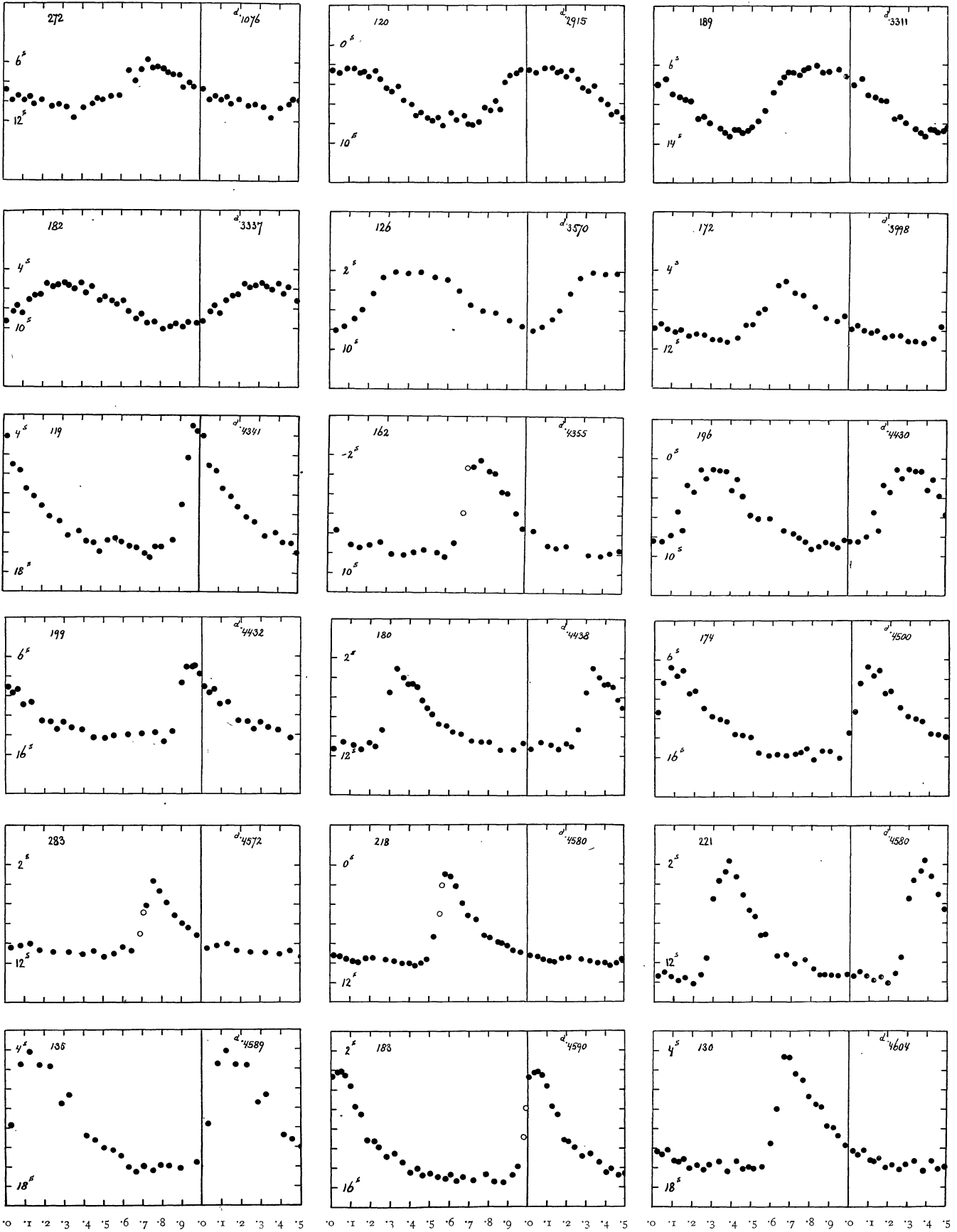


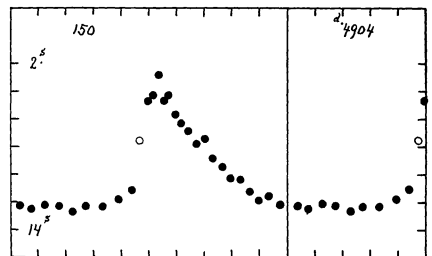
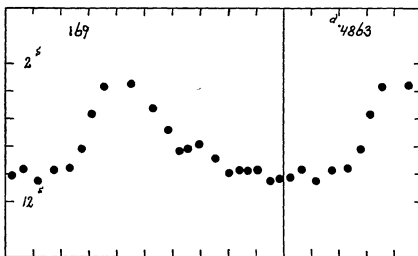
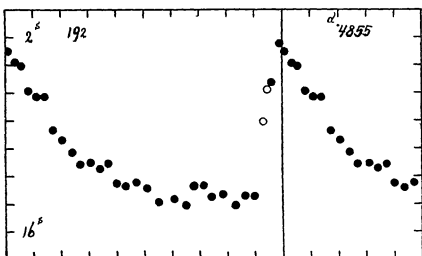
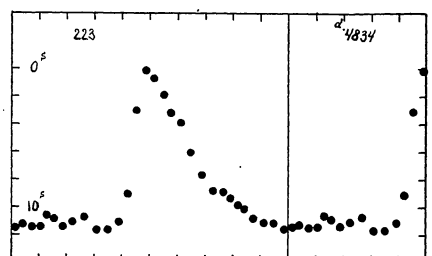
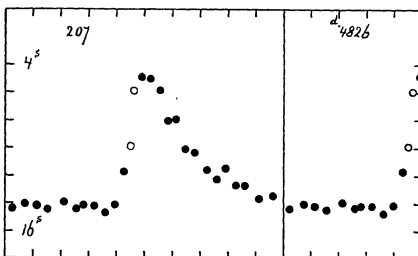
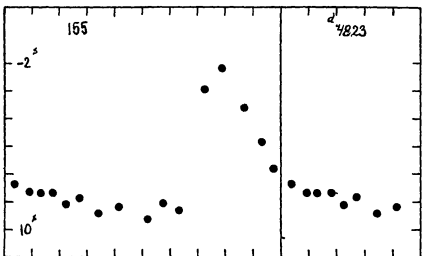
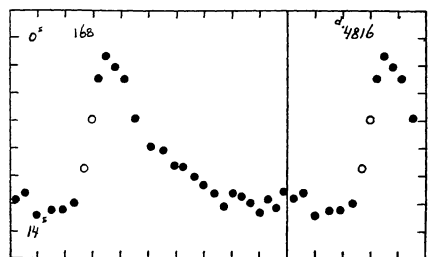
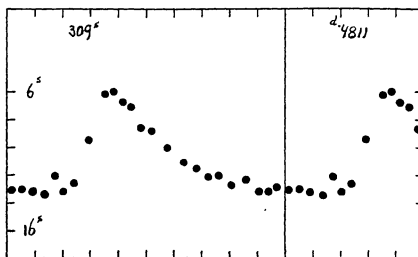
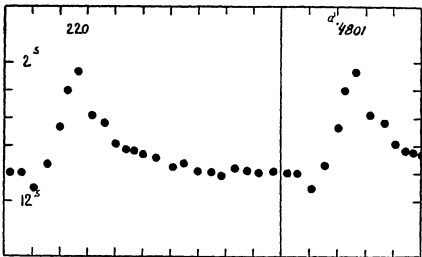
