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Provisional ephemerides of 25 new variable stars in or near the constellation Crux

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COMMUNICATIONS FROM THE OBSERVATORY AT LEIDEN.

Provisional ephemerides of 25 new variable stars in or near the constellation Crux, by *P. P. Bruna.*

Of the 25 new variable stars published in the present paper, 15 are to be found in the constellation Crux, 8 in Centaurus, 2 in Musca. They were discovered by comparing 7 pairs of plates in the blink microscope. The estimates were made on the same plates as used by P. TH. OOSTERHOFF (*B. A. N.* 148) and in the course of the work supplemented by new ones, taken

by H. VAN GENT at Johannesburg. The ordinary centre of the plates is $11^{\text{h}}51^{\text{m}}, -61^{\circ}30'$ (1875), The normal exposure time is 30 min.

The variable stars are listed in Table 1 together with observational data for each object. Of these 25 variables only x occurs in the *C. P. D.* viz: $-62^{\circ}28'17''$ ($9^{\text{m}}.5$) = *H. D.* 109176 ($10^{\text{m}}.2$; Ko).

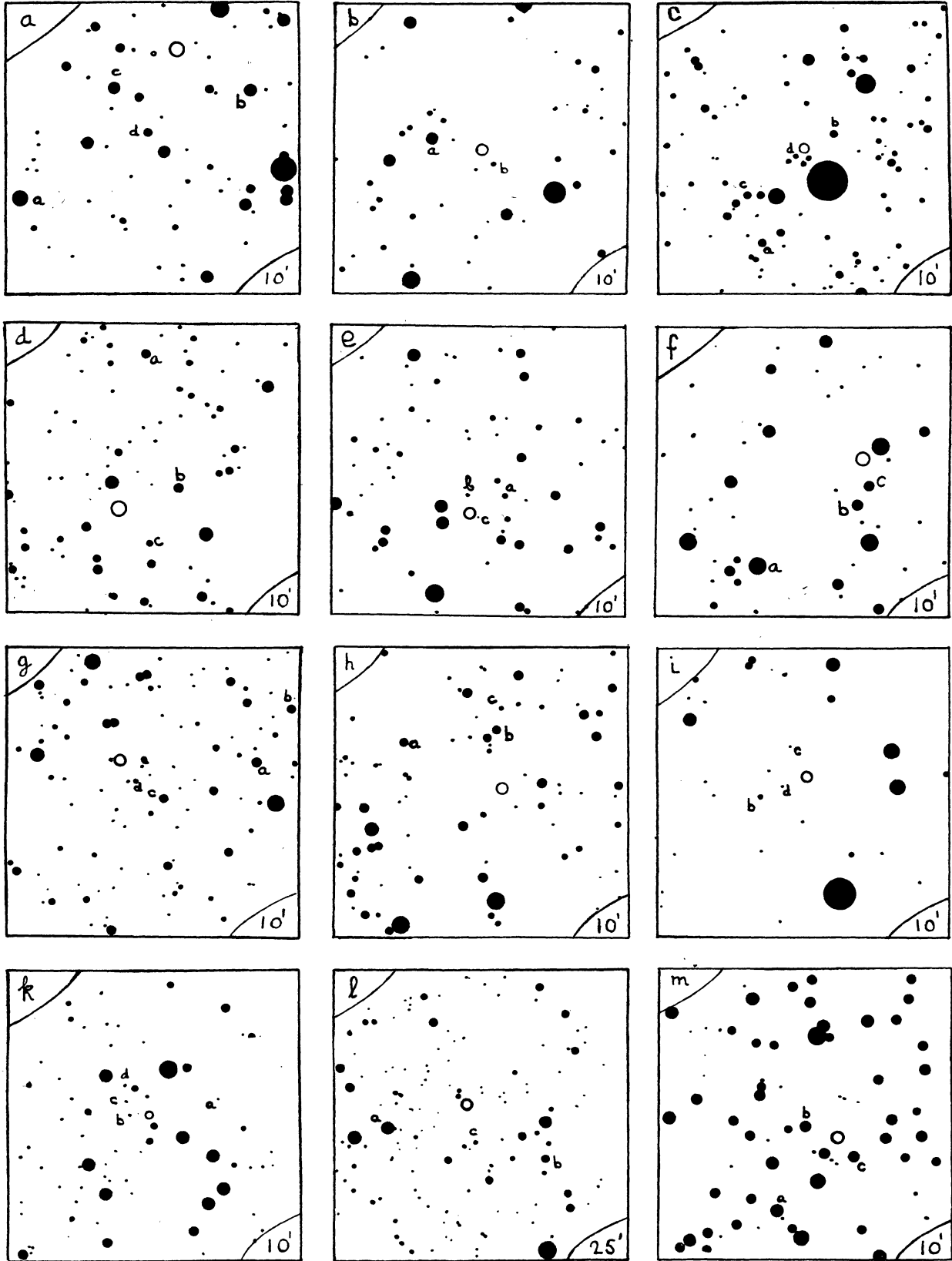
TABLE I.

