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COMMUNICATION FROM THE OBSERVATORY AT LEIDEN.

Provisional ephemerides of 25 new faint variable stars within 6° of τ Sagittarii, by *J. G. Ferwerda*.

The 25 variables forming the subject of this paper have been found and estimated on a series of about 300 plates taken by Dr. H. VAN GENT with the Franklin-Adams instrument at Johannesburg. The centre of the plates is $15^h 58^m - 27^\circ 5'$ (1875). The time of exposure of each plate is 30^m . The extreme dates of the plates are J. D. 2425442.3847 and J. D. 2426573.5093. The size of the plates is 20×20 cm, each plate covering a field of $10^\circ \times 10^\circ$.

The variables have been discovered with the blink microscope. One pair of plates was compared by Dr. H. VAN GENT in the blink microscope at Johannesburg. The J. D. of these plates are 2425792.4344 and 2425326.4790. This comparison produced 30 variables of which as yet 6 have turned out to be short period variables, viz: f , n , r , t , u , x of this paper. The other variables have been discovered by myself on a second plate pair (J. D. 2425862.2866 and 2425863.3492) with the blink microscope at Leiden. The only rediscovery was the variable x .

Of the 25 variables 21 are of the cluster type and 4 are eclipsing variables. The southern part of the plates joins with a field of $10^\circ \times 10^\circ$ in which also a great number of cluster variables have been found, as published in *B. A. N.* No. 227.

The estimates were made on an arbitrary scale of steps, the variable being ordinarily inclosed between two comparison stars. Thus each plate yielded a value for the difference in steps between those two comparison stars, for which difference finally the mean of all plates was adopted. This proceeding generally in use at Leiden is to be preferred to adopting a priori certain brightnesses of the comparison stars. Table 1 shows in column 2 the brightness of the comparison stars in the scale of steps with arbitrary zeropoint (care has been taken that all values of brightness of the variable become positive). Column 3 gives the brightness of the comparison stars in magnitudes found by direct comparison with some stars of the selected area No. 159 which is in the field of the plates. The

situations of the variables and the comparison stars are shown on the accompanying diagrams. The variable is indicated by an open circle. The size of these diagrams is $10' \times 10'$.

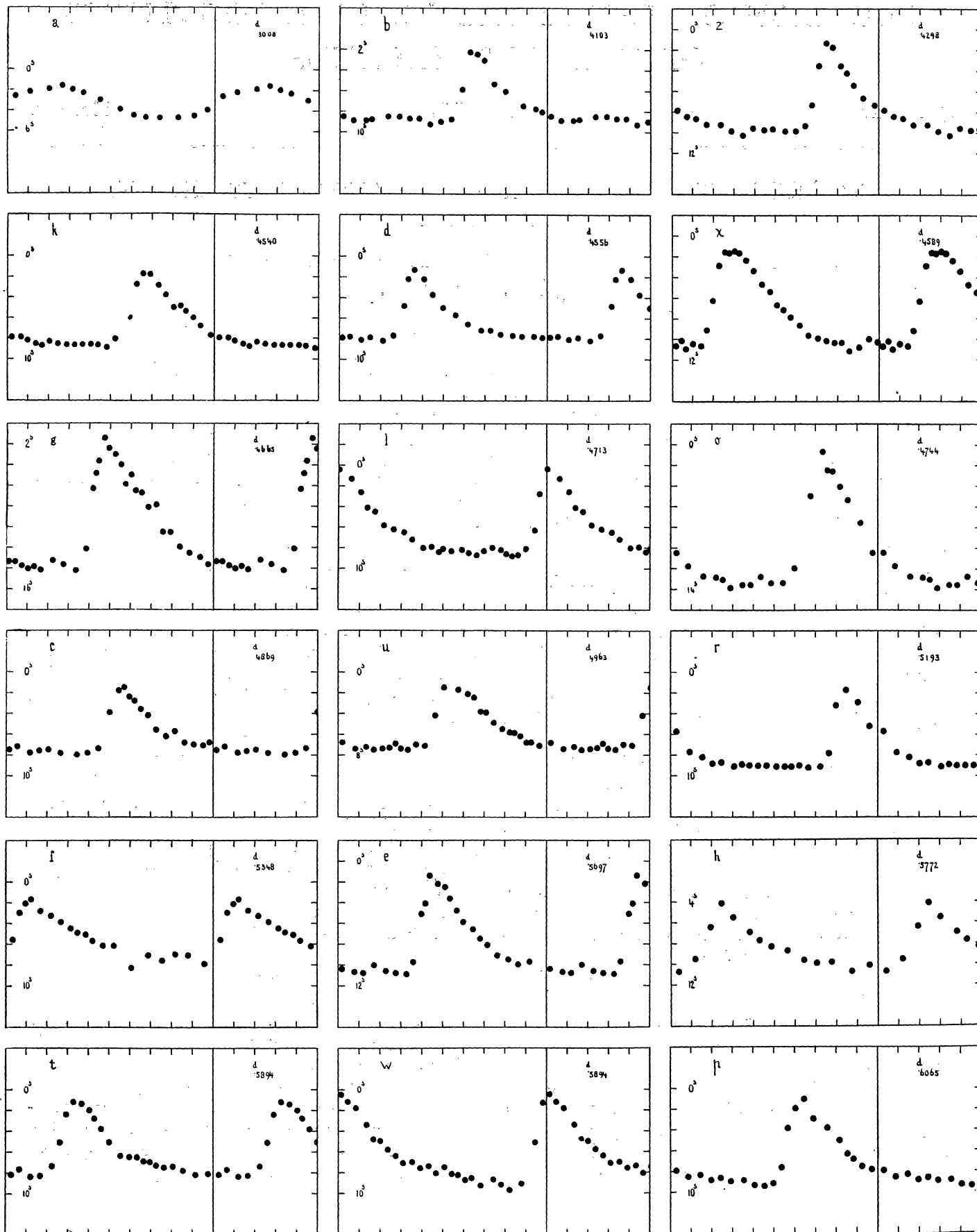
The determination of the period did in no case raise any serious difficulty owing to the circumstance that among the plates there are three sets of 15 or 16 plates taken on one night. The epochs used for the computation of the provisional period are the maxima of the cluster variables and the minima of the eclipsing variables. For variable a both maxima and minima have been used. For several variables this period has been corrected graphically as explained below. For those stars of which the period has been derived by least squares Table 2 contains in column 1 the J. D. Hel. Grw. astr. mean times of the epochs used, in column 2 the number of periods elapsed after the earliest epoch and in column 3 the residuals $O-C$.