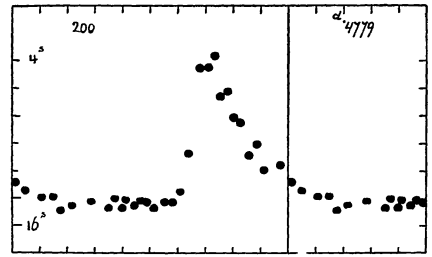
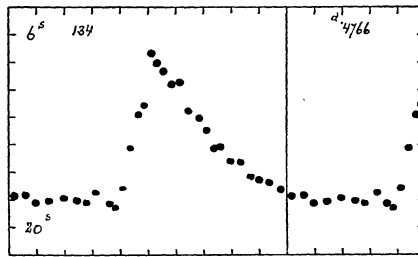
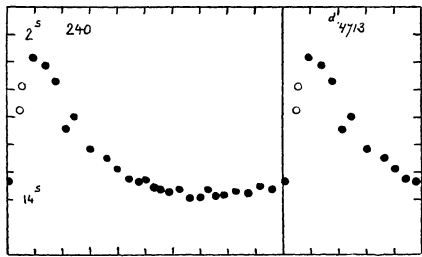
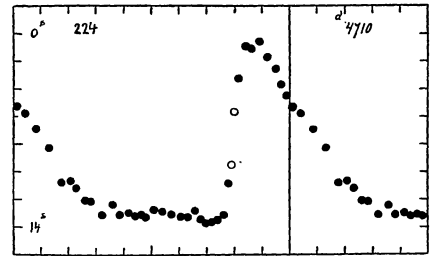
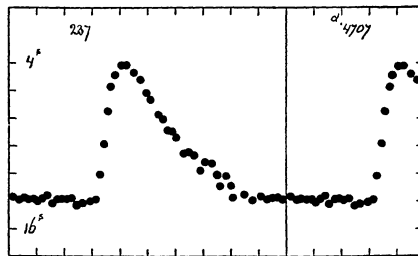
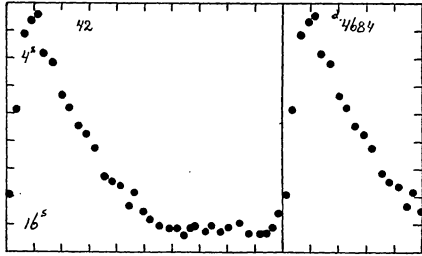
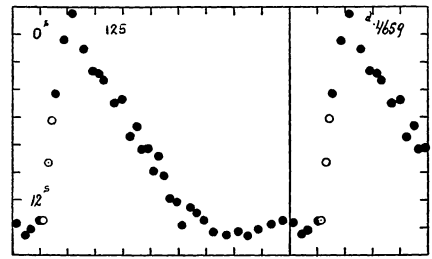
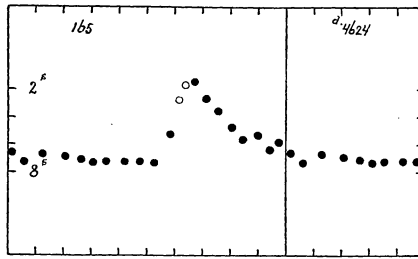
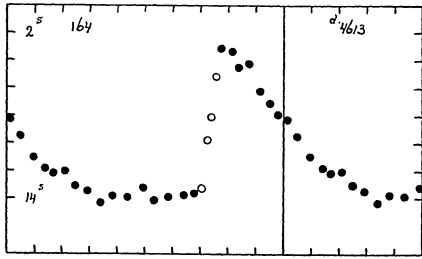








