

# BULLETIN OF THE ASTRONOMICAL INSTITUTES OF THE NETHERLANDS

1943 March 9

Volume IX

No. 354

## COMMUNICATION FROM THE OBSERVATORY AT LEIDEN

Discussion of 121, mostly new, faint variable stars within  $6^\circ$  of  $\tau$  Sagittarii,  
by *J. G. Ferwerda*.

1. The variable stars discussed in the present paper have been studied on Franklin-Adams plates. Particulars about these plates are to be found in

*B.A.N.* VI, No. 231, of which this paper is a continuation.

The series consists of 422 plates, distributed over

FIGURE 1.

The open area in the year 1934 refers to the Mount Wilson plates.

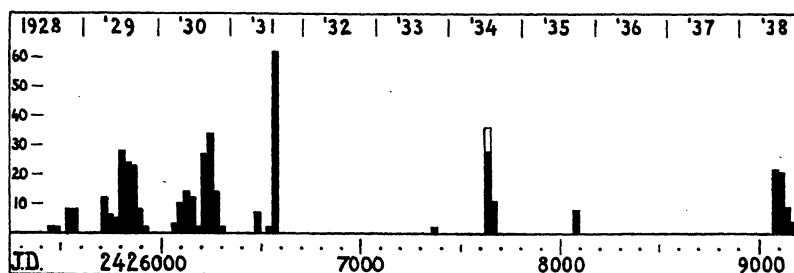


TABLE I.

name	pairs of blink-plates		numbers of new objects		numbers of old variables		total
	J.D. Hel. Grw. — 2420000	M.A.T. — 2420000	dis-cov-eries	redis-cov-eries	found	found again	
A	6123'4131	6129'4656	19	—	2	—	21
B	6120'5890	6125'5549	12	1	3	0	16
C	6118'5888	6125'5770	11	0	3	0	14
D	6120'5669	6123'6146	9	4	2	0	15
E	6118'5666	6123'5928	14	1	0	1	16
F	6156'3460	6161'5159	18	4	0	1	23
G	6155'6157	6156'2995	14	0	1	1	16
H	6153'5422	6155'5939	9	5	1	1	16
I	6210'2931	6212'2654	3	5	0	0	8
J	6122'5698	6129'4438	7	3	0	1	11
K	6093'5276	6103'5684	14	1	3	2	20
L	6569'4069	6570'3175	8	2	0	0	10
M	6561'5043	6568'2708	12	3	0	0	15
N	6559'3129	6570'4893	6	2	1	2	11
O	6565'4942	6571'2587	6	3	2	2	13
P	6570'4249	6573'4449	2	6	1	0	9
Q	6569'2338	6573'4878	3	4	0	1	8
R	6568'2923	6569'4928	2	1	1	0	4
S	6211'2845	6212'2872	5	6	0	2	13
T	6236'2900	6241'3129	7	2	0	1	10
U	5853'3275	5862'2866	2	4	0	3	9
V	5860'2362	5862'2648	1	4	2	1	8
W	5885'2982	5886'2417	4	2	0	1	7
X	5826'4535	5835'3921	0	1	1	1	3
Y	5823'4774	5830'2720	7	4	0	2	13
Z	5797'3076	5799'5077	5	2	0	2	9
	total		200	70	23	25	318

the years of observation as shown in Figure 1. This diagram gives the number of plates taken in each time-interval between two successive full moons. The plates of the years 1928 till 1935 have all been taken by Dr H. VAN GENT, those of 1938 by Dr A. DE SITTER. The time of exposure was usually  $30^m$ . In addition to these plates we have a series of 8 plates taken in one night at Mount Wilson by Dr P. TH. OOSTERHOFF on Imperial Eclipse (backed). The centre of these plates is  $19^h 0, -20^\circ$ , the field  $10^\circ \times 12\frac{1}{2}^\circ$ , time of exposure  $30^m$ ; the scale is nearly that of the Franklin-Adams plates. The two fields have a strip of  $3\frac{1}{2}^\circ$  in common.

2. Most of the variables discussed in this paper have been discovered by Dr H. VAN GENT with the aid of the blink-microscope at Groningen (cf. *B.A.N.* No. 243, p. 21) on 1933 January 16–27. He blinked 26 pairs of plates. We list the results in Table I. This table shows that 200 new and 23 old variable objects have been found. In order of detection they received the current numbers from 1 to 223, which have been used throughout this paper.

Of the 23 old variables 18 have already been studied on these Franklin-Adams plates before. Thus 205 objects remain for further examination. During this examination three objects in addition have been discovered by the author. These have been indicated by the numbers 224, 225 and 227.

3. For particulars about the treatment of the plate material see *B.A.N.* No. 231, 1932 e.g.

All variables have been estimated in 1933 and 1934 on all plates (349) available in 1934. The principal computations followed soon. The plates taken after 1934 are as yet not sufficient in number to justify supplementary estimates and repetition of computations for all variables. So these plates have only been used for a few stars to examine a certain particularity. These particularities are described in the last section of this paper: individual stars. In Table 10, column 4, the J.D. for each variable, up to which the observations are complete, is given. The number of estimates for each star is found in column 5. The total number of estimates made for this investigation is 41441, of which 31260 have been used for the computation.

4. Table 2 gives the number of stars of each type of variability occurring among the 223 objects indicated by VAN GENT.

Of the three variables discovered during examination two turned out to be long-period variables, whereas the third is an eclipsing variable.

TABLE 2.

type of variability	number of new variables	number of old variables found again	
novae . . . . .	1		
long-period . . . . .	3	5	
irregular . . . . .	4		
$\delta$ Cep . . . . .	2	1	
cluster { <i>a</i> - and <i>b</i> -type	95	15	
<i>c</i> -type			88
ultra-short			6
eclipsing . . . . .	18	2	
too faint . . . . .	4		
asteroids . . . . .	4		
not or little variable	69		
total	200	23	

The table shows that a relatively large number of objects appears to be not or little variable. The meaning of "little variable" is that the observable changes of brightness can be ascribed to errors of the plate and to errors of estimation.

The objects denoted by "too faint" are really variable but not bright enough to be estimated.

The most striking feature in the table is the relatively large number of cluster variables. This is in good agreement with the results of other studies in these surroundings<sup>1)</sup>. Moreover a former investigation of the present field suggested already the preponderance of cluster variables<sup>2)</sup>.

<sup>1)</sup> H. VAN GENT, *B.A.N.* No. 227, 1932; *B.A.N.* No. 243, 1933; W. CHR. MARTIN, *B.A.N.* No. 235, 1932.

<sup>2)</sup> *B.A.N.* No. 231, 1932.

Of course Table 2 does not give a true image of the real proportions of the numbers of each type, the chance of discovery being not the same for each of them.

5. The epochs used for the determination of the preliminary period were the observations of maximal or minimal brightness of the variable, the observations of minimum when they are less frequent than those of maximum and the latter in the reverse case. The most probable value of the period of nearly all variables was determined by least squares. For this computation all observations of brightness below or above a certain magnitude have been used rigorously. This critical brightness was chosen for each star in such a way that about 20 epochs were selected. These epochs as well as the counting of the periods and the residuals (O—C) are given in Table 11. For the Algol stars the observations of minimum are given in full (Table 12), because for these stars as a rule the preliminary period has been corrected in a graphical way.

Also among the periods determined with least squares a few appeared to need a small correction. These corrections were determined graphically.

Table 10, column 7, gives the periods definitively adopted and columns 9, 10 and 12 respectively the number of epochs used, the mean error of one epoch and the mean epoch with its mean error. The graphically corrected periods are recognizable because no mean error has been given for them.

6. The determination of the period generally produced no difficulties because VAN GENT took the plates in small and large hour-angles (up to 5½ hours east and west), during all phases of the moon between last and first quarter<sup>1)</sup> and till 3½ months before and after opposition.

Thanks to this favourable distribution of the plates over the time spurious periods are not to be feared. Most valuable for the determination of the period are five long series of plates (each taken during one night). Each series consists of 14 to 21 plates taken in direct succession, thus registering the behaviour of the stars during 7 to 11 hours continuously.

The series of 8 plates taken at Mount Wilson midway between two Johannesburg nights appeared to be of great importance for checking the computed periods of the stars.

By its distribution, outlined above, the material is very suitable for the investigation of variables of short period. On the other hand it is less suitable for the study of long-period stars on account of the small number of years of observation and the irregular yearly distribution of the plates. For this reason the

<sup>1)</sup> Because of the large relative aperture of the Franklin-Adams instrument ( $f : 4.4$ ) practically no plates could be obtained between first and last quarter of the moon.

long-period variables have not been examined: VAN GENT has already taken into account this circumstance in the choice of the pairs of blink-plates, taking small intervals of time between them, thus suppressing the chance of discovery of long-period stars. Moreover discovery of short-period variables has been stimulated by taking the intervals not nearly equal and the plates of each pair in very different hour-angles. In this manner stars having periods very near to 1 day,  $\frac{1}{2}$  day etc. were prevented to be missed. This is of particular importance because galactic cluster variables are known to prefer periods of about half a day.

7. Of all observations phases have been computed with the formula

$$\text{phase} = (\text{J.D.Hel.M.A.T. Grw.} - 2420000) \times P^{-1}.$$

The estimated brightnesses have been reduced to a scale of steps in the usual manner (see *B.A.N.* No. 231 e.g.).

Phase and brightness of all single estimates have been plotted in a diagram, discriminating between observations of different years by dots of various colours. For uncertain estimates a special sign was used. These diagrams are of very great utility, a wrong period as well as changes in period or in shape of the light curve immediately being noticed.

In the case of star No. 221 a change of period is shown and changes in the light curves in the cases of Nos. 64, 74, 100, 163 and 189.

After complete examination of the diagram of estimates the observations have been classed in order of phase and means have been computed of generally ten consecutive observations. Occasionally uncertain estimates have been omitted (see remarks No. 3 to Table 10). The computed means as well as the number of estimates in each group are given in Table 13. They correspond to the mean light curves given at the end of this paper. Of all stars  $1\frac{1}{2}$  period has been drawn, to begin with phase 0.

At the same time the mean error of a single estimate has been computed by first taking the difference in brightness  $\Delta s$  between every two estimates following each other in phase and then using the formula

$$\text{m.e. of a single estimate} = \pm \sqrt{\frac{\sum (\Delta s)^2}{2n}},$$

where  $n$  means the number of estimates used.

The results are tabulated in Table 10, column 16.

8. The positions of the variables in the sky have been obtained by computations according to the principle of the method of dependencies (see *B.A.N.* No. 231) with the aid of photographic enlargements of the surroundings of the variables. The resulting co-ordinates have been checked by reading them

directly on the "Maps of the sky south of  $-19^{\circ}$ " edited by the Union Observatory. The differences between the two determinations were always very small (mean deviation  $\pm 2^s$  in  $\alpha$  and  $\pm 4$  in  $\delta$ ), whereas it may be expected that they have mainly been caused by the errors in the readings on the map. The co-ordinates are found in Table 10, column 2.

The enlargements mentioned above served also to draw the maps of the surroundings (Figure 13) indicating the position of the variable (open circle) and its comparison stars with respect to the neighbouring stars. The current number of the variable is shown in the left-hand upper corner, the right-hand bottom corner indicating the size of each dimension of the square in minutes of arc (usually  $10'$ ). In case of absence of a map the description of individual stars is to be consulted (section 13).

9. The magnitudes of all comparison stars have been determined by direct comparison with a sequence of stars of Selected Area No. 159, which is within the field of the plates. As the quality of the images is not constant over the plate, there are generally great differences in appearance between the images to be compared. Therefore the individual determinations of the magnitude of the comparison stars may be expected to possess rather large errors (up to  $m.5$ ).

In order to get some idea of the accuracy the magnitudes of a number of comparison stars were also determined by means of star counts. For this purpose one of the plates was placed against a plate of glass bearing a photographic reseau (side of each square = 2.210 mm,  $80.11$  squares = one square degree). The comparison star was placed in the centre of a square and then the number of stars brighter than the comparison star was counted. The counts were extended to a number of squares which formed a rectangle with the comparison star in its centre and this figure was expanded until a number of about hundred stars was counted.

The counted numbers were reduced to numbers per square degree and from these numbers the magnitudes of the comparison stars were deduced by the aid of the table of  $\log N_{m,\beta,\lambda}$ , occurring in *Groningen Publications* No. 43. As the counted number was always about hundred the mean error of one determination may be estimated to be about  $m.1$ .

At first these counts have been executed for 14 comparison stars chosen at random. The results are given in Table 3, upper part. When the magnitudes found were compared with the values obtained by direct estimates, considerable systematic differences appeared to exist. This is illustrated in Figure 2a. For the 14 stars the estimated magnitudes (abscissae) have been plotted against the magnitudes from star

counts (ordinates). The broken line gives the theoretical relation. Apart from the systematic deviation the agreement is not worse than could be expected.

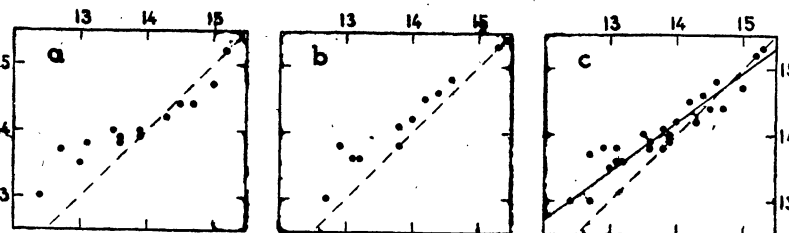
For further examination of the deviation found counts have also been executed for all comparison

stars of the three variables situated nearest to the Selected Area. The results (Table 3, bottom part) are compared with the estimated magnitudes in Figure 2*b*. The scale deviation appears to be a little smaller than in the case of stars chosen at random.

TABLE 3.

variable	name comp. star	number units of area	counted number $n$	$N_{m,\beta,\lambda}$	$\log N_{m,\beta,\lambda}$	$\lambda$	$\beta$	$m_{\text{counted}}$	$m_{\text{estimated}}$
33	a	5 × 5	122	392	2.59	333°	-16°	13.8	13.6
	c	3 × 3	83	740	2.87	333	-16	14.4	14.7
59	a	3 × 5	111	594	2.77	333	-16	14.2	14.3
74	b	3 × 5	81	433	2.64	335	-15	13.9	13.9
95	b	5 × 7	67	154	2.19	329	-11	13.0	12.4
	e	3 × 5	95	509	2.71	329	-11	14.0	13.9
116	a	5 × 5	145	466	2.67	333	-16	14.0	13.5
	c	3 × 3	139	847	2.93	333	-16	14.7	15.0
132	b	5 × 7	95	218	2.34	336	-9	13.5	13.0
	c	5 × 5	98	314	2.50	336	-9	13.9	13.6
152	d	3 × 3	186	1660	3.22	333	-13	15.2	15.2
155	c	3 × 3	91	811	2.91	333	-10	14.4	14.5
189	a	5 × 5	118	379	2.58	329	-13	13.8	13.1
205	A	5 × 5	89	286	2.46	337	-15	13.7	12.7
105	a	5 × 7	114	261	2.42	335	-17	13.6	13.1
	b	5 × 5	115	369	2.57			13.8	13.8
	c	3 × 5	141	754	2.88			14.5	14.2
121	a	6 × 7	78	149	2.17	336	-17	13.0	12.7
	b	5 × 5	90	289	2.46			13.6	13.2
	c	5 × 5	170	545	2.74			14.2	14.0
	d	3 × 3	119	1060	3.03			14.8	14.6
50	a	5 × 5	101	323	2.51	336	-16	13.8	12.9
	b	5 × 5	169	542	2.73			14.1	13.8
	c	3 × 3	89	793	2.90			14.6	14.4
	d	3 × 3	196	1746	3.24			15.3	15.3

FIGURE 2.



A priori there are two possible explanations for the deviations:

1st. A difference in scale between the magnitudes of the Selected Area and those obtained by star counts.

2nd. Errors arising from comparison of star images in consequence of:

a) their difference in size.

b) apart from a), the unknown field correction.

The second explanation seemed the most probable. Therefore it was decided to adopt the system of the star counts and to apply corrections to the estimated magnitudes in order to reduce them to the scale of the counts.

In Figure 2*c*, where the dots of Figures 2*a* and 2*b* are repeated the relation adopted between estimated

and definitive magnitudes is shown by the full line. It is defined by the dots with a rather large uncertainty and therefore an additional correction depending on the position of the star on the plate would be without significance. The magnitudes of the comparison stars obtained in this way are listed in Table 14.

The light curves have now been reduced to the magnitude scale of the star counts and consequently the results derived from them are directly comparable with those obtained by VAN GENT in the Corona Australis region <sup>1)</sup>, VAN GENT having used the same scale as a base for the magnitudes of his variable stars. Nevertheless there will appear to exist a con-

<sup>1)</sup> B.A.N. No. 227, 1932; B.A.N. No. 243, 1933.



siderable difference between the two scales (see section 11).

Table 10, column 14, gives the maximum, the minimum and the range for all variables, expressed in magnitudes. Next, column 15 gives the value of one step in magnitudes for each star. This value appears to be occasionally very different from one star to another.

10. To determine the limiting magnitude of the plates, 12 small square fields have been chosen situated on the diagonals of the plate at three different distances from their point of intersection. Each of the fields has a size equal to four squares of the glass plate with reseau mentioned in the former section. The numbers of stars visible in each of the fields were counted. In this way the numbers of Table 4 have been found. It is evident from this that the limiting magnitude depends on the distance from the centre of the plate, as could be expected.

The values found for the limiting magnitude are in good agreement with VAN GENT's value ( $16^m.4$ )<sup>1)</sup>, but differ considerably from MARTIN's ( $15^m.5$ )<sup>2)</sup>.

TABLE 4.

distance from centre of plate	counted number of stars on 4 areas together	number $N$ per square degree	$\log N$	limiting magnitude
10.6 cm	508	2540	3.40	$15^m.8$
6.4	810	4050	3.61	16.2
1.3	1204	6020	3.78	16.5

Whatever the right value may be, it is beyond doubt that the practical limit is given by a smaller number than  $16^m.4$ ; in our investigation the estimates never surpass  $15^m.4$ .

For estimating the total number of stars registered on a plate the surface of a plate has been divided in three areas by two circles concentric with the plates and having radii of 4 and 9 cm respectively. Each of these areas was supposed to have a star density equal to that of the four fields inclosed. In this way the total number of stars on a plate was found to be 38000.

11. In Figure 3 the range of the cluster variables (except of ultra-short periods) has been plotted against the reciprocal period. The upper part of the figure refers to the variables of this paper completed with the variables published before<sup>3)</sup>. The bottom part gives for comparison the equivalent diagram of the variables of the Corona Australis field<sup>4)</sup>. Stars with light curves of BAILEY's  $c$ -type are indicated by open circles. The figures give rise to the following remarks:

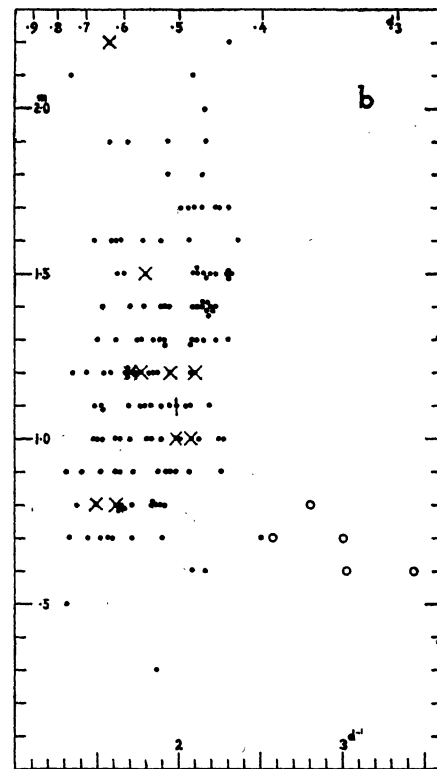
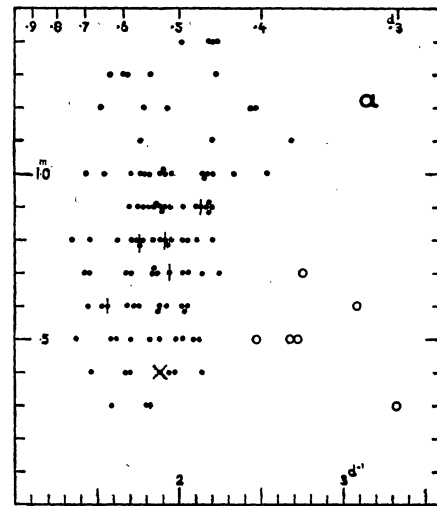
1) B.A.N. No. 227, 1932; B.A.N. No. 243, 1933.

2) B.A.N. No. 235, 1932.

3) B.A.N. No. 231, 1932.

4) B.A.N. No. 227, 1932; B.A.N. No. 243, 1933.

FIGURE 3.

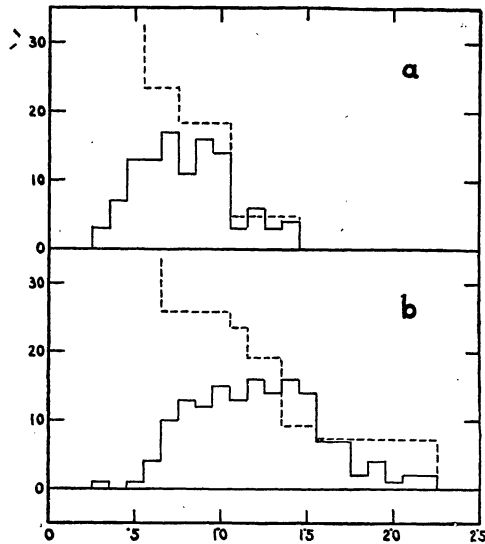


cated by open circles. The figures give rise to the following remarks:

1st. In the field of  $\tau$  Sgr the stars examined have ranges from  $m.3$  up to  $1^m.4$ ; in the CrA field from  $m.5$  to  $2^m.2$ . This difference is still more striking, when the frequency curves of the ranges in the two fields are compared (see Figure 4, full line). It is very improbable that this difference between the two neighbouring fields should be real. The explanation has to be found in an error in at least one of the magnitude scales.

VAN GENT has found ranges which are excep-

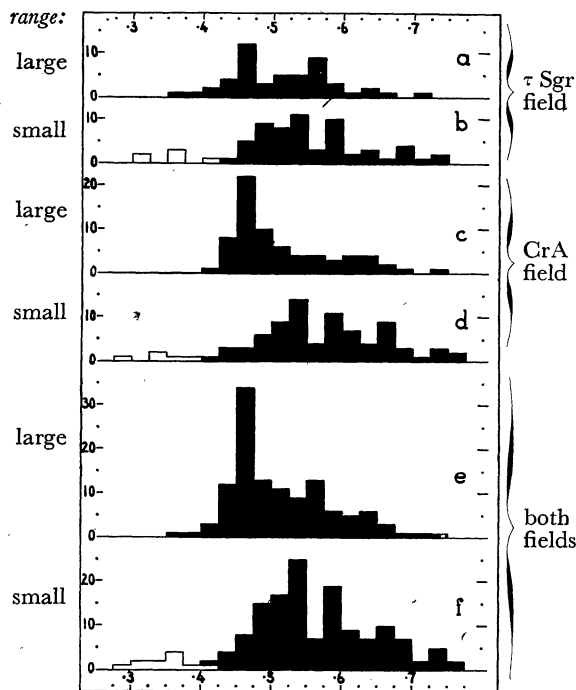
FIGURE 4.



tionally high for cluster variables. E.g. MARTIN <sup>1)</sup> in his investigation of 135 cluster variables in  $\omega$  Centauri has found no ranges above  $1^m.4$ , the scale being fixed by a grating. This is an indication that VAN GENT's scale is too large. The ratio between the present scale and that of VAN GENT is 2 to 3.

2nd. The variables with changing periods (indicated by crosses) appear to prefer neither a definite range nor a definite period; the stars with variable light curves (indicated by vertical dashes) also have

FIGURE 5.



<sup>1)</sup> *Ann. Leiden* XVII, 2e stuk, 1938.

no preference for a distinct period, whereas they may prefer a mean range ( $^m.8$ ).

3rd. The stars with large ranges appear to prefer a period of about  $^d.46$ , longer periods occurring also, but shorter being rare; the small ranges on the other hand have their maximum frequency at a longer period. These particularities appear more clearly in Figure 5, which shows the frequencies of the periods separately firstly for the two fields and secondly for large and small ranges. The separation between large and small range is taken at  $^m.85$  for the  $\tau$  Sgr field and at  $1^m.25$  for the CrA field. Stars of BAILEY's subclass *c* are indicated by open areas.

The difference between the two frequency distributions in the  $\tau$  Sgr field is not so obvious that it could be taken as real. However, the two diagrams show such a great similarity with the corresponding ones in the CrA field, that the observations in the two fields confirm each other and corroborate the conclusion. This is shown in Figures 5e and 5f, being respectively found from the frequency curves a, c and b, d by addition. In the small range curve there is no trace of the preference for periods of  $^d.46$ , which is so obvious in the large range curve.

It may be of interest to compare this result with that of an other investigation. In his study of the variables in  $\omega$  Cen MARTIN <sup>1)</sup> also gives a diagram where ranges are plotted against periods. This diagram also shows a slight decrease of the period of preference for increasing range. His period of preference for large ranges however appears to be  $^d.60$  and differs considerably from  $^d.46$ , the value mentioned above. This period of  $^d.46$  is even avoided in  $\omega$  Cen: whereas the number of cluster variables in this globular cluster is 135, only four of them have a period between  $^d.43$  and  $^d.49$  and three of these are of BAILEY's subclass *c*.

P. TH. OOSTERHOFF <sup>2)</sup> and O. HACHENBERG <sup>3)</sup> independently compared frequency distributions of periods of cluster variables in various globular clusters and in parts of the Milky Way. They concluded that these may be very different for various fields investigated.

12. For the computation of the most probable number of cluster variables in the present field which have not been discovered, the method of VAN GENT has been used <sup>4)</sup>. We denote by:

$N$  the number of cluster variables, discovered or not, present in this field,

$A$  the number of cluster variables, discovered in this field,

<sup>1)</sup> L.C.  
<sup>2)</sup> *The Observatory* 62, p. 104, 1939.  
<sup>3)</sup> *Zs. f. Ap.* 18, p. 49, 1939.  
<sup>4)</sup> *B.A.N.* No. 243, 1933.

$a_k$  the number of cluster variables, which have been discovered  $k$  times,

$\alpha$  the chance of discovery of a cluster variable by comparing one pair of plates,

$n$  the number of plate pairs compared,

$G$  the mean number of times that a variable has been discovered, the undiscovered ones not included.

Between these quantities there are the following relations. The total number of discoveries is  $Nn\alpha$ , consequently

$$G = \frac{Nn\alpha}{A} \quad (1)$$

The number of cluster variables not discovered is equal to  $N-A$ . On the other hand it is equal to  $N(1-\alpha)$ , therefore:

$$N-A = N(1-\alpha)^n \quad (2)$$

$G$  can be expressed in  $\alpha$  by means of (1) and (2):

$$G = \frac{n\alpha}{1-(1-\alpha)^n} \quad (3)$$

$G$  being a known quantity, to be found from

$$G = \frac{\sum k a_k}{A},$$

we may compute  $\alpha$  from (3). This is performed by reading a table giving  $G$  as a function of  $\alpha$ . Finally  $N$  is determined from (1), consequently  $N = \frac{AG}{n\alpha}$ .

In this computation  $\alpha$  is supposed to be the same for all cluster variables, which is not true, the chance of discovery for example being dependent on range and on brightness. The dependence on brightness has been neglected, because this may be supposed to be much smaller than the influence of range, the material being too small to divide it according to both range and brightness. Therefore the cluster variables have been divided into 4 groups according to range, with the assumption that all stars in one group have equal chances of discovery.

In the computations those stars discovered by VAN GENT have also been inserted, which have further not been mentioned in the present paper, because they have been published before (see section 2). The results are tabulated in Table 5.

TABLE 5.

range	m m '3-'55		m m '55-'75				m m '75-'1'05				m m '1'05-'1'45				
	1	2	1	2	3	4	1	2	3	4	1	2	3	4	5
$k$	22	1	18	8	2	4	16	19	4	2	5	4	3	3	1
$a_k$	23		30				41				16				
$A$	24		48				74				39				
$\sum k a_k$	1'04		1'60				1'80				2'44				
$G$	0'004		0'041				0'052				0'087				
$\alpha$	230		45				55				17'2				
$N$	230		45				55				17'2				

This table shows that the computed number of cluster variables with a range larger than m.55 is 117, whereas the number really discovered is 87. Of the group of smallest range only one tenth of the computed number has been discovered. This number of 230 however is very uncertain.

The results are represented graphically in Figure 4a. The full line gives the numbers of cluster variables really observed, whereas the dotted line gives the computed numbers per m.1 interval in range. These numbers have been obtained from the values  $N$  of Table 5 supposing the stars in each group to be equally distributed over the ranges. Figure 4b shows the corresponding frequencies in the CrA field. The discordance between the distributions for large ranges is caused by the fact that VAN GENT has divided his material both according to range and to median magnitude, thus being obliged to bring together stars of very different range in some groups.

Therefore in Figure 5b for ranges greater than m.3 the full line may be supposed to represent the real number of stars better than the dotted line.