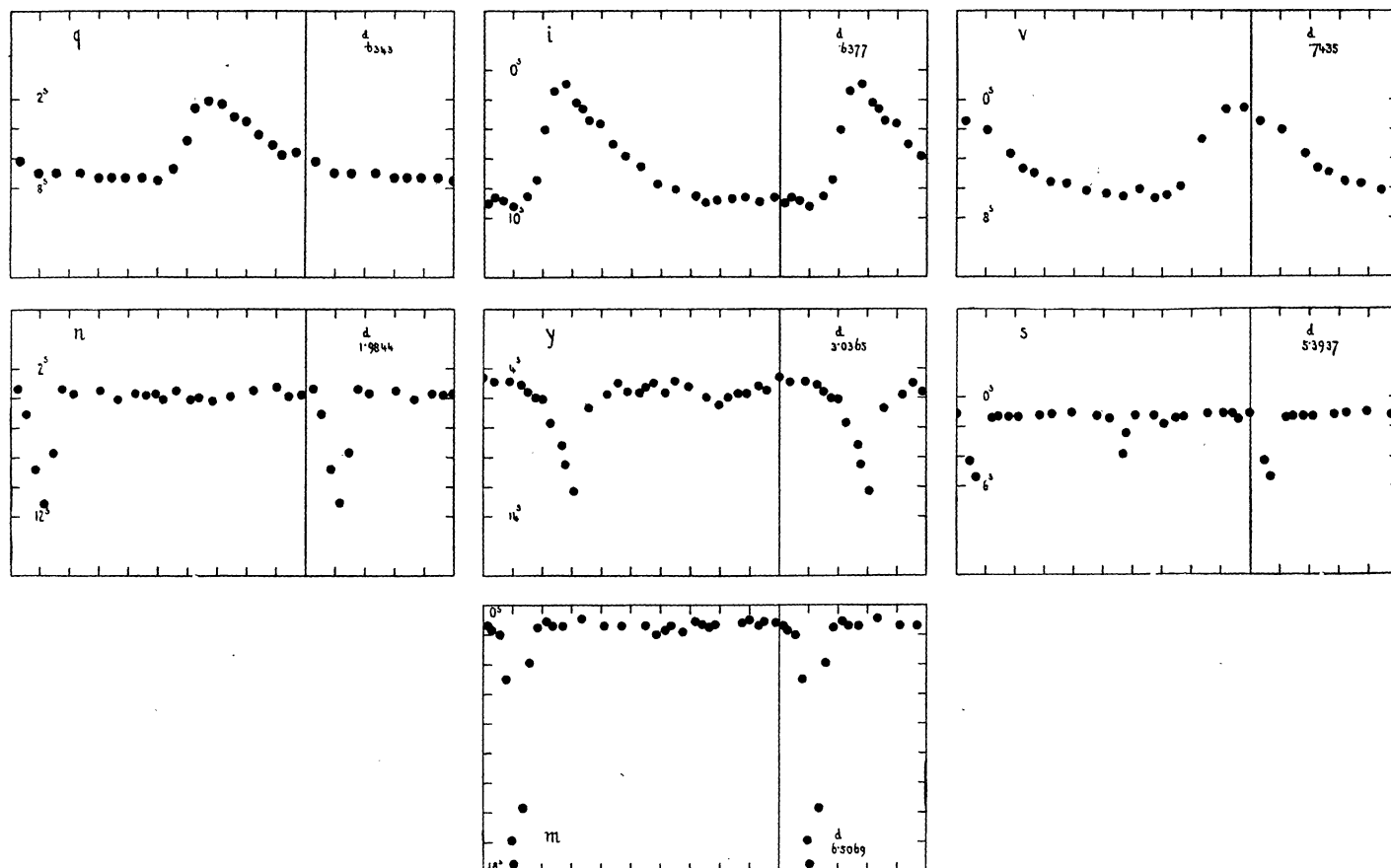
The phases have been computed by the formula