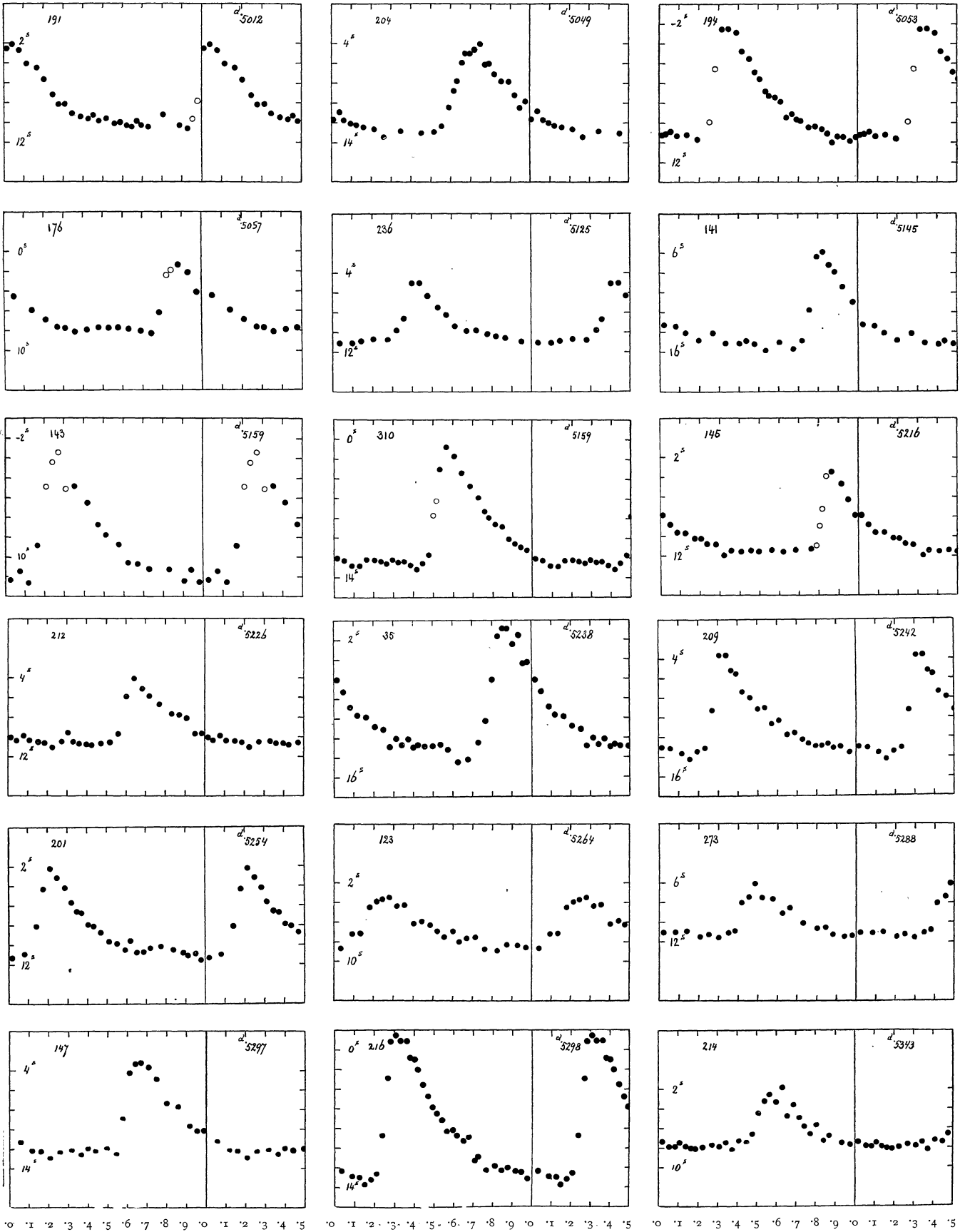


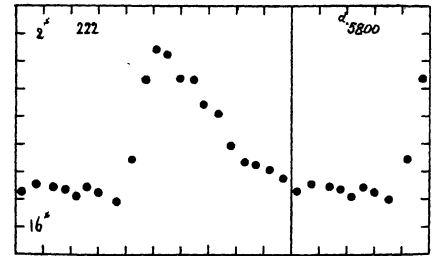
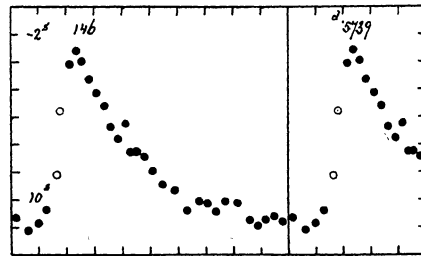
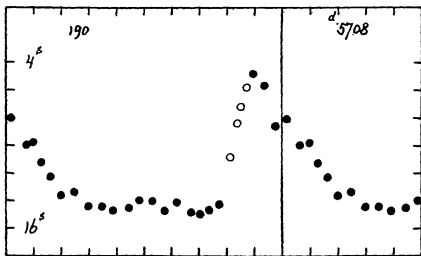
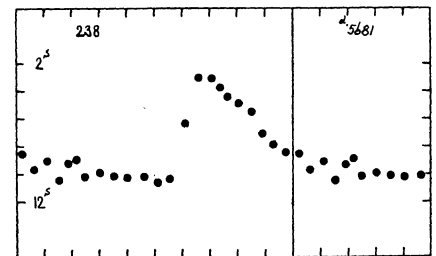
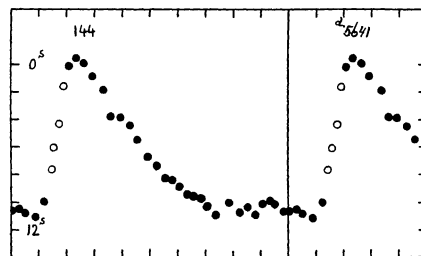
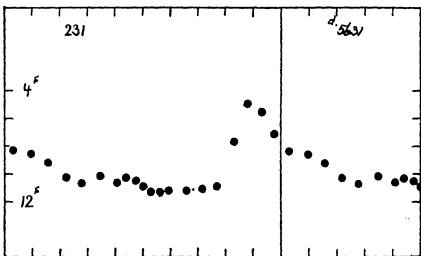
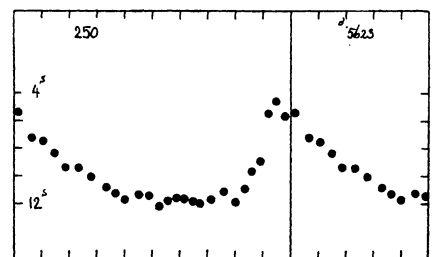
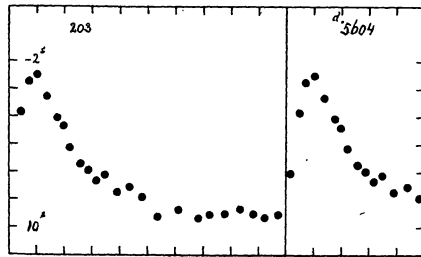
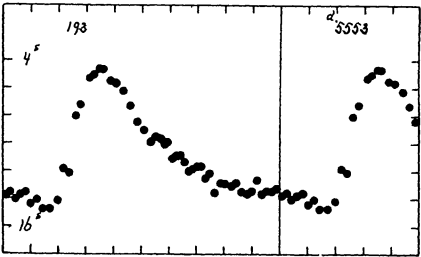
