

TABLE I2.
U Normae.

number of plates	phase	brightness									
n	P	s	n	P	s	n	P	s	n	P	s
II	.015	.17	IO	.238	.72	IO	.588	.74	IO	.811	+.02
II	.032	.22	IO	.265	.78	IO	.612	.66	IO	.830	-.11
II	.048	.27	IO	.292	.76	IO	.628	.60	IO	.860	-.22
II	.072	.29	IO	.317	.70	IO	.656	.54	IO	.878	-.33
IO	.097	.37	IO	.334	.87	IO	.686	.46	IO	.900	-.31
IO	.131	.42	IO	.368	.94	IO	.704	.40	IO	.918	-.25
IO	.152	.48	IO	.410	.91	IO	.732	.35	IO	.945	-.13
IO	.171	.51	IO	.471	.92	IO	.761	.23	IO	.961	-.06
IO	.188	.55	IO	.516	.86	IO	.792	.14	IO	.982	+.07
IO	.212	.64	IO	.559	.82						

Observation of a maximum of EP Carinae, by Ejnar Hertzsprung.

The variable star EP Carinae, $10^h 23^m 49^s.4$, $-58^\circ 19' 6$ (1875) was formerly (B. A. N. 77, 209) only found visible during ten days from J. D. 2423879.35 to 3889.47 and on one old plate from J. D. 2421722.27. In addition to this I found EP Car visible on 3 M F Harvard plates and on 4 Johannesburg plates all from 1926 May. Using the same comparison star as in B. A. N. 77 the estimates are

plate	J. D. hel.	M. astr.	T. Grw.
M F 10162	2424639.548	+ 1.1	s
3300	42°348	— .32	
3301	.374	— .35	
3324	48°240	+ .12	
M F 10226	.552	— .26	
3328	49°244	+ .75	
M F 10233	.519	.00	

The last estimate on M F 10233 is indicated as particularly certain.

The two intervals between the 3 observed maxima are about 2160 and 760 days. No period is as yet apparent.