If we suppose the magnitude scales to be in the ratio of 2 to 3 the computed number of 117 cluster variables with a range greater than m.55 may be compared with the corresponding result in CrA. The resulting number is 189.

13. Individual variables give rise to the following remarks (compare also the remarks to Table 10).

Var. 1. The light curve of this eclipsing variable shows a secondary minimum.

Var. 2. This cluster variable, having a large number of observations on the rising branch, seemed to be a favourable case for determining the period from these observations. For comparison the period has also been determined as usually from the maxima. The results are:

rising branch:  $P = d.5316468 \pm d.0000014$  (m.e.)  $n = 22,$   
 maxima:  $P = d.5316426 \pm d.0000059$  (m.e.)  $n = 18.$

Phases and mean light curves have been computed from the former value (reciprocal  $10^{\text{d}^{-1} \cdot 880948}$ ).

*Var. 5.* The shape of the light curve differs from the usual shape for variables of this period.

*Var. 6.* The variable is situated near to a star which is a little fainter than the variable in its minimum. Their apparent distance is so small that their images overlap on most of the plates, making estimates of brightness difficult. Determination of a period from the estimates on plates until J. D. 7664 appeared to be impossible. This was considered to be the result of the inaccurate estimates. The plate material being extended until J.D. 9165 and some precautions being taken, a new attempt was made to determine the period. The choice of the comparison stars was improved by replacing star c by another star c', situated nearer to the variable.

During estimation distinction has now been made between the case that the variable and its component were visible separately and the case of overlapping images. This occurred 104 and 264 times respectively. The impression having been established during the first series of estimates that also the component showed variability, now the brightness of this component was also estimated in the series of comparison stars adopted when both stars were visible separately.

From the observations of maximum it proved to be possible to determine the period of light variation. The most probable value was determined from the estimates on the rising branch for plates of good quality. A mean light curve was also computed with weights of 4 and 3 respectively assigned to certain and uncertain estimates. (numbers respectively 64 and 40). These weights have been adopted after determining the mean error of a single observation for each group separately; they amounted to  $\pm 1^{\text{s}}.27$  and  $\pm 1^{\text{s}}.46$  respectively.

TABLE 6.

name	period	references
CY Aqr	<sup>d</sup> 0610	e.g. HOFFMEISTER and JENSCH, <i>A.N.</i> 253, No. 6051, 1934; A. J. WESSELINK, <i>B.A.N.</i> No. 341, 1941.
BAILEY 65 in ω Cen	0627	E. HERTZSPRUNG, <i>B.A.N.</i> No. 247, 1933.
VV Pup	0697	H. VAN GENT, <i>B.A.N.</i> No. 214, 1931.
DY Peg	0729	A. SOLOVIEV, <i>Tadjik Circular</i> No. 37, 1938.
KU Cen	0800	W. CHR. MARTIN, <i>B.A.N.</i> No. 232, 1932.
RV Ari	0852	H. GURIEV, <i>Tadjik Circular</i> No. 38, 1938.
V494 Sgr	1076	H. VAN GENT, <i>B.A.N.</i> No. 243, 1933.
BP Peg	1094	*)
AI Vel	1116	E. HERTZSPRUNG, <i>B.A.N.</i> No. 224, 1931.

\*) In the "Katalog und Ephemeriden Veränderlicher Sterne" by H. SCHNELLER this star does not occur among the eclipsing variables for the first time in the edition for 1941.

The light curve obtained in this way is shown by the dots in Figure 14a. Every dot has a weight corresponding to about 10 certain estimates. Open circles are means derived from observations of the combined images of variable and component. These points have to be considered as less certain and systematically in error. They have been obtained as follows.

The observations of the first series of estimates have been reduced to the scale of the second; the means have been taken of both estimates made on each plate. Of the values for combined blackness arrived at in this manner, means have been taken in the ordinary way (in groups of 12). The internal mean error of one (mean) observation of the combined blackness of variable and component proved to be  $\pm 1^{\text{s}}.51$ .

Finally an attempt has been made to prove the variability of the component by establishing a period in its light variation. For that purpose the difference has been taken of all observations of combined blackness with the mean light curve (of the single variable) at the same phase. The largest differences have been considered as the maxima of the component. However neither in these differences, nor in the direct estimates of the component any period could be found. The greatest brightness observed is  $8^{\text{s}}.6$ , corresponding to  $14^{\text{m}}.5$ .

*Var. 14.* This star belongs to the group of cluster variables with ultra-short periods. The only stars of this type known to have shorter periods are listed in Table 6.

The determination of the period did not produce any difficulty, owing to the nights with long series of observations. The estimates from these nights together with the mean light curve repeated periodically are given in Figure 6 (open circles are uncertain estimates).

*Var. 17.* Single observations show a large spread caused by the fact that immediately north of the variable (about 0.04 mm on the plate) a faint star is situated, which is only visible on plates of good quality when the variable is also faint.

To check the period supplementary estimates have been made on the plates after J.D. 7664. The new observations proved to be in accordance with the originals. They have not been used further.

*Var. 20.* This Algol star shows a deep minimum (Figure 14e) and a secondary minimum is not indicated with certainty. The period has been determined graphically from the branches. Open circles in the mean light curve are means of 4 to 8 observations; dots of 10, as usual.

*Var. 23.* This eclipsing variable has two minima of nearly equal depth (Figure 14e). The least squares solution is related to the period of light variation,



FIGURE 6.

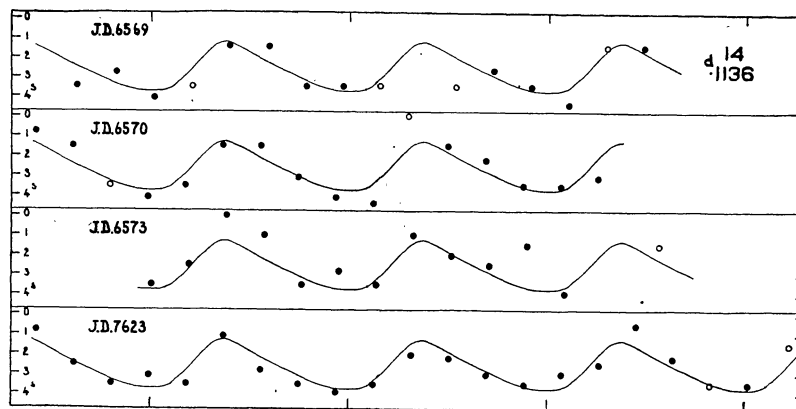
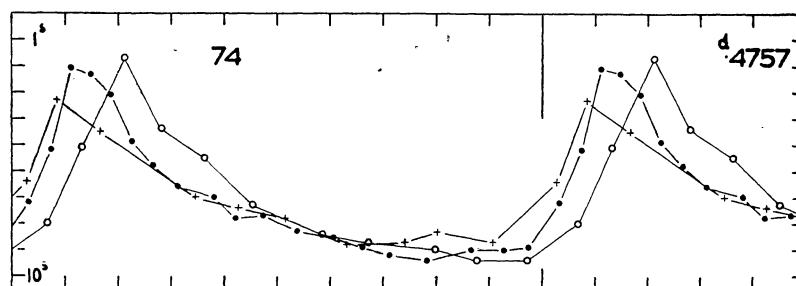


FIGURE 7.



which is half the period of rotation, given in Table 10.

*Var. 31 and 36.* These irregular variables (V729 and V730 Sgr) have already been published before<sup>1)</sup>. The values of the magnitude, given in Table 10, column 14, as well as those of the comparison stars (Table 14) differ slightly from the values in that publication, because they have been determined anew according to the method mentioned in section 9.

*Var. 45.* Upon an indication of variability of the light curve of this star, the plates until J.D. 9165 have been estimated in addition. The presumption was not confirmed.

*Var. 58.* The light curve of this Algol variable (Figure 14e) shows an indication of a secondary minimum. The open circles are means of 6 observations (dots of 10).

*Var. 61.* The light curve of this Algol variable (Figure 14e) shows a secondary minimum.

*Var. 64.* The diagram of single estimates shows a large spread, suggesting a false period or a variable light curve. To obtain further information the plates until J.D. 9165 have been estimated in addition. Checking of the period showed the counting to be correct without doubt. A small correction could be applied with the aid of the new observations. The diagram of single estimates shows:

<sup>1)</sup> B.A.N. No. 256, 1934.

1st. Some maxima are extra high.

2nd. After correction of the period two observations of minimum occur at the time of maximum.

3rd. The observations in 1934 are shifted systematically  $P \cdot 1$  in phase with regard to the others.

Consequently there are indications that this star deviates now and then from its mean light curve. The new observations have not been used further.

*Var. 70.* This bright Algol star (V 523 Sgr = C.P.D. - 29°5846) has two minima of nearly equal depth with a difference in phase of  $P \cdot 414$  (Figure 14e) indicating an orbital eccentricity of at least  $\cdot 14$ <sup>1)</sup>. In connection with the brightness of the variable and its comparison stars and also with their rather large mutual distance the estimates have been executed with the naked eye. For further information reference is made to the publication concerned<sup>2)</sup>, which is based on observations until J.D. 6573. The star has also been estimated on all plates after that date, but appeared in minimum on two plates only (Table 12). These observations occurred at the expected phase, so they did not alter the original value of the period. For the computation of the mean light curve the new observations have not been used.

<sup>1)</sup> J. UITTERDIJK, "A method of deriving limits for the eccentricity of the orbit and for the longitude of periastron of an eclipsing binary." B.A.N. No. 237, 1932.

<sup>2)</sup> B.A.N. No. 256, 1934.

Miss JONES<sup>1)</sup> studied this star more fully on 1650 Harvard plates dating down to 1889, and demonstrated the rotation of the line of apsides. Her observations have been discussed anew by RUSSELL<sup>2)</sup> and DE KORT<sup>3)</sup>. Both concluded to an orbital eccentricity of at least .17 and a period of apsidal revolution of at least 200 years.

*Var. 74.* The diagram of single estimates showed the light curve to be variable. For some time the observations satisfy a certain light curve, whereas during the next interval of time the star has an other light curve. To get further information the star was estimated on all plates available. From the beginning of the series of plates (compare Figure 1) until J.D. 5805 the star proved to satisfy the light curve indicated by crosses (means of 5 or 6 observations) in Figure 7. After this date there are 12 nights without observations and the plates taken since show the star to have the light curve indicated by open circles (means of 9 or 10 observations) in the same figure. The observations fit this curve until J.D. 6214. Then, after an interval of 3 nights without observations, to begin with J.D. 6218 the star has the light curve indicated by black dots<sup>4)</sup> (means of 10 observations). The straight lines connecting consecutive points give only a rough approximation of the light curves; they have been drawn to bring out the principal characteristics of the figure.

These observations show the behaviour of the star to change with jumps. The changes concern the magnitude of maximum and minimum and the phase of the rising branch, but not the period, for this fits all three light curves very well. As the rising branches first advance and then recede, they cannot be made to coincide by correcting the period. A representation of the observations by a uniformly variable period is also impossible, the phases changing abruptly.

The computation of the mean error has been carried out separately for each of the three light curves. Moreover every curve has been divided into a part

TABLE 7.

curve	J.D. — 2420000	mean error of a single estimate		
		phase '0—'5	phase '5—'0	whole period
+++++	5442–5808	± .76 (27)	± .88 (23)	
ooooo	5821–6214	± 1.10 (58)	± .69 (47)	
.....	6218–9165	± 1.28 (108)	± .70 (80)	
all three		± 1.16 (193)	± .73 (150)	± 1.00 (433)
mean curve	5442–9165	± 1.62 (193)	± .74 (150)	

1) *Harvard Bulletin* No. 909, 1938.

2) *Ap. J.* 90, p. 641, 1939.

3) *B.A.N.* No. 348, 273, 1942.

4) The light curve in Figure 14a is also based on the observations after J.D. 6218 only.

where the three curves coincide practically (phase '0 to '5) and a part where they are different (phase '5 to '0).

The results are given in Table 7 (between parentheses: numbers of observations). A remarkable feature is the difference in spread between the variable and the constant part of the light curve, the spread in the former part in two of the three curves being considerably larger than in the latter. This suggests the two changes occurring with an interval of about a year to be not the only changes the light curve has undergone during the years of observation. Whether the remaining variations occurred also with jumps or had the character of a continuous change cannot be settled with the material available. Anyway the observations are well represented by the three light curves mentioned.

*Var. 76.* This star is at the border of the platefield and situated close to a star of magnitude 15<sup>m</sup>.2, which disturbs the image, both circumstances making estimation difficult.

*Var. 80.* The diagram of single estimates shows a large spread of individual observations about the mean light curve.

*Var. 83.* The light curve of this Algol star shows no trace of a secondary minimum. Open circles in the mean light curve (Figure 14c) are means of 5 single observations.

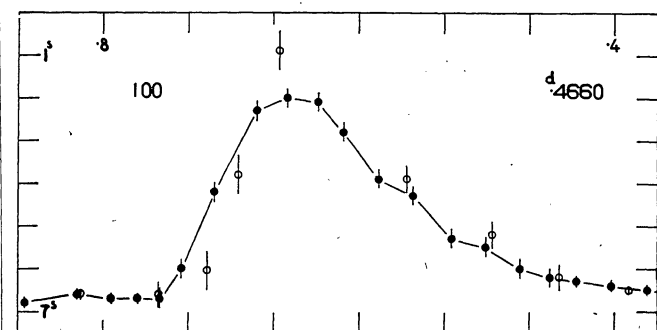
*Var. 86.* The period of this star is 47<sup>d</sup>.50. The light curve proves to be strongly and irregularly variable (minima not equally deep). The star has not further been examined on account of the unsuitability of the plate material for this type of variables. Min. J.D. 5808, max. J.D. 5834.

*Var. 88.* The light curve (Figure 14c) is rather uncertain in minimum because the number of observations is too small as a consequence of the star's faintness.

*Var. 96.* The light curve of this Algol star shows no indication of a secondary minimum. Open circles in the diagram (Figure 14e) are means of 5 observations; dots of 10, as usual.

*Var. 100.* From observations till J.D. 7664 a good

FIGURE 8.



light curve resulted, but two of the latest observations showed a remarkable deviation. For further information also the plates till J.D. 9165 have been estimated. The light curve, to begin with the observations of 1933 or 1934<sup>1)</sup> proved to have a shape differing from that before. Therefore separate light curves have been constructed from plates before and after J.D. 7000 (see Figure 8). The dots are those of Figure 14a; they represent 10 observations before J.D. 7000 each. Open circles are means of 10 observations (except 3 points on the rising branch, which are means of 5) after that date. The lengths of the vertical dashes through the normal points represent the mean errors. The difference between the two curves, though not large, is undoubtedly real.

The mean error of a single observation was determined separately for each of the two curves, this quantity moreover being computed for the coinciding part of the curves (phase .35 till .85) and for the differing part (phase .85 till .35) separately.

The results are given in Table 8 (between parentheses: numbers of observations). Especially in the

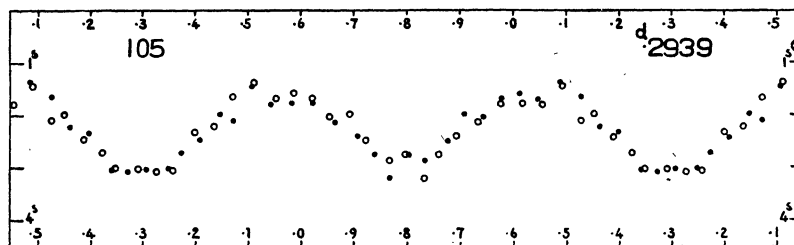
TABLE 8.

curve	J.D. — 2420000	mean error of a single estimate		
		phase .85—'.35	phase .35—'.85	whole period
• • • • •	until 7000	$\pm .74$ (134)	$\pm .42$ (114)	
○ ○ ○ ○ ○	after 7000	$\pm 1.96$ (47)	$\pm .33$ (58)	
both		$\pm .84$ (181)	$\pm .39$ (172)	$\pm .66$ (353)

curve after J.D. 7000 the mean error of the observations in the changing part is much greater than in the constant part. Therefore the changes in the light curve mentioned above are probably not the only changes which occurred during the years of observation and Figure 8 is likely to give only a rough image of the behaviour of the star.

*Var. 105.* This variable of short period has a very small range. Therefore the type of variability cannot be stated with certainty. If it is not a variable of the W UMa type, it must be an ultra-short period cluster variable (<sup>d</sup>1469722). However the W UMa type appears to be the most probable for the following reasons:

FIGURE 9.



1st. The minima are sharper than the maxima.  
2nd. When the light curve is drawn in reverse direction together with a normal one in the same diagram, in such a way that the phases of the minima coincide (see Figure 9), the minima appear to have unequal shape. If the star were a cluster variable then the minima of Figure 9 should have equal shape. The least squares solution corresponds to the period of light variation, which is half the period of rotation, given in Table 10.

*Var. 122.* No secondary minimum of this Algol star has been observed. The period may have to be doubled. The variable being below the limiting magnitude of the plates on a number of minimum plates the magnitude of the minimum is uncertain. In the mean light curve (Figure 14c) means of 4 or 5 observations are indicated by open circles.

*Var. 130.* The light curve of this  $\delta$  Cephei star with a period of  $13^d.51$  resembles very much that of FI Carinae with a period of  $13^d.45$ , which has

been discovered and studied by E. HERTZSPRUNG<sup>1)</sup>. The new plates till J.D. 9165 have been estimated in addition for the purpose of detecting possible changes. The star proved to be perfectly regular. The period could be corrected graphically. The new observations have not been used further.

*Var. 132.* In the diagram of single estimates there are indications of inconstancy of the period (lengthening).

*Var. 140.* In minimum (phase .40 till .60) the estimates are distributed sensibly (over 2 times) thinner than in the remainder of the period. In the former part the mean light curve may be systematically too high.

*Var. 144.* For the map of the surroundings see var. No. 142. The corresponding comparison stars have been marked with an accent.

*Var. 146.* This star (V 522 Sgr) is a Nova of 1931. Particulars are to be found in a separate publication<sup>2)</sup>.

<sup>1)</sup> From the two observations in 1933 no conclusion can be drawn.

<sup>1)</sup> B.A.N. No. 95, 1926.

<sup>2)</sup> B.A.N. No. 269, 1935.

*Var. 148.* The diagram of single estimates shows a gap in the observations between phase .40 and .50. Beyond .50 some observations occur and the density increases a little till .85, but in the whole interval of phase .50 till .85 there are only 25 observations. The cause is possibly that comparison star c is a star close to the variable. The stars are only seen separately when the variable is about equally bright as or fainter than c and only in case the images are sharp; the plates where the variable is a little fainter than b are unsuitable for estimation, the estimates becoming easier when the variable is fainter.

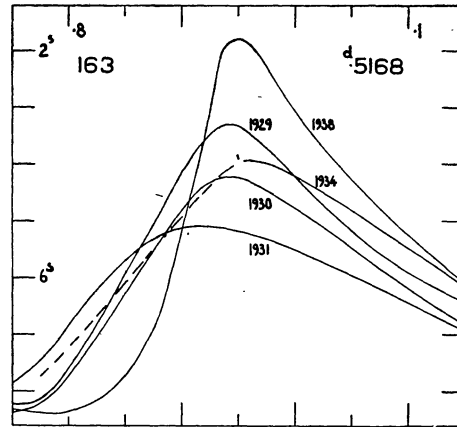
*Var. 152.* The light curve of this Algol star (Figure 14e) shows no trace of a secondary minimum. Open circles are means of 4 or 5 observations (dots of 10 as usual).

*Var. 159.* This variable proves to be an irregular one. It seems to be nearly periodical with a period of about 80 days. The plate material, being unsuitable for this type of variables, has not been further examined.

*Var. 162.* The light curve of this Algol star (Figure 14e) shows an indication of a secondary minimum at phase .3. Open circles are means of 5 or 6 observations (dots of 10 as usual).

*Var. 163.* The diagram of single estimates shows the light curve to differ from year to year. For further information the plates of 1935 and 1938 have been estimated in addition. To begin with this yielded a correction to the period. Next, the observations of

FIGURE 10.

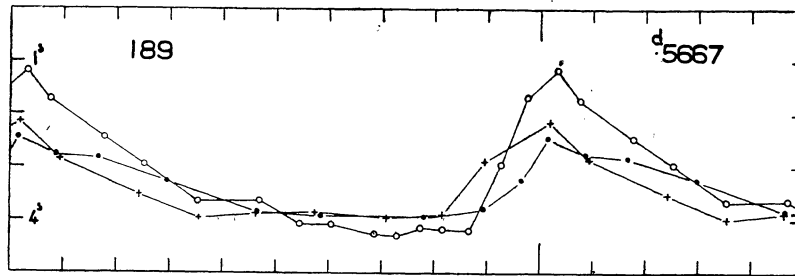


1938 proved to differ very much from the older ones. For this reason a new mean light curve has not been constructed. The mean light curve covering the years 1929 till 1934, computed with the original reciprocal period  $1^d \cdot 934844$  is given in Figure 14b.

Figure 10 gives the mean light curves of the different years separately for the new reciprocal period  $1^d \cdot 934833$ . In 1934 no observations on the rising branch were available: the broken line in the figure is a nearly straight line connecting the last observations in minimum with the first in maximum. The figure shows the greatest range to be at least twice the smallest.

*Var. 169.* For the map of the surroundings see

FIGURE 11.



*var. 69.* The corresponding comparison stars have been marked with an accent.

*Var. 175.* In the diagram of single estimates there are some indications of a possible shortening of the period.

*Var. 188.* This star, having a period of  $1^d \cdot 835$ , is the only  $\delta$  Cep star known to have a period between that of VZ Aql ( $1^d \cdot 668$ ) and SU Cas ( $1^d \cdot 949$ ) according to the *Katalog und Ephemeriden* for 1941.

*Var. 189.* The diagram of single estimates shows the light curve to be variable. The changes are imperceptible in the course of one year of observation. The plates until J.D. 9165 have been estimated in addition. The material has been divided into three groups

and a mean light curve has been computed for each group separately. The results are given in Figure 11. The curve indicated by open circles is identical with that given in Figure 14c. Straight lines of connection between consecutive points serve to emphasize the chief features of the figure.

TABLE 9.

curve	J.D. — 2420000	mean error of a single estimate	number of observations
• • • • •	5442-5926	$\pm$ .70	89
o o o o o	6067-6573	$\pm$ .45	146
+ + + + +	7367-9165	$\pm$ .60	85



Mean errors of a single estimate have been computed separately for each of the three curves. Table 9 shows the results. From this the mean value  $\pm 0.57$  is found, which is also given in Table 10.

*Var. 191.* This eclipsing variable has two unequal minima (Figure 14e).

*Var. 207.* For the map of the surroundings see var. 90. The corresponding comparison stars have an accent.

*Var. 219.* For the map of the surroundings see var. 124. The corresponding comparison stars have an accent.

*Var. 221.* It proved to be impossible to derive a period satisfying all observations of this star. A period can be found however, which represents the observations of the first four years. This has been determined with least squares and came out to be  $d.5310747 \pm 0.000135$  ( $d^{-1} \cdot 882974$ ). The corresponding mean light curve for the first four years is given in Figure 14b. The phase of maximum is .36. With this period, however, the maxima in 1934 have the phase .96, those in 1938 .42, whereas for the observations in 1935, containing no maximum, the

phase of maximum can be estimated to be .65<sup>1</sup>). The shape of the light curve was the same in all years.

These phases of maximum being plotted in a diagram against the mean J.D. it proves to be impossible to bring them on a straight line, whatever the counting of the periods may be. Accurate examination of the observations from 1928 till 1931 showed the phase to increase first and to decrease afterwards, suggesting a shortening of the period. With least squares a parabola has been computed passing through the best determined individual maxima. For this parabola the following equation has been found:

$$ph_{\max} = -674 \times 10^{-10} E^2 + 0.00186 E + 2.298 \\ \pm 33 \qquad \qquad \pm 25 \qquad \pm 28 \text{ (m.e.)}$$

where E means the number of periods elapsed after J.D. 5470228. From this parabola the quadratic term in the ephemeris of the maximum was found to be:

$$-358^d \times 10^{-10} E^2 \\ \pm 18 \text{ (m.e.)}$$

In order to collect all observations in a single light curve phases have been computed with the formula:

$$\text{phase} = (\text{J.D.Hel.M.A.T. Grw. } -2420000) \times d^{-1} \cdot 882974 + \\ (\text{J.D.Hel.M.A.T. Grw. } -2426203)^2 \times 239 d^{-2} \cdot 5 \times 10^{-10}.$$

With these phases the diagram of single estimates has been drawn and a rather good light curve resulted. Two things however drew the attention:

1. in the year 1935 an observation of minimum occurs at the phase of maximum.

2. the spread of the observations in 1938 is too large. From 1, though this is based upon one observation only, it follows that the representation of the progress of the phase of maximum by a quadratic function of the time is erroneous or at least an insufficient approximation.

This becomes also evident when the phase of maximum is derived separately for each month<sup>1</sup>). The phases (from reciprocal period  $d^{-1} \cdot 882974$ ) together with the parabola have been drawn in Figure 12. The diameter of the circles increases, as shown in the figure, with increasing number of observations  $n$ , from which the corresponding phase has been determined.

The figure shows that the parabola passes close to the circles, but also that it cannot describe the behaviour of the star with sufficient accuracy. Moreover the three circles of 1938 indicate that the parabola does not describe approximately the behaviour of the star, but probably gives an entirely wrong representation of it. The situation namely of the

three circles mentioned indicates an increase of the phase of maximum (connected with an increasing period), whereas according to the parabola the phase (and also the period) ought to decrease. Consequently the counting of the periods probably is wrong, some points in the figure having to rise or to descend by a whole number of periods.

Some possibilities for the maxima in 1934, 1935 and 1938 are shown in the figure. The problem is how to determine how the phase of maximum varies with the time. The dotted line gives one of the possibilities, supposing the period to vary gradually. Other solutions are possible.

If it is supposed that the period could vary with a leap, it becomes fully impossible to determine its behaviour. Frequent observation will be necessary.

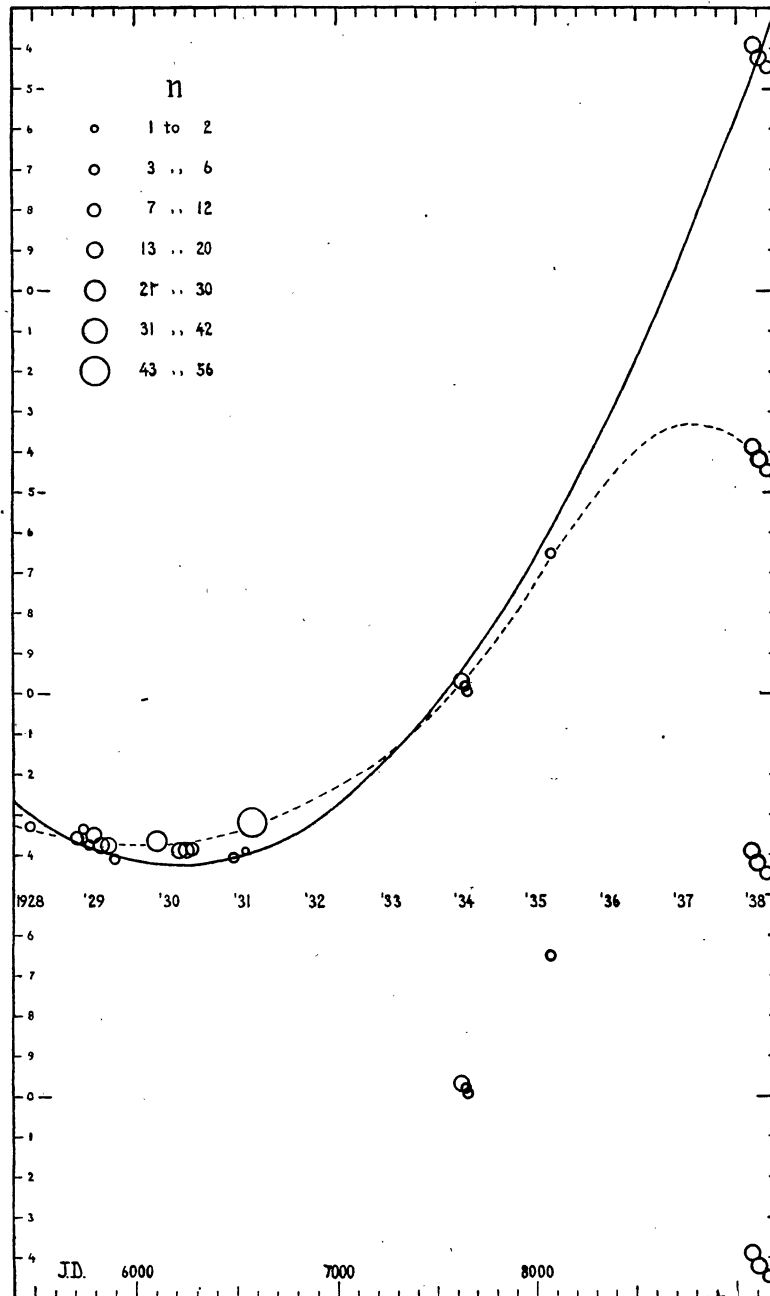
*Var. 224.* This variable has been discovered during examination of var. 83 and is indicated on its map. The type being long-period it has not been further examined.