$$\text{phase} = \text{reciprocal period} \times (\text{J. D. H. Gr. A. M. T.} - 2420000)$$

In some cases the period finally adopted was derived with the aid of the observations made on the steepest part of the lightcurve. The slope of this part of the lightcurve determined for each observation the reduction of its epoch to that of a point on that steep part of the light curve where the variable has a fixed brightness. The normal epochs (often few in number but relatively accurate) thus obtained were used for a graphical determination of the period.

All estimates have been classed according to phase and divided into groups. Table 3 shows in column 1 the numbers of estimates used in each group, in column 2 the mean value of phase, in column 3 the mean value of brightness. The accompanying diagrams give the corresponding light curves. For several variables the uncertainty of the estimates has been taken into account in computing the mean value of brightness. These uncertainties have been specially marked in the





registrations of the estimates. For some stars all uncertain estimates have been omitted, for others half weight has been assigned to uncertain estimates.

The total number of estimates used in the present paper is 6426.

For the computation of the co-ordinates of the variable three C. P. D. stars have been chosen defining a triangle within which the variable is situated. On an enlargement of the neighbourhood of the variable two straight lines have been drawn joining the four stars two by two. The position of the variable has been computed by assuming a linear connection between line segments and co-ordinate differences.

Finally, Table 4 gives a summary of the various results. Column 4 gives the total number of estimates used for computing the points of the mean lightcurve.

Column 5 gives the period of variation together with the mean error of this period computed by means of least squares.

Column 7 gives the phase of a point on the rising branch of the mean light curve defined by the condition that the difference in phase with the point of equal brightness on the descending branch becomes a fixed value ΔP (to be found in column 8).

Column 8 gives the fraction ΔP as explained in the remarks to Column 7

Column 10 gives the epoch nearest to the mean date of the plates which is J. D. 2426119. The epoch of the variable f only is an exception, as at the time when the variable was estimated the set of plates was not yet complete. The mean date in this case is J. D. 2425991.

Column 12 gives the mean error of a single estimate, computed by means of the formula