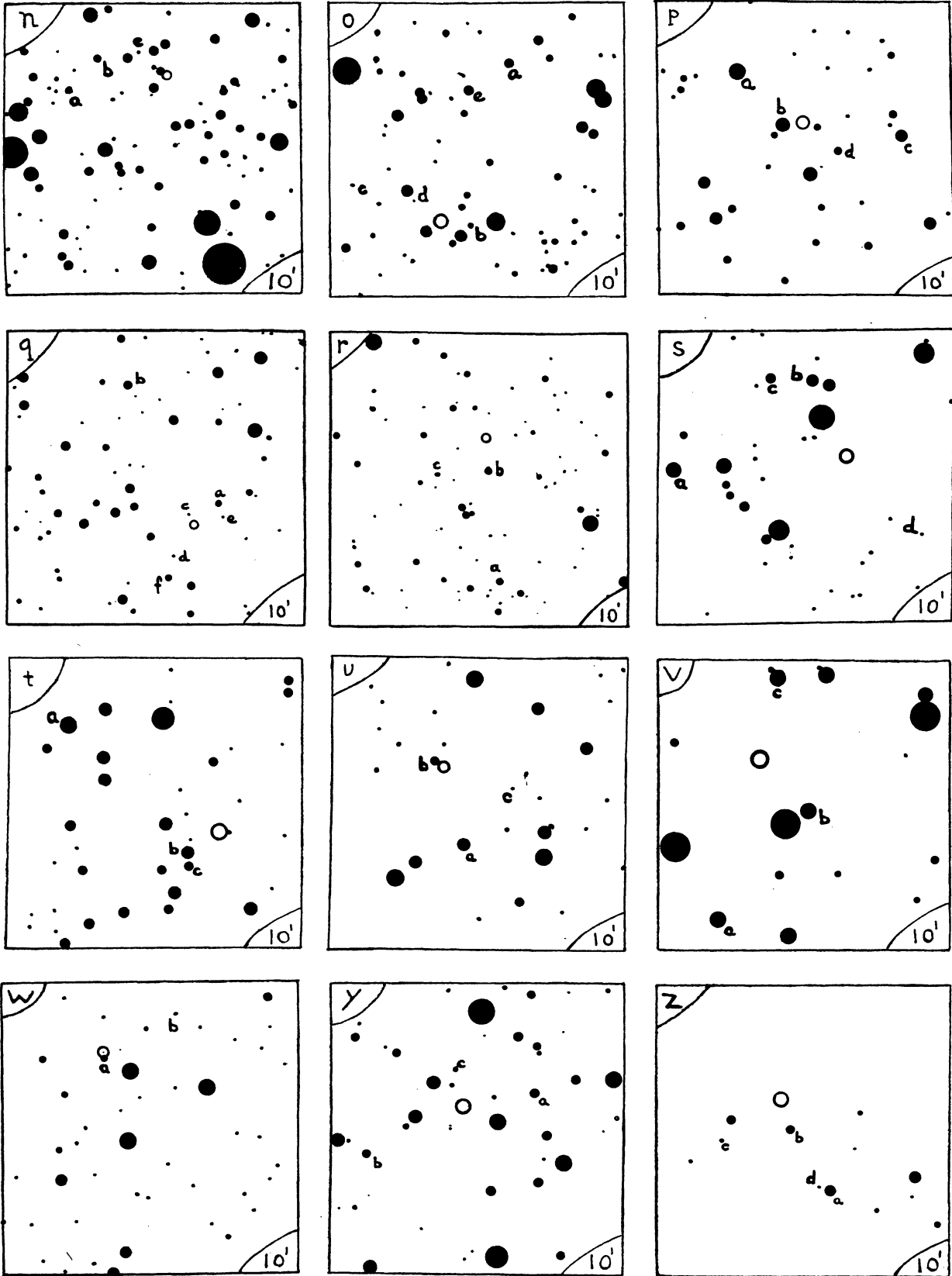
	α (1875)			number of plates	period		m. e.	reciprocal period	phase of epoch	epoch 2420000+	fraction of period occupied by min.	max. min. range			m. e. of single estimate	max. min.			
	h	m	s		d	d						d ⁻¹	P	d		P	s	s	s
<i>a</i>	11	22	30	—	59	28.8	334	19.138	± .003	.052253	.29	4464.62	.11	3.0	8.0	5.0	± .92	11.4	12.1
<i>b</i>	11	27	34		59	8.2	337	1.132963	± .000015	.882641	.580	4776.098		1.8	6.8	5.0		12.0	12.4
<i>c</i>	11	31	48		60	46.4	383	9.0904	± .0004	.110006	.539	4950.085	.09	1.7	5.9	4.2		12.6	13.8
<i>d</i>	11	32	17		62	24.4	329	20.461	± .004	.048873	.70	4679.48	.18	2.1	4.5	2.4		12.0	12.5
<i>e</i>	11	35	25		61	34.2	305	5.89192		.169724	.104*	4773.067		2.0	5.0	3.0	± .52	13.4	13.8
<i>f</i>	11	36	48		58	18.0	334	12.1822	± .0010	.082087	.406*	4780.367		— 1.1	12.2	13.3	± 1.01	11.3	12.3
<i>g</i>	11	40	0		62	3.2	324	5.29654		.188802	.971	4708.483	.045	3.5	8.7	5.2		12.3	13.6
<i>h</i>	11	43	12		61	22.3	324	.773320	± .000006	1.293125	.552	4775.681		0.2	2.0	1.8	± .57	13.0	13.4
<i>i</i>	11	54	5		65	19.0	307	3.705654	± .00006	.2698579	.081	4776.888	.11	1.8	5.1	3.3		13.5	14.1
<i>k</i>	11	54	51		63	2.5	303	1.615336	± .000008	.619066	.459	4691.679	.14	— 0.7	6.7	7.4		13.5	14.4
<i>l</i>	11	55	1		65	11.0	329	3.11175	± .00005	.321362	.527	4775.073	.18	7.3	14.0	6.7		11.0	11.7
<i>m</i>	11	56	53		62	14.5	312	.3881358	± .000002	2.576418	.378	4776.1574		1.8	3.8	2.0	± .45	12.4	12.9
<i>n</i>	11	57	20		60	18.1	348	61.50		.01626	.504	5381.56	.20	1.5	4.0	2.5		13.6	14.1
<i>o</i>	12	0	10		61	54.5	302	4.9898	± .0003	.20041	.770*	4774.063		2.7	7.2	4.5	± .68	12.5	13.2
<i>p</i>	12	3	18		60	7.3	315	1.045570	± .000034	.956416	.232	4791.045	.15	2.3	5.4	3.1		11.8	12.4
<i>q</i>	12	5	45		62	30.3	303	3.55272	± .00005	.281474	.698	4834.187	.11	1.4	12.5	11.1		13.4	14.3
<i>r</i>	12	5	58		61	57.3	309	16.120		.062035	.845	4478.843	.06	1.8	5.5	3.7		13.0	14.0
<i>s</i>	12	7	25		63	59.3	374	6.35447	± .00014	.157370	.849	4949.157	.07	3.5	7.0	3.5		12.3	13.1
<i>t</i>	12	11	20		62	0.4	328	4.0987		.24398	.36	4776.46		6.7	11.4	4.7		11.7	12.1
<i>u</i>	12	14	38		63	40.2	313	1.40443	± .00002	.71203	.317	4917.373	.15	2.7	6.0	3.3		12.7	13.5
<i>v</i>	12	14	40		57	8.3	375	56.252		.017777	.882	4943.58	.05	1.2	2.9	1.7		11.1	11.6
<i>w</i>	12	16	36		63	46.7	245	6.1217		.163353	.662*	4778.988		— 1.0	6.7	7.7	± 1.25	13.0	13.6
<i>x</i>	12	26	12		62	48.9	333	5.26485		.189939	.332*	4776.965		0.4	7.7	7.3	± 1.00	10.3	11.3
<i>y</i>	12	27	20		60	32.6	324	12.2144	± .0017	.081870	.723*	4772.481		2.4	13.0	10.6	± 1.16	12.2	13.7
<i>z</i>	12	30	53		63	12.1	305	.562905	± .000002	1.776498	.380	4775.902	.23	2.5	6.2	3.7		12.9	14.0