*Var. 225.* This variable (C.P.D. — 30°5669, CPD magn. 8<sup>m</sup>.8) was discovered during examination

<sup>1</sup>) See note in the next column.

<sup>1</sup>) This has been done by plotting the observations against phase in a diagram and shifting the mean light curve of the first four years, drawn on a piece of transparent paper, until it coincides as closely as possible with the observed points. By this method it is possible to obtain a rather good value of the phase of maximum even from a few observations.

FIGURE 12.



of var. 109. The type of variation is Algol, the provisionally determined period being  $4^d.15$ . The examination of this star has not been finished as yet.

*Var. 227.* This variable has been discovered during examination of var. 162 and is indicated on

its map. The type being long-period it has not been studied further.

For much help in preparing this paper I am indebted to the computers of the Observatory, especially to Messrs DE HAAS and DE ROOY.

FIGURE 13a.

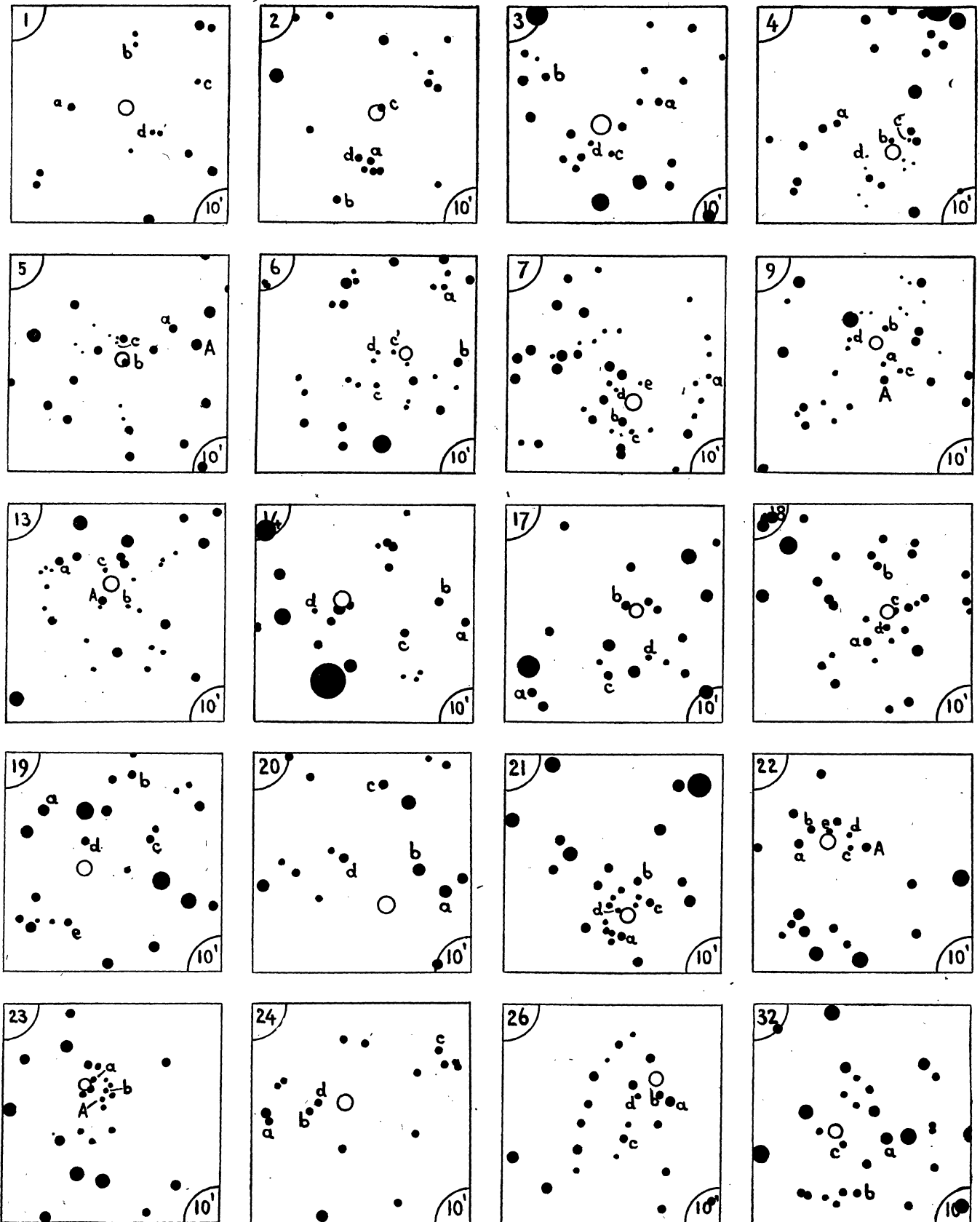


FIGURE 13b.

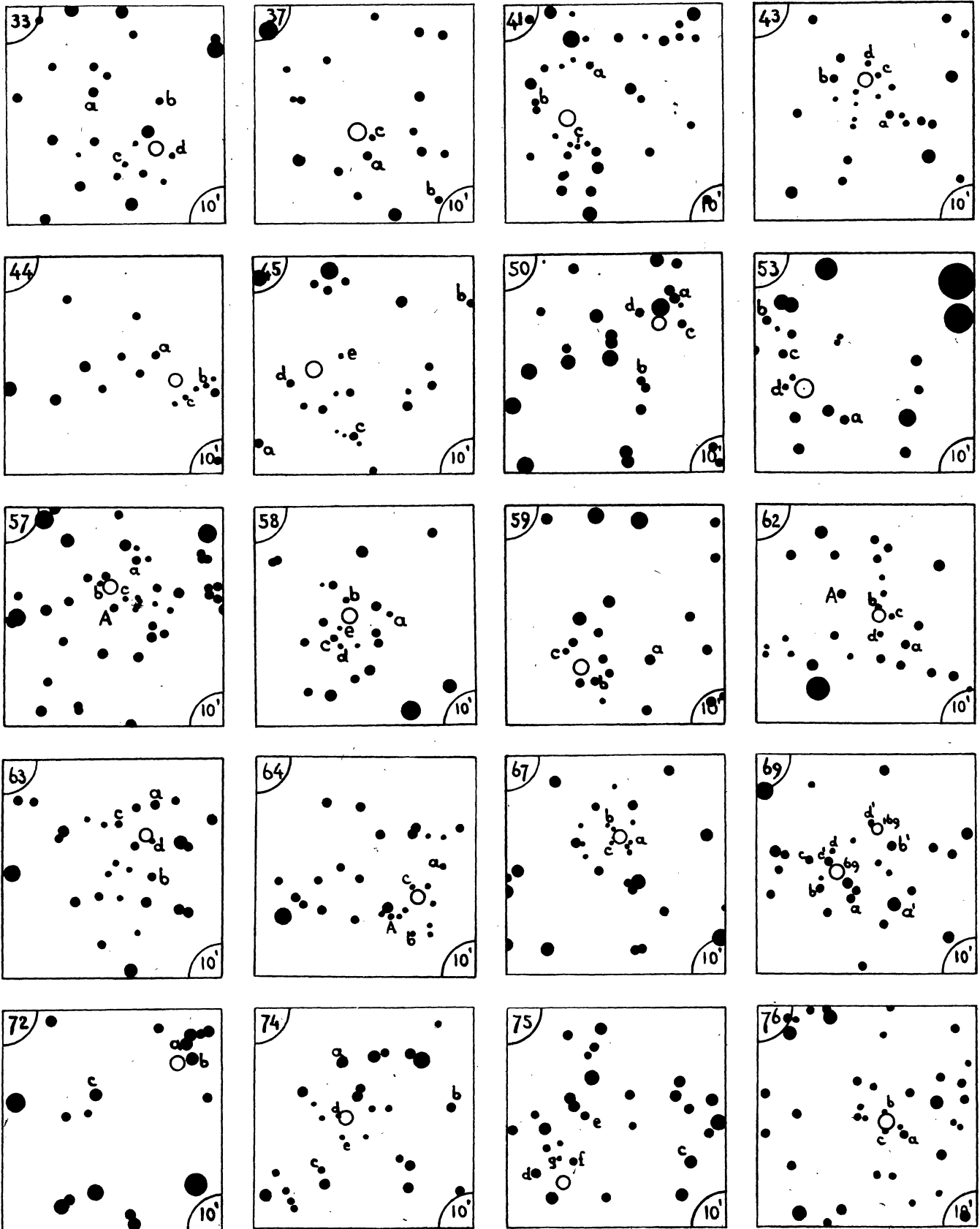




FIGURE 13C.

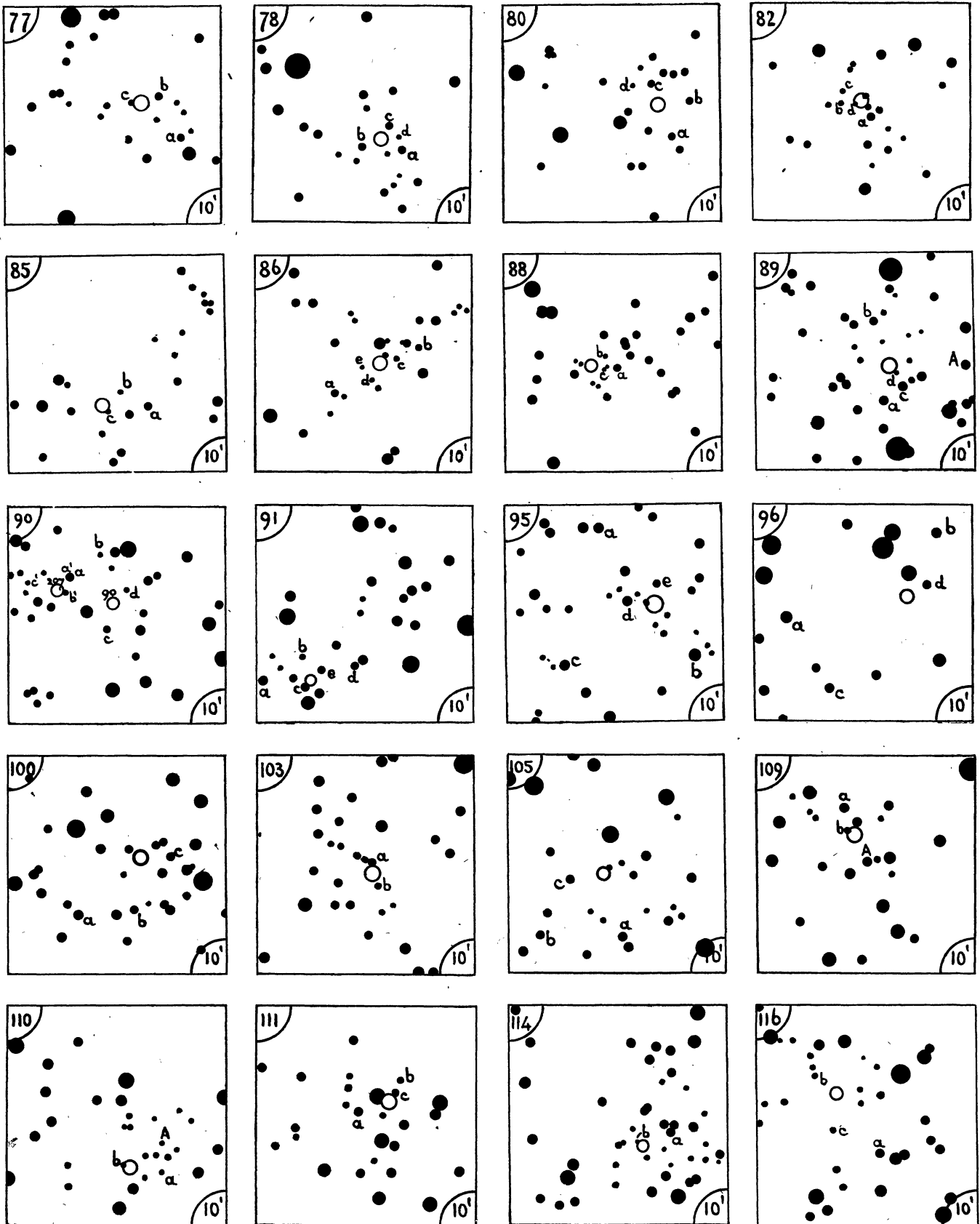


FIGURE 13d.

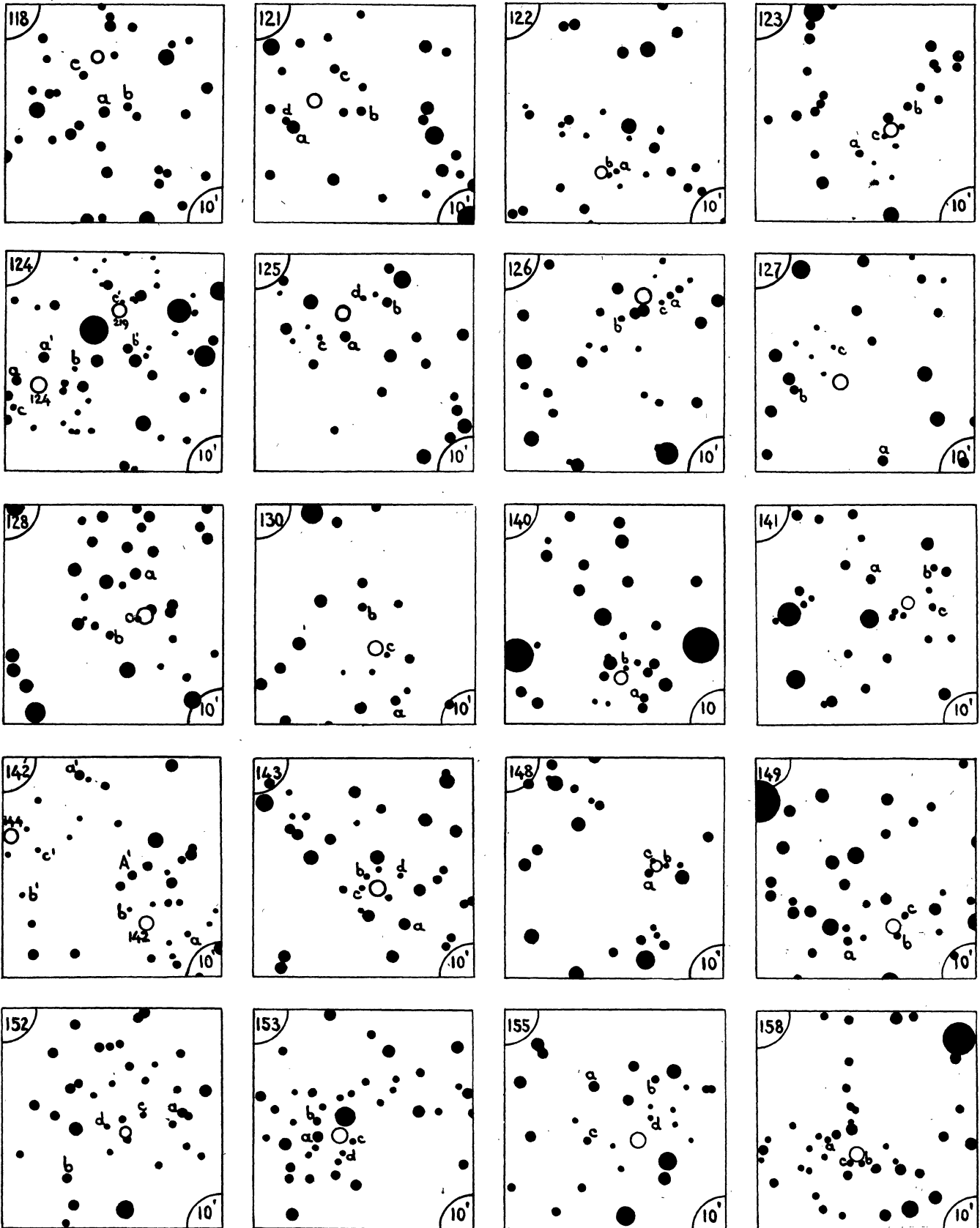


FIGURE 13c.

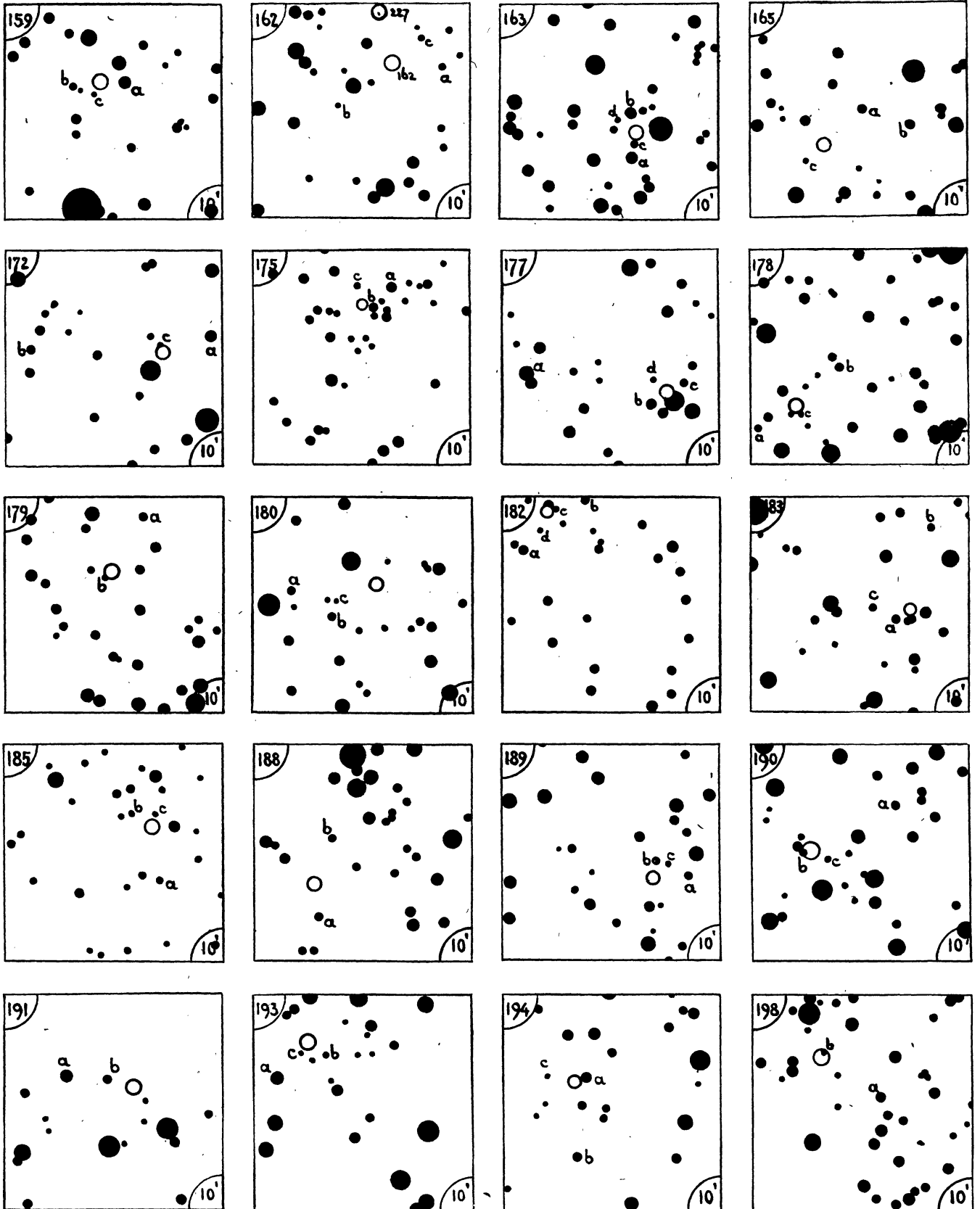


FIGURE 13f.

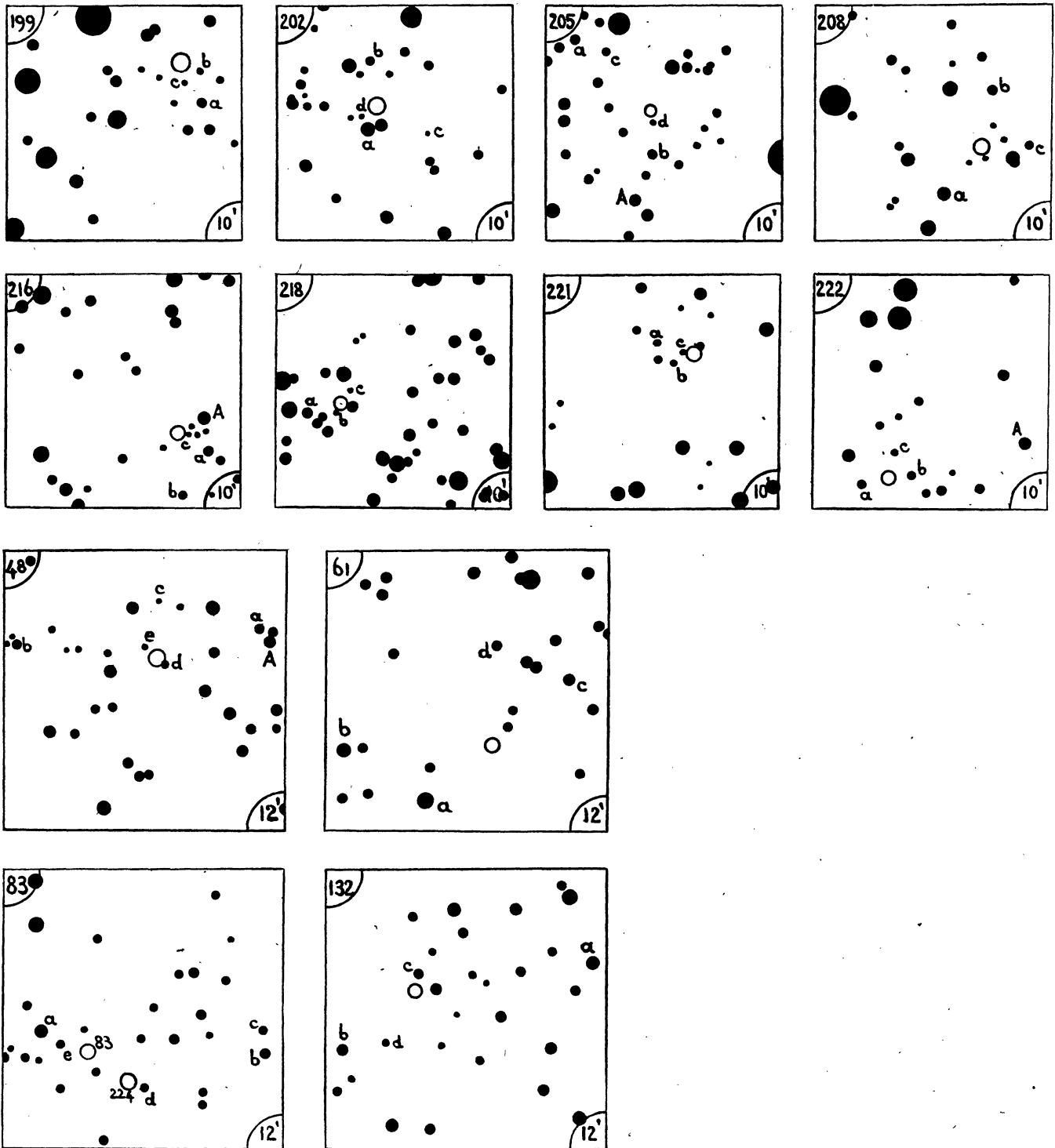




FIGURE 14a.

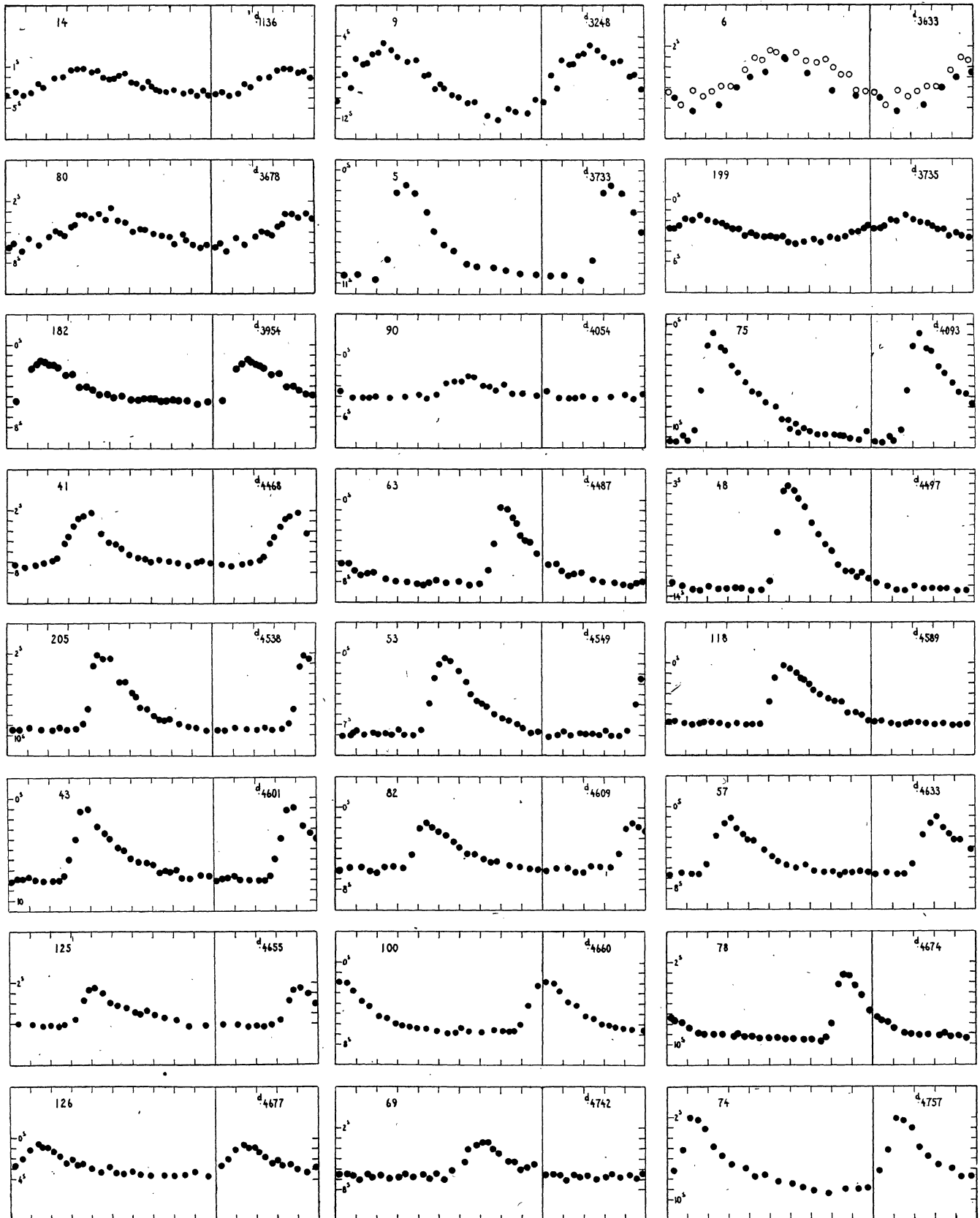


FIGURE 14b.

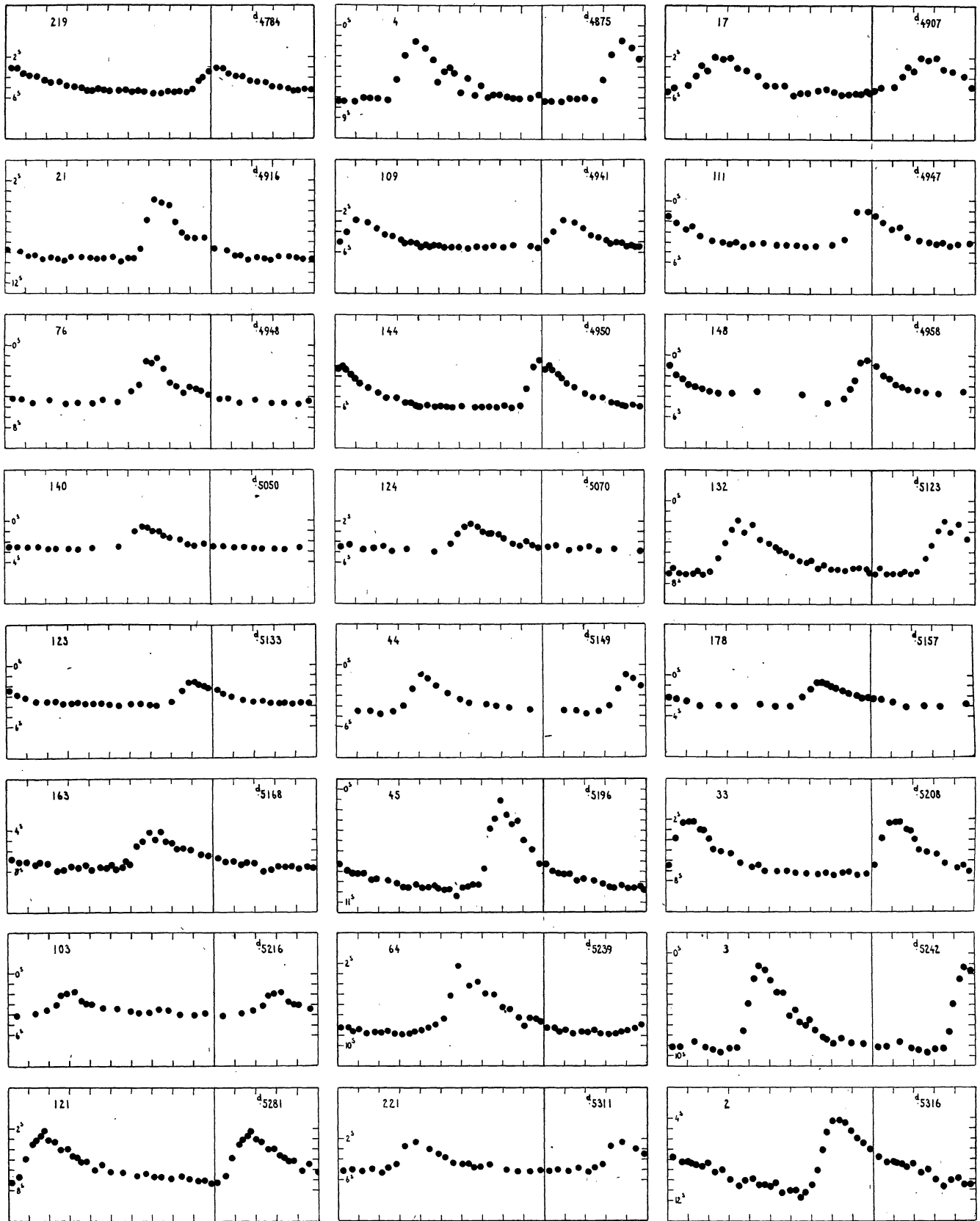


FIGURE 14C.