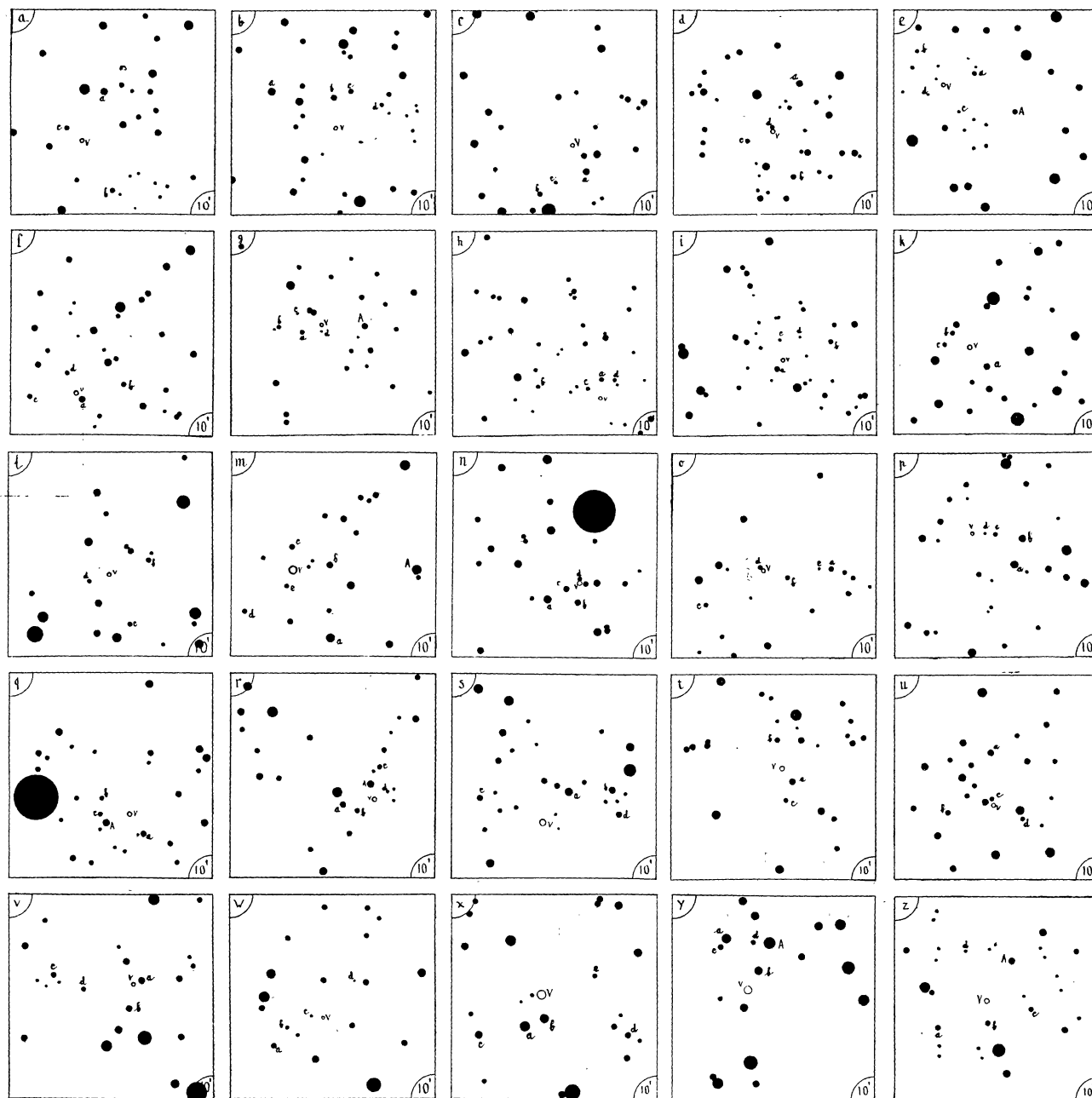
$$\text{m. e.} = \pm \sqrt{\frac{\sum (\Delta s)^2}{2n}}$$

in which Δs means the difference in brightness between two estimates following each other in phase and n the number of estimates used.

Column 13 shows the value of one step in magnitudes which gives together with the dates of column 12 some idea of the accuracy of a single estimate.

Remarks on a few individual variables.

Variable α . A period has been derived from maxima as well as from minima.



period computed from maxima: $^d\cdot3007609 \pm ^d\cdot0000058$
 period computed from minima: $^d\cdot3007738 \pm ^d\cdot0000044$

Or taking the mean value: period = $^d\cdot3007673 \pm ^d\cdot0000051$.

The lightcurve is approximately a sine curve, the dots of the lightcurve being represented with a mean error of $\pm ^s\cdot18$ by the formula

$$s = \begin{matrix} s \\ 3\cdot36 \\ \pm \cdot05 \end{matrix} - \begin{matrix} s \\ 1\cdot59 \\ \pm \cdot07 \end{matrix} \sin 2\pi P - \begin{matrix} s \\ \cdot09 \\ \pm \cdot07 \end{matrix} \cos 2\pi P,$$

in which s is the brightness in steps and P the phase (J. D. - 2420000) $\times 3^d\cdot32483$.

From this formula the phase of maximum is $\cdot250 \pm \cdot018$, which is given in Table 4.

Variable s . The depth of the secondary minimum is about half of that of the primary.

Variable y . The lightcurve of this eclipsing variable shows an indication of a secondary minimum.

I wish to thank Prof. E. HERTZSPRUNG for his help and advice during the preparation of this paper.

TABLE 1.