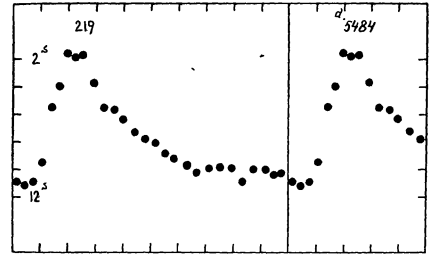
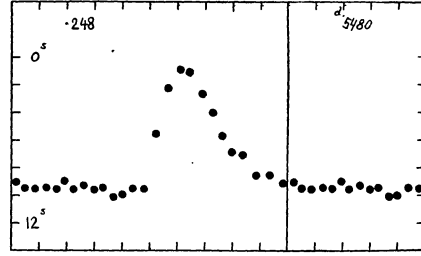
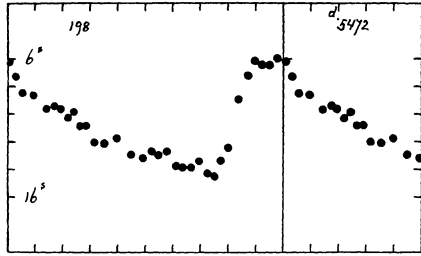
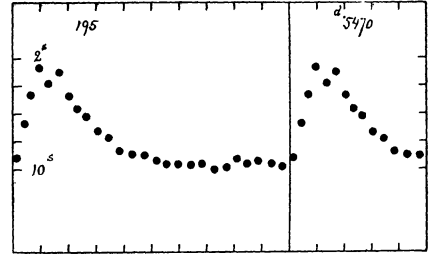
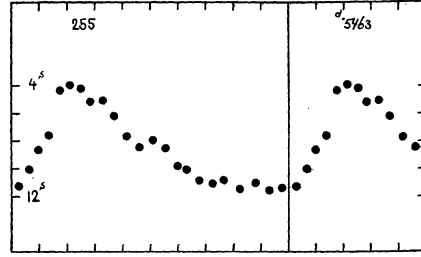
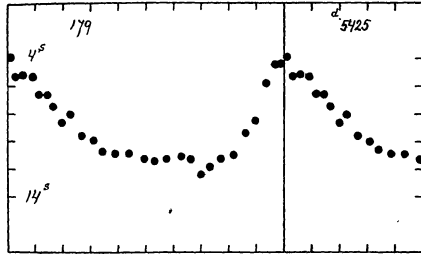
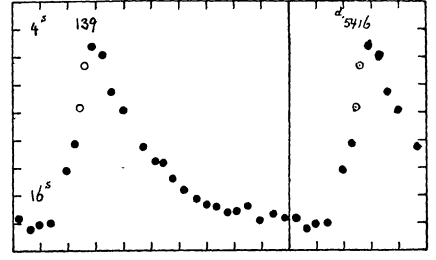
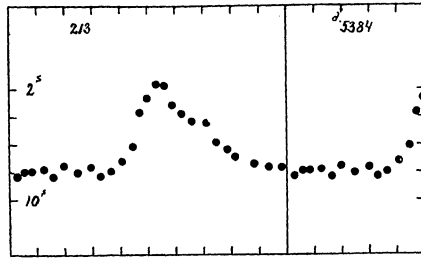
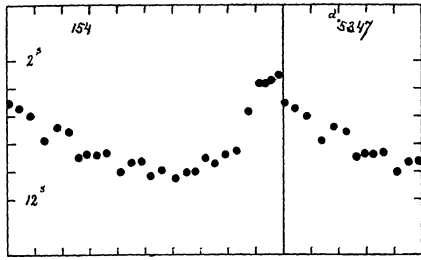