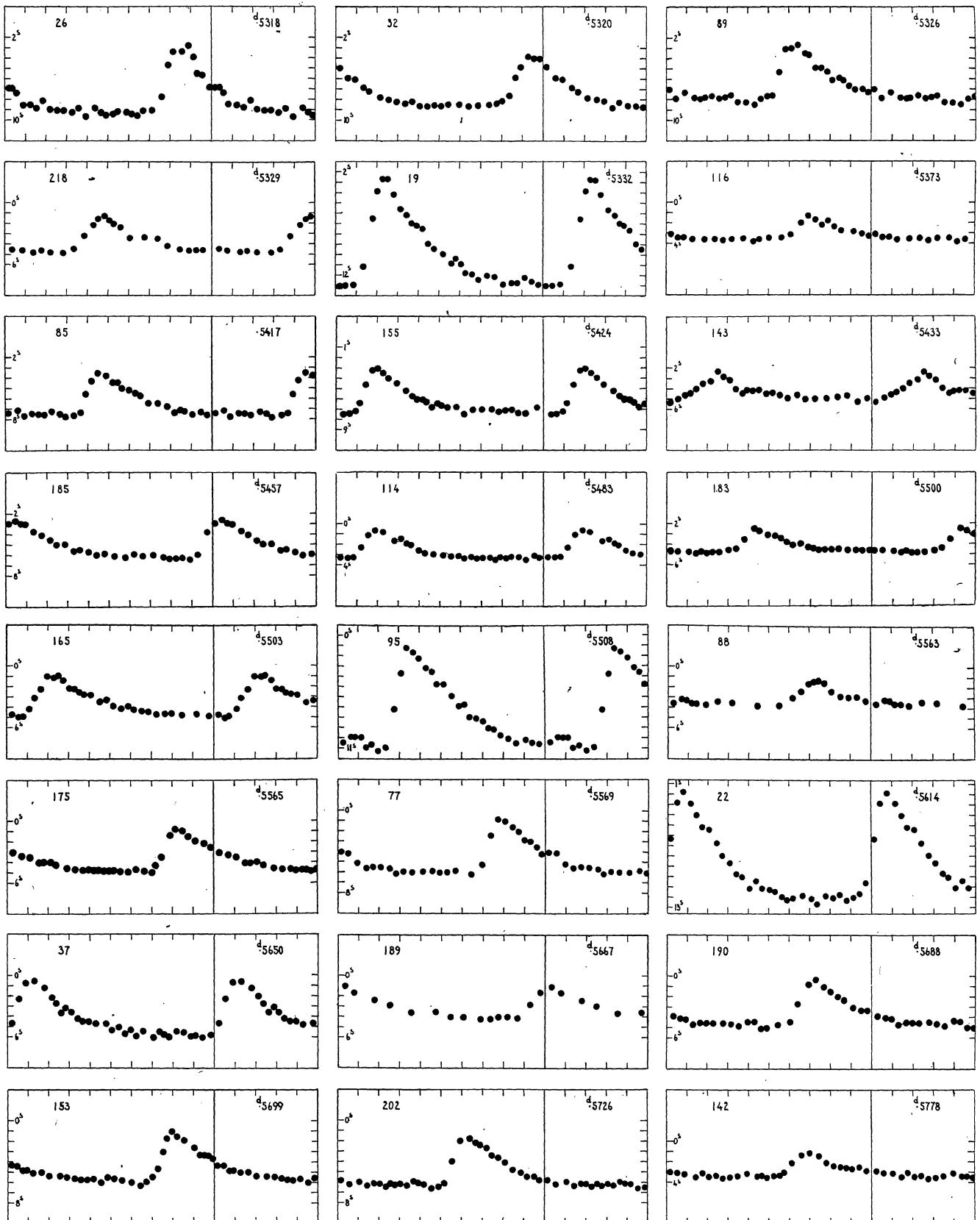
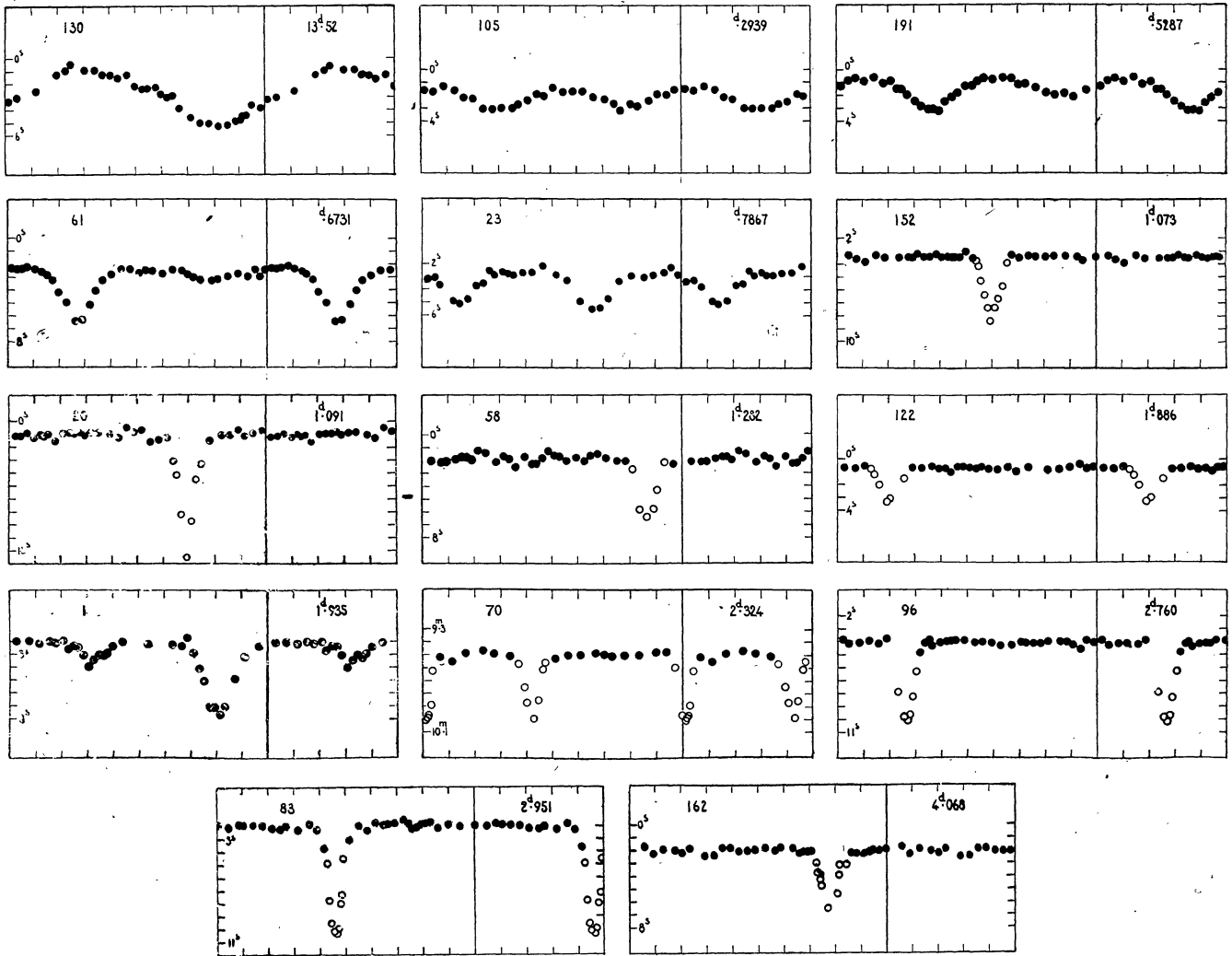






FIGURE 14c.



TABLE

1	2	3	4	5	6	7	8		
No. var.	$\alpha$ (1875)	$\delta$ (1875)	discovered on blink-pairs	plates estimated until JD 242	number of estimates used	type	period $10^7 \times$ m.e. $\pm$	reciprocal period	
109	18 36 53	-30 34'4	H	7664	271	cl	d 4941035	d 55	d <sup>-1</sup> 2'023868
225 <sup>1)</sup>	38 04	-30 36'9	S*)			ecl <sup>1)</sup>			
190	38 37	-28 59'3	S	7664	255	cl	5687622	40	1'758204
76	38 38	-28 15'4	F	7564	218	cl	4948335	46	2'020882
22	39 38	-29 12'9	BLX	7664	301	cl	5613666 <sup>7)</sup>	40	1'78367 <sup>7)</sup>
146	40 18	-25 30'2	LO	7664	35	nova*)			
1	40 52	-27 09'9	A	7664	312	ecl	1'934850		516836
130	40 56	-26 17'2	KY	9165	313	$\delta^*$ )	13'5135 <sup>4) 7)</sup>		0'74000 <sup>7)</sup>
122	40 59	-27 34'1	J	8079	278	ecl	1'885874		530258
123	41 08	-27 57'0	J	7664	246 <sup>3)</sup>	cl	5132710	38	1'948289
110	41 40	-28 53'5	H	7664	238 <sup>3)</sup>	cl	6406812	68	1'560839
62	41 56	-31 13'3	E	7664	254	cl	5859507	75	1'706628
37	42 00	-26 37'2	CI	7664	310	cl	5650292 <sup>4)</sup>		1'769820
111	42 05	-28 21'3	H	7664	213	cl	4947376 <sup>6)</sup>	37	2'021274 <sup>6)</sup>
77	42 20	-27 21'3	F	7664	260	cl	5569029	48	1'795645
155	42 33	-27 48'0	MP	7664	304	cl	5424195	29	1'843592
219	42 37	-30 23'0	Z	7664	307	cl	4784034	37	2'090286
124	42 53	-30 25'4	J	7664	227	cl	5069735	76	1'972490
78	43 19	-27 21'2	F	7664	278 <sup>3)</sup>	cl	4673535	24	2'139708
95	43 41	-31 52'9	GHIN	7664	315 <sup>3)</sup>	cl	5507782	27	1'815613
63	44 18	-30 37'5	EHP	7664	273	cl	4487108	20	2'228607
185	44 43	-28 31'9	QT	7664	264	cl	5456616	42	1'832638
125	45 07	-24 48'4	JY	7664	222 <sup>3)</sup>	cl	4655084	42	2'148189
64	45 21	-28 51'9	E	9165	288	cl	523751 <sup>4) 7)</sup>		1'908852 <sup>7)</sup>
2	46 12	-27 37'9	AU	7664	328	cl	5316468	14 <sup>*)</sup>	1'880948
132	46 18	-26 00'8	K	7664	320	cl	5123466 <sup>*)</sup>	52	1'951804
67	46 52	-28 16'0	E	7664	270	cl	6828666	109	1'464415
53	46 52	-22 56'2 <sup>2)</sup>	DF	7664	301	cl	4548698	24	2'198431
80	46 54	-23 34'0	F	7664	315	cl	3678216	33	2'718709
41	46 55	-29 25'8	CJKW	7664	269 <sup>3)</sup>	cl	4467935 <sup>5)</sup>	29	2'238170 <sup>5)</sup>
23	47 20	-27 50'5	B	7664	285	ecl	7867419	57	1'271065
158	47 50	-29 27'5	M	7664	265	cl	5937966	61	1'684078
4	48 16	-28 23'9	A	7664	283	cl	4874996 <sup>7)</sup>	38	2'051284 <sup>7)</sup>
96	48 36	-29 35'5	GT	8079	340	ecl	2'760319		362277
6	49 22	-30 16'9	A	9165	368	cl	3633023 <sup>7)</sup>	28	2'752529 <sup>7)</sup>
5	49 29	-26 12'1	A	7664	171	cl	3732583 <sup>7)</sup>	39	2'679110 <sup>7)</sup>
182	49 53	-29 31'6	PQ	7664	286	cl	3954471	17	2'528783
3	50 01	-23 18'5	A	7664	280	cl	5241802	32	1'907741
148	50 05	-25 24'0	L	7664	172	cl	4958414	26	2'016774
24	50 16	-27 28'4	B	7664	326	cl	5880825	54	1'700442
7	50 37	-29 25'7	AS	7664	283	cl	5927139	42	1'687155
149	50 49	-25 23'8	L	7664	273	cl	5906682 <sup>4)</sup>		1'692998
159	51 12	-25 12'6	M	6573		irr	80		
202	51 33	-24 56'4	U	7664	313	cl	5726207	37	1'746357
169 <sup>1)</sup>	51 50	-30 01'9	N			ecl <sup>1)</sup>			
175	51 52	-29 01'7	OQ	7664	301	cl	5565165	30	1'796892
69	51 59	-30 03'9	E	7664	279	cl	4741555	53	2'109012
188	52 04	-23 31'8	R	7664	277 <sup>3)</sup>	$\delta$	1'835347 <sup>*)</sup>		544856
82	52 05	-28 24'8	F	7664	255 <sup>3)</sup>	cl	4608651	23	2'169832
189	52 15	-31 53'6	R	9165	319	cl	5666805	65	1'764663
9	52 24	-27 07'6	A	7664	278 <sup>3)</sup>	cl	3247891	61	3'078921
221	52 29	-26 44'9	Z	9165	223	cl	5310747 <sup>*)</sup>	135	1'882974
191	53 12	-23 24'9	S	7664	330	ecl	5287385	23	1'891294
26	53 25	-26 55'2	B	7664	315	cl	5318327	102	1'880291
43	53 53	-27 06'3	CE	7664	296	cl	4600653	27	2'173604
44	54 34	-28 28'8	C	7664	162 <sup>3)</sup>	cl	5149085 <sup>7)</sup>	36	1'942093 <sup>7)</sup>
83	54 35	-31 18'2	F	8079	344	ecl	2'95061		338913

9	10	11	12	13	14	15	16	17	18			
number of epochs	m.e. of single epoch $\pm$	phase of epochs	mean epoch JD 242	m.e. $\pm$	range	max	min	range	value of one step	m.e. of single estimate $\pm$	remarks	No. var.
17	d '024	'10	d 6248'001	d '006	s 2'8	m 13'9	m 14'5	m '6	m '21	s '50	*	109
13	'016	'72	6238'017	'004	4'7	13'5	14'6	1'1	'23	'62	*, I	225 <sup>1)</sup>
17	'017	'72	6329'775	'004	4'4	14'4	15'1	'7	'16	1'00	*	190
14	'018	'07	6408'599	'005	10'7	13'5	14'7	1'2	'11	1'15	7	76
		'815	5928'022		5'4	13'9	15'0	1'1	'20	'96	*	22
		'01 <sup>8)</sup>	6311'07		4'7	13'4	14'6	1'2	'26	'51	*	146
		'195	6385'948		2'6 <sup>11)</sup>	14'3	14'8 <sup>11)</sup>	'5	'19	'40	*, 4, 7, 8	I
12	'013	'90	6215'148	'004	2'4	14'2	15'0	'8	'33	'36 <sup>3)</sup>	*, II	130
24	'032	'90	6203'026	'006	3'3	14'1	14'7	'6	'18	'56 <sup>3)</sup>	3	122
23	'024	'73	6150'563	'005	4'7	13'9	14'7	'8	'17	'79	3	123
		'10	6274'141		5'5	13'8	14'6	'8	'15	'72	4	110
24	'021	'93	6404'358	'004	3'5	14'0	14'8	'8	'23	'47	6	62
22	'024	'78	6309'045	'005	5'4	13'8	14'8	1'0	'19	'83		37
21	'013	'18	6349'125	'003	4'5	13'9	14'7	'8	'18	'62		111
22	'022	'03	6568'022	'005	2'5	13'7	14'2	'5	'20	'46		77
18	'040	'64	6385'677	'009	2'6	14'5	14'9	'4	'15	'48		155
20	'013	'87	6363'427	'003	6'6	13'9	14'9	1'0	'15	'86 <sup>3)</sup>	3	219
19	'012	'34	6282'374	'003	10'0	12'7	14'0	1'3	'13	'76 <sup>3)</sup>	3	124
22	'012	'81	6290'844	'003	7'6	13'6	15'0	1'4	'18	'84		78
16	'020	'05	6378'264	'005	3'7	13'9	14'8	'9	'24	'52		95
22	'018	'41	6244'989	'004	3'8 <sup>11)</sup>	14'1	15'1 <sup>11)</sup>	1'0	'26	'53 <sup>3)</sup>	3, II	185
		'40	5993'338		6'7	14'0	14'8	'8	'11	1'03	*, 4, 7	125
22	'0038 <sup>*</sup> )	'740 <sup>10)</sup>	6081'370	'001 <sup>*</sup> )	7'6	13'5	14'5	1'0	'13	'68	*, 10	64
21	'020	'36	6088'405	'004	5'0	13'4	14'4	1'0	'20	'73	*	2
22	'040	'09	6438'810	'008	3'1	14'6	15'0	'4	'13	'71		132
23	'016	'53	6268'808	'003	7'6	13'0	14'4	1'4	'18	'61	2	67
22	'026	'45	6282'185	'005	3'4	14'0	14'5	'5	'15	1'34	*	53
25	'020	'37	6574'736	'004	5'1	14'2	14'9	'7	'14	'98 <sup>3)</sup>	3, 5	80
30	'029	'66 <sup>8)</sup>	6523'396	'005	3'2	14'3	14'9	'6	'19	1'00	*, 8	41
24	'026	'06	6269'358	'005	3'5	14'3	14'9	'6	'17	'43		23
17	'022	'20	6405'355	'005	5'9	14'3	15'0	'7	'12	1'18	7	158
		'267	6382'594		6'1	13'5	14'5	1'0	'16	'42	*	4
18	'029	'27 <sup>10)</sup>	6326'280	'007	4'5	13'6	14'3	'7	'16	1'27	*, 7, 10	96
22	'021	'43	6199'979	'004	9'2	13'8	14'9	1'1	'12	1'18	*, 7	6
23	'012	'17	6261'577	'003	4'1	13'7	14'7	1'0	'24	'49		5
22	'015	'46	6336'007	'003	8'7	14'0	15'0	1'0	'11	1'23		182
19	'014	'97	6377'001	'003	4'2	14'2	14'7	'5	'12	'59	*	3
20	'031	'03	6353'075	'007	7'9	13'5	14'2	'7	'09	1'25		148
21	'018	'39	6387'903	'004	7'8	14'2	15'1	'9	'12	1'38		24
		'03	6178'407		3'9	14'2	14'9	'7	'18	'59	4	7
		'62	6414'856	'005	4'8	13'9	14'5	'6			*	149
20	'020					13'3	14'2	'9	'19	'47		159
		'82	6277'962	'003	4'2	13'6	14'6	1'0	'24	'47	I	202
20	'015	'71	6293'332	'004	3'8	14'6	15'1	'5	'13	'75	*	169 <sup>1)</sup>
19	'020	'83 <sup>9)</sup>	6170'126		1'8	14'2	14'6	'4	'22	'43 <sup>3)</sup>	*, 3, 9	175
24	'014	'44	6273'043	'003	5'0	14'1	15'0	'9	'18	'56 <sup>3)</sup>	3	69
20	'020	'01	6390'466	'004	3'1	13'7	14'5	'8	'26	'57	*	188
17	'035	'24	6197'055	'008	6'8	14'5	15'1	'6	'09	1'75 <sup>3)</sup>	3	82
14	'033	'36	6007'189	'009	2'8	14'2	14'6	'4	'14	'59	*	189
13	'012	'37 <sup>8)</sup>	6498'926	'003	2'5	13'4	13'7	'3	'12	'53	*, 8	9
17	'025	'84	6118'132	'006	6'5	14'1	14'6	'5	'08	1'26		221
18	'012	'37	6209'675	'003	7'0	14'0	14'9	'9	'13	'74		191
19	'021	'41	6549'337	'005	3'9 <sup>12)</sup>	14'7	15'1 <sup>12)</sup>	'4	'10	'76 <sup>3)</sup>	3, 7, 12	26
		'461	6289'110		8'5	12'9	13'6	'7	'08	'62	*	43
												44
												83

TABLE 10

1	2	3	4	5	6	7	8	
No. var.	$\alpha$ (1875)	$\delta$ (1875)	discovered on blink-pairs	plates estimated until JD 242	number of estimates used	type	period $10^7 \times$ m.e. $\pm$	reciprocal period
70	h m s	° ' "	E	9165	306	ecl *)	d	d <sup>-1</sup>
45	18 55 01	-29 19'1	CQ	9165	326	cl	2'323802	370
85	57 02	-30 35'1	FI	7664	296	cl	'5196097	74
140	57 38	-24 52'7	K	7664	204 <sup>3)</sup>	cl	'5417421	59
141	18 58 50	-30 07'8	K	7664	329	cl	'5049887	20
86	19 00 03	-25 13'8	F	6573		irr *)	'7275312	105
88	00 14	-29 22'8	F	7664	214 <sup>3)</sup>	cl	47'50	
89	01 44	-27 28'2	FJ	7664	318	cl	'5563080	31
13	01 50	-29 39'9	ANP	7664	294	cl	'5325551	34
152	01 52	-29 51'8	L	8079	332	ecl	'6900761	98
100	01 58	-24 36'1	GH	9165	353	cl	'1072605	
108 <sup>1)</sup>	02 05	-26 25'7	T			ecl <sup>1)</sup>	'4659810 <sup>4)</sup>	
183	02 47	-28 48'8	PS	7664	288	cl	'5499901	54
153	02 55	-29 30'0	LMV	7664	316	cl	'5698756	42
177 <sup>1)</sup>	03 20	-23 40'1	O			ecl <sup>1)</sup>		
103	03 31	-25 05'1	GL	7664	202	cl	'5216189	30
178	03 50	-25 58'9	O	7664	209	cl	'5156861	28
57	03 58	-27 03'1	D	7664	264	cl	'4632683	78
14	04 26	-23 52'9	A	7664	325	cl *)	'11359409 <sup>7)</sup>	3'9
72	04 31	-27 07'4	E	7664	336	cl	'5973433	59
90	05 06	-28 43'9	F	7664	222 <sup>3)</sup>	cl	'4053721	68
207	05 18	-28 43'3	W	7664	233	cl	'5812065 <sup>4)</sup>	
114	05 29	-28 25'7	H	7664	298	cl	'5483316	48
48	06 11	-25 39'5	CF	7664	279	cl	'4496928	18
208	06 20	-31 04'9	W	7664	329	cl	'6950270 <sup>4)</sup>	
199	06 30	-31 19'6	T	7664	308	cl	'3735253	37
142	06 40	-25 14'2	K	7664	276	cl	'5778171	72
143	06 52	-23 54'8	KU	7664	265	cl	'5433479	67
144	07 06	-25 10'1	KR	7664	295	cl	'4950461	27
126	07 13	-28 30'8	J	7664	254	cl	'4677259	31
91	07 58	-29 01'2	FO	7664	324	cl	'6189279	52
179	08 00	-26 00'4	O	7664	301	cl	'6212388 <sup>5)</sup>	54
17	08 57	-23 23'6	A	9165	268	cl	'4906629	56
222	09 02	-22 58'0 <sup>2)</sup>	Z	7664	303	cl	'6289945	74
31	09 22	-26 27'6	BF	6573		irr		
116	09 41	-31 26'2	H	7664	244 <sup>3)</sup>	cl	'5373436	134
58	09 58	-30 30'1	DY	7664	326	ecl	'128246	
162	10 06	-24 12'0	M	8079	326	ecl	'406848	
163	10 47	-28 34'6	MV	9165	312	cl	'5168271 <sup>4)</sup> 7)	
18	10 57	-25 32'4	ABDH	7664	226	cl	'6484454	69
193	10 57	-23 45'1	SW	7664	319	cl	'7005342 <sup>4)</sup>	
216	11 02	-24 35'8	Y	7664	308	cl	'6934490	94
32	11 26	-31 35'4	BJ	7664	269 <sup>3)</sup>	cl	'5319859 <sup>7)</sup>	28
74	11 29	-29 21'9	EH	9165	343	cl	'4757204	
19	11 35	-31 46'3	AI	7664	313	cl	'5331563	20
33	12 09	-30 59'9	BDU	7664	256	cl	'5207599	31
59	12 09	-30 36'3	D	7664	298 <sup>3)</sup>	cl	'5865256	129
172 <sup>1)</sup>	12 09	-30 01'0	NP			ecl <sup>1)</sup>		
127	12 23	-24 41'4	JMS	7664	305	cl	'7062917	34
21	12 33	-26 12'8	AO	7664	271	cl	'4916242 <sup>4)</sup>	
205	12 49	-26 38'4	V	7664	275	cl	'4537734 <sup>4)</sup>	
20	12 57	-31 16'6	A	7664	340	ecl	'1091379	
218	13 15	-28 34'7	Y	7664	218	cl	'5329164	36
165	13 45	-30 06'6	M	7664	286	cl	'5502785	36
118	14 18	-27 48'3	HIMV	7664	304	cl	'4589226	16
36	15 11	-28 43'6	BDEP	6573		irr		
75	15 45	-28 24'6	ESUYZ	7664	319	cl	'4092612 <sup>7)</sup>	27

1943BAN.....9...337F

(continued).