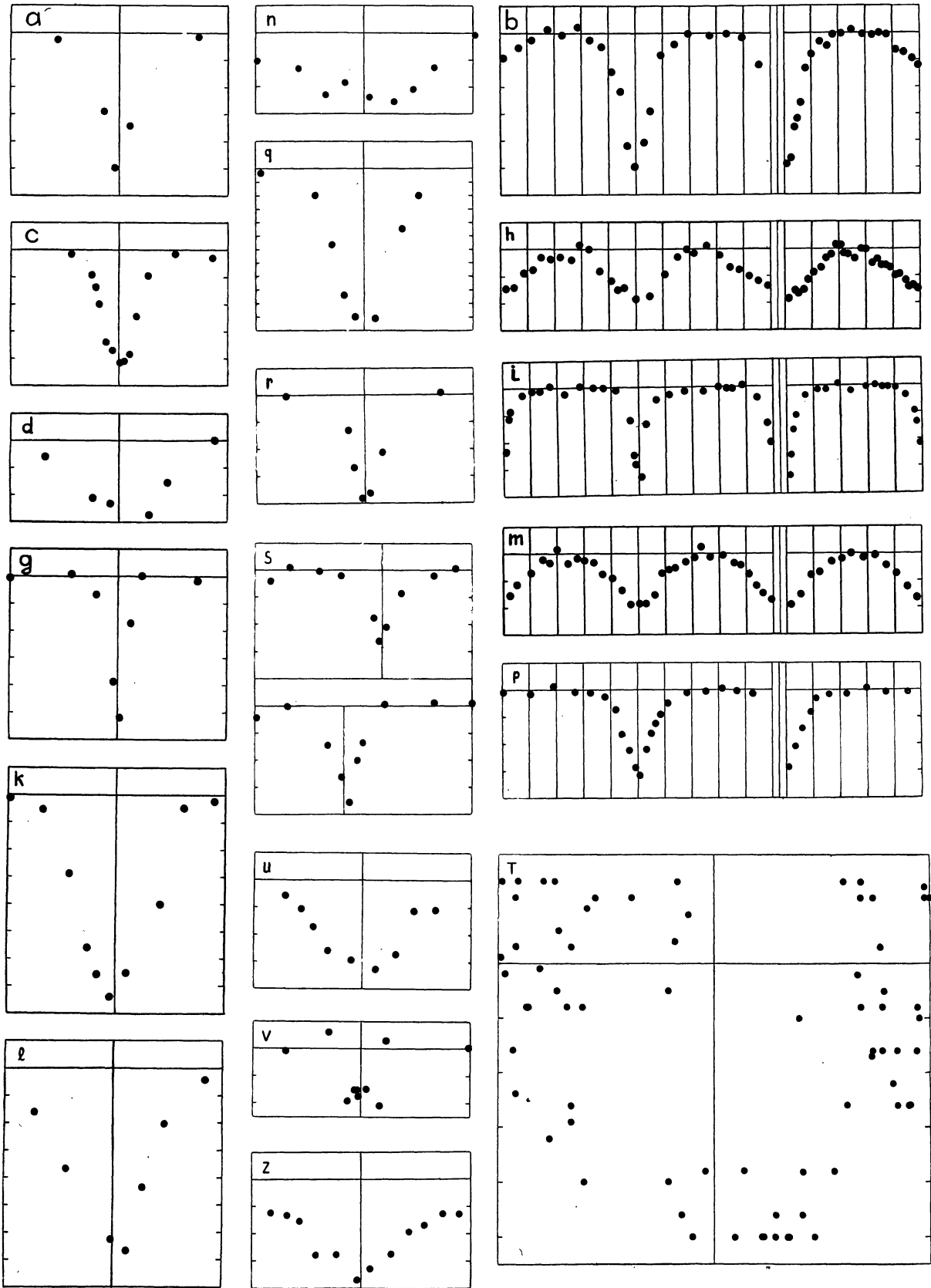
The total number of estimates is 8095. The phases have been computed with J. D. 2420000 as zeropoint and with the reciprocal period given in Table 1. For the Algol stars the epoch of the phase of the principal minimum has been given, for the δ Cep. stars the character of the epoch given in Table 1 is explained