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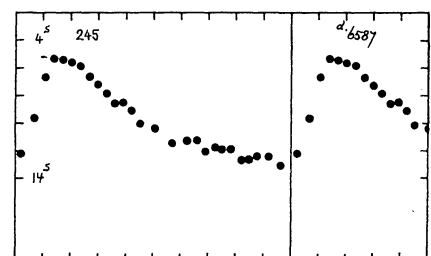
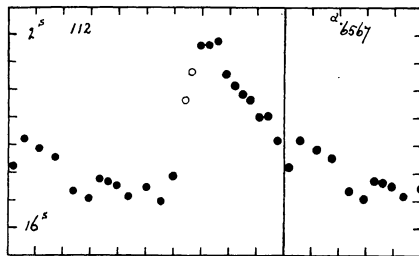
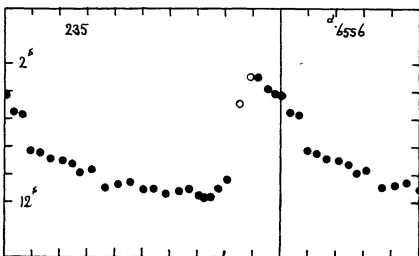
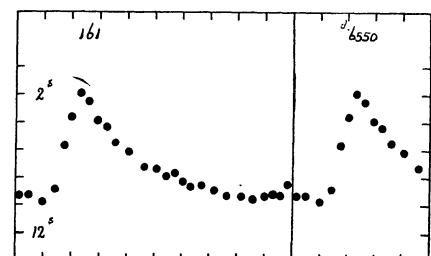
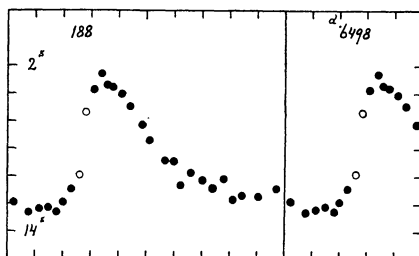
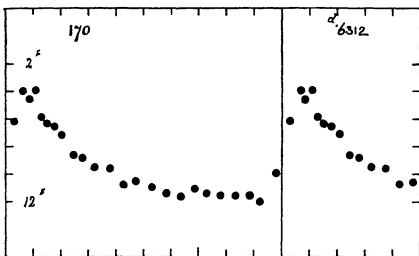
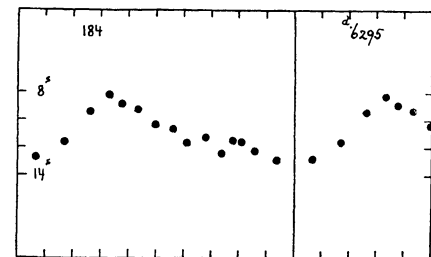
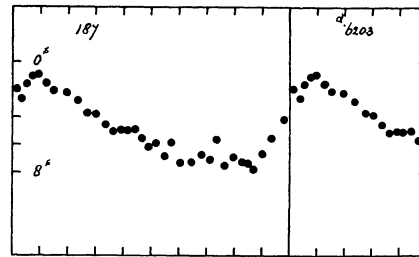
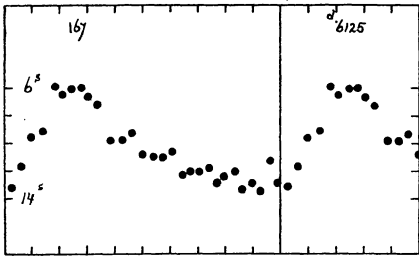