9	10	11	12	13	14	15	16	17	18			
number of epochs	m.e. of single epoch ±	phase of epochs	mean epoch JD 242	m.e. ±	range	max	min	range	value of one step	m.e. of single estimate ±	remarks	No. var.
18	d '037	'011	d 6051'210	d '009	s 2'1	m 9'47	m 10'02	m '55	m '26	s '29	*	70
21	'018	'79	6344'332	'004	9'0	13'3	14'5	1'2	'13	1'10	*	45
17	'023	'44	6444'279	'006	4'2	14'1	14'8	'7	'17	'71	*	85
22	'019	'68	6225'346	'004	2'2	14'5	15'0	'5	'23	'44 <sup>3)</sup>	*, 3	140
20	'040	'60	6292'856	'009	2'4	13'7	14'2	'5	'21	'48	*	141
					6'9	13'5	14'5	1'0	'14		*	86
19	'015	'72	6374'029	'003	2'4	14'8	15'1	'3	'12	'72 <sup>3)</sup>	*, 3	88
26	'018	'63	6279'697	'003	5'6	13'9	14'8	'9	'16	'77		89
20	'035	'30	6255'772	'008	7'7	14'0	14'8	'8	'10	'84		13
		'596	6414'815		4'9	13'9	14'8	'9	'18	'54	*	152
		'01	6320'110		4'8	13'8	14'5	'7	'15	'66	*, 4	100
22	'028	'44	6508'289	'006	2'2	14'3	14'8	'5	'23	'39	I	198 <sup>1)</sup>
18	'020	'79	6360'273	'005	5'3	14'0	14'9	'9	'17	'66		183
											I	153
15	'014	'30	6478'137	'004	2'4 <sup>11)</sup>	14'3	14'9 <sup>11)</sup>	'6	'25	'54	II	177 <sup>1)</sup>
18	'016	'75	6436'157	'004	2'3	14'3	15'0	'7	'30	'43		103
16	'020	'30	6063'870	'005	5'7	14'0	14'9	'9	'16	'82		178
16	'008	'38	6268'732	'002	2'5	13'9	14'5	'6	'24	'63	*, 7	57
20	'030	'42	6356'592	'007	5'1	13'0	13'4	'4	'08	'62		14
20	'034	'63	6129'819	'008	2'2	14'6	15'1	'5	'23	'83 <sup>3)</sup>	3	72
		'53	6265'714		2'5	14'5	15'1	'6	'24	'69	4	90
26	'031	'19	6352'540	'006	2'7	14'2	14'5	'3	'11	'37		207
22	'012	'60	6277'984	'002	10'2	13'3	14'6	1'3	'13	'82		114
		'74	6421'861		5'0	13'1	13'7	'6	'12	'47	4	48
17	'024	'18	6244'291	'006	2'7	13'8	14'3	'5	'19	'54		208
19	'030	'69	6465'018	'007	2'4	14'1	14'9	'8	'33	'48		199
19	'028	'24	6116'595	'006	2'9	14'1	14'8	'7	'24	'73		142
21	'017	'99	6544'014	'004	4'7	13'6	14'5	'9	'19	'48		143
23	'020	'13	6462'639	'004	3'0	14'7	15'1	'4	'13	'58		144
18	'025	'11	6433'206	'006	5'5	14'0	14'8	'8	'15	'72		126
18	'028	'59	6492'309	'007	2'6	14'2	14'7	'5	'19	'48	5	91
26	'038	'27	6464'122	'008	3'6	14'3	14'9	'6	'17	1'10	*	179
20	'030	'30	6149'876	'007	4'3	13'9	14'4	'5	'12	'65	2	17
					7'0	13'6	14'7	1'1	'16		*	222
12	'024	'70	6093'319	'007	2'4	14'2	14'9	'7	'29	'46 <sup>3)</sup>	3	31
		'87	6140'263		5'0	14'0	14'9	'9	'18	'96	*	116
		'778	6716'158		4'5 <sup>11)</sup>	14'0	15'1 <sup>11)</sup>	1'1	'24	'50	*, II	58
		'97	6322'330		6'6	13'7	14'9	1'2	'18	'60	*, 4, 7	162
16	'018	'50	6168'999	'005	7'0 <sup>11)</sup>	14'0	15'0 <sup>11)</sup>	1'0	'14	'93	II	163
		'51	6086'599		4'1	13'3	14'3	1'0	'24	'58	4	18
19	'027	'81	6070'337	'006	4'0	13'9	14'6	'7	'18	'50		193
16	'017	'94	6348'700	'004	4'9	14'2	14'8	'6	'12	'76 <sup>3)</sup>	3, 7	216
		'15	6591'178		7'1	13'9	14'8	'9	'13	1'00	*	32
20	'011	'24	6431'059	'002	10'2	13'7	14'6	'9	'09	1'12		74
19	'018	'11	6429'363	'004	5'3	14'2	15'1	'9	'17	'72		19
19	'042	'48	6102'498	'010	3'9	14'3	14'7	'4	'10	'89 <sup>3)</sup>	3	33
											I	59
18	'014	'65	6215'113	'003	3'3	14'0	14'7	'7	'21	'47		172 <sup>1)</sup>
		'74	6271'522		6'0	14'3	14'9	'6	'10	'85	4	127
		'45	6558'596		7'3	13'5	14'6	1'1	'15	'59	4	21
		'691	5962'957		9'6	12'6	13'5	'9	'09	'85	*	205
23	'018	'47	6249'768	'004	3'5	14'3	15'1	'8	'23	'53		20
25	'016	'22	6187'456	'003	4'1	13'7	14'6	'9	'22	'57		218
22	'011	'57	6557'812	'002	5'8	13'3	14'7	1'4	'24	'45		165
					8'2	13'3	14'7	1'4	'17		*	118
18	'012	'20	6281'023	'003	10'8	13'4	14'6	1'2	'11	'94	7	36
												75



TABLE 10

1	2		3	4	5	6	7		8
No. var.	$\alpha$ (1875)	$\delta$ (1875)	discovered on blink-pairs	plates estimated until JD 242	number of estimates used	type	period	$10^7 \times \text{m.e.} \pm$	reciprocal period
180 <sup>1)</sup>	<sup>h m s</sup> 19 15 50	<sup>° ' "</sup> -25 01'4	OS			ecl <sup>1)</sup>	<sup>d</sup>		<sup>d<sup>-1</sup></sup>
50	17 09	-28 41'9	CFQ	7664	262	cl	5954149	<sup>d</sup> 58	1'679501
105	17 23	-29 23'5	G	7664	279 <sup>3)</sup>	ecl <sup>*</sup> )	2939444	19	3'402004
61	17 59	-27 37'5	D	8079	347	ecl	6730857		1'485695
121	20 17	-28 57'6	IPSV	7664	290	cl	5281197	22	1'893510
128	20 39	-26 59'8	J	7664	262	cl	6547620	69	1'527272
194	20 49	-25 01'9	S	7664	286	cl	6630479 <sup>5)</sup>	94	1'508187 <sup>5)</sup>

## EXPLANATION OF THE COLUMNS:

- Column 6 cl = cluster variable,  $\delta$  =  $\delta$  Cep variable, ecl = eclipsing variable, irr = irregular.
- Column 7 The value of the period definitively adopted. When a mean error has been given the period has been determined with least squares, otherwise graphically.
- Column 8 Always the reciprocal value of column 7. As a rule the value used for computing phases. In exceptional cases remark 7 indicates which value has been used instead.
- Columns 9 and 10 Relate to the determination of the period with least squares.
- Column 11 Phase of the epoch corresponding to the value of column 8. The epoch is normally the maximum for cluster variables and the minimum for eclipsing variables. Otherwise reference has been made to a remark.
- Column 12 The mean epoch is the epoch which differs least from the mean value of the epochs used for the least squares solution or for the graphical correction.

(continued).

9	10	11	12	13	14	15	16	17	18			
number of epochs	m.e. of single epoch ±	phase of epochs	mean epoch JD 242	m.e. ±	range	max	min	range	value of one step	m.e. of single estimate ±	remarks	No. var.
22	d .022	.24	d 6305.587	d .005	s 6.7	m 13.6	m 14.9	m 1.3	m .19	s .79	1	180 <sup>1)</sup>
28	.015	.30 <sup>8)</sup>	6273.155	.003	1.3	13.8	14.3	.3	.23	.49 <sup>3)</sup>	*, 3, 8	50
16	.011	.280	6369.598	.003	4.4	13.0	13.4	.4	.09	.57	*	105
19	.019	.17	6457.932	.004	5.1	13.6	14.6	1.0	.20	.69		61
12	.024	.86	6088.544	.007	2.7	14.3	14.9	.6	.22	.59		121
		.78	6378.386		5.6	13.3	14.5	1.2	.22	.53	5	128
												194

REMARKS.

- \*) Refers to section 13: discussion of individual variables.
- 1) The examination of this eclipsing variable has not been finished as yet (nos. 169, 172, 177, 180, 198 and 225).
- 2) This star being situated north of  $\delta = -23^\circ$  the co-ordinates are given for 1855 conventionally (nos. 53 and 222).
- 3) Uncertain estimates omitted (nos. 9, 32, 41, 44, 59, 78, 82, 88, 90, 95, 105, 110, 116, 123, 125, 140 and 188).
- 4) The period of this star has been obtained by graphical correction of the least squares solution:

var. no.	period from least squares solution	$10^7 \times$ m.e. ±
21	.4916146	38
37	.5650360	60
64	.5238648	134
100	.4659780	31
130	13.5221	26000
149	.5906966	76
163	.5168375	37
193	.7005146	96
205	.4537767	18
207	.5811861	64
208	.6950375	35

- 5) The following reciprocal period may satisfy the observations somewhat better: no. 41:  $2^{d-1}.238151$ ; no. 179:  $1^{d-1}.609659$ ; no. 194:  $1^{d-1}.508169$ .
- 6) The period of this star may be slightly too large (no. 111).
- 7) Phases and light curves of this star have been computed with the provisional reciprocal period:

var. no.	provisional reciprocal period $d^{-1}$
4	2.051314
5	2.67910
6	2.74255
14	8.803283
22	1.781368
32	1.879746
44	1.942096
64	1.908890
75	2.443431
130	.07397
163	1.934844

- 8) The epochs of this star are the minima (nos. 23, 105, 130, 191).
- 9) The epochs of this star are the maxima (no. 188).
- 10) For this star the epochs where the rising branch passes a certain brightness have been used (no. 2: 8<sup>s</sup>.5; no. 6: 5<sup>s</sup>.5).
- 11) This star is invisible on many plates of minimum (nos. 18, 103, 122, 125, 162). Thus the minima may be systematically in error.
- 12) This star is very faint in minimum (no. 44).

TABLE II.

2		d	t	d	d	t	d	17	d	t	d		
from maxima		2425825.248	229	+ 017	2426212.464	700	- 006		2425713.549	385	- 002		
		58.350	297	- 31	14.237	703	- 11		39.587	438	- 20		
		61.286	303	- 20	40.313	747	- 15	2425706.566	0	+ 028	5823.208	608	+ 27
		62.265	305	- 16	43.290	752	- 1	65.430	120	+ 13	.609	438	+ 2
		64.280	309	+ 49	46.235	757	- 20	90.412	171	- 29	.230	608	+ 49
		6067.540	726	+ 22	6570.489	1304	+ 20	5849.413	291	+ 92	6093.549	1158	- 20
		6103.590	800	- 3	72.277	1307	+ 29	50.315	293	+ 13	94.542	1160	- 10
		23.593	841	+ 13	73.402	1309	- 31	6122.570	848	- 50	6120.589	1213	- 19
		29.444	853	+ 14	.466	1309	+ 33	.592	848	- 28	22.592	1217	+ 18
		6212.287	1023	- 18	7656.303	3136	- 18	53.542	911	+ 10	29.444	1231	- 13
		35.228	1070	+ 10	.317	3136	- 4	6212.442	1031	+ 31	.466	1231	+ 9
		6475.540	1563	- 15	.332	3136	+ 11	13.409	1033	+ 16	53.542	1280	- 4
		7623.592	3918	- 25	9			18.345	1043	+ 46	6475.540	1935	- 14
		.614	3918	- 3				40.335	1088	- 44	80.493	1945	+ 29
		56.281	3985	+ 2				41.313	1090	- 47	6540.430	2067	- 17
		.303	3985	+ 24				42.285	1092	- 57	70.425	2128	- 10
		5						.307	1092	- 35	.446	2128	+ 11
		2425740.530	0	+ 032				43.290	1094	- 33	73.380	2134	- 5
		75.566	94	- 18				46.235	1100	- 32	7623.485	4270	+ 11
		95.351	147	- 16				.257	1100	- 10	64.272	4353	- 6
		98.328	155	- 25				48.254	1104	+ 24	22		
		99.486	158	+ 14				73.313	1155	+ 60			
		5822.257	219	+ 16				7623.571	3907	+ 13			
		32.356	246	+ 37				.592	3907	+ 34	2425740.552	0	- 021
		85.298	388	- 24				35.361	3931	+ 27	5825.362	151	+ 22
		6156.300	1114	- 7				36.311	3933	- 4	26.454	153	- 8
		.346	1114	+ 39				58.366	3978	- 29	.479	153	+ 17
		61.516	1128	- 17				64.272	3990	- 11	61.264	215	- 3
		76.444	1168	- 19				18			6213.237	842	- 7
		6212.287	1264	- 9				2425745.548	0	- 016	14.353	844	- 14
		13.409	1267	- 7				.570	0	+ 6	41.291	892	- 21
		37.294	1331	- 10				92.247	72	- 5	63.213	931	+ 8
		40.313	1339	+ 23				5821.434	117	+ 2	68.284	940	+ 26
		6476.576	1972	+ 13				36.339	140	- 7	6570.275	1478	+ 2
		6567.267	2215	+ 2				6090.550	532	+ 13	7622.248	3352	- 26
		69.493	2221	- 11				6123.593	583	- 15	23.399	3354	+ 2
		70.232	2223	- 19				25.577	586	+ 24	.420	3354	+ 23
		73.230	2231	- 7				29.444	592	0	23		
		7636.289	5079	+ 12				60.546	640	- 23			
		6						6218.323	729	+ 42			
		2425716.607	0	- 052				40.313	763	- 15			
		19.524	8	- 41				.335	763	+ 7			
		92.643	209	+ 54				42.285	766	+ 12			
		99.478	228	- 14				6570.360	1272	- 27			
		5823.467	294	- 3				7623.463	2896	+ 1			
		34.371	324	+ 2				19					
		.348	324	+ 21				2425526.288	0	- 005			
		50.371	368	+ 17				5707.550	340	- 16			
		58.358	390	+ 11				16.636	357	+ 6			
		61.277	398	+ 24				5826.479	563	+ 19			
		.257	398	+ 4				64.302	634	- 12			
		6123.540	1120	- 17				6123.413	1120	- 15			
		6266.366	1513	+ 31				25.555	1124	- 6			
		6569.292	2347	- 37				6212.464	1287	- 1			
		70.454	2350	+ 35				18.323	1298	- 7			
		7623.288	5248	+ 19				.345	1298	+ 15			
		56.334	5339	+ 4				65.247	1386	- 1			
		9109.521	9339	- 18				66.327	1388	+ 13			
		7						6486.520	1801	+ 12			
		2425797.556	0	- 004				6570.210	1958	- 3			
		5850.315	89	- 7				71.280	1960	0			
		.337	89	+ 15				73.423	1964	+ 11			
		63.349	111	- 13				7367.288	3453	+ 6			
		85.298	148	+ 6				7657.322	3997	+ 3			
		6094.520	501	0				58.366	3999	- 19			
		.542	501	+ 22				64.251	4010	+ 1			
		6129.466	560	- 24				21					
		6161.516	614	+ 19				2425524.263	0	- 016			
		4						.571	18463	+ 2	2425745.548	0	- 038
2425713.571		0	- 022										

TABLE II (continued).

d	t	d	d	t	d	d	t	d	d	t	d	d	t	d			
2425745'570	o	-016	2426570'468	1498	-022	2426559'313	1669	-012	2427623'463	4176	-002	59					
75'566	51	-12	489	1498	-1	335	1669	+10	56'281	4249	-12	d	t	d			
91'485	78	+28	7623'442	3520	+24	69'450	1691	+4	303	4249	+10	2425524'285	o	+101			
93'251	81	+30	485	3520	+19	70'360	1693	-6	50			5719'518	333	+21			
273	81	+52	56'281	3583	+7	71'280	1695	-6				92'247	457	+20			
97'329	88	-8	58'366	3587	+9	7623'463	3982	+7				94'542	461	-31			
5808'472	107	-39	37			44			2425706'588			o	-012	564	461	-9	
21'434	129	-15	2425442'428	o	-004	2425745'570	o	+005	40'530	57	-8	5803'351	476	-20			
30'272	144	+2	5739'609	526	-32	97'556	101	-15	99'508	156	+24	30'272	522	-79			
34'377	151	-10	64'494	570	-9	5822'257	149	-30	5857'217	253	-23	50'337	556	+44			
6122'570	641	+23	516	570	+13	25'362	155	-14	60'214	258	-3	6123'593	1022	-21			
592	641	+45	97'308	628	+33	93'364	287	+20	6161'516	764	+19	615	1022	+1			
6476'576	1243	+3	5823'230	674	-37	6125'555	738	-13	6210'293	846	-28	29'466	1032	-13			
86'520	1260	-50	25'509	678	-18	577	738	+9	13'320	851	+22	53'474	1073	-52			
7623'292	3193	-41	63'371	745	-13	6235'228	951	-15	342	851	+44	60'524	1085	-41			
313	3193	-20	88'277	789	+31	250	951	+7	14'472	853	-17	6213'342	1175	-10			
334	3193	+1	6125'555	1209	-6	6303'249	1038	+38	19'260	861	+8	40'313	1221	-19			
356	3193	+23	6210'293	1359	-23	6486'520	1439	+1	38'298	893	-7	43'290	1226	+25			
377	3193	+44	14'237	1366	-35	6569'407	1600	-12	41'291	898	+9	311	1226	+46			
26			36'290	1405	-18	428	1600	+9	63'279	935	-34	6573'488	1789	+9			
2425527'273	o	+007	40'246	1412	-17	450	1600	+31	301	935	-12	7636'289	3601	+26			
5794'263	502	+17	48'254	1426	+80	7623'399	3647	-38	66'305	940	+15	62					
5860'236	626	+43	66'305	1458	+50	420	3647	-17	327	940	+37	2425707'572			o	-012	
61'242	628	-15	6570'232	1996	-12	442	3647	+5	6569'321	1449	-35	75'566	116	+12			
264	628	+7	253	1996	+9	463	3647	+26	364	1449	+8	94'263	148	-41			
63'349	632	-35	7623'442	3860	-29	36'311	3672	+2	58'366	3278	-4	286	148	-18			
371	632	-13	463	3860	-8	45			53			5832'377	213	-14			
6067'562	1016	-46	485	3860	+14	2425716'615			o	-028	35'392	218	+71				
6125'577	1125	o	41			636	o	-7	2425478'239	o	-005	49'391	242	+7			
6211'216	1286	+14	2425740'552	o	-021	5893'341	340	+30	5706'588	502	-1	93'341	317	+11			
13'342	1290	+12	5821'434	181	-9	6125'577	787	+1	16'615	524	+19	6123'593	710	-16			
37'294	1335	+32	86'220	326	-8	6235'228	998	+14	92'587	691	+28	615	710	+6			
38'298	1337	-28	242	326	+14	36'290	1000	+37	97'556	702	-7	29'444	720	-24			
6570'210	1961	+20	6118'567	846	+7	37'272	1002	-20	5808'472	726	-8	466	720	-2			
71'259	1963	+6	589	846	+29	294	1002	+2	23'499	759	+8	53'474	761	-18			
73'359	1967	-22	22'570	855	-12	63'279	1052	+6	30'294	774	-20	60'524	773	+1			
380	1967	-1	592	855	+10	6303'271	1129	-12	49'413	816	-5	546	773	+23			
32			60'524	940	-35	6561'526	1626	-3	50'315	818	-13	6237'272	904	-11			
2425526'265	o	+015	6213'320	1058	+39	67'267	1637	+23	337	818	+9	40'224	909	+11			
27'205	2	-19	38'298	1114	-3	68'271	1639	-13	6160'546	1500	-3	246	909	+33			
5706'588	339	-5	42'307	1123	-15	292	1639	+8	76'466	1535	-4	68'306	957	-32			
92'247	500	+4	6475'540	1645	-8	69'299	1641	-24	6214'215	1618	-9	6569'514	1471	-3			
97'556	510	-6	6569'364	1855	-11	321	1641	-2	18'323	1627	+5	71'280	1474	+5			
5825'225	562	-1	385	1855	+10	342	1641	+19	19'260	1629	+33	7635'361	3290	o			
35'370	581	+37	70'253	1857	-16	70'339	1643	-23	6569'471	2399	-6	63					
6122'592	1121	-14	73'402	1864	+6	360	1643	-2	70'382	2401	-5	2425764'494			o	-012	
53'474	1179	+13	423	1864	+27	73'466	1649	-14	71'280	2403	-17	516	o	+10			
6218'345	1301	-18	7623'356	4214	-5	488	1649	+8	7367'288	4153	-31	94'564	67	-6			
42'285	1346	-18	377	4214	+16	48			310	4153	-9	99'486	78	-20			
307	1346	+4	399	4214	+38	2425745'548	o	+000	7623'420	4716	+9	508	78	+2			
7623'356	3942	+18	36'289	4243	-29	570	o	+22	442	4716	+31	5808'472	98	-8			
56'303	4004	-18	311	4243	-7	57			2425478'288			o	-011	32'258	151	-3	
317	4004	-4	57'294	4290	-23	99'508	120	-3	5739'587	564	+4	35'392	158	-10			
332	4004	+11	43			5826'479	180	-14	40'530	566	+21	6122'570	798	-7			
33			2425790'390	o	-001	49'413	231	-14	552	566	+43	592	798	+15			
92'456	4	-18	2425791'462	o	-014	50'315	233	-12	45'570	577	-35	53'542	867	+4			
5823'208	63	+9	485	o	+9	337	233	+10	5888'277	885	-15	6239'267	1058	+25			
50'315	115	+36	94'263	6	+27	60'214	255	-6	6067'562	1272	-15	42'397	1065	+14			
51'311	117	-5	5823'208	69	-12	6067'540	716	+12	6123'615	1393	-17	43'290	1067	+9			
86'220	184	+9	230	69	+10	8120'589	834	-3	76'444	1507	-1	48'232	1078	+16			
88'277	188	-17	25'509	74	-11	25'555	845	+16	466	1507	+21	65'269	1116	+2			
6094'542	584	+27	26'454	76	+13	61'516	925	+2	6241'291	1647	-11	6569'493	1794	o			
6212'222	810	+15	86'242	206	-7	6213'216	1040	-13	42'221	1649	-8	70'382	1796	-9			
13'237	812	-11	6153'542	787	-5	237	1040	+8	48'254	1662	+3	71'280	1798	-8			
39'267	862	-19	56'300	793	-7	35'250	1089	-14	73'291	1716	+23	7367'288	3572	-13			
40'313	864	-15	6212'442	915	+7	6570'296	1834	+11	6570'210	2357	-13	7623'506	4143	-8			
6569'450	1496	+2	37'272	969	-7	73'423	1841	-10	232	2357	+9	7656'281	4216	+10			
						445	1841	+12									









TABLE II (continued).

175	d	t	d	d	t	d	d	t	d	205				
2425478'239	63'371	394	+ 9	2426212'353	275	- 021	2425797'308	111	+ 011	d	t	d		
5825'509	5926'239	553	+ 1	14'237	276	+ 28	5823'208	148	- 8	2425764'494	o	+ 007		
53'328	61'55'594	1133	- 3	65'68'292	469	- 183	25'362	151	+ 45	92'609	62	- 12		
57'217	61'538	1148	+ 9	70'253	470	- 58	53'328	191	- 10	5834'377	154	+ 8		
239	6213'342	1279	+ 30	275	470	- 36	86'220	238	- 42	60'236	211	+ 2		
58'328	19'260	1294	- 4	339	470	+ 28	242	238	- 20	93'364	284	+ 4		
350	40'224	1347	+ 1	360	470	+ 49	6161'538	631	- 26	6161'538	875	- 4		
62'265	66'305	1413	- 17	403	470	+ 92	6211'284	702	- 17	6212'353	987	- 12		
63'349	68'284	1418	- 16	72'255	471	+ 109	13'409	705	+ 7	375	987	+ 10		
371	6475'509	1942	- 5	189			18'345	712	+ 39	42'307	1053	- 7		
6160'546	6570'403	2182	- 18	2425745'570	o	- 014	39'289	742	- 32	48'232	1066	+ 19		
6213'409	425'2182	+ 4	5803'372	102	- 14	6569'234	1213	- 30	70'312	1776	- 13			
46'235	7623'506	4845	+ 9	08'472	111	- 14	65'269	779	+ 28	6540'430	1710	- 16		
66'305	35'361	4875	+ 1	6093'549	614	+ 23	277	1213	+ 13	70'382	1776	- 13		
6475'540	56'317	4928	- 2	6161'538	734	+ 10	7623'549	2718	+ 11	403	1776	+ 8		
6573'466	183			6235'228	864	+ 32	194			7622'248	4094	- 1		
488	2425713'549	o	- 004	36'312	866	- 18	199			23'592	4097	- 19		
7623'592	5825'225	203	+ 24	44'249	880	- 14	2425716'636	o	- 028	614	4097	+ 3		
614	49'413	247	+ 13	65'247	917	+ 17	92'268	114	+ 16	635	4097	+ 24		
635	86'220	314	- 29	6476'576	1290	- 26	97'535	122	- 21	207				
178			242	6559'313	1436	- 24	6156'300	663	+ 35	2425478'239	o	+ 012		
2425775'566	6120'589	740	+ 44	335	1436	- 2	6211'284	746	- 14	288	o	+ 61		
5821'456	56'300	805	+ 5	67'267	1450	- 4	6480'493	1152	- 2	5524'294	81	- 9		
35'370	6211'284	905	- 10	288	1450	+ 17	515	1152	+ 20	5745'570	460	- 3		
50'337	12'375	907	- 19	69'536	1454	- 6	6569'321	1286	- 23	91'462	539	- 24		
52'418	13'476	909	- 18	71'259	1457	+ 21	342	1286	- 2	485	539	- 1		
64'280	6565'472	1549	- 15	280	1457	+ 42	385	1286	+ 41	5823'477	594	+ 25		
6088'562	68'271	1554	+ 34	73'488	1461	- 16	73'316	1292	- 6	25'248	597	+ 53		
6120'567	69'299	1556	- 38	509	1461	+ 5	7622'248	2874	- 16	86'220	702	o		
53'542	321	1556	- 16	7623'549	3314	- 14	199			242	702	+ 22		
55'616	385	1556	+ 48	190			2425713'549	o	+ 037	6213'409	1265	- 18		
6240'224	70'403	1558	- 34	2425790'412	o	+ 011	19'518	16	+ 29	48'254	1325	- 45		
6480'493	425	1558	- 12	98'328	14	- 35	20'584	19	- 25	66'305	1356	- 10		
515	468	1558	+ 31	99'486	16	- 15	97'556	225	o	73'291	1368	+ 1		
6570'210	7623'659	3473	- 9	508	16	+ 7	98'328	227	+ 25	313	1368	+ 23		
7622'227	36'289	3496	- 29	5822'257	56	+ 6	5821'434	289	- 27	6475'509	1716	- 33		
248	311	3496	- 7	93'341	181	- 6	456	289	- 5	6561'504	1864	- 54		
56'281	58'366	3536	+ 48	6093'549	533	- 2	6094'520	1020	+ 12	526	1864	- 32		
57'294	185			6118'589	577	+ 12	6219'260	1354	- 6	70'232	1879	- 44		
179			2425524'285	6212'442	742	+ 20	41'291	1413	- 13	253	1879	- 23		
2425442'385	5706'566	334	+ 11	48'254	805	o	313	1413	+ 9	7622'248	3689	+ 25		
5745'548	5825'509	552	- 1	73'291	849	+ 11	42'397	1416	- 27	23'399	3691	+ 14		
570	64'280	623	+ 28	7623'506	3223	- 15	6303'271	1579	- 38	442	3691	+ 57		
5803'351	88'277	667	+ 16	528	3223	+ 7	6569'256	2291	- 3	208				
21'434	6094'520	1045	- 1	191			277	2291	+ 18	2425525'272	o	- 004		
52'418	6129'444	1109	+ 1	2425791'485	o	+ 011	7623'377	5113	+ 30	5720'584	281	+ 2		
6088'540	6213'476	1263	+ 1	98'328	13	- 19	57'322	5204	- 16	91'462	383	- 14		
6160'546	41'291	1314	- 13	5823'208	60	+ 10	202			485	383	+ 9		
6213'409	313	1314	+ 9	35'370	83	+ 11	2425527'273	o	- 021	94'263	387	+ 7		
41'313	42'397	1316	+ 2	60'214	130	+ 4	295	o	+ 1	286	387	+ 30		
6480'493	6569'234	1915	- 12	63'371	136	- 11	5792'434	463	+ 16	5885'298	518	- 8		
6569'342	70'296	1917	- 42	6156'300	690	- 3	97'556	472	- 15	6094'520	819	+ 8		
364	7623'506	3847	+ 42	6212'353	796	+ 4	5808'472	491	+ 21	6160'524	914	- 16		
7623'549	57'294	3909	- 2	6573'466	1479	- 12	35'370	538	+ 6	6236'290	1023	- 10		
571	58'366	3911	- 21	7623'571	3465	+ 18	62'265	585	- 12	68'284	1069	+ 13		
592	188			29'895	3477	- 2	287	585	+ 10	6486'520	1383	+ 7		
614	2425707'550	o	- 041	56'317	3527	- 17	6123'391	1041	- 2	6569'212	1502	- 10		
35'383	40'552	18	- 79	64'272	3542	+ 7	6240'246	1245	+ 39	71'280	1505	- 28		
182			64'494	193			6475'540	1656	- 14	73'380	1508	- 13		
2425707'550	516	31	+ 22	2425719'518	o	- 022	6569'450	1820	- 14	402	1508	+ 9		
16'636	75'566	37	+ 59	540	o	o	471	1820	+ 7	7623'571	3019	- 23		
39'587	5821'456	62	+ 60	40'530	30	- 25	493	1820	+ 29	56'281	3066	+ 20		
75'566	32'377	68	- 33	552	30	- 3	73'423	1827	- 49	58'345	3069	- 1		
92'587	6076'600	201	+ 59	75'556	80	- 15	466	1827	- 6	366	3069	+ 20		
95'351	6120'567	225	- 28	92'434	104	+ 41	7623'659	3661	o	216				
5821'434	55'594	244	+ 123	456	104	+ 63	56'281	3718	- 17	2425525'272	o	- 014		
			6212'222	275	- 152				303	3718	+ 5	5764'494	345	- 31
									317	3718	+ 19			

TABLE II (continued).

d	t	d	d	t	d	d	t	d	d	t	d	222
2425764'516	345 -	'009	2425791'462	655 +	'001	2425864'280	140 -	'010	2425706'566	430 -	'035	
92'247	385 -	16	97'308	666 -	15	6090'550	613 -	25	'588	430 -	13	
'268	385 +	5	'329	666 +	6	6123'593	682 +	8	64'494	539 +	5	2425525'272
5803'351	401 -	8	5823'477	715 +	41	25'555	686 +	56	92'609	592 -	27	'294
'372	401 +	13	61'242	786 -	31	55'616	749 -	22	95'351	597 +	60	5792'587
21'434	427 +	46	85'274	831 +	20	6218'323	880 +	14	99'508	605 -	31	'609
'456	427 +	68	6076'600	1190 +	29	63'279	974	0	5823'477	650 +	39	94'542
60'214	483 -	7	6212'442	1445 -	22	73'291	995 -	34	6090'550	1153 -	18	99'486
6067'540	782 -	23	42'307	1501 -	1	'313	995 -	12	6214'331	1386 +	22	'508
6156'300	910 -	24	66'305	1546 +	16	6480'493	1428 +	19	39'289	1433 +	20	5832'258
76'466	939 +	32	67'350	1548 -	5	6569'428	1614 -	29	6540'430	2000 +	41	35'370
6240'224	1031 -	7	6480'515	1948 -	6	'450	1614 -	7	73'273	2062 -	42	'392
'246	1031 +	15	6561'504	2100 -	21	'471	1614 +	14	'295	2062 -	20	64'302
42'285	1034 -	27	'526	2100 +	1	70'403	1616 -	11				6160'546
'307	1034 -	5	69'514	2115 -	4	'446	1616 +	32	epochs not used:			6237'294
6570'296	1507 -	17	73'273	2122 +	24	7623'377	3817 -	3	7623'549			42'307
7658'345	3076 +	10	7367'288	3612 -	6	35'361	3842 +	21	'592			44'249
			'310	3612 +	16	36'289	3844 -	8	36'311			66'327
218			7635'339	4115 -	12	'311	3844 +	14	9079'454			6569'407
						57'322	3888 -	25	'476			71'280
2425442'385	0 -	'015	219						'498			7656'303
5719'518	520 +	1				221			9104'480			'332
20'563	522 -	20	2425797'308	0 -	'006				06'538			3388 -
'584	522 +	1	'329	0 +	15	2425478'239	0	'000	'560			3388 +

TABLE 12.