s	m	s	m	s	m	s	m	s	m	s	m	s	m
<i>a</i>		<i>b</i>	6.5 13.8	<i>c</i>	7.1 14.8	<i>o</i>		<i>s</i>		<i>b</i>	5.4 13.7		
<i>a</i> — 2.0 13.0		<i>c</i>	9.1 14.4	<i>d</i>	9.6 15.6	<i>a</i>	1.0 13.4	<i>a</i> — 2.2 12.4		<i>c</i>	8.1 14.2		
<i>b</i> 1.7 13.5		<i>d</i>	11.0 14.6			<i>b</i> 7.0 13.8		<i>b</i> 2.0 12.8		<i>d</i>	9.9 14.5		
<i>c</i> 5.0 13.8				<i>k</i>		<i>c</i> 11.7 14.0		<i>c</i> 3.2 13.0					
<i>b</i>		<i>a</i>	0.0 13.2	<i>a</i> — 1.0 12.9		<i>d</i> 13.6 14.2		<i>d</i> 4.2 13.1		<i>x</i>			
<i>a</i> — 1.0 12.9		<i>b</i>	3.5 13.6	<i>b</i> 4.8 13.5		<i>e</i> 15.1 14.5				<i>a</i>	0.0 12.0		
<i>b</i> 2.9 13.4		<i>c</i>	6.6 13.8	<i>c</i> 9.2 14.3				<i>t</i>		<i>b</i>	4.2 12.4		
<i>c</i> 5.5 13.8		<i>d</i>	9.2 14.4					<i>a</i>	0.0 13.4	<i>c</i>	7.1 12.7		
<i>d</i> 8.8 14.8				<i>l</i>				<i>b</i>	3.2 13.8	<i>d</i>	9.5 13.0		
		<i>g</i>		<i>a</i> 2.0 13.4				<i>c</i>	8.0 14.2	<i>e</i>	13.2 13.8		
		<i>A</i>	0.0 13.0	<i>b</i> 4.3 13.7									
		<i>a</i>	4.3 13.6	<i>c</i> 7.5 14.2				<i>u</i>					
		<i>b</i>	8.8 13.9	<i>d</i> 9.6 14.6				<i>a</i>	2.0 13.2	<i>y</i>			
		<i>c</i>	12.2 14.3					<i>b</i>	4.0 13.8	<i>A</i>	0.0 11.5		
		<i>d</i>	15.5 14.8	<i>m</i>				<i>c</i>	6.0 14.5	<i>a</i>	3.5 12.4		
				<i>A</i>	1.0 11.9			<i>d</i>	7.8 15.1	<i>b</i>	6.4 12.7		
		<i>h</i>		<i>a</i>	4.5 12.4					<i>c</i>	10.0 13.4		
		<i>a</i>	1.0 13.2	<i>b</i> 8.1 12.9				<i>v</i>		<i>d</i>	13.1 13.9		
		<i>b</i>	4.6 14.0	<i>c</i> 11.1 13.1				<i>a</i>	1.0 13.7				
		<i>c</i>	9.0 14.5	<i>d</i> 14.6 14.0				<i>b</i>	2.4 14.0				
		<i>d</i>	11.3 14.7	<i>e</i> 17.6 14.6				<i>c</i>	5.0 14.3	<i>A</i> — 1.0 13.1			
								<i>d</i>	7.6 15.1	<i>a</i>	3.0 13.6		
		<i>i</i>		<i>n</i>						<i>b</i>	6.5 14.0		
		<i>a</i>	1.0 13.8	<i>a</i> 1.0 13.1				<i>w</i>		<i>c</i>	9.0 14.4		
		<i>b</i>	5.1 14.3	<i>b</i> 5.4 13.6				<i>a</i>	1.0 13.2	<i>d</i>	11.4 14.8		
				<i>c</i> 9.8 14.0									
				<i>d</i> 12.5 14.5									

TABLE 2.

d	t	d	d	t	d	d	t	d	d	t	d
<i>a</i>			<i>b</i>			<i>c</i>			<i>d</i>		
Max.											
2425478.263	0	+	.003	2425707.572	0	—	.008	2425706.566	0	—	.002
5792.230	1044	—	24	5716.615	22	+	7	5764.494	119	—	13
5799.497	1068	+	25	5765.430	141	—	7	5821.456	236	—	16
5823.230	1147	—	2	5825.362	287	+	16	5825.362	244	—	6
.488	1148	—	45	5860.214	372	—	10	5463.349	322	+	4
5849.402	1234	+	3	5862.287	377	+	11	5886.242	369	+	14
5852.429	1244	+	23	6094.520	943	—	4	6076.600	760	0	
5860.225	1270	—	1	6210.226	1225	—	12	6153.542	918	+	14
6094.531	2049	+	12	6212.287	1230	—	2	6212.442	1039	+	1
6103.579	2079	+	38	6476.554	1874	+	10	6213.409	1041	—	6
6118.578	2129	—	2	6569.277	2100	—	2	6236.312	1088	+	14
6212.353	2441	—	64	6570.511	2103	+	1	6237.272	1030	0	
6213.331	2444	+	12	6573.381	2110	—	1	.294	1090	+	22
6214.237	2447	+	16					6561.526	1756	—	11
6241.302	2537	+	12					6569.321	1772	—	7
6565.483	3615	—	27					6570.296	1774	—	5
6569.407	3628	—	13					6571.259	1776	—	16
6570.339	3631	+	17					6572.255	1778	+	6
6573.348	3641	+	18					6573.230	1780	+	7
Min.											
2425525.283	0	—	.009								
5707.561	606	0									
5791.474	885	—	3								
5793.262	891	—	19								
5435.381	1031	—	9								
5861.240	1117	—	16								
5864.290	1127	+	26								
5486.230	1200	+	10								
6160.535	2112	+	9								
6212.254	2284	—	5								
6240.235	2377	+	4								
6265.247	2460	+	52								
6569.280	3471	+	2								
.550	3472	—	29								
6570.479	3475	—	2								
6573.209	3484	+	21								
.456	3485	—	33								

TABLE 3 (continued).