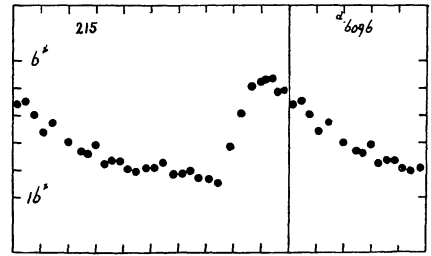
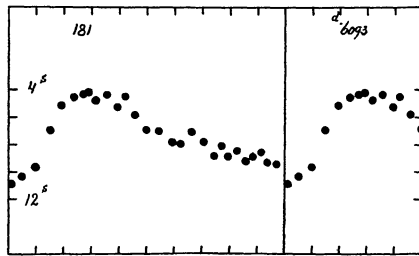
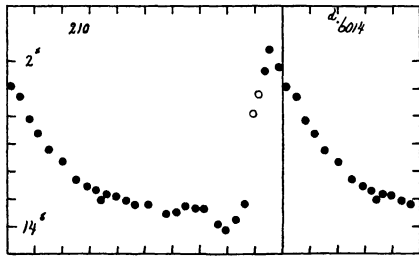
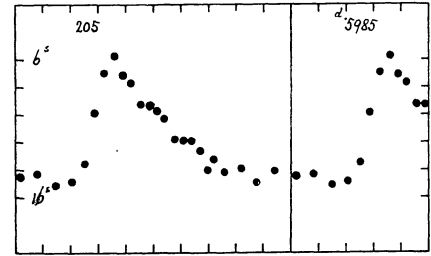
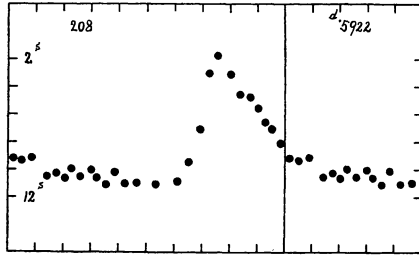
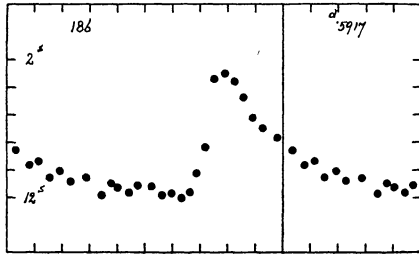
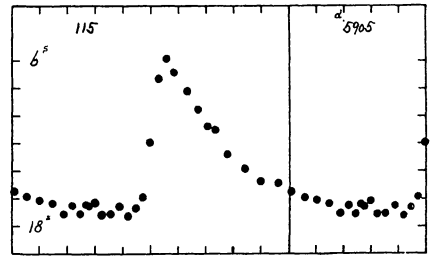
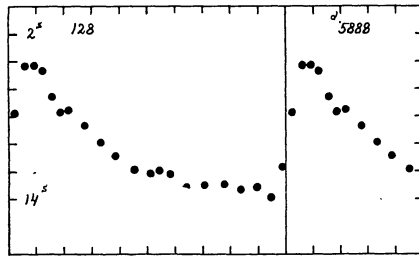
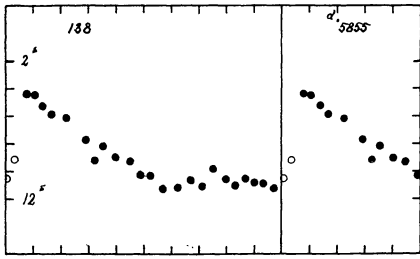