J.D.	bright-ness	J.D.	bright-ness	J.D.	bright-ness	J.D.	bright-ness	J.D.	bright-ness	J.D.	bright-ness
I		d	s	d	s	d	s	6I		d	s
2425442'3847	7'0	2426129'4438	1'7	2427636'2894	2'5	2426161'5377	2'4	2425524'2627	2'5	2426240'3128	3'8
'4276	6'3	56'2995	8'0	'3107	2'5	6218'3233	7'4	'2848	2'9	3347	4'3
5707'5497	3'9	'3460	7'0	20		'3451	9'7	42'3073	2'5	3973	5'7
'5719	3'9	6210'2931	2'1			40'2239	3'3	26'2653	4'1	63'2127	4'7
40'5302	6'0	12'2217	1'2			'2461	2'0	'2882	3'8	2786	5'7
65'4302	8'0	'2435	4'1	2425442'3847	8'3	41'2911	4'9	5706'5663	4'7	3010	2'5
92'4344	2'5	'2654	3'4	'4276	4'9	'3129	3'3	'5882	6'1	65'2467	6'4
'4555	5'2	'2872	5'2	78'2876	4	42'2851	2'4	16'6365	2'5	'2686	6'4
'5868	8'0	'3533	6'3	5524'2627	5'4	'3073	4'5	39'5871	5'3	67'2251	3'6
6089	6'3	'3751	7'5	'2848	3'3	'3973	4'1	'6091	6'4	3276	4'7
94'5422	8'0	'4420	8'0	25'2723	4'5	65'2467	7'4	64'4943	4'7	'3495	2'5
'5642	6'7	'4036	7'5	'2945	6'1	'2686	10'2	'5165	4'7	73'2913	4'3
98'3281	7'0	14'2154	4'1	5706'5663	2'9	66'3052	4'1	91'4624	4'0	'3130	4'7
5821'4345	3'9	'2373	5'2	07'5497	6'9	'3270	6'6	'4846	2'5	6476'5544	2'5
'4563	5'2	'3311	7'5	'5719	11'1	6565'4721	2'9	97'5347	2'5	'5760	4'7
23'4774	7'5	'3533	7'5	19'5181	2'4	'4942	1'0	99'4857	6'4	4942	5'1
'4992	8'0	'4724	3'4	'5401	4'9	7652'3391	1'2	'5077	6'1	69'4284	3'6
52'4176	3'9	41'2911	5'2	20'5844	2'0	58		5826'4535	4'3	65'4721	7'0
'4397	3'4	43'2896	6'3	'4455	4'9			'4790	3'6	4942	5'1
58'3280	7'5	'3114	7'0	92'6089	0	2425527'2729	5'0	51'3110	7'0	4498	4'7
'3498	7'5	6475'5094	7'5	5822'2574	4	5764'4943	6'4	53'3275	7'0	'4713	5'7
60'2143	7'5	'5395	7'5	23'2084	3'3	'5105	6'4	63'3711	4'7	'4928	6'9
'2362	7'5	6568'2708	2'5	'2302	7'0	91'4624	6'4	88'2772	4'0	'5143	6'4
62'2648	7'5	'2932	4'3	25'3621	2'4	'4846	4'5	'5888	4'7	'5357	4'7
'2866	6'0	6570'2102	3'4	'5089	2'0	5822'2574	4'5	20'5669	7'6	70'2102	4'7
64'2796	5'2	'2317	3'4	26'4535	1'2	23'4774	5'7	'5890	4'7	'2317	4'3
'3021	2'5	'2531	4'3	'4790	3'3	'4992	6'4	22'5698	7'6	'2531	4'1
85'2740	3'9	'2746	3'4	36'3171	4'9	36'3171	5'0	'5916	6'6	72'2552	2'5
'2982	4'3	'2961	5'2	'3391	7'9	'3391	5'0	53'4740	4'0	'2766	3'8
93'3414	2'5	'3175	6'7	49'3914	2'9	6123'5928	8'4	'5422	7'0	73'4664	3'2
6088'5405	7'5	'3390	6'7	'4132	4'1	'6146	6'4	55'5939	4'7	'4878	4'0
'5623	5'2	'3005	7'5	6088'5405	1'2	55'6157	6'4	'6157	4'7	'5093	5'3
90'5496	8'5	'3819	7'5	'5623	0	6213'3197	5'5	56'2995	2'5	7622'2267	4'3
'5713	7'5	'4034	7'5	90'5496	1'2	'3419	5'9	61'5377	2'5	'2485	3'9
94'5200	6'3	'4249	7'5	'5713	2'4	'4090	5'0	76'4437	7'0	23'4632	3'2
'5418	4'7	'4464	7'5	6123'3913	11'1	40'2461	6'4	'4658	4'3	'4847	3'4
6123'3913	6'7	'4678	5'2	'4131	7'4	'3128	5'3	6213'4090	5'3	'5062	4'1
'4131	8'0	'4893	6'7	25'5549	7'9	'3347	5'7	'4762	4'7	'5276	5'3
'5928	4'3	'5108	6'7	'5770	10'2	67'3495	5'0	36'2900	4'7	'5491	6'1
'6146	2'0	'5322	6'3	60'5245	6'8	6486'5199	6'4	'3118	6'1	'5706	6'4
28'5549	2'0	72'2552	3'9	'5404	3'3	6540'4296	5'0	38'2767	2'5	'5920	4'0
'5770	2'0	'2766	6'7	61'5159	1'0			'2985	4'7	'6135	3'8



TABLE I2 (continued).

J.D.	bright- ness	J.D.	bright- ness	J.D.	bright- ness	J.D.	bright- ness	J.D.	bright- ness	J.D.	bright- ness
2427623 <sup>d</sup> 6350 <sup>s</sup>	2.1	2426088 <sup>d</sup> 5623 <sup>s</sup>	3.6	2425808 <sup>d</sup> 4502 <sup>s</sup>	9.4	2425886 <sup>d</sup> 2417 <sup>s</sup>	1.8	2426303 <sup>d</sup> 2710 <sup>s</sup>	8.0	2426240 <sup>d</sup> 2461 <sup>s</sup>	F
36 <sup>d</sup> 2894 <sup>s</sup>	4.3	6153 <sup>d</sup> 4740 <sup>s</sup>	3.5	4720 <sup>d</sup> 10.1	1.1	6129 <sup>d</sup> 4438 <sup>s</sup>	3.8	6569 <sup>d</sup> 2338 <sup>s</sup>	6.0	3347 <sup>d</sup> 3.5	3.5
31 <sup>d</sup> 07 <sup>s</sup>	5.3	56 <sup>d</sup> 2995 <sup>s</sup>	9.6	22 <sup>d</sup> 2574 <sup>s</sup>	10.1	4656 <sup>d</sup> 3.8	F	2559 <sup>d</sup> 7.4	7.4	48 <sup>d</sup> 2325 <sup>s</sup>	4.6
57 <sup>d</sup> 2941 <sup>s</sup>	4.0	3460 <sup>d</sup> 10.5	10.5	6153 <sup>d</sup> 4740 <sup>s</sup>	10.1	6151 <sup>d</sup> 5159 <sup>s</sup>	F	2774 <sup>d</sup> 7.7	7.7	2543 <sup>d</sup> 4.6	4.6
70		6212 <sup>d</sup> 2435 <sup>s</sup>	3.8	5422 <sup>d</sup> 8.6	8.6	5377 <sup>d</sup> 3.3	F	2989 <sup>d</sup> 6.4	6.4	6561 <sup>d</sup> 5043 <sup>s</sup>	F
See B.A.N. No. 256		2654 <sup>d</sup> 4.1	4.1	56 <sup>d</sup> 2995 <sup>s</sup>	8.7	6212 <sup>d</sup> 4420 <sup>s</sup>	F	3210 <sup>d</sup> 5.1	5.1	5259 <sup>d</sup> F	F
additional minima:		2872 <sup>d</sup> 5.3	5.3	3460 <sup>d</sup> 7.2	7.2	4636 <sup>d</sup> 1.8	F	70 <sup>d</sup> 3175 <sup>s</sup>	6.7	65 <sup>d</sup> 4721 <sup>s</sup>	4.0
2429082 <sup>d</sup> 4125 <sup>s</sup>	3.6	3533 <sup>d</sup> 10.1	10.1	6214 <sup>d</sup> 2154 <sup>s</sup>	10.6	14 <sup>d</sup> 3311 <sup>s</sup>	1.8	3390 <sup>d</sup> 7.9	7.9	4942 <sup>d</sup> 4.6	4.6
4343 <sup>d</sup> 3.8	3.8	3751 <sup>d</sup> 10.3	10.3	2373 <sup>d</sup> 9.5	9.5	3533 <sup>d</sup> 1.8	F	3605 <sup>d</sup> 7.4	7.4	69 <sup>d</sup> 4928 <sup>s</sup>	3.5
83.		4420 <sup>d</sup> 8.4	8.4	36 <sup>d</sup> 2900 <sup>s</sup>	10.1	48 <sup>d</sup> 2325 <sup>s</sup>	1.8	3819 <sup>d</sup> 6.4	6.4	5357 <sup>d</sup> 4.6	4.6
2425442 <sup>d</sup> 3847 <sup>s</sup>	4.1	4636 <sup>d</sup> 6.3	6.3	3118 <sup>d</sup> 10.1	10.1	63 <sup>d</sup> 3010 <sup>s</sup>	2.8	73 <sup>d</sup> 5093 <sup>s</sup>	5.0	7367 <sup>d</sup> 2882 <sup>s</sup>	4.0
5716 <sup>d</sup> 6147 <sup>s</sup>	7.5	18 <sup>d</sup> 3233 <sup>s</sup>	10.1	6476 <sup>d</sup> 5544 <sup>s</sup>	5.5	65 <sup>d</sup> 2467 <sup>s</sup>	1.8	7367 <sup>d</sup> 2882 <sup>s</sup>	8.3	7623 <sup>d</sup> 2271 <sup>s</sup>	3.8
6365 <sup>d</sup> 10.6	10.6	3451 <sup>d</sup> 8.8	8.8	6559 <sup>d</sup> 3129 <sup>s</sup>	8.5	2686 <sup>d</sup> 2.3	F	3100 <sup>d</sup> 7.0	7.0	2485 <sup>d</sup> 3.7	3.7
19 <sup>d</sup> 5181 <sup>s</sup>	4.6	6569 <sup>d</sup> 2774 <sup>s</sup>	3.6	3347 <sup>d</sup> 7.6	7.6	7636 <sup>d</sup> 2894 <sup>s</sup>	3.8	7623 <sup>d</sup> 6135 <sup>s</sup>	6.3	2700 <sup>d</sup> 4.1	4.1
5401 <sup>d</sup> 5.8	5.8	2989 <sup>d</sup> 5.5	5.5	70 <sup>d</sup> 2102 <sup>s</sup>	8.4	3107 <sup>d</sup> F	F	6350 <sup>d</sup> 8.3	8.3	2915 <sup>d</sup> 4.6	4.6
75 <sup>d</sup> 5660 <sup>s</sup>	3.6	3210 <sup>d</sup> 5.6	5.6	2317 <sup>d</sup> 9.6	9.6	8064 <sup>d</sup> 3317 <sup>s</sup>	F	6592 <sup>d</sup> 6.3	6.3	3129 <sup>d</sup> 5.6	5.6
90 <sup>d</sup> 3896 <sup>s</sup>	8.6	3425 <sup>d</sup> 7.0	7.0	2531 <sup>d</sup> 9.2	9.2	3531 <sup>d</sup> 2.8	F			3344 <sup>d</sup> 6.6	6.6
4116 <sup>d</sup> 10.5	10.5	3640 <sup>d</sup> 8.6	8.6	2746 <sup>d</sup> 10.1	10.1			162		3559 <sup>d</sup> 6.6	6.6
93 <sup>d</sup> 2513 <sup>s</sup>	3.6	3854 <sup>d</sup> 9.8	9.8	2961 <sup>d</sup> 10.6	10.6	F = invisible on		2425524 <sup>d</sup> 2627 <sup>s</sup>	3.8	3774 <sup>d</sup> F	F
5849 <sup>d</sup> 3914 <sup>s</sup>	7.2	4069 <sup>d</sup> 10.5	10.5	3175 <sup>d</sup> 10.1	10.1	good plate		2848 <sup>d</sup> 4.0	4.0	3988 <sup>d</sup> 6.6	6.6
4132 <sup>d</sup> 8.0	8.0	4284 <sup>d</sup> 10.5	10.5	3390 <sup>d</sup> 8.6	8.6			5719 <sup>d</sup> 5181 <sup>s</sup>	5.1	4203 <sup>d</sup> F	F
52 <sup>d</sup> 4176 <sup>s</sup>	11.1	4498 <sup>d</sup> 9.9	9.9	3605 <sup>d</sup> 8.4	8.4	152		39 <sup>d</sup> 5871 <sup>s</sup>	4.6	4418 <sup>d</sup> F	F
4397 <sup>d</sup> 10.5	10.5	4713 <sup>d</sup> 7.5	7.5	3819 <sup>d</sup> 7.1	7.1	2425745 <sup>d</sup> 5484 <sup>s</sup>	5.1	6091 <sup>d</sup> 4.6	4.6	4632 <sup>d</sup> 6.1	6.1
58 <sup>d</sup> 3280 <sup>s</sup>	10.5	4928 <sup>d</sup> 6.3	6.3	4034 <sup>d</sup> 6.7	6.7	75 <sup>d</sup> 5660 <sup>s</sup>	6.6	92 <sup>d</sup> 4344 <sup>s</sup>	3.5	4847 <sup>d</sup> 6.6	6.6
3498 <sup>d</sup> 9.0	9.0	5143 <sup>d</sup> 4.6	4.6	4249 <sup>d</sup> 5.5	5.5	5830 <sup>d</sup> 2720 <sup>s</sup>	7.0	4555 <sup>d</sup> 4.1	4.1	5062 <sup>d</sup> F	F
61 <sup>d</sup> 2420 <sup>s</sup>	10.5	72 <sup>d</sup> 2766 <sup>s</sup>	4.6	73 <sup>d</sup> 2087 <sup>s</sup>	5.5	2938 <sup>d</sup> 5.9	F	5868 <sup>d</sup> F	F	5276 <sup>d</sup> 5.6	5.6
2642 <sup>d</sup> 9.5	9.5	7652 <sup>d</sup> 3391 <sup>s</sup>	9.0	2302 <sup>d</sup> 5.7	5.7	32 <sup>d</sup> 3555 <sup>s</sup>	5.9	5825 <sup>d</sup> 2250 <sup>s</sup>	F	5491 <sup>d</sup> 5.6	5.6
2860 <sup>d</sup> 10.3	10.3	58 <sup>d</sup> 3446 <sup>s</sup>	4.1	7652 <sup>d</sup> 3391 <sup>s</sup>	10.1	3773 <sup>d</sup> 7.4	F	2482 <sup>d</sup> F	F	5706 <sup>d</sup> 5.6	5.6
64 <sup>d</sup> 2796 <sup>s</sup>	7.5	96		122		60 <sup>d</sup> 2362 <sup>s</sup>	5.1	49 <sup>d</sup> 3914 <sup>s</sup>	3.7	5920 <sup>d</sup> 4.6	4.6
3021 <sup>d</sup> 4.6	4.6	2425794 <sup>d</sup> 5422 <sup>s</sup>	6.7	2425797 <sup>d</sup> 5347 <sup>s</sup>	2.8	6129 <sup>d</sup> 4656 <sup>s</sup>	5.5	4132 <sup>d</sup> 4.6	4.6	64 <sup>d</sup> 2510 <sup>s</sup>	4.6
5926 <sup>d</sup> 2392 <sup>s</sup>	6.3	5642 <sup>d</sup> 9.1	9.1	5565 <sup>d</sup> 2.8	2.8	56 <sup>d</sup> 2995 <sup>s</sup>	7.4	86 <sup>d</sup> 2199 <sup>s</sup>	F	2725 <sup>d</sup> 3.8	3.8
2624 <sup>d</sup> 8.0	8.0	97 <sup>d</sup> 3076 <sup>s</sup>	7.6	99 <sup>d</sup> 4857 <sup>s</sup>	1.8	3460 <sup>d</sup> 6.6	F	2417 <sup>d</sup> F	F	8079 <sup>d</sup> 3395 <sup>s</sup>	4.6
6076 <sup>d</sup> 5998 <sup>s</sup>	8.6	3294 <sup>d</sup> 7.6	7.6	5850 <sup>d</sup> 3148 <sup>s</sup>	1.8	6213 <sup>d</sup> 2155 <sup>s</sup>	6.1	6093 <sup>d</sup> 5276 <sup>s</sup>	4.1	F = invisible on	good plate
88 <sup>d</sup> 5405 <sup>s</sup>	6.6	5347 <sup>d</sup> 5.7	5.7	3368 <sup>d</sup> 1.8	1.8	14 <sup>d</sup> 2154 <sup>s</sup>	6.7	5493 <sup>d</sup> 4.6	4.6		
				86 <sup>d</sup> 2199 <sup>s</sup>	1.8	2373 <sup>d</sup> 9.1	F	6236 <sup>d</sup> 2900 <sup>s</sup>	3.1		
						6303 <sup>d</sup> 2491 <sup>s</sup>	7.4	40 <sup>d</sup> 2239 <sup>s</sup>	F		

TABLE I3.

n phase	bright- ness	n phase	bright- ness	n phase	bright- ness	n phase	bright- ness	n phase	bright- ness	n phase	bright- ness	n phase	bright- ness
I										variable + component			
10	026	10	834	10	668	10	504	10	436	10	446	10	059
10	072	10	870	10	701	10	535	10	473	10	480	10	098
10	112	11	910	10	727	10	569	10	498	10	525	12	020
10	153	11	966	10	752	10	597	10	524	10	574	12	076
10	178			10	775	10	624	10	551	10	636	12	135
10	209	2		10	802	10	644	10	577	10	689	12	185
10	224			10	839	10	672	10	605	10	765	12	229
10	245	10	025	10	864	10	692	10	638	10	825	12	273
10	269	10	069	10	894	10	720	10	674	10	895	12	322
10	285	10	100	10	922	10	751	10	705	11	971	12	392
10	304	10	114	9	951	10	778	10	732			12	440
10	322	10	139	9	983	10	806	10	766	6		12	474
10	343	10	165			10	845	10	796			12	510
10	362	10	196	3		10	899	10	832	single variable		12	545
10	377	10	231			10	955	10	863	weight		12	588
10	400	10	266	10	023			11	897			12	636
10	437	10	303	10	068	4		11	944	39	048	12	689
10	538	10	344	10	132			11	988	37	132	12	736
10	628	10	375	10	189	10	015	5		38	263	12	782
10	667	10	414	10	224	10	045			39	348	12	820
10	690	10	448	10	261	10	096			37	415	12	859
10	713	10	474	10	301	10	133	10	042	36	490	12	898
10	732	10	500	10	342	10	170	10	108	36	580	12	932
10	754	10	526	10	371	10	207	10	192	37	695	12	978
10	779	10	555	10	397	10	256	10	249	39	814		
10	792	10	593	10	425	10	297	10	300	38	930	7	
10	812	10	621	10	449	10	334	10	337				
		10	647	10	478	10	478	10	389				





TABLE 13 (continued).

n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness			
IO	'163	8.9	IO	'376	5.1	IO	'807	6.4	IO	'885	4.4	IO	'300	2.7	IO	'489	1.3	IO	'364	5.5			
IO	'193	8.7	IO	'418	5.2	IO	'844	6.6	IO	'909	4.7	IO	'030	6.2	IO	'516	1.8	IO	'454	7.3			
IO	'246	8.9	IO	'449	5.7	II	'870	6.5	IO	'939	4.6	IO	'064	6.2	IO	'367	3.1	IO	'535	1.8	IO	'588	8.4
IO	'289	9.2	IO	'479	6.2	II	'904	6.4	9	'963	4.5	IO	'093	6.9	IO	'396	3.4	IO	'561	2.4	IO	'674	8.7
IO	'319	9.5	IO	'505	6.4	II	'943	6.3	9	'990	5.0	IO	'122	7.3	IO	'435	3.2	IO	'588	2.4	IO	'799	9.0
IO	'344	9.6	IO	'539	6.5	II	'983	6.4				IO	'149	7.2	IO	'478	3.2	IO	'641	2.9	9	'877	9.4
IO	'381	9.3	IO	'574	6.6				61			IO	'187	7.1	IO	'522	3.0	IO	'684	3.3	9	'972	9.4
IO	'410	9.6	IO	'618	7.3				IO	'011	2.4	IO	'243	7.8	IO	'568	3.4	IO	'715	3.8			
IO	'441	9.6	IO	'658	7.1	58			IO	'037	2.4	IO	'293	8.0	IO	'615	3.8	IO	'750	4.0	J.D. 2426218		
IO	'471	9.4	IO	'694	7.5	IO	'032	2.1	IO	'074	2.2	IO	'348	8.1	IO	'677	3.9	IO	'792	4.4	- 9165		
IO	'491	9.7	IO	'740	7.5	IO	'074	2.2	IO	'052	2.4	IO	'392	8.2	IO	'729	3.9	IO	'824	4.9			
IO	'518	9.8	IO	'785	7.6	IO	'096	2.1	IO	'074	2.2	IO	'424	8.3	IO	'766	3.9	IO	'853	5.0	IO	'031	7.2
IO	'542	9.7	IO	'840	7.8	IO	'124	1.9	IO	'105	2.4	IO	'450	8.1	IO	'800	4.1	IO	'883	5.0	IO	'075	5.2
IO	'577	10.4	IO	'882	7.5	IO	'153	1.8	IO	'131	2.6	IO	'486	7.9	IO	'857	3.6	9	'911	5.2	IO	'111	2.1
IO	'606	9.6	II	'940	7.7	IO	'170	1.8	IO	'152	2.8	IO	'534	8.1	IO	'896	4.0	9	'935	4.8	IO	'149	2.3
IO	'632	9.5	II	'988	7.4	IO	'190	2.0	IO	'175	3.3	IO	'600	7.9	IO	'927	3.7	9	'904	5.0	IO	'186	3.1
IO	'656	9.2				IO	'217	1.3	IO	'200	4.2	IO	'648	8.4	IO	'965	3.1	9	'990	5.5	IO	'226	4.9
IO	'682	9.3				IO	'248	1.4	IO	'229	5.0	IO	'694	8.1	IO	'988	2.2						
IO	'711	7.8	53			IO	'286	2.2	IO	'262	6.5	IO	'739	6.8									
IO	'741	3.8	IO	'931	8.1	IO	'314	1.7	IO	'293	6.4	IO	'764	4.3	69								
IO	'765	2.9	IO	'073	8.0	IO	'337	1.9	IO	'319	5.2	IO	'797	.8	mean curve								
IO	'793	1.2	IO	'103	7.6	IO	'363	2.5	IO	'342	4.1	IO	'828	.9	IO	'012	6.5	IO	'023	9.0	IO	'314	6.6
IO	'821	2.6	IO	'140	8.0	IO	'400	1.8	IO	'370	3.3	IO	'853	1.8	IO	'054	6.5	IO	'069	9.0	IO	'382	7.0
IO	'846	3.4	IO	'181	7.8	IO	'426	2.2	IO	'401	2.9	IO	'875	2.3	IO	'086	6.6	IO	'104	8.8	IO	'424	7.8
9	'874	3.1	IO	'208	7.9	IO	'443	2.2	IO	'439	2.5	IO	'895	3.4	IO	'115	7.1	IO	'150	9.5	IO	'474	7.7
9	'901	5.0	IO	'242	7.8	IO	'460	1.9	IO	'476	2.5	II	'917	4.0	IO	'151	6.6	IO	'193	9.0	IO	'540	8.3
9	'947	5.9	IO	'274	8.0	IO	'487	1.3	IO	'508	2.7	II	'940	4.1	IO	'180	6.8	IO	'233	8.4	IO	'608	8.5
9	'982	7.2	IO	'304	7.6	IO	'511	1.6	IO	'535	2.6	II	'976	5.2	IO	'216	6.6	IO	'274	8.9	IO	'661	8.9
	48		IO	'334	8.0	IO	'531	1.7	IO	'561	2.6				IO	'260	7.0	IO	'303	7.0	IO	'712	9.2
			IO	'376	8.1	IO	'556	2.0	IO	'599	2.8				IO	'301	6.9	IO	'333	5.6	IO	'784	9.4
			IO	'414	7.6	IO	'630	2.1	IO	'672	2.6				IO	'336	6.5	IO	'359	6.4	IO	'866	9.0
			IO	'451	5.0	IO	'630	2.1	IO	'672	2.6	64			IO	'379	6.8	IO	'387	4.2	IO	'928	9.0
IO	'021	12.7	IO	'480	2.6	IO	'653	1.6	IO	'696	2.9	IO	'018	8.3	IO	'420	6.6	IO	'416	2.2	75		
IO	'074	13.1	IO	'504	1.1	IO	'678	1.4	IO	'717	3.2	IO	'051	8.3	IO	'455	6.9	IO	'446	2.9	IO	'019	11.4
IO	'124	13.4	IO	'530	.6	IO	'707	1.8	IO	'745	3.3	IO	'074	8.6	IO	'483	6.5	IO	'472	1.9	IO	'051	11.5
IO	'162	13.4	IO	'557	.8	IO	'747	2.2	IO	'784	3.4	IO	'104	8.4	IO	'524	7.1	IO	'503	4.1	IO	'083	11.0
IO	'205	13.1	IO	'597	1.8	IO	'785	2.1	IO	'811	3.2	IO	'142	8.8	IO	'562	6.2	IO	'530	4.5	IO	'108	11.4
IO	'253	13.3	IO	'633	2.8	6	'809	2.8	IO	'848	3.0	IO	'183	8.7	IO	'603	5.4	IO	'557	5.9	IO	'142	10.3
IO	'296	13.3	IO	'656	4.1	6	'838	5.8	IO	'893	2.8	IO	'212	8.7	IO	'643	4.1	IO	'593	6.0	IO	'173	6.4
IO	'336	13.3	IO	'684	4.7	6	'866	6.4	IO	'923	3.0	IO	'243	8.5	IO	'680	3.6	IO	'628	6.3	IO	'203	2.1
IO	'365	13.4	IO	'707	5.0	6	'890	5.8	IO	'950	2.5	IO	'274	8.8	IO	'712	3.5	IO	'655	6.8	IO	'234	.8
IO	'416	13.5	IO	'735	5.2	6	'907	4.4	IO	'975	3.0	IO	'311	8.9	IO	'739	3.4	IO	'681	7.1	IO	'269	2.3
IO	'463	13.4	IO	'767	6.0	6	'936	2.2	7	'993	2.5	IO	'349	8.8	IO	'761	4.1	IO	'704	8.0	IO	'291	2.6
IO	'505	12.6	IO	'805	6.4	IO	'968	2.2				IO	'374	8.6	IO	'790	4.5	IO	'741	7.4	IO	'325	4.1
IO	'544	7.8	IO	'839	6.6				62			IO	'407	8.5	IO	'833	5.2	IO	'781	8.0	IO	'353	4.8
IO	'572	3.8	IO	'878	7.0	59			IO	'442	8.3	IO	'442	8.3	IO	'867	5.3	IO	'824	8.4	IO	'391	5.7
IO	'596	3.3	IO	'906	7.3				IO	'477	8.0	IO	'477	8.0	IO	'900	6.0	IO	'862	8.3	IO	'422	6.6
IO	'625	3.7	IO	'943	7.8	IO	'032	4.8	IO	'022	5.0	IO	'516	7.4	IO	'925	5.9	IO	'900	8.5	IO	'455	6.8
IO	'644	4.6	II	'978	7.7	IO	'094	4.4	IO	'055	4.9	IO	'548	5.1	9	'962	5.5	II	'933	9.0	IO	'490	7.7
IO	'677	5.4				IO	'129	4.4	IO	'085	5.5	IO	'588	2.3				IO	'533	8.1			
IO	'700	6.9				IO	'150	4.9	IO	'118	5.3	IO	'637	4.2	70			IO	'563	9.4			
IO	'740	8.0	57			IO	'178	4.9	IO	'149	5.0	IO	'678	3.8	see B.A.N.			IO	'591	9.5			
IO	'776	9.0	IO	'016	6.7	IO	'209	5.1	IO	'186	5.5	IO	'716	4.9	No. 256			IO	'609	10.2			
IO	'805	9.6	IO	'071	6.5	IO	'238	4.9	IO	'228	5.3	IO	'756	5.0				IO	'633	9.8			
IO	'840	11.0	IO	'120	6.6	IO	'275	4.7	IO	'284	6.0	IO	'796	6.3	72			IO	'649	10.6			
IO	'871	11.6	IO	'157	6.6	IO	'302	4.7	IO	'324	5.7	IO	'834	6.5				IO	'676	10.2			
IO	'901	11.6	IO	'196	5.6	IO	'328	4.1	IO	'352	5.9	IO	'877	7.3	IO	'015	5.3	IO	'704	10.4			
IO	'929	12.2	IO	'245	2.9	IO	'350	3.7	IO	'396	6.2	IO	'906	8.1	IO	'047	5.5	5	'166	4.5	IO	'737	10.8
IO	'952	11.7	IO	'284	1.6	IO	'382	2.6	IO	'440	5.7	IO	'932	7.3	IO	'072	5.2	5	'346	7.0	IO	'778	10.8
9	'981	12.4	IO	'312	1.1	IO	'409	2.6	IO	'486	6.0	9	'961	7.4	IO	'093	4.8	5	'426	7.4	IO	'818	10.8
	50		IO	'344	2.1	IO	'441	1.8	IO	'528	6.0	9	'987	7.6	IO	'119	5.4	5	'516	7.8	IO	'844	10.8
			IO	'371	2.7	IO	'468	1.4	IO	'565	5.9				IO	'147	5.3	5	'634	8.8	IO	'868	10.9
IO	'036	7.8	IO	'397	3.3	IO	'508	1.8	IO	'600	5.4	67			IO	'178	5.5	5	'741	8.7	IO	'898	11.1
IO	'076	7.8	IO	'426	3.3	IO	'545	2.2	IO	'653	2.9	IO	'012	1.7	IO	'208	5.7	5	'803	8.3	IO	'938	11.3
IO	'106	7.3	IO	'473	4.2	IO	'573	2.5	IO	'691	2.1	IO	'032	1.4	IO	'233	5.9	6	'907	8.7	9	'978	10.4
IO	'137	6.0	IO	'510	4.9	IO	'605	2.7	IO	'719	1.4	IO	'056	1.0	IO	'259	5.7						
IO	'171	3.8	IO	'544	5.4	IO	'649	3.2	IO	'751	2.0	IO	'										

