

Epochs of minima. Middle diagram: epochs observed. Upper and lower diagrams: epochs reduced to the centre of gravity for negative and positive inclination respectively.

the inclination of the visual orbit is positive. To obtain the period reduced to the centre of gravity a least squares solution has been made for a straight line through the points of the lowest diagram (positive inclination) of Figure 4. The result is:

Period =
$${}^{d} \cdot 2678082 \pm {}^{d} \cdot 0000006$$
 (m.e.).

The O-C's of this solution (column 10) show that there is some deviation from linearity caused by the first two epochs. The mean error of an epoch is found to be \pm d·0058 for a unit of weight or three times the estimated internal mean error given in column 2 of Tables 7 and 8. A second solution of the same kind has been made with $\frac{M_A}{M_A+M_B}=$:50. The new O-C's (column 11) show less deviation from linearity. An attempt to find the mass ratio from the fact that, if the true mass ratio is used in the reduction of the epochs to the centre of gravity, there should be no deviation of the epochs from linearity, had no success. After some fifteen years, as already mentioned by Strand, it will be possible to obtain in this manner an accurate value of the mass ratio.

It should be mentioned that from meridian observations B. Boss has derived a mass ratio $\frac{M_{\rm A}}{M_{\rm A}+M_{\rm B}}=$ 209 (General Catalogue, Vol. 1, App. II, 1937)

For further observations an ephemeris of the primary minimum may be computed from the formula:

Primary minimum = J.D. 2428635.435 + d.26780800 E.

I wish to express my thanks to Prof. Hertzsprung for many suggestions during the preparation of this paper.

Photographic observations of six minima of 44 i Bootis B, by P. Th. Oosterhoff.

During the years 1928, 1929 and 1930 a number of photometric observations was made at the Leiden Observatory of the faint component of 44 i Bootis (Σ 1909), with the special purpose to derive accurate epochs of minimum. As the distance between the components at the time of the observations was about 3" only, a Barlow lens was used in combination with the 33 cm photographic refractor. The equivalent focal length obtained in this way amounted to 22.7 metres, the distance between the components on the plate being one third of a millimetre. A photometric scale was provided by the use of an

29.3 mm, and the distance on the plate between the first order spectra two thirds of a millimetre; d and l were practically equal.

Under favourable seeing conditions the photo-

objective grating constructed from card-board and

wooden bars. The grating constant, (d + l), was

graphic images thus obtained are well adapted for photometry, but with poor seeing the images of the two components and the central and first order images partly overlap. A change in the seeing during the observations may result in an asymmetric light curve which in its turn will impede the determination of

the epoch of minimum. Therefore two plates were rejected on account of rapidly changing seeing.

The following plates were measured by the writer and Mr. C. J. KOOREMAN in the Schilt microphotometer.

plate	date	hour angle of first and last exposure	exposure time	number of ex- posures
1477 1634 1682 1698 1900	31 May 1928 28 Febr. 1929 1 May 1929 27 May 1929 8 Febr. 1930 29 April 1930	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	90 90 90 90 90 60	37 40 45 46 50 54

Eastman 40 plates, 9 cm \times 12 cm, were used. The diameter of the projected image of the diaphragm on the plate in the photometer was '13 or '16 mm.

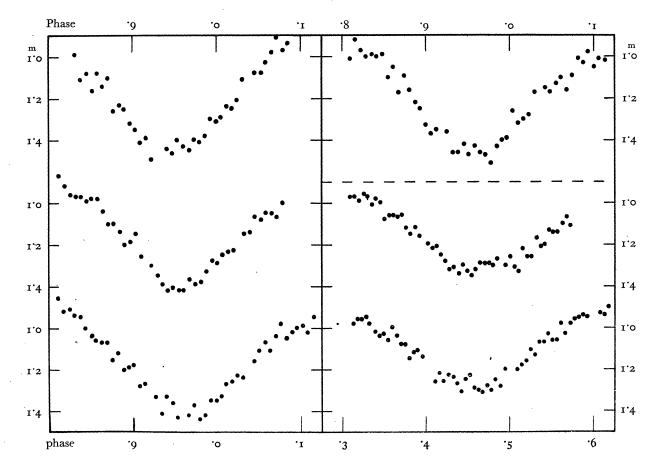
The Julian Day, the phase computed with the formula:

$$phase = {}^{d^{-1}}_{3.7340321} \ (J.D. - 2420000)$$

and the difference in magnitude between the two components are given for each exposure in the table. For the magnitudes the difference between the central and the first order images is taken as unit. The minima are graphically represented in chronological order in the accompanying figure. The first four are primary, the other two are secondary minima. The difference in depth is found to be 13 magnitude. The internal mean error of a single observation derived for all individual plates from the differences in magnitude between succeeding observations is \pm $^{\rm m}$ ·034.

The epochs of minimum computed with the method of B.A.N. No. 147, 179 are:

The first of these epochs has been used by G. P. Kuiper in B.A.N. No. 165, 33. The mean error given there was estimated from the inclination of the branches of the minimum and from the number of observations and their internal mean error. The true mean error will be larger on account of systematic errors. Although the observations do not allow an accurate determination of the systematic errors,



									·		
J.D.			J.D.			J.D.			J.D.		
Hel. M.A.T. Gr.	phase	Δm	Hel. M.A.T. Gr.	phase	Δm	Hel. M.A.T. Gr.	phase	Δm	Hel. M.A.T. Gr.	phase	Δm
2420000			2420000			-2420000			2420000		
				<u> </u>							
d	D	m ·	d	р	m	d	р	m	a .	p .	m
5398-4085	831	.99	5671 6222	.010	1.53	5759'4291	.893	1.52	6016.6879	.206	1.31
'4106	.838	1,11	6257	.032	1.12	'4308	.899	1.33	.6893 .6906	.211	1.33
4123	.845	1.08	·6274	.039	1'14	·4325	'906 '912	1.32	16920	'516	1.52
'4141	·852 ·858	1.08	·6291 ·6309	'045 '052	1.08	'4343 '4377	912	1.32:	6934	.526	1.50
.4158 .4175	.864	1.14	6326	.028	1.02	4395	932	1.46	.6948	532	1.12
4193	871	1,10	6343	.062	1.02	4412	.938	1.46	6962	537	1.51
4210	.877	1.56	.6360	.071	1.02	'4429	945	1'42	.6976	'542	1.50
4227	.884	1.53	.6378	.028	1,00	4446	.951	1.47	.6990	547	1,13
·42 <u>4</u> 4	.890	1.25	0	.0	.06	4464	958	1'43	.4003 .4003	552	1'14 1'14
·4262	.897	1.35	5733.4298	·811	.86	·4481 ·4498	'964 '971	1.46	7017	.557 .563	1,10
`4279 `4296	.903	1 35	'4316 '4333	.824	'92 '91	4516	971	1.21	7045	.568	1.02
4314	.916	1.39	4350	.830	'94	4533	.984	1.43	7059	573	1.11
4331	922	1.49	'4368	.837	.95	4550	.990	1.40			
·438o	'941	1.44	'4385	.843	1,00	4568	.996	1,30	6096:4427	.313	.98
·4397	'947	1.46	'4402	·850	1.04	4585	.003	1.26	4441	.318	.96
4414	953	1'40	4420	.856	1.06	.4602 .4620	.000	1.35	4455	·323 ·328	·96 ·95
4431	.960	1'43	4437	·863	1.02	4637	016	1,30	'4469 '4483	334	.98
*4449 *4466	967	1.42	.4454 .4471	.876	1.02	4654	'029	1.12	4496	339	1,05
·4483	973	1.41	4489	882	1.13	4689	'042	1.12	4510	344	1.04
4501	.986	1.38	4506	.889	1'20	'4706	.048	1.12	4524	349	1.03