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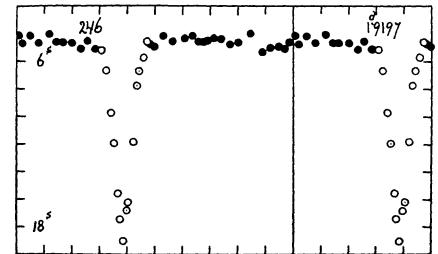
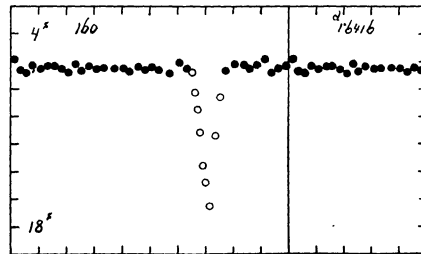
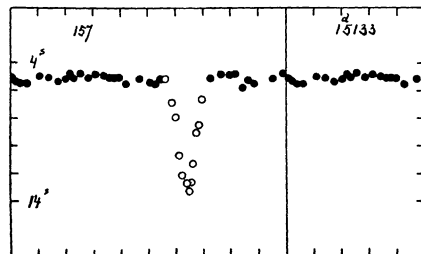
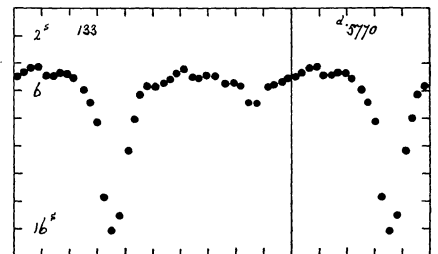
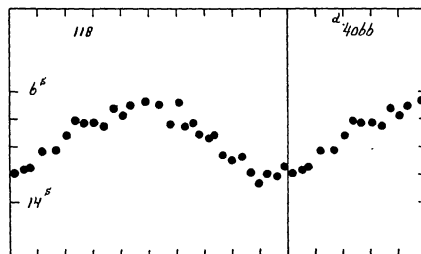
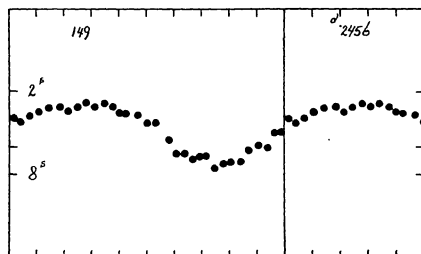
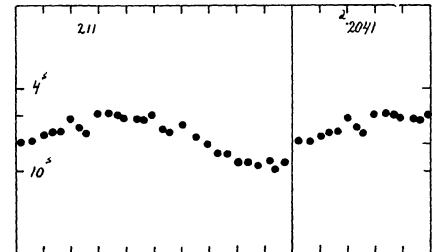
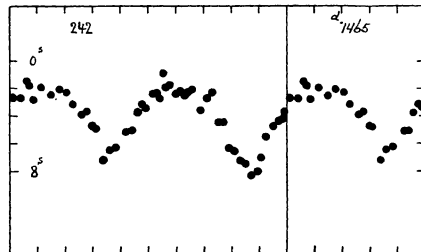
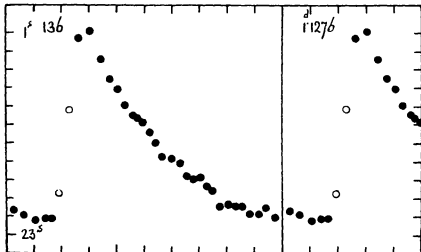
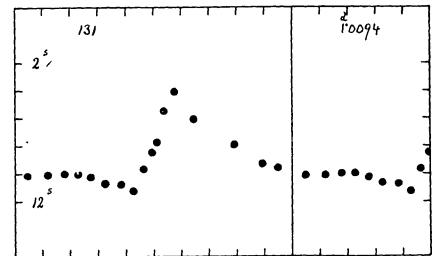