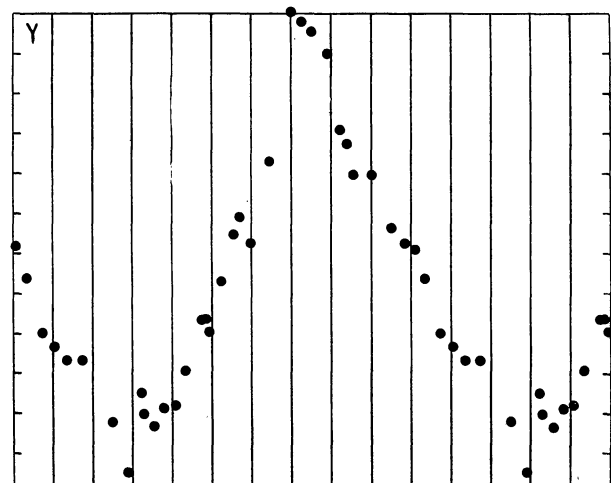
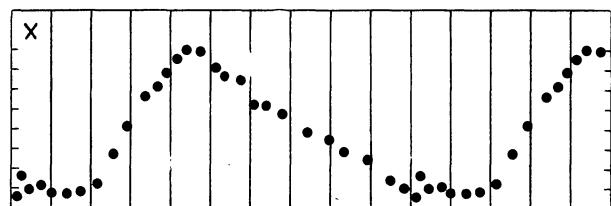
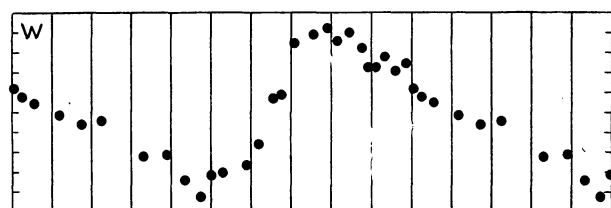
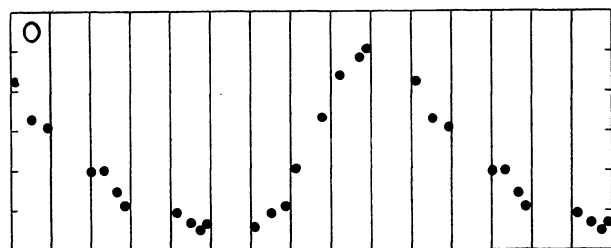
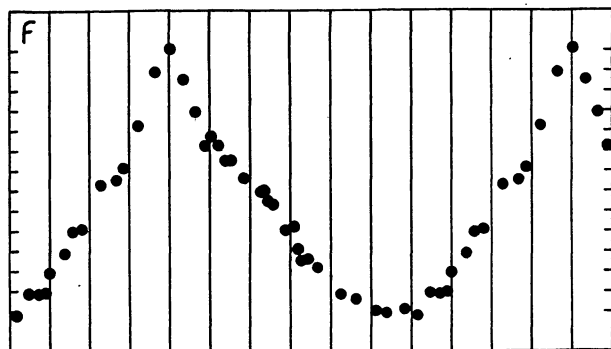
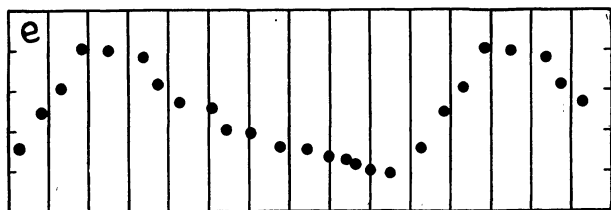
in the remarks. The normal epoch has been chosen near to the mean date of all the observations, or of the observations in the minimum.

The steps used are very different from star to star. In each case the magnitudes of the comparison stars were determined by comparison with the Harvard









sequence of the Selected Area 193, which is in the field of the plates used. The results have been given in Table 3. The corresponding magnitudes of the variables at maximum and minimum have been given in Table 1.

In Table 2 the types of the variables have been indicated with the J. D.'s of the plates by the comparison of which the variability was discovered. Of the 25 variables 19 are eclipsing systems, 6 are δ Cep stars. Of the eclipsing variables 2 are of the W UMa type, one of them shows a notable excentricity of the orbit.

Table 4 gives for each of 21 variables the epochs used for computing the period, together with the counting of epochs and the residuals $O-C$. For the eclipsing variables epochs of minimum were used, or a fixed brightness on the descending or the ascending branch of the lightcurve. For two of the δ Cep stars with sharp maximum (f and y) epochs of maximum were used. The periods of the four other δ Cep stars have been determined in the following way. Using a provisional period, the observations on the ascending branch of the lightcurve were plotted separately for the years 1924, 1925 and 1928. The shifting of the phase of a fixed brightness on the ascending branch gave a correction to the provisional period.

The estimates of each variable have been arranged according to phase and divided into groups, the means of which are given in Table 5. For the eclipsing variables only those mean values are shown on the diagrams which are marked with an asterisk in Table 5. In addition new phases have been computed for 5 eclipsing variables, taking the principal minimum as zeropoint and disregarding the sign of the phase.

On all the diagrams of the lightcurves vertical lines have been drawn for each tenth of the period. For eclipsing variables the principal minimum has for convenience been placed in the middle of the diagram. The brightness is indicated in steps, counted from the maximum light.

The size of the diagrams of the surroundings of each of 23 variables is $10' \times 10'$. For the variable l the size is $25' \times 25'$. On each of these diagrams the variable and the comparison stars used have been indicated. For the variable x no diagram of the surroundings has been given, since the variable and the 2 comparison stars occur in the $C. P. D.$ The $C. P. D.$ numbers of these stars are given below.