<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness	<i>n</i>	phase	brightness			
10	147	8.5	10	108	2.6		<i>n</i>		14	355	8.8	14	722	9.1	10	854	7.8	10	045	1.2	<i>y</i>		
10	179	7.4	10	138	4.1				14	407	9.3	8	764	7.8	10	913	8.2	10	082	1.8			
10	205	4.0	10	174	4.5	11	025	3.4	14	455	9.4	8	799	3.2	10	973	8.1	10	133	3.4			
9	238	1.4	10	217	5.8	8	053	5.1	14	499	9.1	15	848	1.7				10	168	4.8	11	033	4.9
9	279	0.9	10	264	6.2	7	086	8.8	10	538	7.6	15	904	2.9				10	200	5.0	9	086	4.9
10	311	2.2	10	313	6.5	9	116	11.1	6	568	3.8	15	960	5.2		<i>u</i>		10	235	5.8	12	125	5.1
10	333	2.6	10	351	7.2	10	147	7.7	10	604	1.9							10	272	6.4	13	148	5.6
10	356	3.4	10	404	8.0	10	177	3.4	10	646	1.0				10	021	6.8	10	309	7.1	13	173	6.0
10	393	3.6	10	446	7.9	15	214	3.7	11	690	2.9		<i>s</i>		10	083	7.4	10	351	7.0	13	199	6.1
10	434	5.0	10	480	8.4	15	305	3.5	14	758	3.8				10	135	7.2	10	390	7.6	13	223	7.7
10	479	5.8	10	500	8.1	15	363	4.1	13	816	5.0	7	044	4.3	10	171	7.5	10	431	7.4	8	264	9.2
8	530	6.5	10	541	8.3	15	422	3.7	14	853	6.3	9	065	5.4	10	212	7.4	10	467	8.1	8	274	10.5
10	587	7.7	10	591	8.2	15	460	3.8	14	884	6.8	15	119	1.4	10	247	7.3	10	507	7.5	9	303	12.3
10	646	8.1	10	625	8.5	15	492	3.7	15	926	7.5	15	140	1.3	10	275	6.9	10	541	8.2	7	353	6.7
10	716	8.5	10	662	8.7	15	519	4.1	15	972	7.8	15	176	1.3	10	300	7.4	10	571	8.3	10	417	5.8
10	748	9.0	10	699	8.3	15	562	3.5				15	209	1.3	10	335	7.5	10	606	8.8	10	452	5.0
10	789	8.8	10	739	8.0	15	610	4.1				15	280	1.2	10	374	7.0	10	638	8.6	11	485	5.6
10	838	8.7	10	780	8.2	15	638	4.0		<i>q</i> 3)		15	321	1.1	10	416	7.1	10	681	9.3	11	525	5.7
10	882	8.6	10	804	8.6	15	684	4.2				14	390	1.0	10	467	4.2	10	742	8.7	11	548	5.3
10	930	8.9	10	832	8.8	15	744	3.9	15	030	6.2	11	475	1.2	7	509	1.5	10	780	9.2	11	572	5.0
9	980	8.6	10	861	8.7	15	822	3.5	15	094	7.0	6	518	1.4	8	578	1.7	9	820	9.7	12	611	5.7
			10	900	8.1	15	901	3.3	15	152	7.0	4	563	3.8	10	621	2.1	6	878	9.1	12	645	4.9
			9	941	6.3	11	942	3.9	15	234	7.0	6	574	2.4	10	651	2.5	9	942	5.1	11	690	5.3
	<i>k</i>		9	966	2.8	11	987	3.8	17	298	7.3	15	605	1.2	10	684	3.8	10	980	1.3	8	751	6.0
									17	340	7.3	15	670	1.2	10	710	3.9				10	795	6.5
							<i>o</i> 2)		15	388	7.3	15	703	1.8	10	748	4.9				10	823	6.0
10	021	7.9							15	444	7.3	15	743	1.4	10	788	5.5		<i>x</i>		10	859	5.7
10	066	7.9	<i>m</i>						15	497	7.5	14	770	1.3	10	823	5.8				10	887	5.7
10	098	8.2							14	550	6.7	15	852	1.1	10	848	5.9				11	927	5.2