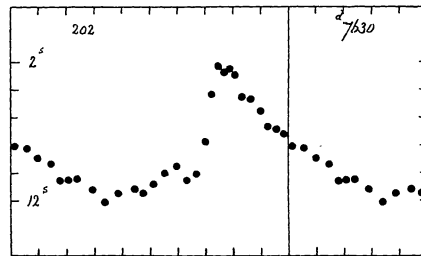
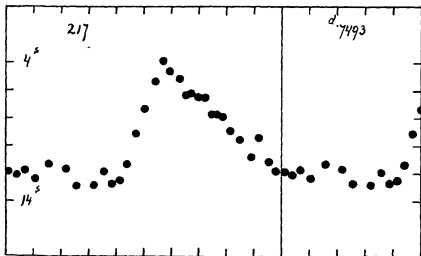
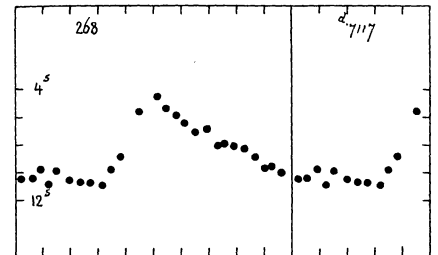
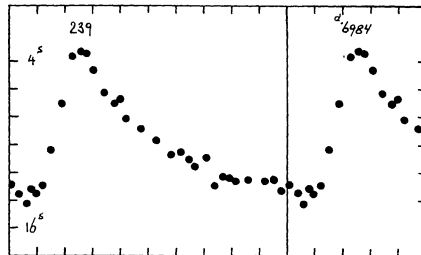
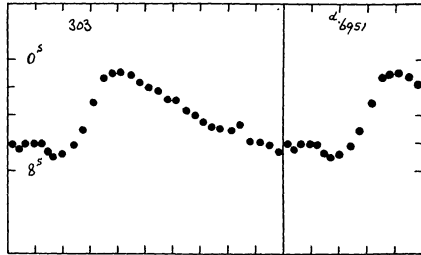
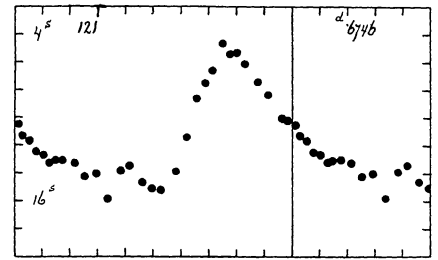
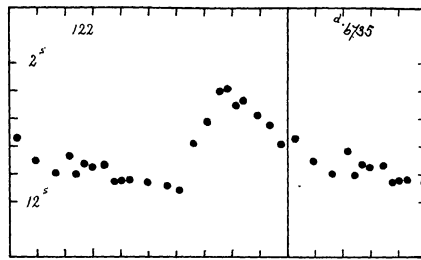
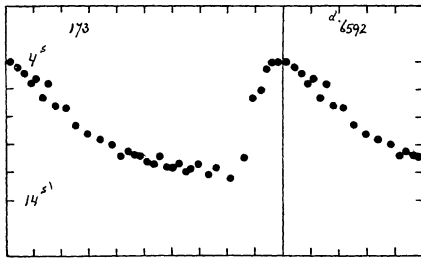




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