The following remarks relate to the individual objects.

a. The comparison star a is $C. P. D.$ — $59^{\circ}3509$ ($10^m 0$).

e. At two phases differing by $\cdot 35$ the brightness is the same on the ascending and the descending

some difficulties. The nights in which more than one observation near minimum was made, and two nights in which only one plate is taken, on which the star is in minimum, gave an apparently satisfactory period. But on a few nights the star was estimated faint between two bright estimates. And in the minimum some estimates of maximum brightness occur. The explanation of this lies probably in the fact that the star is one component of a double star, so that the two components are never completely separated on the plate. The individual estimates near minimum have been given on the diagram.

u. The period has been obtained both from the epochs of minimum and from points on the descending branch.

v. An eclipsing variable of long period: 56 days, and small range: $^m.5$. The duration of the minimum is 3 days. Of this star 10 plates in one night are taken near minimum, covering .2 days, so that a shorter period seems to be excluded. The descending and ascending branches of the lightcurve are probably steep. The comparison star *a* is *C. P. D.* - 57°5417 ($9^m.9$).

w. At two phases differing by .4 the brightness is

the same on the ascending and the descending branches. The epoch of the first of these two phases has been given in Table 1.

x. This δ Cep. star is *C. P. D.* - 62°2817 ($9^m.5$) = H. D. 109176 ($10^m.2$; Ko). The comparison stars are: *a* = *C. P. D.* - 62°2813 ($9^m.4$) and *b* = - 62°2820 ($10^m.0$). At two phases, differing by .3, the brightness is the same on the ascending and the descending branches. The epoch of the first of these two phases has been given in Table 1.

In the course of the work this star was found to be previously marked on one of the plates.

y. The period of this δ Cep star is nearly the same as that of the star *f*. The lightcurves are very similar, but the range of this variable is larger. The epoch of maximum has been given in Table 1.

z. The duration of the minimum is .23 of the adopted period, but the period has probably to be doubled. The period has been obtained both from epochs of minimum and from points on the ascending branch.

I want to thank Prof. E. HERTZSPRUNG for his help and advice in preparing this paper.

TABLE 4.

d	e	d	d	e	d	d	e	d	d	e	d		
			2424290.8	17	0	2423886.24	7	+	8	2423898.40	1073	+	6
			5354.7	69	+	3940.34	147	+	5	4586.51	1818	+	10
2423967.3	0	+	5415.9	72	0	45.34	160	+	2	.54	1818	+	13
86.2	1	0				59.28	196	+	4	5329.49	3740	—	8
4292.3	17	—			<i>f</i>	64.28	209	+	1	.52	3740	—	5
4560.5	31	+			2420991.27	0	—	8	67.28	5330.32	3742	—	2
5383.4	74	0			3940.34	242	+	2	71.25	.39	3742	+	5
					64.28	244	+	6	74.26	31.43	3745	—	7
					76.25	245	—	0	76.25	.55	3745	+	5
2423791.539	0	+			88.22	246	—	4	90.21	50.40	3794	—	5
99.528	7	+			.24	246	—	8	98.21	54.31	3804	—	1
3943.337	134	—			89.22	246	+	3	4259.26	.57	3805	—	13
68.265	156	—			4000.21	247	—	1	.28	57.35	3812	—	6
76.254	163	+			4281.23	247	—	2	.31	.37	3812	—	4
4258.351	412	+			93.20	270	+	5	60.40	62.32	3825	—	12
59.445	413	—			.22	271	+	9	62.29	.48	3825	+	4
91.204	441	+			.29	271	+	6	.32	81.31	3874	—	7
.227	441	+			.39	271	+	7	63.47	85.23	3884	—	2
92.295	442	—			.41	271	+	5	64.37	86.37	3887	—	4
.319	442	—			4560.52	293	—	9	86.27	.44	3887	+	3
4566.504	684	+			5206.53	346	—	3	.39	88.35	3892	+	1
5332.360	1360	—			5816.21	396	+	6	.42	93.35	3905	—	2
.384	1360	0			6010.42	412	—	5	89.40	.38	3905	+	1
.407	1360	+			.45	412	—	0	90.22	5415.34	3962	—	7
					.56	412	—	3	91.35	.36	3962	—	5
								2	.40	17.39	3967	+	5
2424259.075	0	+						4	.42	41.30	4029	—	1
86.292	3	—						4	92.20	53.25	4060	—	5
5386.277	124	+			2423887.5	0	—	6	.22	5702.41	4704	+	10
6013.504	193	—			3972.3	16	+	3	93.29	14.26	4735	—	4
95.313	202	—			4258.3	70	0	7	.39	.33	4735	+	3
					5386.5	283	—	9	94.39	5807.21	4975	+	11
					5418.3	289	+	7	.41	.26	4975	+	16
								2	96.39	32.21	5040	—	2
2423942.8	0	—						3	97.21	36.23	5050	+	13
63.8	1	+						5	.23	6007.42	5493	+	3
					2423883.52	0	+	3	98.37	10.45	5501	—	3

TABLE 5.