TABLE 13 (continued).

n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness
IO	'295	5.7	IO	'835	4.2	IO	'081	1.8	88			IO	'608	2.5	IO	'753	9.2	IO	'840	6.7	IO	'907	2.0
IO	'354	5.6	IO	'855	3.3	IO	'103	1.9	IO	'025	3.6	IO	'643	2.0	IO	'785	9.8	IO	'866	6.7	IO	'945	2.0
IO	'422	5.7	IO	'888	3.3	IO	'138	2.0	IO	'064	3.2	IO	'674	2.0	II	'822	10.1	IO	'892	6.0	9	'978	1.7
IO	'473	5.4	IO	'912	4.3	IO	'178	2.0	IO	'090	3.3	IO	'714	2.9	II	'865	10.5	9	'930	4.2			
IO	'548	5.6	9	'945	5.3	IO	'212	2.2	IO	'115	3.6	IO	'747	3.0	II	'907	10.2	9	'980	2.3			
IO	'614	4.5	9	'986	6.8	IO	'245	2.3	IO	'142	3.6	IO	'777	3.4	II	'940	10.5						
IO	'650	3.9				IO	'267	2.1	IO	'187	3.8	IO	'819	2.8	II	'976	10.6						
IO	'682	1.6		80		IO	'318	2.3	IO	'246	3.4	IO	'856	3.8									
IO	'713	1.7				IO	'357	1.9	IO	'318	3.6	II	'914	3.8									
IO	'738	1.3	IO	'014	6.6	IO	'390	2.3	IO	'443	3.8	II	'975	4.0									
IO	'770	2.3	IO	'038	6.2	9	'413	3.7	IO	'548	3.8												
IO	'802	3.7	IO	'078	6.9	5	'426	4.9	IO	'615	3.1	91											
IO	'832	4.0	IO	'111	5.7	5	'437	7.8	IO	'659	2.4												
IO	'866	4.6	IO	'159	6.3	5	'444	9.6	IO	'692	1.6	IO	'021	4.5	IO	'017	3.8	5	'007	.9			
IO	'897	4.1	IO	'210	5.5	5	'453	10.1	IO	'717	1.4	IO	'057	2.8	IO	'041	4.2	IO	'156	3.9			
IO	'929	4.3	IO	'241	5.0	5	'463	10.4	IO	'740	1.4	IO	'089	1.6	IO	'080	4.1	IO	'256	5.2			
9	'951	4.4	IO	'267	5.2	5	'470	10.0	IO	'768	1.6	IO	'126	2.0	IO	'335	6.2	IO	'335	6.2			
9	'985	4.8	IO	'290	5.3	5	'478	8.0	IO	'807	2.4	5	'234	7.9	IO	'417	6.5	IO	'417	6.5			
			IO	'316	4.6	5	'484	7.1	II	'850	2.9	5	'256	9.8	IO	'511	6.7	IO	'511	6.7			
			IO	'335	4.2	5	'491	4.5	II	'890	2.9	5	'268	10.2	IO	'583	6.8	IO	'583	6.8			
			IO	'355	3.3	IO	'512	3.1	II	'930	2.9	5	'276	9.7	IO	'664	6.8	IO	'664	6.8			
IO	'013	4.1	IO	'384	3.3	IO	'547	2.0	II	'973	3.3	IO	'246	4.4	5	'289	8.3	IO	'762	6.6			
IO	'051	4.2	IO	'416	3.6	IO	'582	2.4				IO	'267	4.7	5	'305	6.3	IO	'865	6.6			
IO	'093	5.2	IO	'455	3.2	IO	'615	1.8	89			IO	'291	5.2	IO	'318	4.8	IO	'922	6.1			
IO	'136	5.6	IO	'487	3.7	IO	'643	1.9				IO	'313	4.9	IO	'337	4.1	5	'958	3.8			
IO	'171	5.5	IO	'512	2.7	IO	'664	1.8	IO	'013	7.0	IO	'332	5.5	IO	'350	3.9	IO	'043	4.1			
IO	'209	5.6	IO	'542	4.0	IO	'687	1.8	IO	'045	7.9	IO	'353	5.4	IO	'363	4.3	IO	'134	3.8			
IO	'254	5.7	IO	'582	4.1	IO	'725	1.5	IO	'092	7.4	IO	'384	5.7	IO	'392	4.1	IO	'189	3.6			
IO	'284	6.2	IO	'616	5.0	IO	'743	1.8	IO	'137	7.8	IO	'423	6.4	IO	'415	4.1	IO	'234	3.1			
IO	'316	6.0	IO	'653	4.8	IO	'756	2.2	IO	'168	7.9	IO	'456	6.3	IO	'437	4.0	IO	'259	2.2			
IO	'359	6.0	IO	'679	4.8	IO	'776	2.1	IO	'187	7.8	IO	'486	6.3	IO	'454	3.9	IO	'289	1.9			
IO	'409	6.0	IO	'717	5.2	IO	'791	1.9	IO	'222	7.7	IO	'516	6.2	IO	'490	3.9	IO	'323	1.8			
IO	'455	5.9	IO	'759	5.4	IO	'810	1.8	IO	'259	7.9	IO	'543	6.7	IO	'526	4.1	IO	'355	2.7			
IO	'488	6.1	IO	'796	5.6	IO	'829	1.8	IO	'288	7.8	IO	'504	6.8	IO	'557	4.1	IO	'381	3.0			
IO	'527	6.1	IO	'820	6.2	IO	'853	2.2	IO	'316	7.6	IO	'595	6.9	IO	'594	4.1	IO	'409	3.0			
IO	'569	5.9	II	'857	5.3	IO	'900	1.9	IO	'350	8.3	IO	'660	6.9	IO	'627	4.2	IO	'463	3.4			
IO	'642	6.2	II	'873	5.9	IO	'940	2.0	IO	'389	8.3	IO	'694	7.0	IO	'671	4.1	IO	'531	3.4			
IO	'693	5.3	II	'908	6.2	IO	'000	2.0	IO	'430	8.5	IO	'720	6.7	IO	'704	4.0	IO	'596	3.7			
IO	'737	2.5	II	'943	6.6				IO	'463	8.0	IO	'751	6.7	IO	'729	4.0	IO	'636	3.8			
IO	'771	.9	II	'975	6.2	85			IO	'496	7.7	IO	'780	6.5	IO	'754	4.1	IO	'691	3.8			
IO	'806	1.1							IO	'520	7.6	IO	'812	6.6	IO	'780	4.1	IO	'733	3.6			
IO	'842	1.7		82		IO	'013	7.4	IO	'554	5.4	II	'873	7.2	IO	'814	3.9	IO	'781	3.6			
IO	'871	2.1				IO	'056	7.2	IO	'583	3.1	II	'916	6.9	IO	'856	3.9	IO	'835	4.0			
IO	'890	2.9	IO	'018	6.2	IO	'087	7.7	IO	'610	3.0	II	'953	7.1	IO	'883	3.9	II	'902	4.1			
IO	'927	3.0	IO	'066	5.9	IO	'125	7.5	IO	'642	2.8	II	'987	6.0	IO	'909	4.1	II	'955	3.9			
IO	'957	3.6	IO	'125	5.9	IO	'156	7.6	IO	'675	3.5				IO	'936	4.5						
IO	'983	4.3	IO	'162	6.3	IO	'186	7.6	IO	'695	3.7	95			IO	'960	3.8						
			IO	'196	6.4	IO	'232	7.3	IO	'725	5.0				IO	'987	4.0						
			IO	'236	5.8	IO	'262	7.5	IO	'756	4.9	IO	'025	10.5									
			IO	'280	5.8	IO	'291	7.8	IO	'785	5.3	IO	'063	10.0	IO	'016	2.0	IO	'012	5.0			
			IO	'326	5.9	IO	'331	7.6	IO	'808	6.2	IO	'086	10.0	IO	'047	1.7	IO	'055	4.0			
IO	'021	7.4	IO	'370	4.6	IO	'393	7.4	IO	'840	6.0	IO	'112	9.9	IO	'089	1.4	IO	'102	2.8			
IO	'040	7.7	IO	'408	2.0	IO	'391	5.6	IO	'861	6.2	IO	'135	11.0	IO	'129	1.6	IO	'157	3.1			
IO	'072	7.8	IO	'437	1.5	IO	'417	4.3	IO	'890	6.7	IO	'162	10.8	IO	'164	2.2	IO	'200	3.6			
IO	'102	8.5	IO	'468	1.9	IO	'449	3.6	IO	'919	7.0	IO	'198	11.3	IO	'164	2.2	IO	'237	4.4			
IO	'151	8.9	IO	'499	2.4	IO	'488	3.8	9	'955	7.1	IO	'236	10.9	IO	'164	2.2	IO	'278	4.5			
IO	'188	9.0	IO	'538	2.7	IO	'518	4.4	9	'981	7.3	IO	'278	7.2	IO	'123	3.9	IO	'308	3.0			
IO	'226	9.1	IO	'571	3.4	IO	'544	4.5				IO	'307	3.8	IO	'163	4.3	IO	'350	3.0			
IO	'268	9.2	IO	'597	3.9	IO	'567	5.0	IO	'339	1.2	IO	'307	3.8	IO	'208	5.3	IO	'375	2.7			
IO	'322	9.2	IO	'634	4.5	IO	'596	5.3	IO	'366	1.6	IO	'339	1.2	IO	'248	5.5	IO	'410	2.5			
IO	'343	9.0	IO	'674	4.5	IO	'624	5.6	IO	'023	3.5	IO	'366	1.6	IO	'289	6.0	IO	'448	2.0			
IO	'377	9.3	IO	'716	5.1	IO	'652	5.8	IO	'081	4.2	IO	'397	2.2	IO	'324	6.2	IO	'471	2.1			
IO	'413	9.3	IO	'752	5.4	IO	'688	6.5	IO	'130	4.2	IO	'431	3.2	IO	'355	6.3	IO	'507	1.4			
IO	'450	9.5	II	'780	5.3	IO	'735	6.5	IO	'164	4.2	IO	'457	3.5	IO	'396	6.4	IO	'544	1.8			
IO	'498	9.4	II	'841	5.7	IO	'780	6.8	IO	'199	4.0	IO	'480	4.8	IO	'439	6.5	IO	'584	1.8			
IO	'533	9.4	II	'894	5.8	IO	'816	7.4	IO	'262	4.2	IO	'510	4.8	IO	'492	6.7	IO	'624	1.8			
IO	'574	9.5	II	'944	6.0	IO	'848	7.2	IO	'335	4.0	IO	'552	5.9	IO	'543	6.9	IO	'665	2.1			
IO	'616	9.5	II	'981	6.1	9	'867	7.3	IO	'401	3.8	IO	'587	6.9	IO	'581	6.9	IO	'706	2.4			
IO	'665	9.6				9	'902	7.5	IO	'441	4.2	IO	'614	6.8	IO	'609	6.4	IO	'740	2.8			
IO	'706	9.5				9	'945	7.4	IO	'488	3.8	IO	'643	8.0	IO	'648	6.7	IO	'767	3.2			
IO	'745	9.8				9	'979	7.6	IO	'534	2.7	IO	'678	8.2	IO	'708	6.8	IO	'803	2.8			
IO	'770	9.4							IO	'576	2.6	IO	'729	9.1	IO	'769	6.6	IO	'834	2.9			
IO	'800	8.0	IO	'044	2.0										IO	'809	6.7	IO	'877	2.5			



TABLE 13 (continued).

<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness			
		<i>s</i>			<i>s</i>			<i>s</i>			<i>s</i>			<i>s</i>			<i>s</i>			<i>s</i>			
				I18																			
IO	'377	4.4	IO	'016	5.8	4	'190	3.3	IO	'789	3.4	IO	'388	5.1	IO	'669	3.9	IO	'018	2.8	IO	'147	4.1
IO	'414	4.3	IO	'048	5.7	4	'205	3.0	IO	'821	3.6	IO	'417	5.1	IO	'708	4.6	IO	'059	2.6	IO	'177	3.6
IO	'472	4.1	IO	'092	5.9	5	'255	1.5	IO	'857	4.2	IO	'446	5.0	IO	'747	5.1	IO	'093	2.5	IO	'218	3.2
IO	'539	4.3	IO	'130	6.0	IO	'291	.7	IO	'891	4.4	IO	'489	5.2	IO	'782	5.1	IO	'121	2.7	IO	'251	2.5
IO	'588	4.4	IO	'165	6.0	IO	'325	.7	9	'922	4.0	IO	'530	4.4	IO	'817	5.4	IO	'155	2.7	IO	'279	2.9
IO	'637	4.4	IO	'189	5.8	IO	'363	.6	9	'952	4.3	IO	'553	3.7	IO	'851	5.2	IO	'197	3.0	IO	'306	3.2
IO	'678	4.5	IO	'224	5.8	IO	'392	.8	9	'985	4.6	IO	'586	2.7	IO	'887	4.8	IO	'237	3.0	IO	'338	4.1
IO	'728	4.4	IO	'263	5.9	IO	'418	.8				IO	'620	1.9	IO	'900	4.8	IO	'273	3.0	IO	'370	4.5
IO	'805	4.3	IO	'304	6.1	IO	'443	1.0	125			IO	'642	1.8	IO	'911	4.5	IO	'309	3.2	IO	'393	4.2
II	'866	3.8	IO	'349	6.0	IO	'468	.6	IO	'034	6.1	IO	'661	1.8	IO	'929	4.4	IO	'345	3.2	IO	'419	4.3
II	'922	1.0	IO	'392	6.1	IO	'487	.6	IO	'690	2.4	IO	'690	2.4	IO	'945	3.7	IO	'375	3.2	IO	'450	4.2
II	'977	1.0	IO	'426	6.0	IO	'512	.7	IO	'718	2.8	IO	'718	2.8	13	'985	3.8	IO	'406	2.9	IO	'484	4.6
			IO	'462	6.0	IO	'538	.8	IO	'740	3.1							IO	'440	2.7	IO	'515	4.5
			IO	'462	6.0	IO	'558	.6	IO	'770	3.1	132						IO	'440	2.7	IO	'553	4.7
			IO	'505	3.8	IO	'585	.8	IO	'800	3.2							IO	'468	2.0	IO	'593	5.0
			IO	'535	1.5	IO	'618	.8	IO	'820	3.3	IO	'018	7.0	IO	'018	7.0	IO	'497	1.5	IO	'632	4.7
			IO	'573	.2	IO	'661	.7	II	'845	3.7	IO	'044	6.6	IO	'519	1.8	IO	'519	1.8	IO	'673	5.1
			IO	'603	.6	IO	'691	1.0	II	'876	4.0	IO	'072	7.0	IO	'545	1.2	IO	'545	1.2	IO	'724	5.0
			IO	'636	1.0	IO	'737	.7	II	'904	4.0	IO	'103	7.0	IO	'568	1.0	IO	'568	1.0	IO	'776	5.1
			IO	'654	1.5	IO	'814	.8	II	'939	4.2	IO	'135	7.1	IO	'608	.8	9	'832	4.9	9	'832	4.9
			IO	'674	1.7	IO	'857	.8	II	'978	4.2	IO	'158	6.9	IO	'643	1.3	9	'879	4.8	9	'879	4.8
			IO	'699	2.1	IO	'897	.6				IO	'190	7.1	IO	'668	1.0	9	'929	5.4	9	'929	5.4
			IO	'720	2.7	IO	'933	.4	128			IO	'219	6.9	IO	'689	1.4	9	'978	5.0	9	'978	5.0
			IO	'750	3.0	IO	'965	.7	IO	'570	4.5	IO	'258	5.6	IO	'714	1.5						
			IO	'789	3.5	IO	'987	.6	IO	'609	5.0	IO	'291	4.1	IO	'752	1.7						
			IO	'826	3.7				IO	'637	5.1	IO	'322	2.8	IO	'778	1.9	144					
			IO	'856	3.8	123			IO	'670	4.8	IO	'350	1.9	IO	'797	1.9	IO	'011	2.3	IO	'011	2.3
			II	'887	4.8	IO	'015	2.4	IO	'708	5.2	IO	'108	7.0	IO	'821	2.0	IO	'033	2.0	IO	'033	2.0
			II	'924	4.9	IO	'050	2.9	IO	'754	5.5	IO	'143	7.0	IO	'853	2.0	IO	'047	2.4	IO	'047	2.4
			II	'950	5.1	IO	'093	3.2	II	'812	5.7	IO	'186	6.9	IO	'878	2.0	IO	'075	2.8	IO	'075	2.8
			II	'981	5.6	IO	'149	3.6	II	'873	6.2	IO	'220	7.1	IO	'907	2.4	IO	'094	3.2	IO	'094	3.2
						IO	'199	3.7	II	'951	6.2	IO	'257	7.1	IO	'939	2.4	IO	'120	3.7	IO	'120	3.7
						IO	'242	3.6				IO	'291	7.0	IO	'965	2.8	IO	'156	4.1	IO	'156	4.1
						IO	'280	3.8	126			IO	'552	4.8	9	'985	2.4	IO	'203	4.7	IO	'203	4.7
			IO	'014	7.2	IO	'320	3.8	IO	'022	2.7	IO	'322	7.2	IO	'583	5.0	IO	'246	5.1	IO	'246	5.1
			IO	'052	6.7	IO	'349	3.7	IO	'057	2.1	IO	'401	7.1	IO	'613	5.2	IO	'296	5.2	IO	'296	5.2
			IO	'087	5.0	IO	'384	3.8	IO	'097	1.2	IO	'434	7.3	IO	'650	5.8	IO	'336	5.6	IO	'336	5.6
			IO	'117	3.6	IO	'428	3.8	IO	'135	.7	IO	'486	7.4	IO	'708	5.7	IO	'366	5.6	IO	'366	5.6
			IO	'136	3.1	IO	'466	3.8	IO	'159	.9	IO	'530	7.2	IO	'743	6.5	IO	'386	5.9	IO	'386	5.9
			IO	'158	2.8	IO	'504	3.8	IO	'183	1.0	IO	'573	7.2	IO	'767	6.2	IO	'406	5.9	IO	'406	5.9
			IO	'175	2.3	IO	'548	4.0	IO	'207	1.4	IO	'625	7.2	IO	'801	6.7	IO	'441	5.8	IO	'441	5.8
			IO	'197	3.1	IO	'607	3.8	IO	'238	1.8	IO	'676	7.3	IO	'833	6.7	IO	'478	6.0	IO	'478	6.0
			IO	'225	3.4	IO	'656	3.8	IO	'273	2.4	IO	'729	7.0	IO	'872	6.8	IO	'505	6.0	IO	'505	6.0
			IO	'257	4.1	IO	'696	4.0	IO	'303	2.2	IO	'775	6.9	IO	'910	6.5	IO	'534	6.0	IO	'534	6.0
			IO	'287	4.0	IO	'732	4.0	IO	'325	2.6	IO	'813	6.0	IO	'940	6.6	IO	'563	6.1	IO	'563	6.1
			IO	'310	4.7	IO	'805	3.6	IO	'364	2.6	IO	'839	4.7	IO	'972	6.6	IO	'614	5.9	IO	'614	5.9
			IO	'334	4.9	IO	'855	2.6	IO	'398	3.0	IO	'859	4.8	IO	'991	7.0	IO	'677	6.0	IO	'677	6.0
			IO	'356	5.3	IO	'886	1.7	IO	'443	3.4	IO	'886	4.9	IO	'400	3.2	IO	'708	6.0	IO	'708	6.0
			IO	'380	5.2	9	'916	1.7	IO	'483	2.8	II	'924	6.0	IO	'432	3.4	IO	'741	6.0	IO	'741	6.0
			IO	'418	6.1	9	'934	1.9	IO	'516	3.5	II	'972	6.2	IO	'461	3.4	IO	'781	6.1	IO	'781	6.1
			IO	'453	5.5	9	'961	2.1	IO	'556	3.5				IO	'490	3.6	IO	'816	5.9	IO	'816	5.9
			IO	'499	6.3	9	'980	2.3	IO	'596	3.4	130			IO	'522	3.4	9	'854	6.1	9	'854	6.1
			IO	'557	6.4	IO	'627	6.6	IO	'635	3.6	IO	'009	3.2	IO	'111	2.6	IO	'892	5.9	IO	'892	5.9
			IO	'627	6.6				IO	'688	3.7	IO	'162	2.6	IO	'162	2.6	IO	'923	4.2	IO	'923	4.2
			IO	'670	6.4	124			IO	'750	3.7	IO	'042	3.1	IO	'207	2.8	IO	'958	2.1	IO	'958	2.1
			IO	'705	6.8	IO	'026	4.5	II	'802	3.7	IO	'113	2.6	IO	'249	2.8	9	'988	1.4	9	'988	1.4
			IO	'741	6.8	IO	'070	4.3	II	'848	3.6	IO	'195	1.3	IO	'306	2.8	IO	'698	1.2			
			IO	'790	6.9	IO	'137	4.8	II	'900	3.3	IO	'227	.9	IO	'357	2.8	IO	'744	1.4	148		
			IO	'838	6.7	IO	'183	4.6	II	'966	3.8	IO	'248	.6	IO	'428	2.6	IO	'780	2.1	IO	'020	1.1
			IO	'879	7.0	IO	'229	4.5				IO	'303	.9	IO	'556	2.5	IO	'814	2.4	IO	'053	2.0
			IO	'920	7.1	IO	'274	4.9															

TABLE 13 (continued).