4518	.992	1.30	4523	.895	1,10	'4724	.022	1.13	'4538	354	1.06
·4535	.999	1.31	4541	.001	1.18	4741	.001	1,10	4552	359	1,00
[•] 4553	.002	1.50	4558	.908	1.58	4758	:068	1.19	'4566	365	1.04
4570	'012	1'24	4575	'914	1.52	·4776	.04 .081	1,01	'4580 '4593	370	1.08
·45 ⁸ 7	018	1'25	'4610 '4627	927	1.33	.4793 .4810	087	1.03	4393	.380	1.12
·4604 ·4622	.031	1.11	4645	934	1.33	4827	.093	.98	4621	.385	1.15
4660	.045	1.08	4662	947	1.36	4845	.100	1.02	4635	.390	1,11
4677	'052	1.08	4679	953	1.43	4862	.109	1,01	'4649	.396	1'14
4695	.028	1.03	4714	'966	1'42	4879	,113	1.03	'4690	411	1.56
4712	.062	.98	'4731	'973	1.32	66-6			4704	416	1'22
4729	.071	.01	4748	979	1'44	6016.6352	.309	.97	'4718 '4732	421	1.52
4747	·078	.97	.4766 .4783	986	1'42	6380	314	.97 .99	4732	432	1.54
.4764	004	.94	4800	.999	1,32	6394	325	.96	4760	437	1.27
5671.5668	.812	.87	4818	.002	1.33	.6408	.330	.97	4773	'442	1.31
.5685	.819	.92	4835	.011	1.52	.6422	'335	1.01	4781	447	1.22
.5703	.826	.96	4852	.018	1.56	.6436	'340	.98	4801	452	1.53
.5720	.832	'97	'4870	'024	1,53	.6449	345	1,00	'4815	458	1.50
5737	.838	.97	'4887	.031	1.24	·6463	350	1.09	.4829 .4843	'463 '468	1,31
:5755	·845 ·851	.99 .98	4922	050	1,11	.6477 .6491	356	1.00	4857	473	1.58
·5772 ·5789	.858	.98	.4939 .4956	.057	1.02	.6505	366	1.02	4870	478	1,30
·5807	.864	1.04	4974	.063	1,11	6519	'37I	1.06	·4884 ·4898	'483	1.52
·5807 ·5824	.871	1.10	4991	.070	1.04	6532	376	1,15	· ₄ 898	489	1.58
•5841	.877	1,10	.5008	.076	.98	.6546	.381	1.12	4912	494	1'20
.2829	.884	1'14	.5026	.083	1.02	6560	:387	1'12	4953	509	1,18
.5876	.890	1.30	5043	.089	1.03	·6574 ·6602	'392 '402	1,10	.4967 .4981	514	1.19
.2893	.897	1,12:	·5060 ·5077	102	1,00	.6616	408	1.55	4995	525	1,11
•5910 •5928	.910	1.56	5095	.108	1.05	6629	413	1,51	.5009	.230	1.13
.5962	'922	1.30	5112	1115	.95	.6643	418	1.22	.5023	535	1.02
.5980	929	1.35				.6657	'423	1.58	5037	'541	1.02
.5997 .6014	935	1.39	5759.4066	.809	1,01	.6671	428	1.35	.2020	.546	1.03
.6014	'942	1.42	4083	815	'92	.6685	433	1,31	5064	551	1.00
6032	948	1'41	4100	822	.97	.6699	439	1.34	.5078	556	1.06
·6049 ·6066	955	1'42:		828	1,00	·6713 ·6726	444	1,33	.5092 .5106	566	1.03
6084	967	1.42	'4135 '4152	841	1.00	6740	449	1.32	'5120	572	1.98
.6101	907	1,30	4170	848	.99	6754	459	1.33	5134	577	.96
6118	.981	1,38	4187	.854	1,10	•6768	464	1.50	·5147	'582	'95
.6135	.987	1.33	4204	.861	1.02	6782	'470	1.50	.2161	.587	'94
.6153	994	1.58	4221	.867	1.12	·6796	475	1.50	5175	592	'95
.6170	.000	1.50	4239	874		.6810	:480	1,30	5217	.608	.93
.6187	.006	1	4256	.880	1,16	·6823 ·6851	485	1.30	·5231 ·5244	.618	94
6205	.013	1,54	4273	.887	1.55	6865	'495 '501	1,30	3444	1 313	95
	1	1	1	1	i	. 0005	, 501	, _ ~ ~ ~	•	1	1