<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness
	P	s		P	s		P	s		P	s
	<i>a</i>		22	'076	1'72				12	'101	1'63
			22	'160	1'65		<i>f</i> *		12	'139	1'09
20	'039	3'12	22	'258	1'62	10	'015	12'22	12	'172	0'94
20	'104	3'14	22	'296	1'80	9	'046	11'10	12	'201	0'50
20	'172	3'32	22	'362	1'78	9	'071	11'13	12	'237	0'58
20	'233	3'26*	21	'428	1'45	9	'089	11'04	12	'273	0'48
2	'276	5'90*	20	'495	1'86*	9	'100	10'08	12	'314	0'60
6	'286	7'98*	5	'514	2'62*	9	'137	9'12	12	'344	0'04
5	'300	6'44*	3	'518	3'07*	9	'157	8'03	12	'380	0'22
20	'363	3'14*	4	'521	3'70*	9	'180	7'91	12	'419	1'02
20	'431	2'98	3	'527	5'10*	9	'228	5'70	12	'461	1'37
20	'480	2'98	4	'533	5'40*	9	'267	5'46	12	'485	1'71
20	'536	3'10	4	'540	5'85*	9	'285	4'87	12	'509	1'63
20	'570	2'82	3	'544	5'80*	9	'322	2'77	12	'552	2'05
20	'597	2'68	4	'549	5'55*	9	'364	0'09	12	'603	1'99
20	'627	2'81	2	'555	4'15*	9	'403	1'11	12	'659	1'14
20	'706	3'00	4	'566	2'65*	9	'434	0'44	12	'706	0'50
20	'784	3'18	21	'591	1'85*	9	'465	2'07	12	'741	0'20
20	'836	2'92	21	'625	2'00*	9	'488	3'76	12	'768	0'35
20	'918	3'16	22	'655	1'86	9	'503	3'28	12	'815	0'08
21	'975	3'05	22	'708	1'57	9	'521	3'73	12	'863	0'43
	<i>b</i> *		22	'757	1'70	9	'538	4'50	12	'901	0'87
20	'037	2'99	22	'841	1'59	9	'553	4'51	12	'936	0'96
20	'092	2'72	22	'890	1'75	9	'584	5'39	12	'972	1'20
20	'150	2'34		'955	1'59	9	'627	6'07			
20	'198	2'06				9	'636	6'01	12	'015	2'05
20	'256	1'66		<i>d</i>		9	'644	6'52	12	'039	1'75
20	'309	1'86	20	'036	2'15	9	'658	6'70	12	'051	1'86
20	'368	1'58	20	'114	2'12	9	'688	7'98	12	'071	1'72
20	'411	2'26	20	'184	2'19	9	'708	7'80	12	'086	1'35
16	'455	2'37	20	'241	2'39	9	'718	8'93	12	'108	1'08
5	'492	3'24	20	'300	2'29	9	'726	9'51	12	'135	0'91
5	'523	3'98	20	'402	2'00	9	'744	9'41	12	'153	0'55
5	'548	6'00	20	'473	2'14	9	'767	9'84	12	'173	0'42
10	'575	6'77	20	'521	2'12	9	'825	11'16	12	'190	0'06
10	'610	5'87	20	'566	2'27	9	'863	11'41	12	'208	0'07
10	'633	4'71	20	'598	2'29	9	'912	11'99	12	'219	0'37
16	'674	2'62	12	'632	2'70*	9	'938	12'09	12	'235	0'41
20	'725	2'22	10	'676	4'21*	9	'982	11'92	12	'260	0'57
20	'775	1'82	12	'692	4'42*				12	'285	0'22
20	'854	1'90	11	'728	4'84*	18	'044	3'54*	12	'303	0'23
20	'918	1'83	5	'746	3'64*	18	'144	3'65	12	'323	0'75
20	'974	1'97	20	'788	2'11*	18	'179	3'63	12	'343	0'60
7	'008	6'67	20	'835	2'17	18	'214	3'56	12	'357	0'82
11	'024	6'45	20	'885	2'12	18	'258	3'43	12	'374	0'83
8	'037	5'31	19	'964	1'98	18	'311	3'24	12	'390	0'91
7	'048	4'99				18	'352	3'21	12	'410	1'19
6	'060	4'40		<i>e</i> *		18	'382	3'50	12	'423	1'17
9	'080	3'12	17	'027	4'44	18	'424	3'46	12	'447	1'39
19	'101	2'60	17	'083	3'54	18	'487	3'58	12	'459	1'62
18	'132	2'13	17	'132	2'94	18	'565	3'68	12	'476	1'56
22	'159	2'28	17	'184	1'95	18	'609	3'62	12	'491	1'68
22	'180	1'86	17	'250	2'01	18	'691	3'52			
24	'210	1'83	17	'338	2'18	18	'759	3'73	7	'050	3'04
24	'247	1'69	17	'374	2'85	18	'813	3'66	5	'063	4'32
23	'289	1'85	17	'428	3'29	18	'871	3'56	6	'069	4'67
23	'324	1'87	17	'508	3'44	18	'928	3'43*	5	'090	5'14
17	'350	1'81	17	'544	3'99	6	'951	4'17*	5	'109	3'18
22	'377	1'87	17	'605	4'06	5	'967	7'38*	15	'147	2'27
23	'413	2'43	17	'677	4'41	2	'973	8'70*	15	'195	2'09
24	'442	2'51	17	'744	4'47	3	'983	5'23*	15	'253	1'96
18	'473	2'74	17	'799	4'64	2	'993	3'50*	15	'322	1'97
10	'495	2'99	17	'841	4'72				15	'380	1'81
	<i>c</i>		17	'864	4'85				15	'410	1'86
22	'022	1'81	17	'900	4'99	12	'006	1'38	15	'430	1'87
			16	'950	5'06	12	'042	1'56	15	'467	1'75
						12	'072	1'67	15	'521	2'21

TABLE 5 (continued).