10	137	8.5	11	012	1.4	10	028	10.5	9	596	4.8	15	907	1.1	10	875	6.2	10	019	10.7	11	955	5.5
10	167	8.7	11	026	1.7	10	083	11.8	9	598	4.8	15	938	1.1	10	904	6.8	10	045	10.2	11	999	4
10	201	8.3	11	054	2.0	10	158	12.8	9	622	2.6	15	959	1.5	10	928	6.8	10	066	11.0			
10	242	8.5	4	078	5.0	10	219	12.9	14	669	2.1	15	997	1.1	10	968	7.1	10	100	10.5			
10	285	8.6	5	095	15.9	10	251	13.1	16	699	2.1							10	140	10.7	<i>z</i> 3)		
10	324	8.6	4	101	17.5	10	289	13.9	15	713	2.3							9	168	9.2			
10	362	8.6	4	132	13.7	10	345	13.6	15	754	3.2		<i>t</i> 2)					9	197	6.3			
10	404	8.6	8	155	3.9	10	385	13.6	15	797	3.5					<i>v</i>		9	229	2.9	13	026	7.9
10	439	8.7	12	181	1.5	10	433	12.8	14	839	4.4	10	025	8.2	15	031	1.5	9	256	1.6	12	071	8.5
10	481	8.9	12	211	1.1	10	484	13.4	15	885	5.1	10	065	7.7	15	105	2.1	9	276	1.7	12	117	8.7
10	521	8.1	12	232	1.4	9	541	13.4	15	918	4.8	10	120	8.4	15	184	3.7	10	302	1.5	12	168	9.3
8	591	6.0	13	269	1.4	6	598	12.0	15	965	5.6	10	166	8.3	15	226	4.7	10	324	1.7	14	235	9.3
8	626	2.8	17	331	0.9	8	733	0.7				8	222	7.4	15	265	5.0	10	358	2.4	14	288	9.9
8	658	1.8	14	409	1.4	8	758	2.5		<i>r</i> 3)		8	260	5.1	15	320	5.6	10	393	3.4	12	342	10.3
10	690	1.8	14	468	1.4	8	781	2.6	14	029	5.7	8	290	2.4	15	374	5.7	10	433	4.7	13	391	9.6
10	731	2.9	10	549	1.4	10	817	4.1	12	092	7.7	8	326	1.2	15	442	6.2	9	473	5.4	14	445	9.8
10	767	3.8	10	584	2.0	10	854	5.4	14	153	8.2	10	366	1.4	15	510	6.4	10	507	6.7	12	487	9.7
10	805	5.0	10	613	1.7	10	913	7.6	14	202	8.8	10	402	2.0	15	568	6.6	10	538	7.2	12	546	9.9
10	838	4.9	10	633	1.4	10	973	10.5	13	246	8.7	10	427	2.8	15	622	6.1	10	571	7.9	10	599	9.9
10	862	5.4	10	673	1.8				14	308	9.1	10	460	3.8	15	675	6.7	10	616	8.7	10	643	9.4
10	899	6.0	10	714	1.1				14	345	8.9	10	499	5.1	15	714	6.5	10	659	9.7	10	676	7.4
10	932	6.8	10	739	1.3		<i>p</i> 3)		14	381	9.0	10	552	6.4	15	762	5.9	10	703	10.0	10	710	3.6
10	981	7.7	10	761	1.5				14	422	9.0	10	596	6.5	15	832	2.7	10	744	10.2	12	747	1.4
			11	782	1.3	13	030	7.9	15	464	9.0	10	632	6.5	15	916	0.7	10	783	10.4	12	777	1.8
			11	871	1.2	15	088	8.5	15	511	9.1	10	664	6.9	15	978	0.6	10	818	10.4	13	815	3.6
	<i>l</i>		10	897	1.0	15	146	8.3	14	552	9.1	10	694	7.0				10	854	11.2	12	844	4.3
			10	929	1.4	15	198	8.8	14	585	9.1	10	725	7.3		<i>w</i>		10	903	10.8	13	878	5.5
9	003	0.4	10	947	1.1	15	243	8.6	14	622	9.0	10	761	7.5				10	951	10.0	13	923	6.7
10	061	1.3	12	987	1.2	15	291	8.9	14	667	9.2	10	805	7.4	10	012	0.5	10	992	10.3	13	980	7.4