n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	n	phase	bright- ness	
		s			s			s			s			s			s			s			s	
IO	'897	3.4	IO	'880	3.3			158	IO	'056	7.1	IO	'286	4.6	IO	'700	9			185	IO	'964	3.2	
IO	'914	2.4	IO	'921	3.5	IO	'026	2.9	IO	'096	7.2	IO	'328	4.8	IO	'740	1.4	IO	'013	3.0				
II	'942	7	IO	'950	3.6	IO	'066	1.8	IO	'132	7.4	IO	'366	4.7	IO	'784	1.5	IO	'045	2.7				from J.D.
II	'976	4	IO	'998	3.4	IO	'100	2.0	IO	'159	7.3	IO	'396	4.8	IO	'829	1.6	IO	'073	3.0				6000 to 7000
						IO	'133	3.1	IO	'196	7.3	IO	'422	4.8	IO	'861	2.2	IO	'096	3.0				
	149			153		IO	'155	3.0	IO	'241	8.0	IO	'444	4.8	IO	'904	2.2	IO	'135	3.8	IO	'032	1.1	
IO	'018	7	IO	'017	4.4	IO	'181	2.9	IO	'274	7.9	IO	'469	4.8	IO	'954	2.4	IO	'177	4.1	IO	'077	1.7	
IO	'060	9	IO	'044	4.4	IO	'210	3.3	IO	'310	7.6	IO	'492	4.9	II	'982	2.6	IO	'216	4.6	IO	'179	2.5	
IO	'087	1.2	IO	'072	4.9	IO	'242	4.1	IO	'346	7.7	IO	'516	4.8				IO	'246	5.0	IO	'252	3.0	
IO	'118	1.5	IO	'091	4.9	IO	'270	4.3	IO	'377	7.5	IO	'552	4.8		182		IO	'287	5.0	IO	'356	3.6	
IO	'141	2.0	IO	'123	5.1	IO	'303	4.4	IO	'410	7.8	IO	'587	4.9				IO	'326	5.6	IO	'470	3.6	
IO	'162	2.6	IO	'162	5.1	IO	'347	4.6	IO	'450	7.7	IO	'623	4.8	IO	'048	5.4	IO	'358	5.6	IO	'545	4.1	
IO	'184	2.5	IO	'210	5.4	IO	'401	4.9	IO	'481	7.8	IO	'664	4.8	IO	'117	2.3	IO	'401	5.8	IO	'605	4.1	
IO	'219	3.2	IO	'252	5.4	IO	'447	4.8	IO	'505	7.5	IO	'699	5.0	IO	'145	1.8	IO	'439	6.0	IO	'687	4.3	
IO	'249	3.1	IO	'287	5.6	IO	'477	4.9	IO	'527	7.9	IO	'716	4.4	IO	'169	1.5	IO	'480	6.0	IO	'730	4.3	
IO	'280	3.7	IO	'322	5.7	IO	'504	5.0	IO	'554	7.7	IO	'747	3.6	IO	'188	1.6	IO	'529	6.1	IO	'773	4.1	
IO	'321	3.9	IO	'352	5.8	IO	'527	5.0	IO	'574	7.1	IO	'786	1.4	IO	'206	2.0	IO	'580	6.3	IO	'815	4.2	
IO	'356	3.9	IO	'382	5.8	IO	'558	5.1	IO	'595	7.3	IO	'814	8	IO	'226	2.0	IO	'622	6.0	IO	'868	4.2	
IO	'391	4.1	IO	'415	5.7	IO	'620	5.1	IO	'624	5.6	IO	'844	1.0	IO	'249	2.2	IO	'664	6.2	IO	'925	3.0	
IO	'433	4.0	IO	'452	6.0	IO	'679	5.0	IO	'660	5.1	IO	'874	1.5	IO	'283	3.0	IO	'712	6.1	IO	'978	1.7	
IO	'486	4.2	IO	'485	5.6	IO	'772	5.2	IO	'693	4.3	IO	'909	2.0	IO	'320	2.8	IO	'762	6.3				after
IO	'531	4.4	IO	'516	5.7	IO	'812	5.2	IO	'719	4.9	IO	'949	2.2	IO	'356	4.1	IO	'792	6.4	IO	'019	2.1	
IO	'570	4.4	IO	'551	5.9	IO	'841	5.1	IO	'745	4.2	II	'982	2.6	IO	'388	4.1	IO	'822	6.4	II	'889	6.4	
IO	'603	4.4	IO	'598	6.1	9	'870	5.3	IO	'770	5.2			178	IO	'419	4.4	II	'856	6.3	IO	'974	3.8	
IO	'641	4.3	IO	'643	6.4	9	'897	5.1	IO	'799	5.3				IO	'452	4.7	II	'926	6.1	IO	'094	2.9	
IO	'674	4.3	IO	'673	6.0	9	'924	5.2	IO	'822	5.9				IO	'486	4.8				IO	'241	3.5	
IO	'699	4.0	IO	'703	5.5	9	'957	5.0	IO	'852	5.8	IO	'013	2.3	IO	'519	5.1	II	'974	3.8	IO	'356	4.0	
IO	'739	4.6	IO	'727	4.7	9	'985	5.0	IO	'887	6.0	IO	'053	2.4	IO	'556	5.0				IO	'463	3.9	
IO	'785	4.4	IO	'752	3.1			162	II	'934	6.4	IO	'108	2.6	IO	'605	5.3				IO	'574	3.8	
IO	'839	4.6	IO	'770	1.7				II	'978	6.6	IO	'171	3.0	IO	'640	5.3	IO	'030	2.0				
II	'883	4.5	IO	'794	1.1	IO	'060	1.8				IO	'262	3.0	IO	'666	5.2	IO	'078	2.2				
II	'919	4.0	IO	'819	1.6	IO	'091	2.3			165	IO	'334	3.1	IO	'699	5.2	IO	'114	2.2				
II	'964	2.2	IO	'855	1.9	IO	'129	2.0	IO	'020	4.8	IO	'403	2.8	IO	'716	5.3	IO	'149	2.4				
			IO	'903	2.6	IO	'174	2.1	IO	'053	5.1	IO	'538	3.0	IO	'747	5.5	IO	'207	2.5				
	152		9	'931	3.3	IO	'202	2.2	IO	'077	5.0	IO	'609	3.1	IO	'775	5.4	IO	'237	2.6				
IO	'046	3.3	9	'954	3.4	IO	'230	1.9	IO	'107	4.2	IO	'668	2.2	IO	'806	5.3	IO	'276	2.8				
IO	'074	3.6	9	'970	3.5	IO	'291	2.4	IO	'133	3.2	IO	'708	1.4	9	'833	5.4	IO	'316	3.0	IO	'024	3.9	
IO	'109	3.8	9	'990	3.8	IO	'326	2.4	IO	'165	2.3	IO	'736	7	9	'878	5.4	IO	'369	3.3	IO	'067	4.1	
IO	'151	3.3				IO	'357	1.9	IO	'197	1.0	IO	'763	8	9	'927	5.7	IO	'412	3.0	IO	'095	4.2	
IO	'184	3.5		155		IO	'392	1.9	IO	'223	1.1	IO	'808	1.1	9	'974	5.6	IO	'468	3.3	IO	'127	4.8	
IO	'242	3.5	IO	'032	7.6	IO	'422	2.1	IO	'248	1.0	IO	'829	1.3				IO	'523	3.3	IO	'165	4.6	
IO	'277	3.5	IO	'062	7.5	IO	'454	2.1	IO	'274	1.4	IO	'862	1.5	IO	'013	4.8	IO	'567	3.2	IO	'196	4.6	
IO	'292	3.6	IO	'092	7.2	IO	'483	2.0	IO	'304	2.3	IO	'894	1.8	IO	'052	4.7	IO	'606	2.8	IO	'229	4.6	
IO	'311	3.3	IO	'114	6.5	IO	'525	2.0	IO	'325	2.3	IO	'931	2.0	IO	'107	4.8	IO	'642	3.0	IO	'277	4.6	
IO	'334	3.5	IO	'145	4.6	IO	'550	2.1	IO	'349	2.6	IO	'956	2.2	IO	'140	4.9	IO	'678	3.1	IO	'313	4.7	
IO	'357	3.5	IO	'175	3.3	IO	'586	1.8	IO	'373	2.8	9	'990	2.2	IO	'165	4.8	IO	'716	3.0	IO	'352	4.9	
IO	'382	3.3	IO	'202	3.1	IO	'633	1.9	IO	'403	2.9			179	IO	'195	4.9	IO	'753	2.2	IO	'397	4.5	
IO	'405	3.5	IO	'227	3.6	IO	'660	2.2	IO	'448	3.6				IO	'221	4.9	IO	'784	1.7	IO	'428	4.5	
IO	'428	3.5	IO	'254	4.0	IO	'676	2.1	IO	'480	3.4				IO	'255	4.8	IO	'819	1.5	IO	'460	5.1	
IO	'446	3.5	IO	'292	4.6	IO	'693	2.2	IO	'515	4.0	IO	'012	2.4	IO	'255	4.8	IO	'848	1.5	IO	'492	5.1	
IO	'458	3.6	IO	'335	5.3	IO	'707	2.1	IO	'552	4.2	IO	'037	2.6	IO	'301	4.7	IO	'877	1.6	IO	'546	4.8	
IO	'473	3.5	IO	'368	5.8	5	'726	3.1	IO	'586	4.0	IO	'064	2.4	IO	'338	4.5	IO	'895	1.8	IO	'604	4.5	
IO	'499	3.0	IO	'390	6.1	5	'736	3.8	IO	'616	4.3	IO	'089	2.6	IO	'379	4.0	IO	'920	1.8	IO	'639	2.7	
IO	'524	3.6	IO	'415	6.1	5	'741	4.3	IO	'651	4.4	IO	'117	2.6	IO	'426	2.6	IO	'920	1.8	IO	'692	8	
4	'540	3.8	IO	'436	6.4	5	'747	4.8	IO	'688	4.5	IO	'149	2.7	IO	'461	2.7	9	'936	2.0	IO	'725	4	
4	'546	4.2	IO	'462	6.8	5	'773	6.5	IO	'726	4.8	IO	'184	2.8	IO	'500	3.0	9	'950	1.9	IO	'764	1.1	
4	'558	5.3	IO	'489	6.5	5	'809	5.3	IO	'764	4.7	IO	'222	2.8	IO	'531	3.2	9	'977	2.0	IO	'799	1.6	
4	'570	6.4	IO	'509	6.7	5	'819	3.9	9	'799	4.7	IO	'252	2.8	IO	'558	3.5				IO	'832	2.0	
4	'582	7.4	IO	'537	6.8	5	'825	3.2	9	'847	4.8	IO	'273	2.8	IO	'584	3.8				IO	'866	2.3	
4	'594	8.4	IO	'579	6.8	6	'835	3.2	9	'913	4.8	IO	'298	2.7	IO	'616	4.1				IO	'907	3.1	
4	'608	7.4	IO	'620	7.4	IO	'861	2.2	9	'978	4.9	IO	'336	2.7	IO	'653	4.0				IO	'949	3.4	
5	'620	6.7	IO	'656	7.1	IO	'883	2.2				IO	'377	2.7	IO	'690	4.3				IO	'980	3.4	
5	'639	5.7	IO	'703	7.1	IO	'914	2.2			175	IO	'419	2.8	IO	'716	4.4							
4	'655	3.9	IO	'748	7.1	IO	'933	2.1				IO	'447	2.6	IO	'744	4.6				IO	'014	2.4	
IO	'674	3.3	IO	'785	7.3	IO	'948	2.0	IO	'022	3.1	IO	'481	1.8	IO	'772	4.6	IO	'088	2.8	IO	'016	1.3	
IO	'701	3.5	IO	'817	7.1	IO	'971	2.0	IO	'068	3.4	IO	'507											





TABLE 14 (continued).

31	57	75		121		193
a 1 <sup>o</sup> 13 <sup>m</sup> 7	A s 5 14 <sup>o</sup> 0	c s 0 13 <sup>m</sup> 3	b s 2 <sup>o</sup> 6 14 <sup>m</sup> 7	a s 0 13 <sup>m</sup> 2	b s 2 <sup>o</sup> 7 14 <sup>m</sup> 1	a s 5 13 <sup>m</sup> 3
b 5 <sup>o</sup> 14 <sup>m</sup> 3	a b 2 <sup>o</sup> 5 14 <sup>m</sup> 1	d 4 <sup>o</sup> 3 13 <sup>m</sup> 6	c 4 <sup>o</sup> 3 15 <sup>m</sup> 1	b 2 <sup>o</sup> 7 13 <sup>m</sup> 6	c 4 <sup>o</sup> 7 14 <sup>m</sup> 7	b 3 <sup>o</sup> 8 13 <sup>m</sup> 9
c 8 <sup>o</sup> 14 <sup>m</sup> 6	b c 5 <sup>o</sup> 2 14 <sup>m</sup> 6	e 8 <sup>o</sup> 0 14 <sup>m</sup> 2	d 5 <sup>o</sup> 4 15 <sup>m</sup> 2	c 5 <sup>o</sup> 2 14 <sup>m</sup> 2	d 5 <sup>o</sup> 9 14 <sup>m</sup> 9	c 6 <sup>o</sup> 1 14 <sup>m</sup> 6
d 10 <sup>o</sup> 6 15 <sup>m</sup> 2	c 6 <sup>o</sup> 7 15 <sup>m</sup> 0	f 10 <sup>o</sup> 7 14 <sup>m</sup> 6	91	d 7 <sup>o</sup> 6 14 <sup>m</sup> 6	144	172
		g 13 <sup>o</sup> 4 14 <sup>m</sup> 8	a 0 13 <sup>o</sup> 9		A' 0 13 <sup>o</sup> 4	a 1 <sup>o</sup> 0 13 <sup>o</sup> 3
32	58	76	b 4 <sup>o</sup> 0 14 <sup>o</sup> 3	122	a' 2 <sup>o</sup> 3 13 <sup>o</sup> 6	b 3 <sup>o</sup> 7 13 <sup>o</sup> 6
a 0 13 <sup>o</sup> 7	a 0 13 <sup>o</sup> 7	a 1 <sup>o</sup> 0 14 <sup>o</sup> 3	c 5 <sup>o</sup> 5 14 <sup>o</sup> 5	a 0 14 <sup>o</sup> 1	b' 3 <sup>o</sup> 9 14 <sup>o</sup> 0	c 6 <sup>o</sup> 1 14 <sup>o</sup> 7
b 4 <sup>o</sup> 1 14 <sup>o</sup> 2	b 2 <sup>o</sup> 5 14 <sup>o</sup> 2	b 3 <sup>o</sup> 1 14 <sup>o</sup> 4	d 7 <sup>o</sup> 1 14 <sup>o</sup> 8	b 1 <sup>o</sup> 8 14 <sup>o</sup> 5	c' 6 <sup>o</sup> 0 14 <sup>o</sup> 6	175
c 9 <sup>o</sup> 0 14 <sup>o</sup> 9	c 5 <sup>o</sup> 0 14 <sup>o</sup> 6	c 5 <sup>o</sup> 7 15 <sup>o</sup> 2	e 8 <sup>o</sup> 3 15 <sup>o</sup> 0			194
	d 6 <sup>o</sup> 4 14 <sup>o</sup> 9			123	see B.A.N. No. 269	a 0 13 <sup>o</sup> 4
	e 8 <sup>o</sup> 9 15 <sup>o</sup> 2	77	95	a 0 13 <sup>o</sup> 7		b 2 <sup>o</sup> 7 14 <sup>o</sup> 0
33		a 0 13 <sup>o</sup> 6	a 0 12 <sup>o</sup> 7	b 1 <sup>o</sup> 9 14 <sup>o</sup> 2	148	c 5 <sup>o</sup> 5 14 <sup>o</sup> 8
a 0 13 <sup>o</sup> 9		b 3 <sup>o</sup> 6 14 <sup>o</sup> 3	b 4 <sup>o</sup> 0 13 <sup>o</sup> 0	c 3 <sup>o</sup> 6 15 <sup>o</sup> 0	a 0 14 <sup>o</sup> 2	a 13 <sup>o</sup> 7
b 3 <sup>o</sup> 0 14 <sup>o</sup> 3	a 1 <sup>o</sup> 0 14 <sup>o</sup> 4	c 6 <sup>o</sup> 4 14 <sup>o</sup> 9	c 6 <sup>o</sup> 1 13 <sup>o</sup> 3		b 3 <sup>o</sup> 7 14 <sup>o</sup> 6	b 14 <sup>o</sup> 7
c 5 <sup>o</sup> 2 14 <sup>o</sup> 7	b 3 <sup>o</sup> 2 14 <sup>o</sup> 5	d 9 <sup>o</sup> 2 13 <sup>o</sup> 7	d 9 <sup>o</sup> 2 13 <sup>o</sup> 7	124	c 5 <sup>o</sup> 6 14 <sup>o</sup> 8	199
d 7 <sup>o</sup> 6 15 <sup>o</sup> 2	c 5 <sup>o</sup> 4 14 <sup>o</sup> 8	e 12 <sup>o</sup> 1 14 <sup>o</sup> 1	e 12 <sup>o</sup> 1 14 <sup>o</sup> 1	a 5 14 <sup>o</sup> 2		a 0 13 <sup>o</sup> 5
36	61	78	96	b 2 <sup>o</sup> 8 14 <sup>o</sup> 5	149	b 3 <sup>o</sup> 4 14 <sup>o</sup> 0
a 0 13 <sup>o</sup> 1	a 0 12 <sup>o</sup> 7	a 0 13 <sup>o</sup> 6	a 0 13 <sup>o</sup> 1	c 4 <sup>o</sup> 8 14 <sup>o</sup> 9	a 0 14 <sup>o</sup> 1	c 5 <sup>o</sup> 6 14 <sup>o</sup> 7
b 4 <sup>o</sup> 2 13 <sup>o</sup> 3	b 2 <sup>o</sup> 5 13 <sup>o</sup> 0	b 3 <sup>o</sup> 4 13 <sup>o</sup> 9	b 4 <sup>o</sup> 3 13 <sup>o</sup> 5		b 1 <sup>o</sup> 4 14 <sup>o</sup> 5	202
c 7 <sup>o</sup> 2 14 <sup>o</sup> 0	c 4 <sup>o</sup> 7 13 <sup>o</sup> 2	c 6 <sup>o</sup> 5 14 <sup>o</sup> 3	c 6 <sup>o</sup> 7 13 <sup>o</sup> 7	125	c 4 <sup>o</sup> 1 14 <sup>o</sup> 8	a 0 13 <sup>o</sup> 1
d 9 <sup>o</sup> 9 14 <sup>o</sup> 6	d 7 <sup>o</sup> 6 13 <sup>o</sup> 5	d 9 <sup>o</sup> 2 14 <sup>o</sup> 9	d 10 <sup>o</sup> 1 14 <sup>o</sup> 7	a 0 13 <sup>o</sup> 5		b 3 <sup>o</sup> 2 13 <sup>o</sup> 5
37	62	80	100	b 2 <sup>o</sup> 7 14 <sup>o</sup> 1	152	c 6 <sup>o</sup> 0 13 <sup>o</sup> 9
a 5 13 <sup>o</sup> 8	A 5 13 <sup>o</sup> 7	a 0 13 <sup>o</sup> 6	a 0 13 <sup>o</sup> 6	c 4 <sup>o</sup> 9 14 <sup>o</sup> 7	a 0 13 <sup>o</sup> 4	d 8 <sup>o</sup> 5 14 <sup>o</sup> 6
b 4 <sup>o</sup> 6 14 <sup>o</sup> 3	a b 1 <sup>o</sup> 9 14 <sup>o</sup> 0	b 2 <sup>o</sup> 5 13 <sup>o</sup> 7	b 2 <sup>o</sup> 5 13 <sup>o</sup> 7	d 6 <sup>o</sup> 5 15 <sup>o</sup> 1	b 4 <sup>o</sup> 0 14 <sup>o</sup> 0	205
c 6 <sup>o</sup> 4 14 <sup>o</sup> 7	b c 3 <sup>o</sup> 6 14 <sup>o</sup> 3	c 5 <sup>o</sup> 0 14 <sup>o</sup> 3	c 3 <sup>o</sup> 9 14 <sup>o</sup> 1		c 7 <sup>o</sup> 4 14 <sup>o</sup> 5	A 0 13 <sup>o</sup> 2
41	c d 5 <sup>o</sup> 1 14 <sup>o</sup> 6	d 7 <sup>o</sup> 7 14 <sup>o</sup> 7	d 6 <sup>o</sup> 8 14 <sup>o</sup> 5	a 1 <sup>o</sup> 0 14 <sup>o</sup> 7	d 9 <sup>o</sup> 1 15 <sup>o</sup> 1	a 2 <sup>o</sup> 8 13 <sup>o</sup> 5
a 5 14 <sup>o</sup> 0	d 6 <sup>o</sup> 6 14 <sup>o</sup> 8	82	103	b 2 <sup>o</sup> 6 14 <sup>o</sup> 9		b 4 <sup>o</sup> 8 13 <sup>o</sup> 8
b 4 <sup>o</sup> 4 14 <sup>o</sup> 5		a 0 13 <sup>o</sup> 8	a 0 13 <sup>o</sup> 8	c 4 <sup>o</sup> 0 15 <sup>o</sup> 1	153	c 7 <sup>o</sup> 2 14 <sup>o</sup> 3
c 7 <sup>o</sup> 1 14 <sup>o</sup> 8	a 5 13 <sup>o</sup> 6	b 2 <sup>o</sup> 9 14 <sup>o</sup> 4	b 2 <sup>o</sup> 8 14 <sup>o</sup> 6		a 5 14 <sup>o</sup> 3	d 9 <sup>o</sup> 5 14 <sup>o</sup> 7
43	b 3 <sup>o</sup> 0 13 <sup>o</sup> 9	c 5 <sup>o</sup> 7 14 <sup>o</sup> 9	105	127	b 2 <sup>o</sup> 8 14 <sup>o</sup> 7	207
a 5 13 <sup>o</sup> 9	c d 5 <sup>o</sup> 9 14 <sup>o</sup> 6	d 7 <sup>o</sup> 4 15 <sup>o</sup> 1	a 0 13 <sup>o</sup> 5	a 0 13 <sup>o</sup> 6		a' 0 14 <sup>o</sup> 4
b 4 <sup>o</sup> 0 14 <sup>o</sup> 3	d 8 <sup>o</sup> 3 14 <sup>o</sup> 9	83	b 2 <sup>o</sup> 2 14 <sup>o</sup> 0	b 3 <sup>o</sup> 0 14 <sup>o</sup> 3	155	b' 1 <sup>o</sup> 3 14 <sup>o</sup> 6
c 6 <sup>o</sup> 2 14 <sup>o</sup> 6		a 0 13 <sup>o</sup> 5	c 4 <sup>o</sup> 5 14 <sup>o</sup> 3	c 5 <sup>o</sup> 3 14 <sup>o</sup> 7	a 0 13 <sup>o</sup> 5	c' 3 <sup>o</sup> 3 15 <sup>o</sup> 1
d 8 <sup>o</sup> 5 15 <sup>o</sup> 0	64	b 2 <sup>o</sup> 2 14 <sup>o</sup> 0	109	128	b 4 <sup>o</sup> 2 14 <sup>o</sup> 0	208
44	A-3 <sup>o</sup> 0 13 <sup>o</sup> 7	c 2 <sup>o</sup> 0 13 <sup>o</sup> 0	a 0 13 <sup>o</sup> 4	a 0 13 <sup>o</sup> 4	c 6 <sup>o</sup> 6 14 <sup>o</sup> 6	a 0 13 <sup>o</sup> 1
a 1 <sup>o</sup> 0 14 <sup>o</sup> 7	a b 1 <sup>o</sup> 0 13 <sup>o</sup> 8	d 4 <sup>o</sup> 6 13 <sup>o</sup> 1	A 0 13 <sup>o</sup> 4	b 4 <sup>o</sup> 4 14 <sup>o</sup> 3	d 8 <sup>o</sup> 1 14 <sup>o</sup> 8	b 3 <sup>o</sup> 2 13 <sup>o</sup> 4
b 3 <sup>o</sup> 6 15 <sup>o</sup> 0	b c 5 <sup>o</sup> 5 14 <sup>o</sup> 5	e 8 <sup>o</sup> 0 13 <sup>o</sup> 3	a 2 <sup>o</sup> 0 13 <sup>o</sup> 6	c 7 <sup>o</sup> 1 14 <sup>o</sup> 8		c 6 <sup>o</sup> 1 13 <sup>o</sup> 8
c 5 <sup>o</sup> 4 15 <sup>o</sup> 2	c d 8 <sup>o</sup> 9 14 <sup>o</sup> 7	85	b 5 <sup>o</sup> 3 14 <sup>o</sup> 5	130	158	216
45	67	a 0 13 <sup>o</sup> 5	110	a 0 13 <sup>o</sup> 2	a 0 14 <sup>o</sup> 0	A 0 13 <sup>o</sup> 3
a 5 13 <sup>o</sup> 2	a 0 14 <sup>o</sup> 5	b 4 <sup>o</sup> 9 14 <sup>o</sup> 3	A 0 13 <sup>o</sup> 7	b 2 <sup>o</sup> 8 14 <sup>o</sup> 0	b 2 <sup>o</sup> 8 14 <sup>o</sup> 5	a 2 <sup>o</sup> 8 14 <sup>o</sup> 0
b 3 <sup>o</sup> 6 13 <sup>o</sup> 5	b 3 <sup>o</sup> 0 14 <sup>o</sup> 8	c 7 <sup>o</sup> 4 14 <sup>o</sup> 8	a 2 <sup>o</sup> 9 14 <sup>o</sup> 3	c 5 <sup>o</sup> 3 14 <sup>o</sup> 6	c 5 <sup>o</sup> 0 14 <sup>o</sup> 9	b 5 <sup>o</sup> 6 14 <sup>o</sup> 3
c 5 <sup>o</sup> 6 13 <sup>o</sup> 7	c 4 <sup>o</sup> 4 15 <sup>o</sup> 1	86	b 5 <sup>o</sup> 1 14 <sup>o</sup> 7	132	a 13 <sup>o</sup> 2	c 7 <sup>o</sup> 5 14 <sup>o</sup> 6
d 9 <sup>o</sup> 2 14 <sup>o</sup> 3	69	a 0 13 <sup>o</sup> 4	111	a 0 13 <sup>o</sup> 2	b 14 <sup>o</sup> 2	218
e 10 <sup>o</sup> 9 14 <sup>o</sup> 7	a 0 14 <sup>o</sup> 0	b 2 <sup>o</sup> 9 13 <sup>o</sup> 7	a 1 <sup>o</sup> 0 14 <sup>o</sup> 0	b 2 <sup>o</sup> 8 14 <sup>o</sup> 0	c 14 <sup>o</sup> 7	a 0 14 <sup>o</sup> 0
48	b 2 <sup>o</sup> 9 14 <sup>o</sup> 6	c 5 <sup>o</sup> 1 14 <sup>o</sup> 0	a 1 <sup>o</sup> 0 14 <sup>o</sup> 0	c 5 <sup>o</sup> 3 14 <sup>o</sup> 6	a 0 13 <sup>o</sup> 5	A 0 13 <sup>o</sup> 3
A 0 13 <sup>o</sup> 1	c d 5 <sup>o</sup> 0 14 <sup>o</sup> 9	d 7 <sup>o</sup> 0 14 <sup>o</sup> 5	b 2 <sup>o</sup> 0 14 <sup>o</sup> 1	138	b 2 <sup>o</sup> 4 14 <sup>o</sup> 1	a 2 <sup>o</sup> 8 14 <sup>o</sup> 0
a 4 <sup>o</sup> 2 13 <sup>o</sup> 4	d 7 <sup>o</sup> 2 15 <sup>o</sup> 1	e 8 <sup>o</sup> 8 14 <sup>o</sup> 7	c 4 <sup>o</sup> 3 14 <sup>o</sup> 8	140	c 4 <sup>o</sup> 6 14 <sup>o</sup> 6	b 5 <sup>o</sup> 6 14 <sup>o</sup> 3
b 6 <sup>o</sup> 1 13 <sup>o</sup> 6	70	88	114	a 5 14 <sup>o</sup> 5		c 7 <sup>o</sup> 5 14 <sup>o</sup> 6
c 9 <sup>o</sup> 7 14 <sup>o</sup> 0	see B.A.N. No. 256	a 5 14 <sup>o</sup> 6	a 1 <sup>o</sup> 0 14 <sup>o</sup> 2	b 2 <sup>o</sup> 5 15 <sup>o</sup> 0	159	219
d 13 <sup>o</sup> 1 14 <sup>o</sup> 5	72	b 3 <sup>o</sup> 3 15 <sup>o</sup> 1	b 3 <sup>o</sup> 2 14 <sup>o</sup> 5	141	a 13 <sup>o</sup> 2	a' 0 13 <sup>o</sup> 2
e 15 <sup>o</sup> 0 14 <sup>o</sup> 8	a 0 13 <sup>o</sup> 0	c 4 <sup>o</sup> 7 15 <sup>o</sup> 2	89	a 0 13 <sup>o</sup> 5	b 14 <sup>o</sup> 2	b' 3 <sup>o</sup> 8 13 <sup>o</sup> 7
50	b 3 <sup>o</sup> 1 13 <sup>o</sup> 2	A 0 13 <sup>o</sup> 6	a 0 13 <sup>o</sup> 8	c 6 <sup>o</sup> 6 14 <sup>o</sup> 6	c 14 <sup>o</sup> 7	c' 6 <sup>o</sup> 0 14 <sup>o</sup> 3
a 0 13 <sup>o</sup> 4	b 3 <sup>o</sup> 1 13 <sup>o</sup> 2	a 3 <sup>o</sup> 0 13 <sup>o</sup> 9	b 1 <sup>o</sup> 8 14 <sup>o</sup> 2	d 8 <sup>o</sup> 3 14 <sup>o</sup> 8	a 0 13 <sup>o</sup> 4	221
b 3 <sup>o</sup> 3 14 <sup>o</sup> 0	c 6 <sup>o</sup> 0 13 <sup>o</sup> 4	b 5 <sup>o</sup> 3 14 <sup>o</sup> 3	c 3 <sup>o</sup> 3 14 <sup>o</sup> 9	162	b 3 <sup>o</sup> 9 14 <sup>o</sup> 1	a 0 13 <sup>o</sup> 9
c 6 <sup>o</sup> 4 14 <sup>o</sup> 5		c d 7 <sup>o</sup> 0 14 <sup>o</sup> 5	116	a 0 13 <sup>o</sup> 1	c 6 <sup>o</sup> 6 14 <sup>o</sup> 9	b 2 <sup>o</sup> 5 14 <sup>o</sup> 6
d 8 <sup>o</sup> 6 15 <sup>o</sup> 2	74	e 8 <sup>o</sup> 8 14 <sup>o</sup> 7	a 0 13 <sup>o</sup> 8	b 3 <sup>o</sup> 0 13 <sup>o</sup> 4		c 4 <sup>o</sup> 9 15 <sup>o</sup> 1
53	a 0 13 <sup>o</sup> 7	90	b 1 <sup>o</sup> 8 14 <sup>o</sup> 2	c 7 <sup>o</sup> 2 14 <sup>o</sup> 5	a 0 13 <sup>o</sup> 5	218
a 5 13 <sup>o</sup> 1	b 2 <sup>o</sup> 8 14 <sup>o</sup> 1	a 5 13 <sup>o</sup> 4	c 3 <sup>o</sup> 3 14 <sup>o</sup> 9	a 1 <sup>o</sup> 0 14 <sup>o</sup> 1	b 2 <sup>o</sup> 4 14 <sup>o</sup> 1	a 0 14 <sup>o</sup> 0
b 2 <sup>o</sup> 6 13 <sup>o</sup> 4	c 5 <sup>o</sup> 0 14 <sup>o</sup> 3	b 3 <sup>o</sup> 1 13 <sup>o</sup> 8	a 0 13 <sup>o</sup> 8	b 3 <sup>o</sup> 2 14 <sup>o</sup> 8	c 4 <sup>o</sup> 6 14 <sup>o</sup> 6	b 2 <sup>o</sup> 5 14 <sup>o</sup> 6
c 6 <sup>o</sup> 1 14 <sup>o</sup> 0	d 7 <sup>o</sup> 5 14 <sup>o</sup> 5	c 6 <sup>o</sup> 7 14 <sup>o</sup> 9	118	142		c 4 <sup>o</sup> 9 15 <sup>o</sup> 1
d 8 <sup>o</sup> 5 14 <sup>o</sup> 6	e 10 <sup>o</sup> 0 15 <sup>o</sup> 1	a 0 14 <sup>o</sup> 4	a 5 13 <sup>o</sup> 4	a 1 <sup>o</sup> 0 14 <sup>o</sup> 1	a 0 13 <sup>o</sup> 8	219
			b 3 <sup>o</sup> 1 13 <sup>o</sup> 8	b 3 <sup>o</sup> 2 14 <sup>o</sup> 8	b 2 <sup>o</sup> 6 14 <sup>o</sup> 4	a' 0 13 <sup>o</sup> 2
			c 6 <sup>o</sup> 7 14 <sup>o</sup> 9	143		b' 3 <sup>o</sup> 8 13 <sup>o</sup> 7
				a 0 13 <sup>o</sup> 6	163	c' 6 <sup>o</sup> 0 14 <sup>o</sup> 3
				b 3 <sup>o</sup> 8 13 <sup>o</sup> 8	a 0 13 <sup>o</sup> 6	221
				c 6 <sup>o</sup> 6 14 <sup>o</sup> 6	b 2 <sup>o</sup> 5 14 <sup>o</sup> 0	a 0 13 <sup>o</sup> 9
				d 8 <sup>o</sup> 3 14 <sup>o</sup> 8	c 4 <sup>o</sup> 7 14 <sup>o</sup> 7	b 2 <sup>o</sup> 9 14 <sup>o</sup> 3
				165		c 5 <sup>o</sup> 1 14 <sup>o</sup> 6
				a 5 13 <sup>o</sup> 6	190	222
				b 3 <sup>o</sup> 3 14 <sup>o</sup> 2	a 1 <sup>o</sup> 0 13 <sup>o</sup> 7	A 0 13 <sup>o</sup> 7
				c 5 <sup>o</sup> 7 14 <sup>o</sup> 8	b 4 <sup>o</sup> 0 14 <sup>o</sup> 2	a 3 <sup>o</sup> 6 14 <sup>o</sup> 0
				169	c 5 <sup>o</sup> 4 14 <sup>o</sup> 8	b 5 <sup>o</sup> 9 14 <sup>o</sup> 2
				a 0 13 <sup>o</sup> 2	191	c 8 <sup>o</sup> 1 14 <sup>o</sup> 6
				a' 0 13 <sup>o</sup> 2	a 0 13 <sup>o</sup> 3	
				b 3 <sup>o</sup> 0 13 <sup>o</sup> 7	b 3 <sup>o</sup> 0 13 <sup>o</sup> 7	