<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness
	P	s		P	s		P	s		P	s
5	.557	3.18	14	.436	7.55*	31	.198	1.84	43	.299	2.23
4	.571	3.88	5	.466	8.90*	27	.309	1.49	43	.371	2.40
2	.592	4.15	5	.495	10.96*	3	.406	2.53*	43	.452	2.40
4	.603	2.95	5	.524	13.56*	6	.445	2.83*			
3	.608	2.67	5	.538	13.98*	5	.469	3.80*		<i>q</i>	
14	.653	2.07	5	.553	11.64*	8	.486	3.35*	21	.024	1.25
14	.689	1.92	6	.573	9.32*	5	.509	3.90*	21	.082	1.74
14	.718	1.92	14	.610	7.73*	5	.532	4.04*	21	.114	1.37
14	.755	1.74	15	.642	7.53	4	.549	3.60*	21	.164	1.83
14	.809	2.04	15	.671	7.81	6	.568	2.78*	21	.269	1.55
14	.866	1.77	15	.706	7.06	31	.608	1.57*	21	.333	1.50
14	.916	1.79	15	.740	7.45	31	.674	1.39	21	.381	1.45
14	.952	1.82	15	.774	7.77	31	.702	1.28	21	.430	1.53
14	.998	1.91	15	.857	7.48	31	.753	1.22	21	.504	1.42
			15	.939	7.05	31	.808	1.52	20	.604	1.76*
7	.009	5.13	15	.976	7.28	31	.895	1.41	5	.653	3.38*
7	.015	4.37							4	.669	7.00*
7	.024	3.44		<i>m</i> *			<i>o</i> *		7	.680	10.76*
7	.035	2.91				17	.011	3.45	4	.690	12.40*
23	.070	2.17	10	.027	2.02	17	.054	4.40	4	.709	12.50*
27	.114	1.93	10	.052	2.14	18	.092	4.62	5	.732	5.88*
21	.146	1.95	10	.078	1.64	15	.202	5.71	3	.747	3.40*
22	.191	1.74	10	.119	2.17	16	.234	5.69	20	.811	1.14
27	.238	2.00	10	.154	1.97	16	.266	6.26	21	.879	1.74
29	.296	1.85	10	.181	2.04	16	.285	6.58	21	.932	1.42
20	.329	1.80	10	.214	2.13	17	.415	6.76			
26	.354	1.87	10	.247	2.56	17	.450	6.99		<i>r</i>	
21	.376	1.86	10	.283	2.70	17	.473	7.19	20	.068	1.76
21	.402	1.90	10	.318	3.16	17	.488	7.02	20	.135	1.83
23	.440	2.17	10	.349	3.67	17	.610	7.12	20	.173	1.89
6	.473	2.77	10	.381	3.64	18	.650	6.76	20	.221	1.76
6	.479	3.18	10	.406	3.66	18	.686	6.59	20	.271	1.69
6	.490	3.97	10	.439	3.33	4	.713	5.65	22	.328	2.07
			10	.465	2.53	5	.778	4.36	18	.418	1.85
	<i>k</i>		10	.493	2.39	19	.822	3.32	20	.498	1.74
20	.040	— 0.60	10	.515	2.34	19	.870	2.87	20	.552	1.56
20	.138	— 0.78	10	.555	2.11	19	.889	2.64	20	.592	1.66
20	.199	— 0.73	10	.588	1.95				20	.647	1.85
20	.295	— 0.70	10	.612	1.56		<i>p</i> *		20	.714	1.64
20	.361	— 0.59*	10	.646	1.94	21	.001	2.38	19	.772	1.81*
6	.391	— 0.17*	10	.694	1.87	20	.058	2.42	2	.827	3.09*
6	.416	2.17*	10	.735	2.14	20	.112	2.57	2	.835	4.50*
5	.433	4.90*	10	.759	2.23	7	.152	3.03	3	.843	5.62*
4	.442	5.88*	10	.790	2.57	8	.175	3.95	3	.849	5.43*
5	.454	6.70*	10	.817	2.99	10	.201	4.54	1	.862	3.93*
5	.469	5.82*	10	.840	3.27	9	.225	5.18	20	.916	1.72*
5	.500	3.30*	19	.872	3.52	7	.239	5.46	19	.976	1.55
7	.521	— 0.22*	10	.905	3.34	10	.264	4.50			
20	.551	— 0.48*	11	.932	2.94	8	.282	3.91		<i>s</i>	
20	.604	— 0.71	11	.983	2.50	7	.299	3.56	11	.016	3.55
20	.653	— 0.68				8	.317	3.21	11	.032	3.39
20	.728	— 0.66	24	.016	3.71	20	.353	2.80	11	.086	3.75
20	.773	— 0.83	24	.054	3.32	20	.416	2.41	11	.114	3.73
20	.817	— 0.69	24	.091	2.60	20	.487	2.39	11	.152	3.77
20	.884	— 0.53	24	.122	2.50	20	.546	2.26	11	.172	3.56
20	.950	— 0.49	24	.167	2.11	20	.601	2.36	11	.213	3.17
			24	.206	2.00	20	.660	2.46	11	.261	3.62
	<i>l</i>		24	.240	1.80	20	.741	2.42	11	.282	3.92*
15	.015	8.16	24	.284	1.96	20	.836	2.46	11	.300	3.40*
15	.052	7.54	24	.328	1.88	20	.925	2.18	11	.326	3.54*
15	.096	8.09	24	.369	2.25				10	.348	3.71*
15	.144	7.80	24	.406	2.54	20	.011	5.17	3	.377	5.27*
15	.211	7.12	24	.447	3.01	20	.036	4.38	6	.383	6.13*
15	.240	6.99	24	.481	3.45	20	.061	3.73	5	.389	5.60*
15	.280	7.29				20	.088	3.11	5	.402	4.36*
15	.326	7.53		<i>n</i>		20	.114	2.62	12	.432	3.71*
15	.354	7.24	31	.021	1.41	43	.159	2.47	11	.452	3.45*
15	.407	7.93	31	.099	2.00	43	.227	2.47	11	.480	3.39

TABLE 5 (continued).