²) Uncertain estimates omitted.

3) Uncertain estimates half weight.

TABLE 4.

1	2	3	4	5	6	7	8	9	10	11	12	13	14
var.	$\alpha(1875)$ $\delta(1875)$	type	number of estimates	period m. e.	reciprocal period	phase of rising branch	ΔP	phase of epoch	epoch J. D. 2420000	max. min. range	m. e. of single estimate	value of step	max. min.
	^h ^m ^s [°] ['] ^{''}			^d ^d	^{d⁻¹}				^d	^s ^s ^s	^s	^m	^m ^m
<i>a</i>	18 40 22 — 30 45.2	Cl.	281	.3007673 ± .0000051	3.32483			.2504)	6118.884	1.6 4.7 3.1	.88	.10	13.5 13.8
<i>b</i>	18 41 18 — 26 25.4	Cl.	254	.4103325 ± .0000032	2.437048	.57	.25	.64	6119.14	2.3 9.3 7.0		.23	13.3 14.9
<i>c</i>	18 44 21 — 23 27.9	Cl.	234	.4868848 ± .0000041	2.053874	.49	.20	.56	6118.95	1.5 8.0 6.5		.15	13.8 14.8
<i>d</i>	18 45 31 — 30 40.4	Cl.	195 ²⁾	.4555582 ± .0000032	2.195169	.32	.20	.37	6119.23	1.3 8.2 6.9	.69	.19	13.2 14.5
<i>e</i>	18 48 59 — 30 31.0	Cl.	212	.5697395 ¹⁾	1.755188	.38	.25	.44	6119.25	1.4 10.9 9.5	.90	.15	13.2 14.6
<i>f</i>	18 51 57 — 30 29.0	Cl.	178 ⁵⁾	.5348452 ± .0000099	1.869700	.05	.25	.12	5990.86	1.7 8.3 6.6		.12	13.4 14.2
<i>g</i>	18 52 12 — 27 52.1	Cl.	279	.4664729 ± .0000038	2.143747	.41	.30	.47	6118.95	1.4 14.2 12.8	1.2	.11	13.2 14.6
<i>h</i>	18 53 18 — 29 57.5	Cl.	275	.5771542 ± .0000098	1.732639	.18	.20	.25	6119.13	4.1 10.7 6.6	1.3	.12	13.8 14.6
<i>i</i>	18 56 15 — 31 22.1	Cl.	245	.637673 ± .000015	1.56820	.20	.25	.26	6119.28	0.9 9.2 8.3	1.0	.21	13.8 15.5
<i>k</i>	18 57 33 — 31 47.0	Cl.	264	.4540303 ± .0000042	2.202496	.59	.30	.67	6118.82	1.8 8.9 7.1	.83	.14	13.2 14.2
<i>l</i>	18 59 22 — 32 20.0	Cl.	257	.471289 ¹⁾	2.12184	.94	.30	.00	6119.22	0.4 8.8 8.4	.78	.14	13.2 14.4
<i>m</i>	19 00 02 — 23 48.7	Ecl.	296	.650692 ¹⁾	.153683			.110	6117.20	0.9 17.5 16.6		.16	11.9 14.6
<i>n</i>	19 00 39 — 24 54.2	Ecl.	287	.1984422 ¹⁾	.503925			.110	6118.19	3.3 11.1 7.8		.12	13.3 14.2
<i>o</i>	19 05 05 — 30 22.1	Cl.	189 ²⁾	.4743880 ¹⁾	2.107979	.66	.25	.73	6119.00	0.7 13.9 13.2	1.2	.08	13.3 14.3
<i>p</i>	19 05 19 — 24 52.6	Cl.	276	.6065108 ± .000014	1.648775	.55	.30	.64	6118.87	1.0 9.4 8.4		.20	12.9 14.6
<i>q</i>	19 05 51 — 27 06.0	Cl.	289	.6343265 ± .000014	1.576475	.61	.20	.67	6119.14	2.1 7.5 5.4		.07	13.1 13.5
<i>r</i>	19 06 27 — 26 27.1	Cl.	284	.5193201 ± .0000047	1.925594	.77	.30	.84	6119.07	1.7 9.2 7.5		.13	13.3 14.3
<i>s</i>	19 07 49 — 27 50.4	Ecl.	296	.53937 ¹⁾	.18540			.062	6116.84	1.0 5.4 4.4		.11	12.7 13.2
<i>t</i>	19 09 51 — 29 06.6	Cl.	232 ²⁾	.5893533 ¹⁾	1.696775	.27	.20	.34	6118.87	1.2 8.4 7.2	.71	.10	13.5 14.2
<i>u</i>	19 11 33 — 30 33.0	Cl.	255	.4963279 ± .0000106	2.01480	.47	.25	.52	6118.98	1.5 7.5 6.0	.71	.32	13.1 15.0
<i>v</i>	19 12 31 — 31 52.1	Cl.	255	.743543 ± .000043	1.34491	.83	.30	.95	6119.33	0.6 6.7 6.1	1.0	.16	13.7 14.7
<i>w</i>	19 12 33 — 31 00.7	Cl.	244	.5894137 ± .0000071	1.69660	.94	.30	.01	6119.30	0.5 9.7 9.2	.97	.14	13.1 14.4
<i>x</i>	19 15 08 — 32 12.7	Cl.	285	.4589099 ± .0000045	2.179077	.19	.30	.29	6119.24	1.5 11.2 9.7	.91	.12	12.1 13.3
<i>y</i>	19 15 16 — 23 39.1	Ecl.	296	.303653 ¹⁾	.329385			.310	6118.40	4.6 12.3 7.7		.16	12.5 13.7
<i>z</i>	19 17 29 — 27 08.3	Cl.	268	.4298087 ± .0000027	2.326616	.69	.20	.75	6119.08	1.4 10.3 8.9		.15	13.3 14.6

1) Period obtained graphically.

2) Uncertain estimates omitted.

4) Phase of maximum of the sine curve.

5) Estimates until J. D. 2426273.