<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness	<i>n</i>	phase	bright- ness
II	P	s	I	P	s	25	P	s	12	P	s
II	.515	3.73	I	.396	10.5 *	25	.388	1.25	12	.678	3.62
II	.550	3.65	I	.396	11.3 *	24	.554	1.37	12	.742	4.56
II	.002	3.37	I	.402	11.7 *	25	.661	1.32	12	.796	4.92
II	.633	3.50	I	.411	10.5 *	24	.722	1.29	12	.834	5.54
II	.649	3.75	I	.415	5.2 *	25	.746	1.30	11	.893	5.95
II	.674	3.84	I	.417	9.3 *	24	.772	1.32	11	.950	6.97
II	.695	3.65	I	.422	6.9 *	25	.812	1.30*	11	.986	7.38
II	.711	3.58	I	.423	7.5 *	4	.852	0.60*			
II	.746	3.51	I	.423	5.5 *	4	.869	3.12*		y*	
II	.769	3.95*	I	.423	5.2 *	4	.876	2.72*			
12	.797	3.53*	I	.428	8.4 *	4	.878	2.72*	10	.008	8.22
4	.834	4.95*	I	.429	8.3 *	4	.880	2.95*	10	.034	9.03
6	.847	6.13*	I	.429	5.5 *	4	.887	2.68*	10	.073	10.40
5	.855	7.00*	I	.432	6.4 *	4	.900	3.32*	10	.103	10.75
4	.861	5.50*	I	.433	7.5 *	4	.905	0.90*	10	.137	11.08
6	.867	4.87*	I	.433	8.3 *	4	.981	1.17*	10	.173	11.07
10	.888	3.43*	I	.434	7.2 *	24			10	.251	12.63
II	.930	3.40*	I	.438	8.9 *		w*		10	.291	13.89
II	.967	3.42*	I	.440	9.3 *	9	.006	1.83	10	.323	11.91
II	.996	3.58	I	.440	8.3 *	9	.026	2.23	10	.332	12.43
			I	.446	9.3 *	9	.056	2.53	10	.357	12.76
			I	.446	9.3 *	5	.118	3.13	10	.381	12.28
24	.049	6.65	I	.449	7.5 *	9	.173	3.62	10	.408	12.21
24	.151	6.68	I	.449	8.3 *	9	.224	3.43	10	.436	11.35
24	.207	6.40	I	.450	7.7 *	9	.331	5.23	10	.476	10.07
12	.239	6.15	I	.453	5.5 *	9	.390	5.12	10	.486	10.05
I	.256	6.6 *	I	.453	5.8 *	9	.435	6.44	10	.496	10.37
I	.257	5.2 *	I	.455	5.2 *	9	.473	7.27	10	.526	9.12
I	.258	6.9 *	26	.487	6.83	9	.498	6.16	10	.555	7.94
I	.262	8.3 *	24	.578	6.48	9	.532	6.04	10	.569	7.51
I	.263	9.1 *	24	.691	7.32	9	.580	5.66	10	.598	8.15
I	.263	6.4 *	24	.749	7.34	9	.622	4.63	10	.647	6.11
I	.263	5.5 *	25	.823	6.75	9	.657	2.32	10	.699	2.38
I	.264	5.2 *	25	.906	6.37	9	.678	2.17	10	.727	2.62
I	.268	7.5 *	25	.986	6.45	9	.708	—	10	.752	2.87
I	.269	7.5 *				9	.757	—	10	.791	3.42
I	.274	6.8 *		u		9	.790	—	10	.821	5.33
I	.276	5.2 *	19	.048	2.60	9	.814	—	10	.840	5.66
I	.279	9.9 *	19	.107	2.60	9	.845	—	11	.857	6.43
I	.281	5.2 *	19	.151	2.73	9	.877	—	11	.900	6.45
I	.282	7.2 *	22	.197	2.74	9	.893	0.22	11	.949	7.77
I	.283	6.1 *	5	.246	3.26*	9	.913	0.76	11	.985	8.15
I	.287	7.5 *	6	.261	3.76*	9	.935	0.22			
I	.289	6.4 *	5	.272	4.42*	10	.959	0.90			
I	.289	9.3 *	5	.288	5.28*	10	.987	0.54	21	.037	2.55
I	.289	9.6 *	4	.307	5.63*				20	.107	2.62
I	.294	7.5 *	4	.329	5.98*		x*		20	.178	2.56
I	.295	10.7 *	4	.348	5.13*	12	.014	7.64	22	.243	2.60
I	.296	5.7 *	5	.365	3.84*	12	.026	6.77	5	.297	3.74*
I	.300	5.5 *	6	.384	3.80*	12	.046	7.43	5	.313	3.82*
I	.317	5.5 *	19	.421	2.82	12	.078	7.32	5	.324	4.04*
I	.334	7.2 *	19	.479	2.81	12	.102	7.62	5	.339	5.28*
I	.334	10.7 *	19	.547	2.66	12	.140	7.65	5	.358	5.26*
I	.337	6.3 *	19	.605	2.85	12	.176	7.57	5	.371	6.20*
I	.338	5.2 *	19	.665	2.79	12	.217	7.16	5	.388	5.78*
I	.340	11.3 *	19	.722	2.94	12	.256	5.66	5	.408	5.24*
I	.343	5.8 *	19	.761	2.77	12	.292	4.24	5	.424	4.44*
I	.345	11.7 *	19	.826	2.87	12	.338	2.76	5	.437	4.18*
I	.361	10.5 *	19	.893	2.73	12	.369	2.27	5	.454	3.78*
I	.365	11.7 *	19	.961	2.80	12	.389	1.57	5	.471	3.80*
I	.369	10.5 *				12	.414	0.88	20	.517	2.33
I	.378	11.7 *		v		12	.441	0.44	20	.585	2.62
I	.378	11.7 *	25	.097	1.22	12	.475	0.50	21	.653	2.56
I	.384	11.3 *	24	.174	1.24	12	.514	1.32	20	.717	2.48
I	.384	11.7 *	25	.215	0.97	12	.535	1.76	20	.801	2.54
I	.390	11.7 *	24	.259	1.22	12	.577	1.94	22	.858	2.57
I	.390	11.7 *	25	.308	1.43	12	.610	3.19	19	.906	2.47
I	.394	7.7 *	24	.349	1.33	12	.642	3.25	20	.972	2.43